

# Comparison of the Smith-Purcell radiation yield for different models

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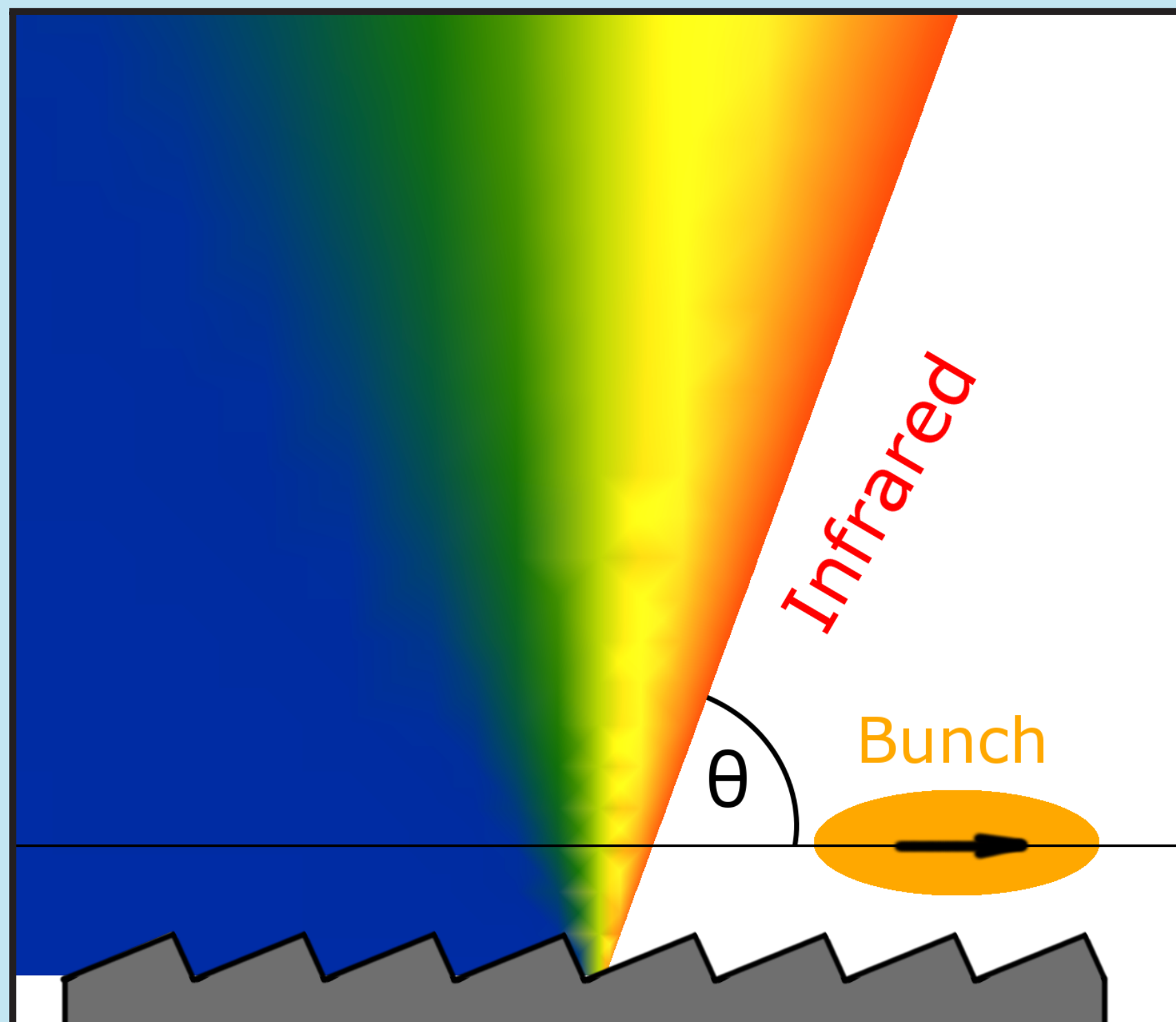
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## Smith-Purcell radiation

