



NuMI Off-Axis ν_e Appearance

The NOvA Experiment Fermilab P929

<http://www-nova.fnal.gov>

Mark Messier
Indiana University
for the NOvA Collaboration

The NOvA (FNAL P929) Collaboration

*160 authors from 34 institutions
US - UK - Greece - Brazil - Canada*

Proposal to Build an Off-Axis Detector
to Study $\nu_\mu \rightarrow \nu_e$ Oscillations in the NuMI Beamline

NOvA
NuMI Off-Axis ν_e Appearance Experiment
(P929)

March 15, 2004

The NOvA Collaboration

Argonne, Athens, Caltech, Cambridge, Fermilab,
College de France, Harvard, Indiana, ITEP, Lebedev,
Louisiana State, Michigan State, Minnesota/Duluth,
Minnesota/Minneapolis, Munich, Stony Brook,
Northern Illinois, Ohio, Oxford, Rio de Janeiro, Rochester,
Rutherford, South Carolina, Stanford, Texas A&M,
Texas/Austin, Tufts, Virginia, Virginia Tech, Washington,
William & Mary, Wisconsin, York

I. Ambats, D. S. Ayres, J. W. Dawson, G. Drake, M. C. Goodman, J. J. Grudzinski, V. J. Guarino, T. Joffe-Minor, D. E. Reyna, R. L. Talaga, J. L. Thron, R. G. Wagner
Argonne National Laboratory, Argonne, IL

D. Drakoulakos, N. Giokaris, P. Stamoulis, S. Stiliaris, G. Tzanakos, M. Zois
University of Athens, Athens, Greece

B. Barish, J. Hanson, D. G. Michael, H. B. Newman, C. W. Peck, C. Smith, J. Trevor, H. Zheng
California Institute of Technology, Pasadena, CA

D. B. Cline, H. Wang, Y. Chen, K. Lee, Y. Seo, X. Yang
University of California, Los Angeles, CA

M. Thomson
University of Cambridge, Cambridge, United Kingdom

J. Beacom, R. Bernstein, G. Bock, S. Brice, L. Camilleri*, S. Childress, B. Choudhary, J. Cooper**, R. Hatcher, D. Harris, J. Hylen, H. Jostlein, D. Koolbeck, J. Kilmer, P. Lucas, V. Makeev, A. Marchionni, O. Mena, S. Mishra, R. Plunkett, S. Parke, S. Pordes, R. Rameika, R. Ray, R. Schmitt, P. Shanahan, P. Spentzouris, R. Wands, R. Yarema
Fermi National Accelerator Laboratory, Batavia, IL

T. Patzak, R. Piteira
College de France, Paris, France

G. J. Feldman**, N. Felt, A. Lebedev, J. Oliver, M. Sanchez, S.-M. Seun
Harvard University, Cambridge, MA

C. Bower, M. Gebhard, M. D. Messier, S. Mufson, J. Musser, B. J. Rebel, J. Urheim
Indiana University, Bloomington, IN

I. Trostин
Institute For Theoretical And Experimental Physics, Moscow, Russia

V. Ryabov, A. Y. Terekhov
P. N. Lebedev Physical Institute, Moscow, Russia

W. Metcalf
Louisiana State University, Baton Rouge, LA

C. Bromberg, J. Huston, R. Miller, R. Richards
Michigan State University, East Lansing, MI

A. Habig
University of Minnesota, Duluth, MN

T. Chase, K. Heller, P. Litchfield, M. Marshak, W. Miller, L. Mualem, E. Peterson, D. Petyt, K. Ruddick, R. Rusack
University of Minnesota, Minneapolis, MN

P. Huber, M. Lindner, W. Winter
Technische Universität München, Munich, Germany

R. Shrock
State University of New York, Stony Brook, NY

C. Albright
Northern Illinois University, DeKalb, IL

C. R. Brune, D. S. Carman, K. H. Hicks, S. M. Grimes, A. K. Opper
Ohio University, Athens, OH

G. Barr, J. H. Cobb, K. Grzelak, N. Tagg
University of Oxford, Oxford, United Kingdom

H. Nunokawa
Pontifícia Universidade Católica do Rio de Janeiro, Rio de Janeiro, Brazil

A. Bodek, H. Budd, S. Manly, K. McFarland, W. Sakumoto
University of Rochester, Rochester, NY

T. Durkin, R. Halsall, T. Nicholls, G. F. Pearce, A. Weber
Rutherford Appleton Laboratory, Chilton, Didcot, United Kingdom

