

SPL-Fréjus

- Some performances
- Fréjus site
- French Photodetector R&D

Thanks: A. Cazes, M. Mezzetto, L. Mosca, Th. Schwetz
and IPNO & LAL engeneers.

Some ingredients for physics analysis

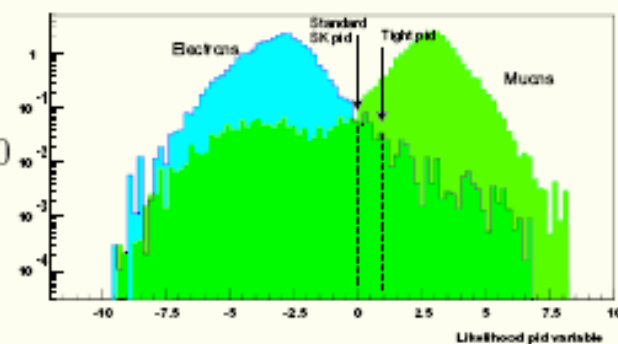
- 440kT Water Č located 130km from CERN (see site later)
- Essentially SK analysis with tighter cuts for e/μ id (cf. hep-ph/0105297)
- Use energy resolution dominated by Fermi motion* (200MeV bins)
- 2% systematics on signal & bkgd (see effect later)
- Optimized machine versions: βB (M. Mezzetto) and SB (A.C + J.E.C)
- Use Atmospheric Neutrinos (Th. Schwetz)
- GLOBES & NUANCE

*: migration matrix for βB

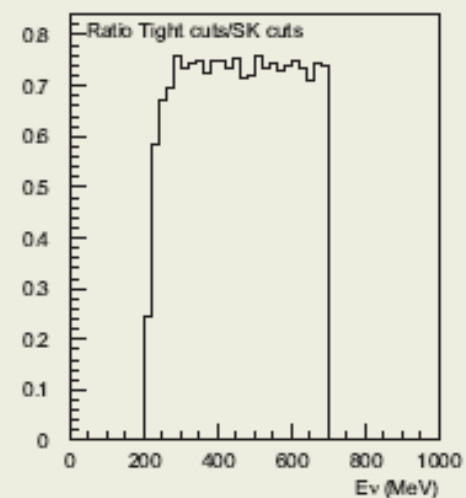
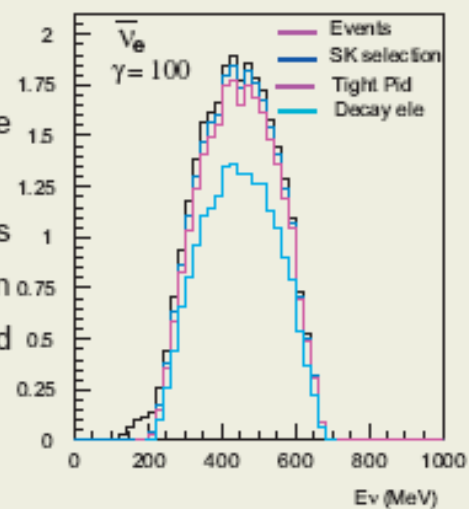
Particle Id (from M. Mezzetto)

Electron/muon misidentification must be suppressed much more than in standard SK analysis to guarantee a negligible background level.

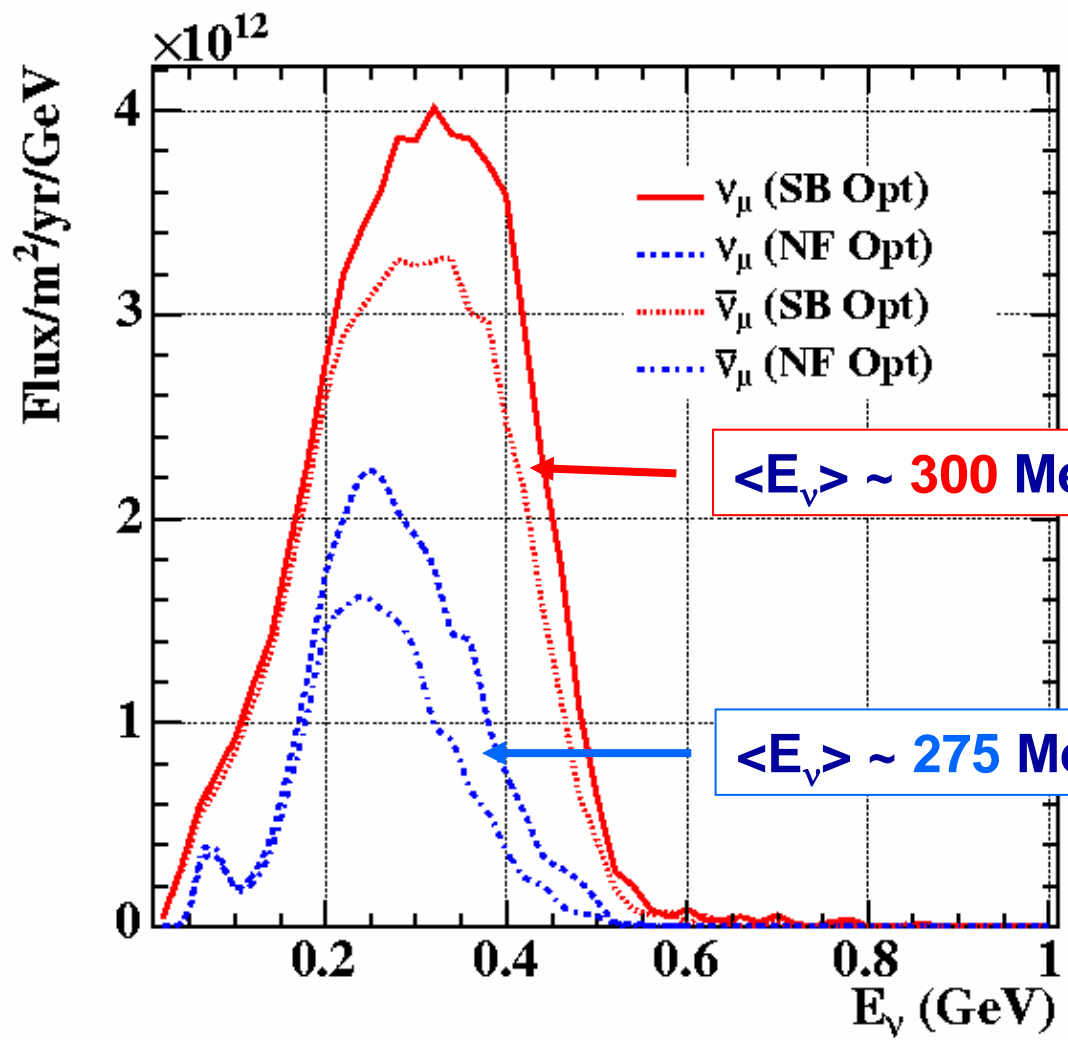
Pid in SK is performed through a Likelihood, $P_{id} > 0$ identifies muons. Use $P_{id} > 1$