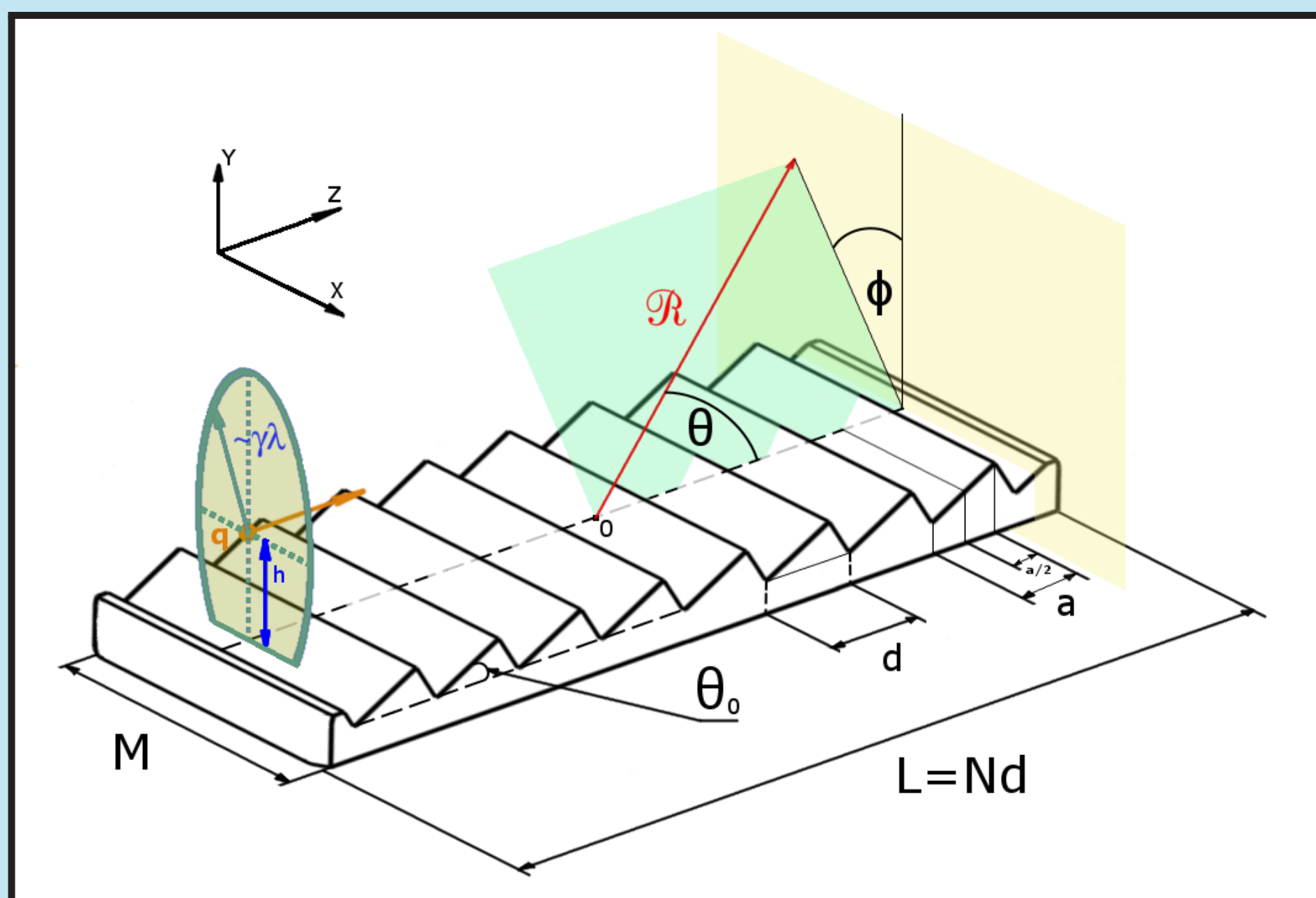
Smith-Purcell radiation is a phenomenon observed, when a charged particle moves near a periodical grating



It has a specific wavelength dependency:

$$\lambda = \frac{d}{n} \left( \frac{1}{\beta} - \cos \theta \right)$$

## Parameters



Symb.	SPESO	E203	Units	Description
$\gamma$	200	$4 \times 10^4$	1	The Lorentz factor (E=100 MeV)
$d$	10	0.25	mm	The grating period
$a$	7.5	0.187	mm	The width of one strip
$R_0$	310	220	mm	The distance between detector and grating
$L$	90	40	mm	The length of the grating
$M$	20	20	mm	The width of the grating
$h$	5	1	mm	The beam-grating separation
$\theta_0$	30	30	deg	The blaze angle
$C'_1$	400	6395	mm <sup>-3</sup>	The normalization constant for the RRR model

## References

- [1] D. V. Karlovets and A. P. Potylitsyn., *Phys. Rev. ST Accel. Beams*, vol. 9, p. 080701, 2006.
- [2] J. H. Brownell, J. Walsh, G. Doucas, *Phys. Rev. E* vol. 57, pp. 1075–1080, 1998.
- [3] D. V. Karlovets and A. P. Potylitsyn., *JETP Letters*, vol. 84, no. 9, pp. 489–493, 2006.