T. Bergfeld, K. Bhskaran, A. Godley, S. R. Mishra, C. Rosenfeld, K. Wu
University of South Carolina, Columbia, SC

S. Avvakumov, G. Irwin, S. Murgia, S. Wojciecki, T. Yang
Stanford University, Stanford, CA

E. Tetteh-Lartey, M. Watabe, R. Webb
Texas A&M University, College Station, TX

J. Klein, S. Kopp, K. Lang, M. Proga
University of Texas, Austin, TX

H.R. Gallagher, T. Kafka, W.A. Mann, J. Schneps, A. Sousa
Tufts University, Medford, MA

C. Dukes, L. Lu, K. Nelson, A. Norman
University of Virginia, Charlottesville, VA

F. F. Chen, K. Creehan, N. K. Morgan, L. Piilonen, R. H. Sturges, R. B. Vogelaar
Virginia Polytechnic Institute and State University, Blacksburg, VA

J. Rothberg, T. Zhao
University of Washington, Seattle, WA

J. K. Nelson
The College of William and Mary, Williamsburg, VA

A.R. Erwin, C. Velissaris
University of Wisconsin, Madison, WI

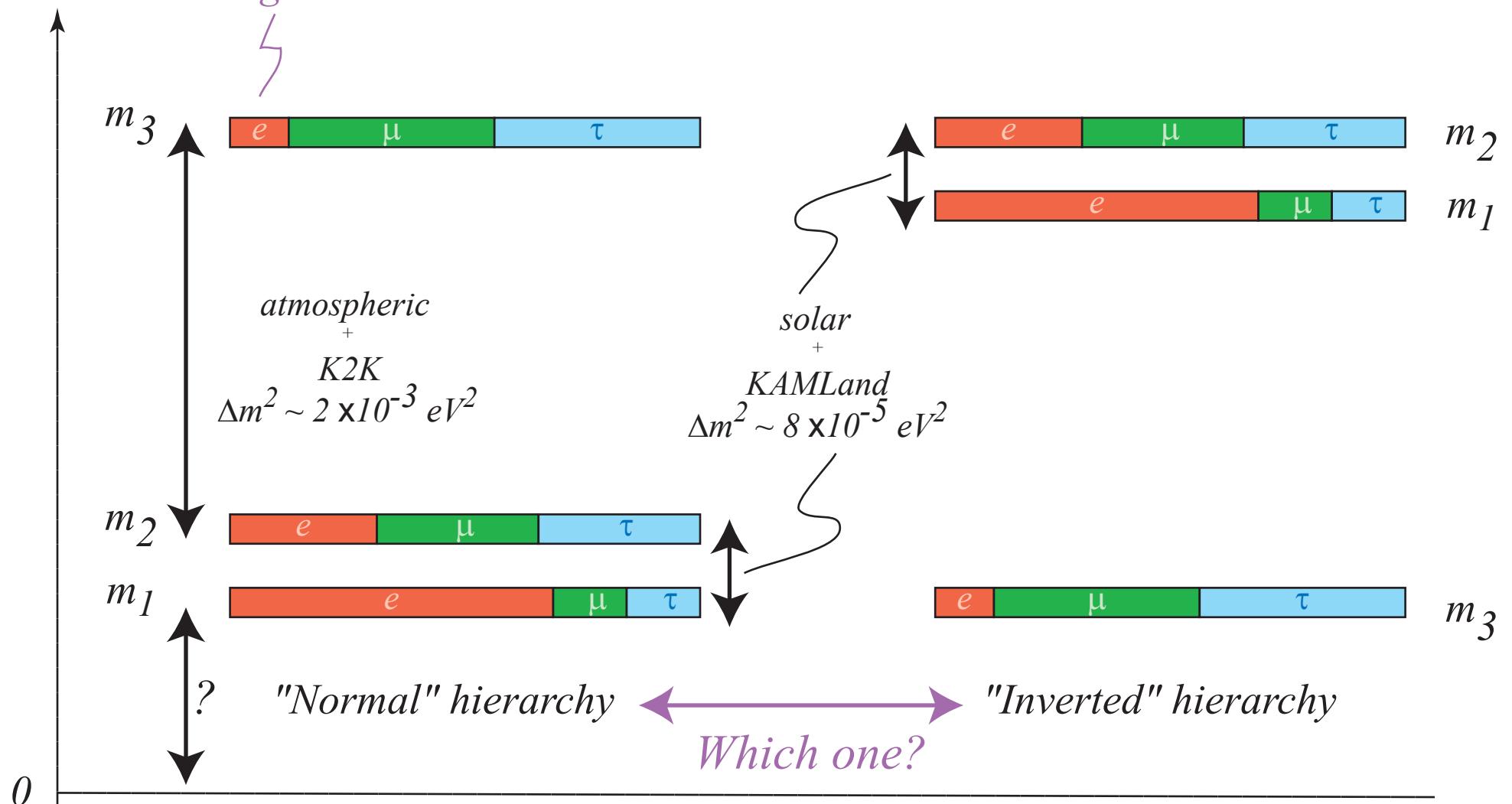
S. Menary
York University, Toronto, Ontario, Canada