To further suppress electron background ask for the signal of the Michel electron from μ decay.
Final efficiency for positive muons. Negative muons have an efficiency smaller by $\sim 22\%$ because can be absorbed before decaying. Electron background suppressed to $\sim 10^{-5}$.



Fluxes comparison @ 130km



$\sim 95 \nu_{\mu}^{CC}/kT/yr^*$



$\langle E_{\nu} \rangle \sim 300 \text{ MeV}, 1.2 \cdot 10^{12}/m^2/yr$

3.5GeV SPL optimum

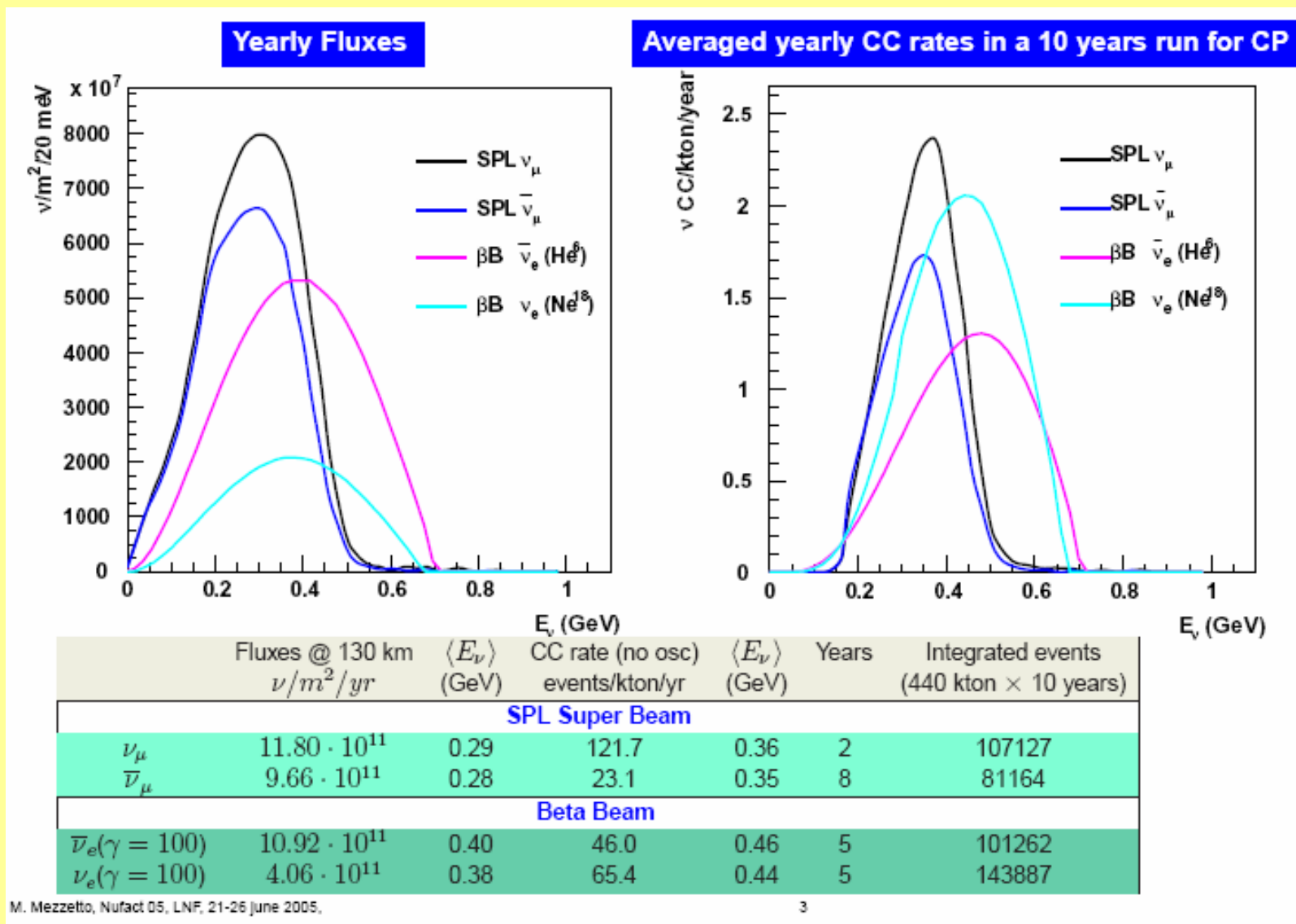
$\langle E_{\nu} \rangle \sim 275 \text{ MeV}, 4.5 \cdot 10^{11}/m^2/yr$

Old ν Fact optimum

Reflector: 50% of the Flux

*: Lipari x-sect. (see later)