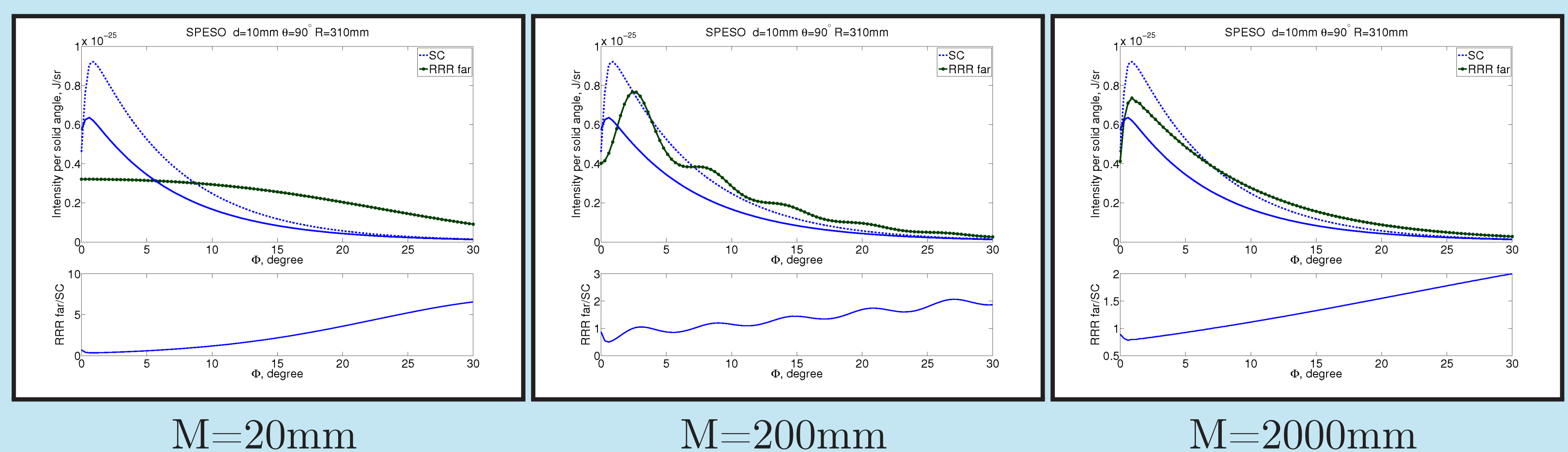
## Models

The models, that were used: **R**esonance **D**iffraction **R**adiation (**RDR**) [1], **S**urface **C**urrent (**SC**, **GFW**)[1,2], **R**esonance **R**eflection **R**adiation (**RRR**) [3].

By SC we mean the Surface current model, with the assumption, that the width of the grating is infinite, whether the GFW model takes it into account.

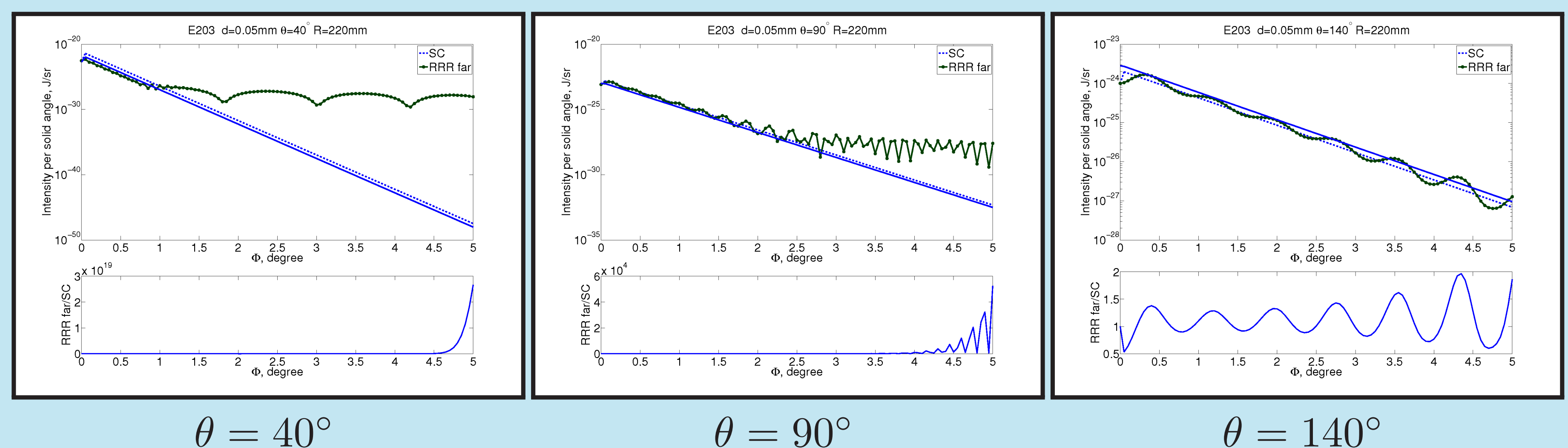
## Width dependance

With the increase of the grating width, the RRR model tends to be like the SC and RDR models. The parameters are from the SPESO experiment.



## Angle dependance

Increasing the observation angle means the decrease of the wavelength, as result the influence of the grating width will be more intense at larger phi. The parameters are from the E203 experiment, the pitch is  $d=50\mu\text{m}$ .



## Comparison of the models