* On leave from CERN
** Temporary Co-Spokespersons

*Gary Feldman (Harvard) and John Cooper (FNAL)
co-spokespersons*

What we know, what we don't know...

How large is this?



Would like to have more precise knowledge of mixing
Do ν_e 's participate in oscillations at atmospheric scale?

$Is \Delta m_{23}^2 > 0$ or < 0 ?
Is CP violated?

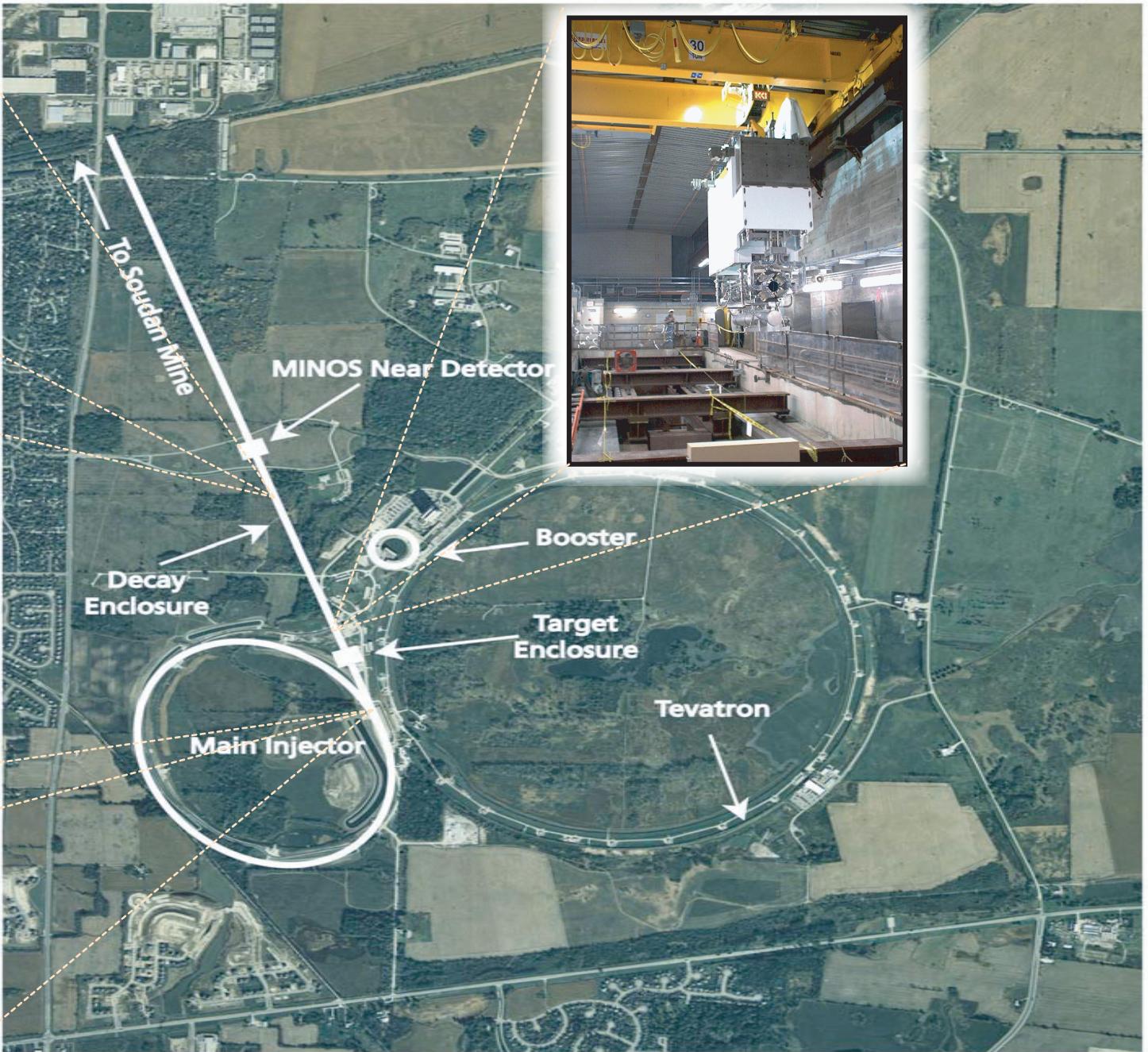
Goals of the NOvA Experiment

- [1] Sensitivity to $\sin^2(2\theta_{13})$ factor 10 below MINOS sensitivity, ie. down to ~ 0.01
- [2] $\sin^2(2\theta_{23})$ measurement to 2% accuracy
- [3] Resolve or contribute to resolution of mass hierarchy via matter effect (*unique contribution*)
- [4] Begin to study CP violation in neutrino sector

How will NO_A meet its goals?