β B and SB fluxes



Analysis: GLoBES + M. Mezzetto's parameterization file

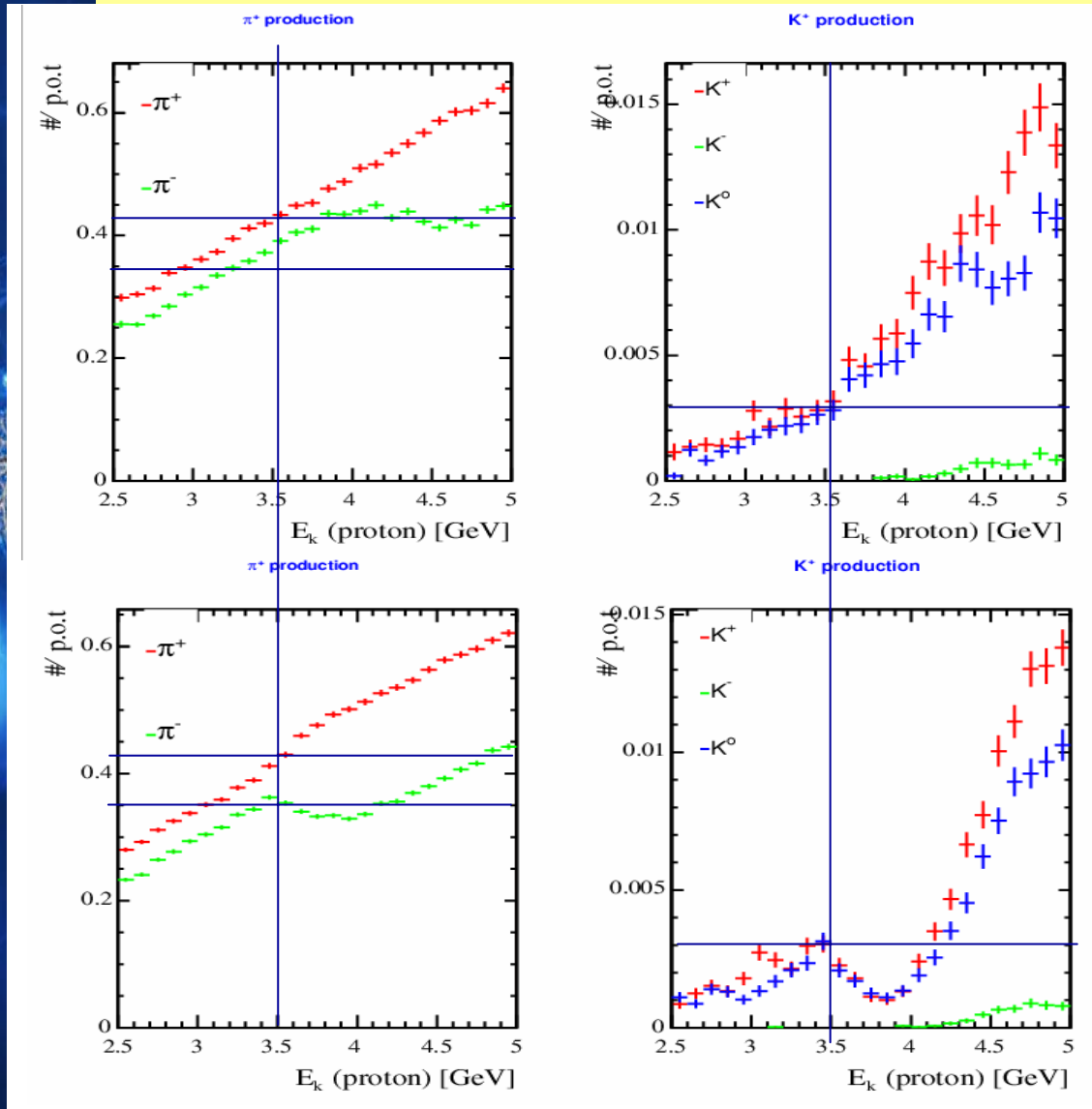
440kT x 5yrs: **2,2 Mt.yrs (+)**

	$\theta_{13} = 1^\circ$	$\theta_{13} = 3^\circ$	$\sin^2 2\theta_{13} = 0.05$	
$\nu_\mu \rightarrow \nu_e$ (Sig)	33 ($\delta = \pi/2$)	330 ($\delta = \pi/2$)	2200 ($\delta = \pi/2$)	3670 ($\delta = 0^\circ$)
$\nu_\mu \rightarrow \nu_e$ (Bkg)	1500			
	$\nu_e \rightarrow \nu_e$ CC	π^0 from NC	$\nu_\mu \rightarrow \nu_\mu$ CC (μ missId)	$\bar{\nu}_e \rightarrow \bar{\nu}_e$ CC
Frac. of Bkg	90%	6%	3%	1%
Reduction Factor	0.707(1060)	$6.5 \cdot 10^{-4}$ (90)	$5.4 \cdot 10^{-4}$ (45)	0.677(15)
$\nu_\mu \rightarrow \nu_\mu$ (Sig)	64950 ($\delta = \pi/2$)		64414 ($\delta = 0^\circ$)	
$\nu_\mu \rightarrow \nu_\mu$ (Bkg)	3 (4.310^{-5} $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$ CC)			

$$\sin^2 2\theta_{12} = 0.82, \theta_{23} = \pi/4, \Delta m^2_{21} = 8.1 \cdot 10^{-5} eV^2, \Delta m^2_{31} = 2.2 \cdot 10^{-3} eV^2$$

Reduction factor and efficiencies taken from SK simulation (D. Casper) and a tight cut for e/ μ misId. (cf. hep-ph/0105297)

Kaon/pion production?