- [1] Reduce NC background to ν_e appearance search by going off the NuMI beam axis for a narrow band beam
- [2] Increase detector mass ∼factor of 9 over MINOS while reducing cost/kilo-ton by factor of 3
- [3] Sandwich detector sampling at 1/3 X₀ (*compare to 1.5 in MINOS*)
 - electrons showers seen as "fuzzy" tracks w/ 1-4 hits/plane/view
 - allow separation of γ's from π⁰ decays
 - good energy resolution to focus on signal energy region
- [4] Choose long baseline to enhance matter effects

Neutrinos At the Main Injector (NuMI)



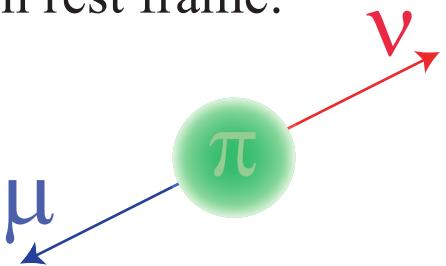
*NuMI beam
set to
commission
start of 2005*



FERMILAB #98-765D

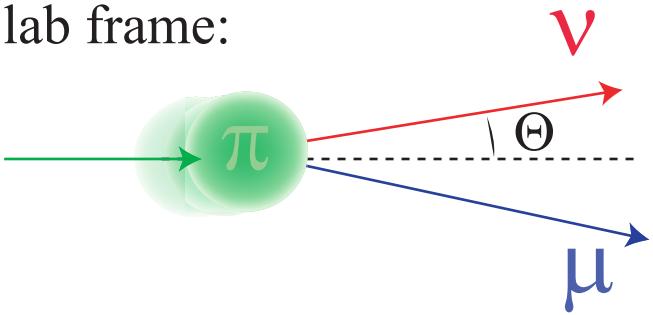
Off-Axis Neutrino Beams

In pion rest frame:

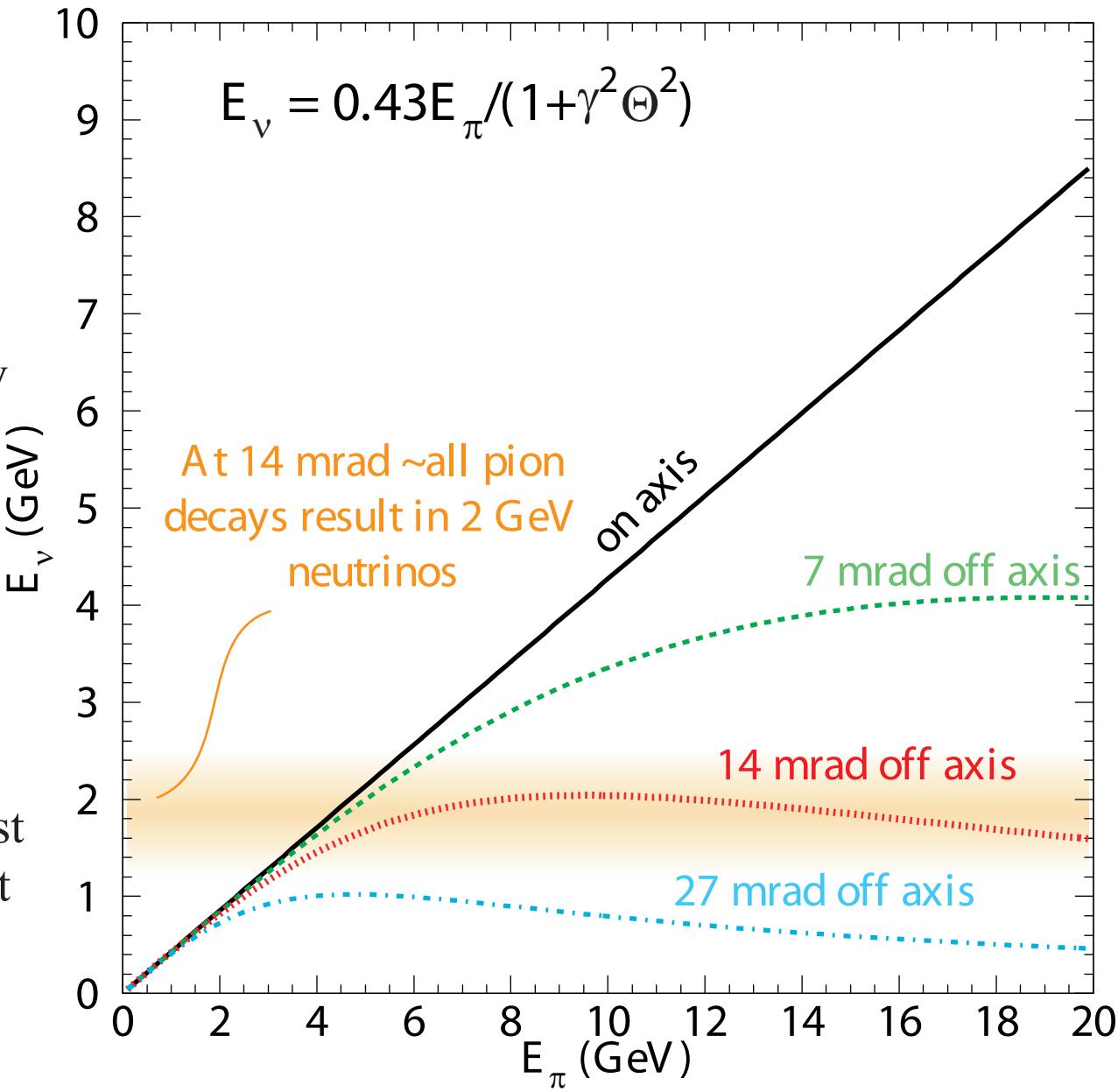


Neutrino and muon energy completely determined by energy conservation

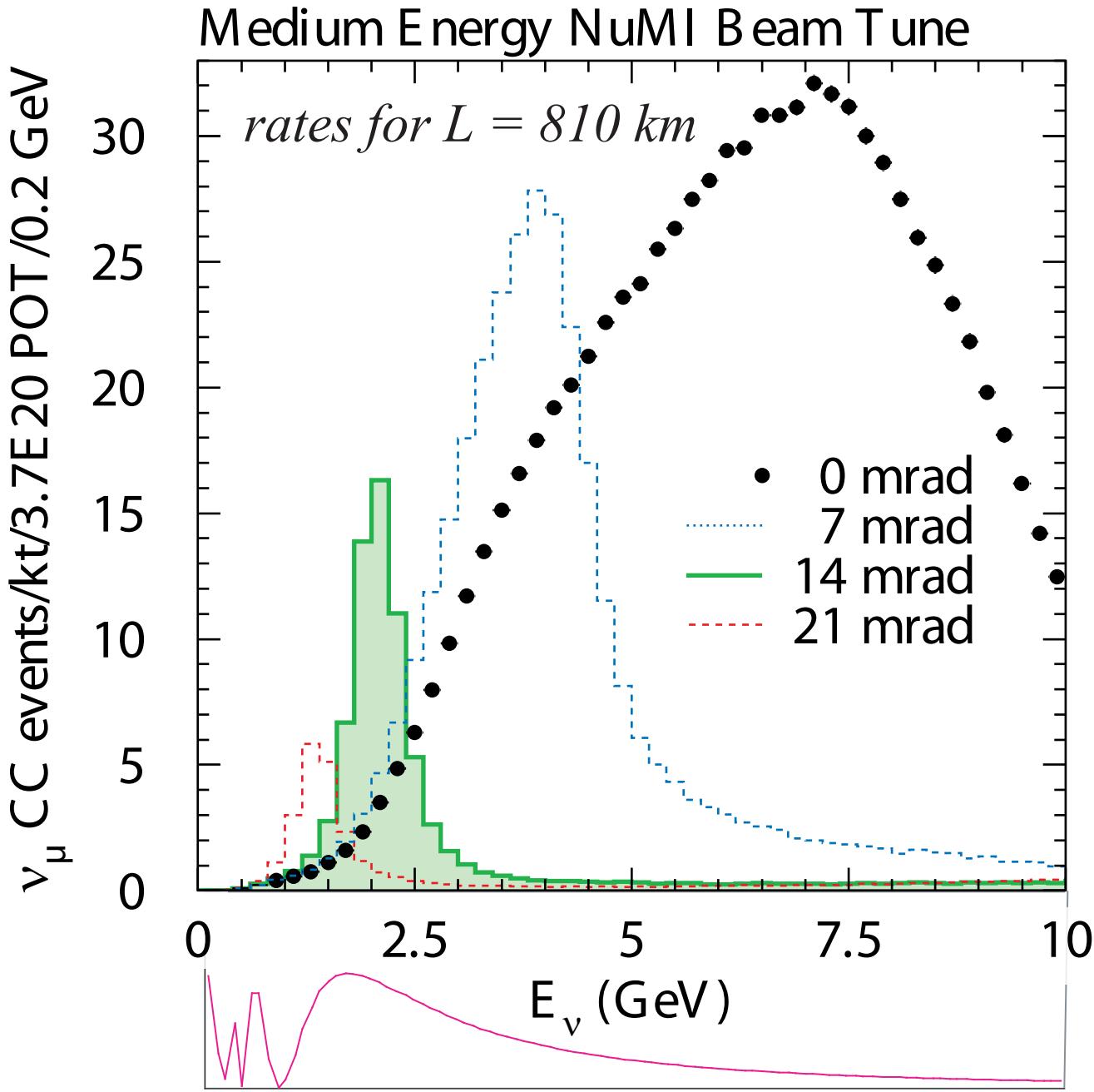
In lab frame:



Neutrino energy depends on boost and angle between neutrino boost direction



Neutrino spectra off the NuMI axis



Using NuMI ME tune
beam at 14 mrad
peaks at $\sim 2 \text{ GeV}$
and has $\sim 20\%$ width

High energy flux
suppressed

Sits just above oscillation
maximum

$$\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$$

Event rates off NuMI beam axis

Event rates for:

$$L = 810 \text{ km}, T = 12 \text{ km}$$

$$\Delta m^2_{23} = 2.5 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta_{23} = 1, \sin^2 2\theta_{13} = 0.01$$

Sets goals for detector

Most ν_μ oscillate away.