FLUKA 2005.6

+12.5% π^- @3.5GeV

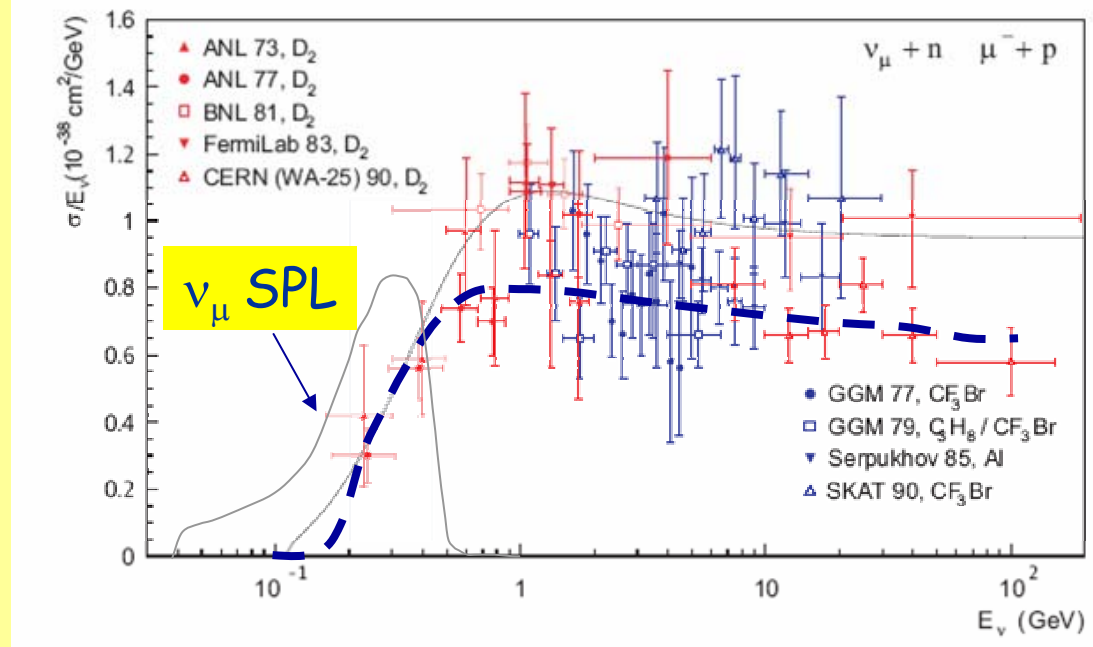
Big difference
[3.5 ÷ 4.5] GeV

FLUKA 2002.4

HARP?

The X-sections

V.V. Lyubushkin et al., internal NOMAD memo



---: Lipari et al.
PRL74(95)4384
on H₂O

βB is an ideal tool to measure these cross-sections and a 2% systematic error on both signal and background are used.

Require close position

Comparison with other facilities

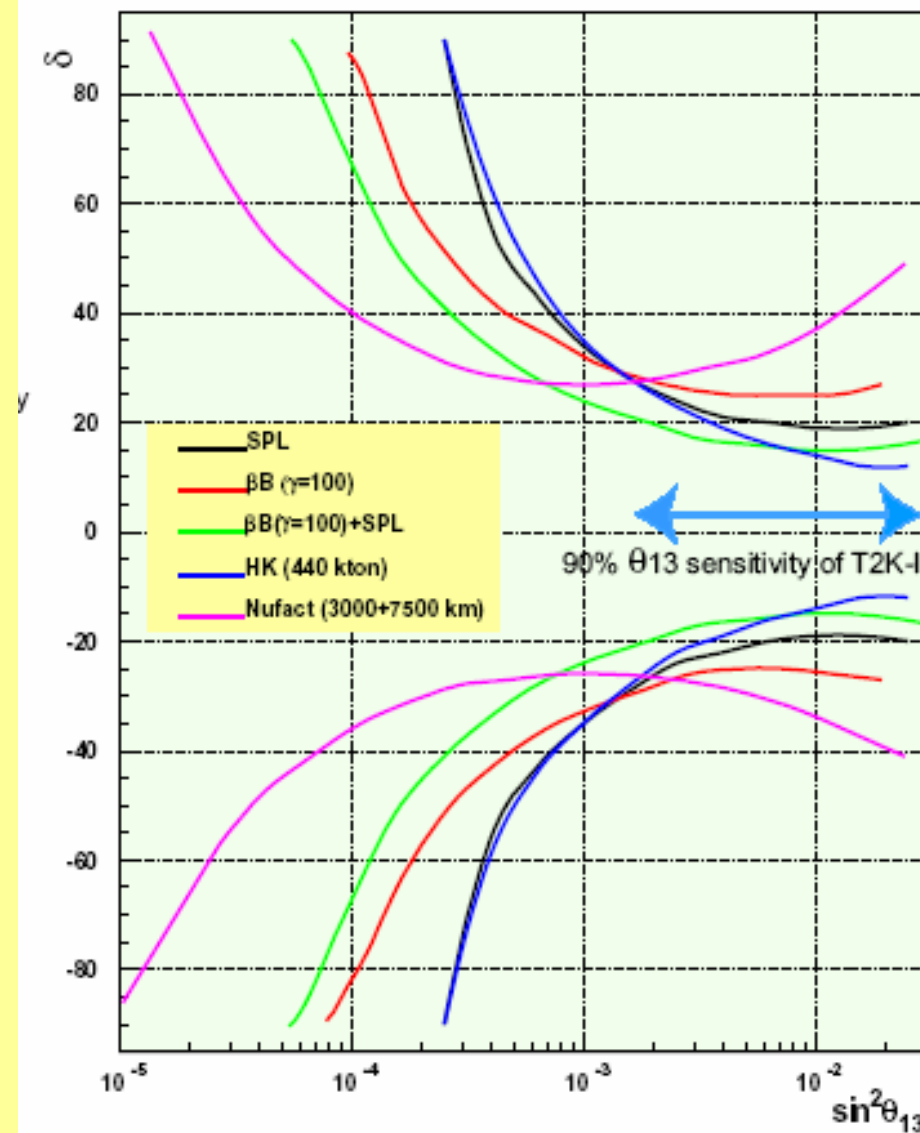
Everything computed with the identical program.
Thanks to the GLoBES experiment library.

HK taken from Huber, Lindner and Winter, hep-ph/0204352, with a fiducial of 440 kton (it was 1 Mton), 2% systematics on QE signal and backgrounds (it was 5%) and 2+8 years running (it was 2+6).

NUFACT taken from Huber, Lindner and Winter, hep-ph/0204352, changing the systematics from 0.1% to 2% and the running time to 5+5 years (it was 4+4).

Other parameters: two iron magnetized detectors, 50 kton, at 3000 and 7500 km, 50 GeV muons, $1E21$ useful decays/year, 5% systematics on matter profile, threshold at 4 GeV, 20 bins from 4 to 50 GeV.

SPL 3.5 GeV (see J.E. Campagne talk) with $2 \nu + 8 \bar{\nu}$ years, 2% systematic error, 200 MeV binning, 440 kton fiducial.

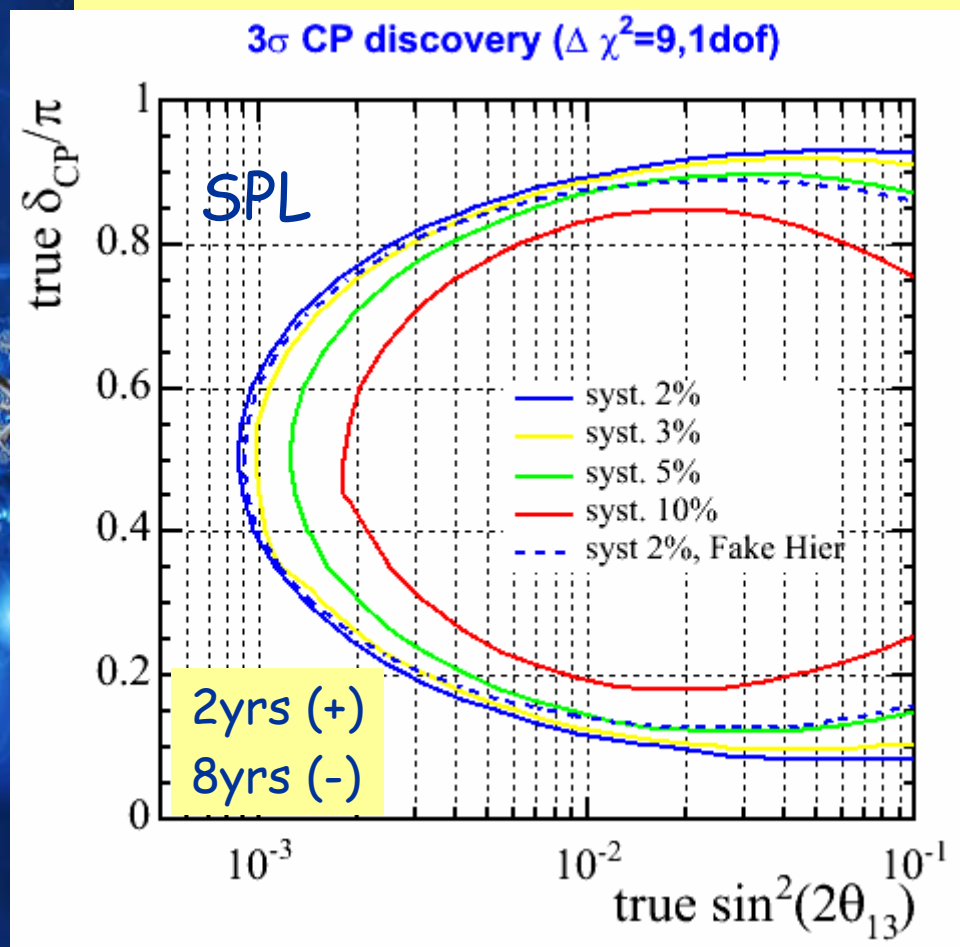


M.M@NuFact05

Systematics...

ISS CERN 05 J.E Campagne LAL

Effect of the systematic (sig. & bkg)

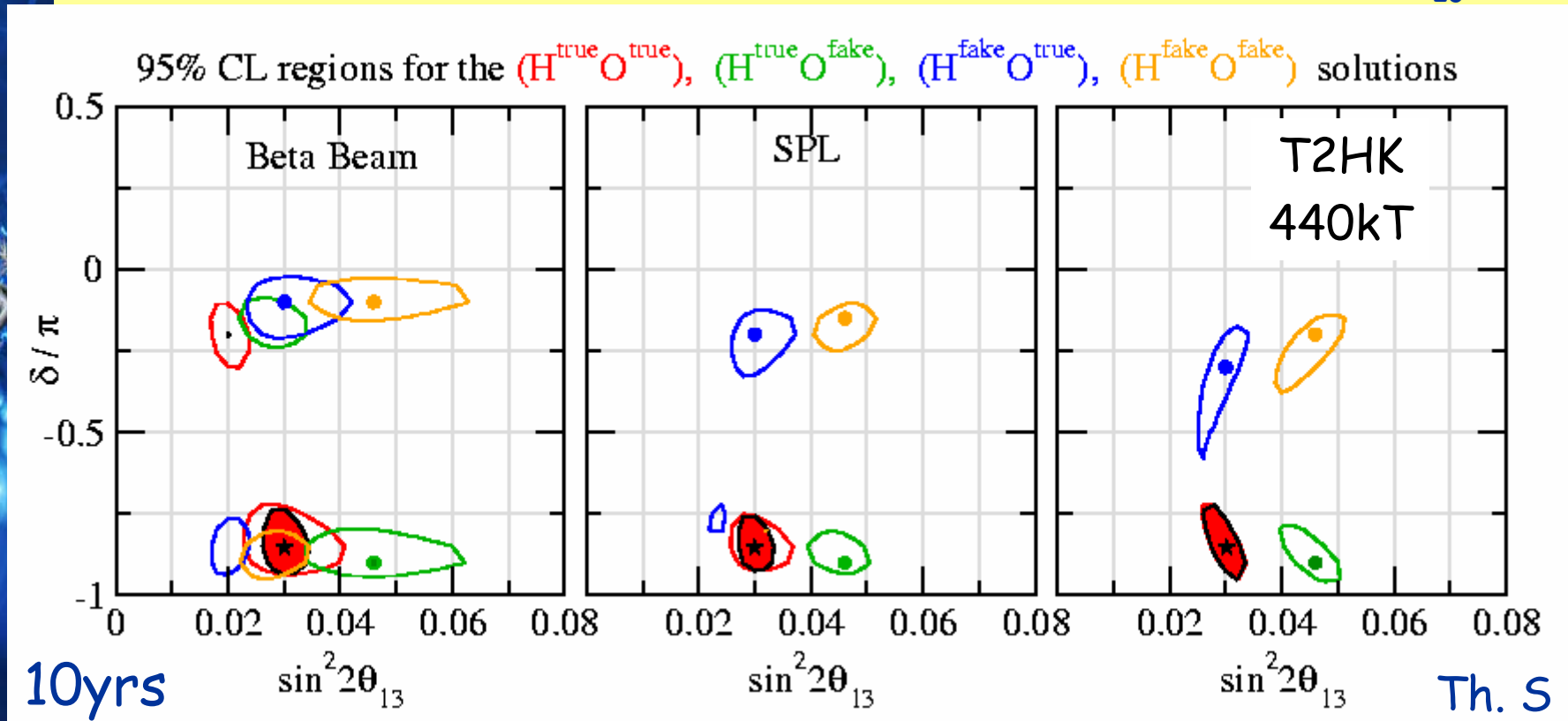


Much more dramatic than ambiguities at small θ_{13}

True values: ($\delta/\pi, \sin^2 2\theta_{13}$)
 $\sin^2 2\theta_{12}=0.82, \sin^2 \theta_{23}=0,4$
 $\Delta m^2_{21}=7.9 \cdot 10^{-5} \text{eV}^2, \Delta m^2_{31}=2.4 \cdot 10^{-3} \text{eV}^2$
 5% external precision on θ_{12} & Δm^2_{21}
 use SPL disappearance channel and spectrum analysis

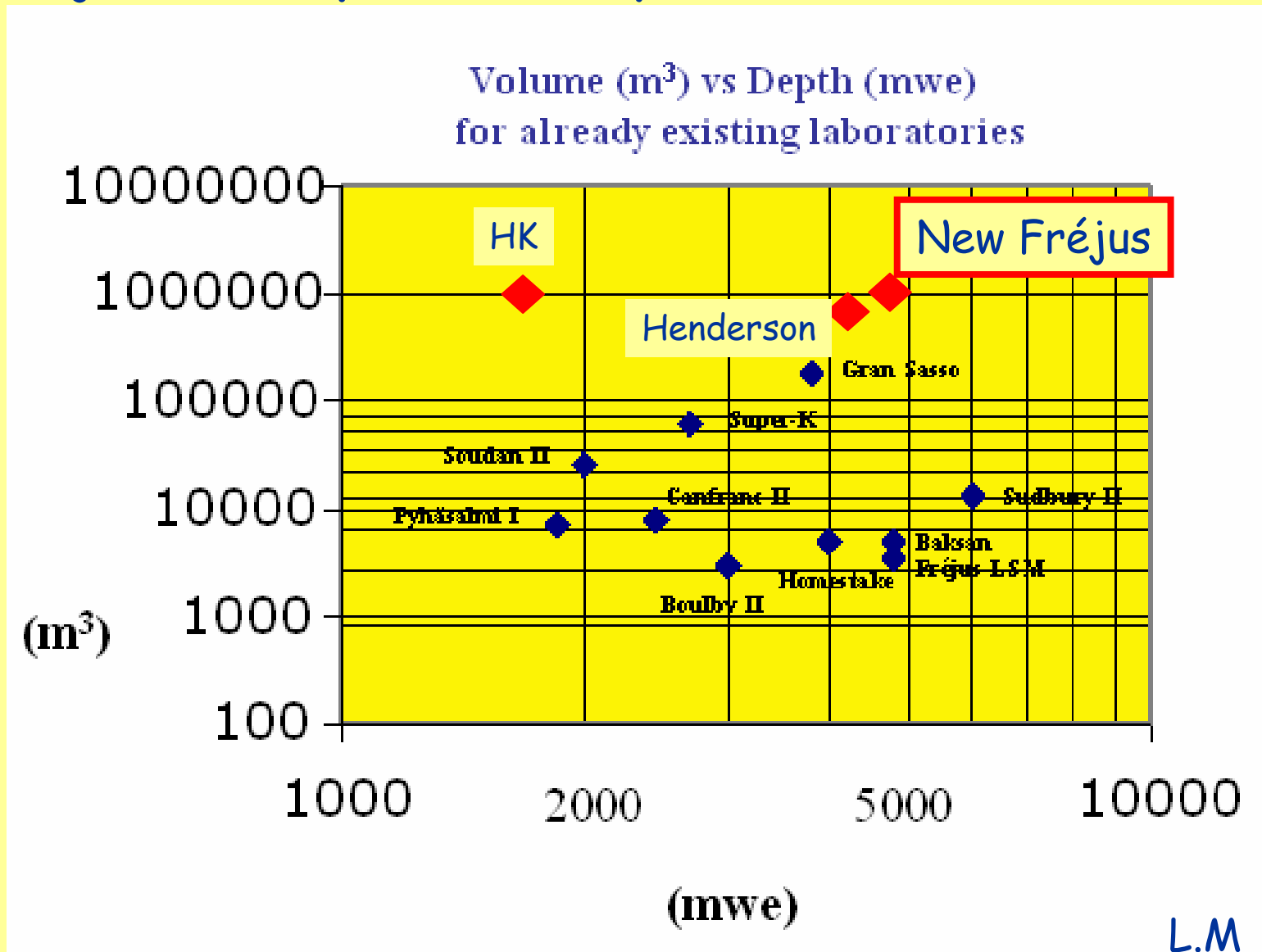
Remove ambiguities with ATM ν

Favorable case $\sin^2\theta_{23}=0.6$



- Contour after ATM combination
- *: true value

Fréjus site possibility

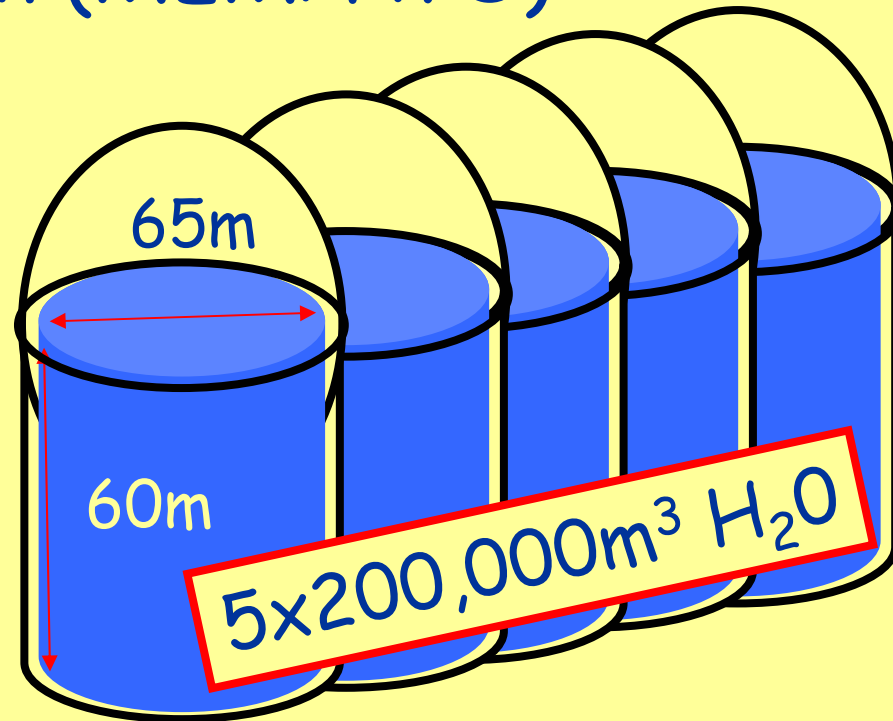


New Fréjus Cavern (MEMPHYS)



Fréjus*

*: Modane 4800mwe



Based on well experienced civil engineer studies.

First cost and time estimate will come soon for a dedicated operation.

Beyond that a Design Study is needed

PMT size \Leftrightarrow cost

Photonis @ NNN05

Diameter	20"	\Leftrightarrow	(20")17"	\Leftrightarrow	12"	
projected area	1660		1450		615	cm ²
QE(typ)	20		20		24	%
CE	60		60		70	%

Cost	2500		2500		800	€
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- *Cost/cm² per useful PE_U = cost/(cm²xQExCE)*

12.6

14.4

7.7

€/PE_U/cm²

Quantities and total cost

Photonis @ NNN05

$$20'' \quad 200,000 \quad \times \quad \text{€ } 2500 = 500M$$

$$12'' \quad 540,000 \quad \times \quad \text{€ } 800 = 432M$$

Comment:

one should integrate the electronic + HV price

Photodetector R&D in France

- R&D launched after NNN05 but based on on-going R&D with Photonis
- **IPN-Orsay, LAL & Photonis** together in an official GIS to develop **Smart-Photodetectors** (*ie electronic up to ADC/TDC included*): 6 engineers + 2 post-docs + Photonis engineers
- 200k€/3yrs has been asked at the new National Research Agency (ANR)

Photonis @ NNN05: 500,000 PMT -12"- 800€/u

New pump capacity needed?

Delivery over 6 years

Photonis @ NNN05

300 working days/year

1. 20" tube

$50,000/6/300 \Rightarrow 28 \text{ good tubes} \times \text{yield } 0.7 = 40$
starts/day

(1 start/pump/day) \Rightarrow **40 pumps** (€ 7M or so)

1. 12" tube

$135,000/6/300 \Rightarrow 75 \text{ good tubes} \times \text{yield } 0.7 = 110$
starts/day.

A multi-array computerised pump at Photonis handles
20 starts/day

\Rightarrow **6 pumps** (€ 2M or so)

+ Sub-conclusions

12" seems much better than 20"/17"

- *cost per useful photoelectron & total PMT cost*
- *Timing*
- *single-electron resolution (17" equal)*
- *granularity*
- *weight and handling*
- *implosion risk*
- *investments and start-up*

Photonis @ NNN05



Photonis has all the technical capability needed!

R&D cooperation: detailed & intensive talks are going on with the MEMPHYS collaboration to define a balanced programme

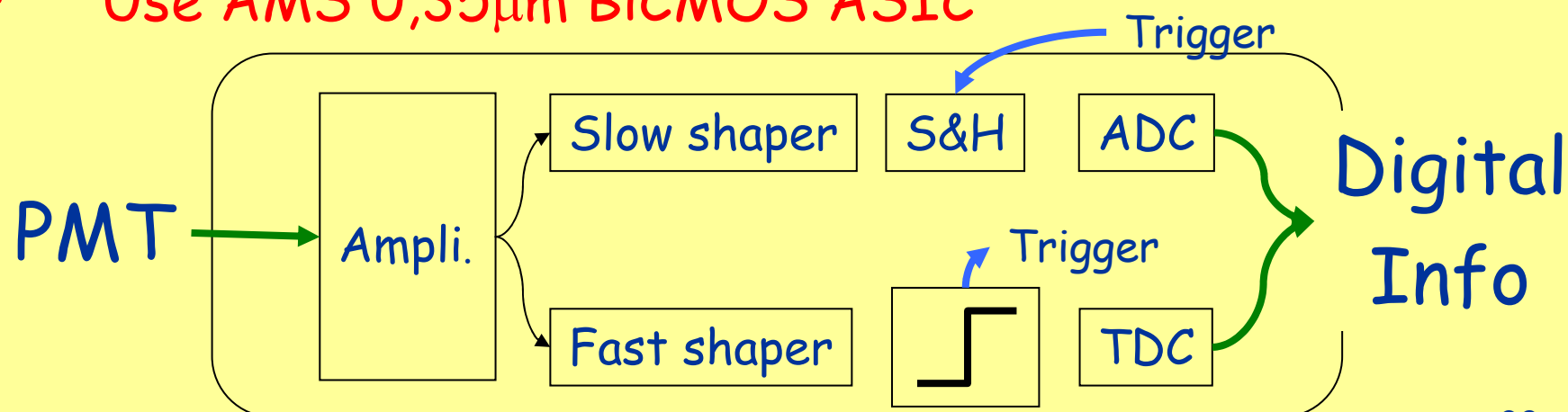
Workshop planned in the spring

Photonis @ NNN05

Electronics

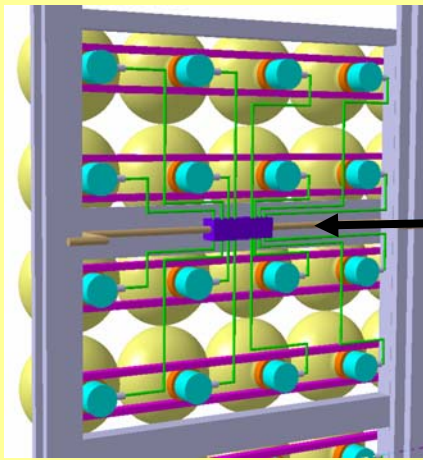
Taken in charge by **LAL**: from amplifier up to ADC/TDC based on past experience with similar state of the art front-end electronics developed for OPERA, W-Si ILC prototype, LHCb...

- Trigger @ $\frac{1}{4}$ p.e (3kHz from SK)
- TDC: 12bits 0,4ns/c
- ADC: 12bits 0,15pC/c with 1 p.e @ 20-30 adc channels.
- High speed digital readout
- **Cost reduction** thanks to high level of integration
- **Use AMS 0,35 μ m BiCMOS ASIC**



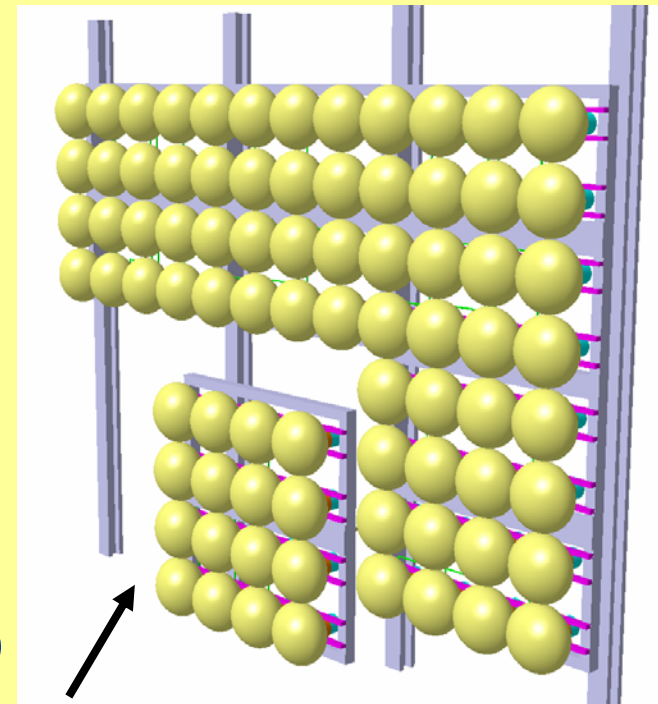
Mechanics & PMT tests

Taken in charge by **IPNO**: well experienced in photodetectors (last operation: Auger). With **PHOTONIS** tests of **PMT** 8", 9" → **12"** and Hybrid-PMT and HPD



Electronic box
water tight

Basic unit that we want to
build and test under water



IPNO

Some PMT characteristics measurements

XP1806 8"

Not the best

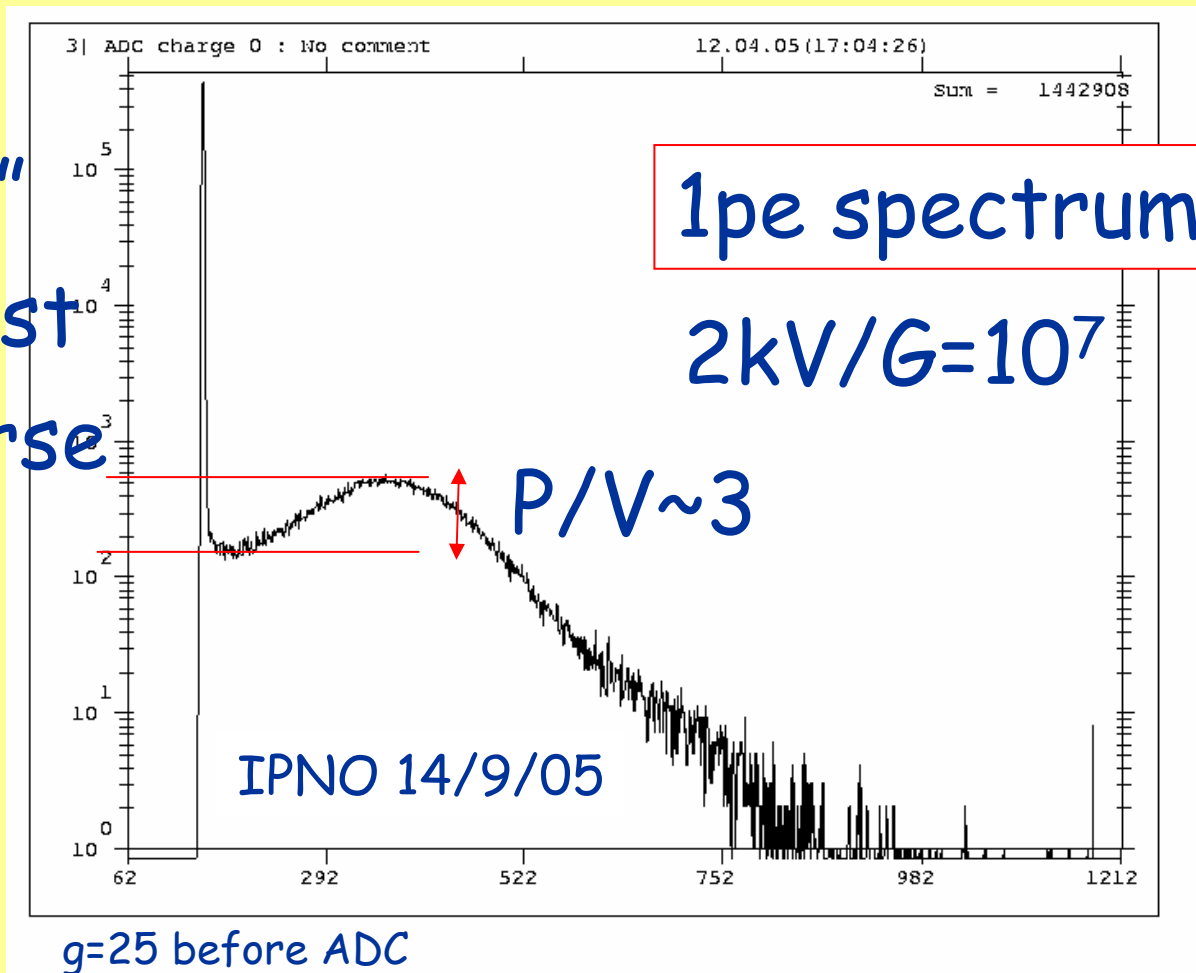
Not the worse

1pe spectrum

2kV/G=10⁷

P/V~3

IPNO 14/9/05



No diff. 5",8",10" so 12" should be identical

Summary

The MEMPHYS Mt-scale Water Cherenkov detector has a quite good accelerator neutrino program (not exposed here the Pdk and SN ν)

The R&D on photodetector is started in France and will come in 2006 with first version of "SmartDetector"

The Civil engineer pre-study for new Fréjus Lab. has been performed and seems encouraging (first costing will come soon)

Thank you