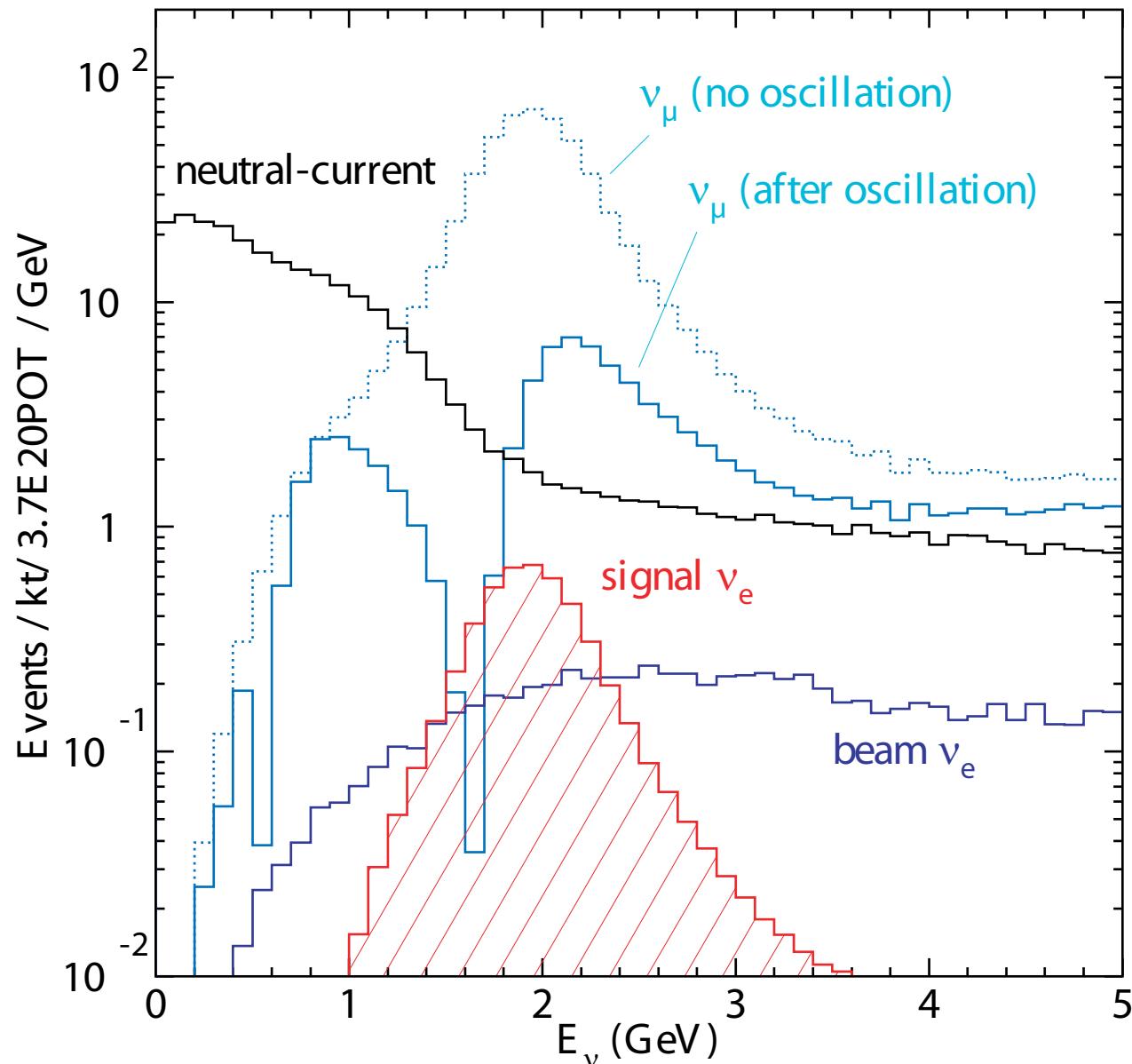
- ☞ need only 50:1 rejection of ν_μ CC
(Easy!)

Need $\sim > 100:1$ NC rejection

- ☞ fine grained,
low density detector

Good energy resolution

- ☞ reject beam ν_e



Baseline detector design

50.7 kilo-ton total mass

43.8 kilo-ton passive

6.9 kilo-ton active (14%)

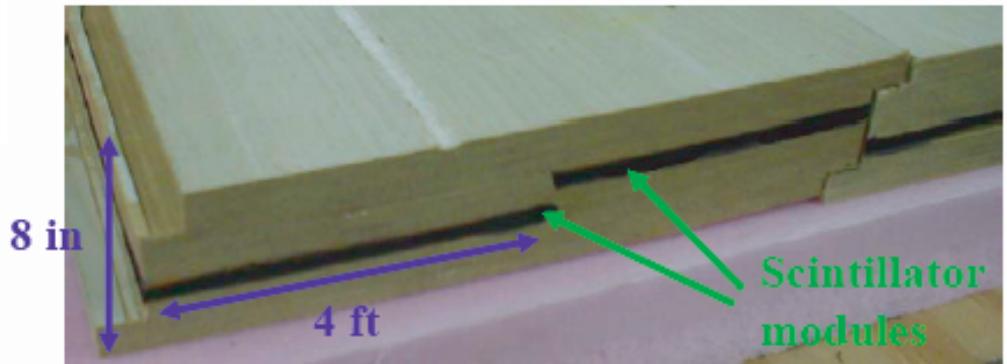
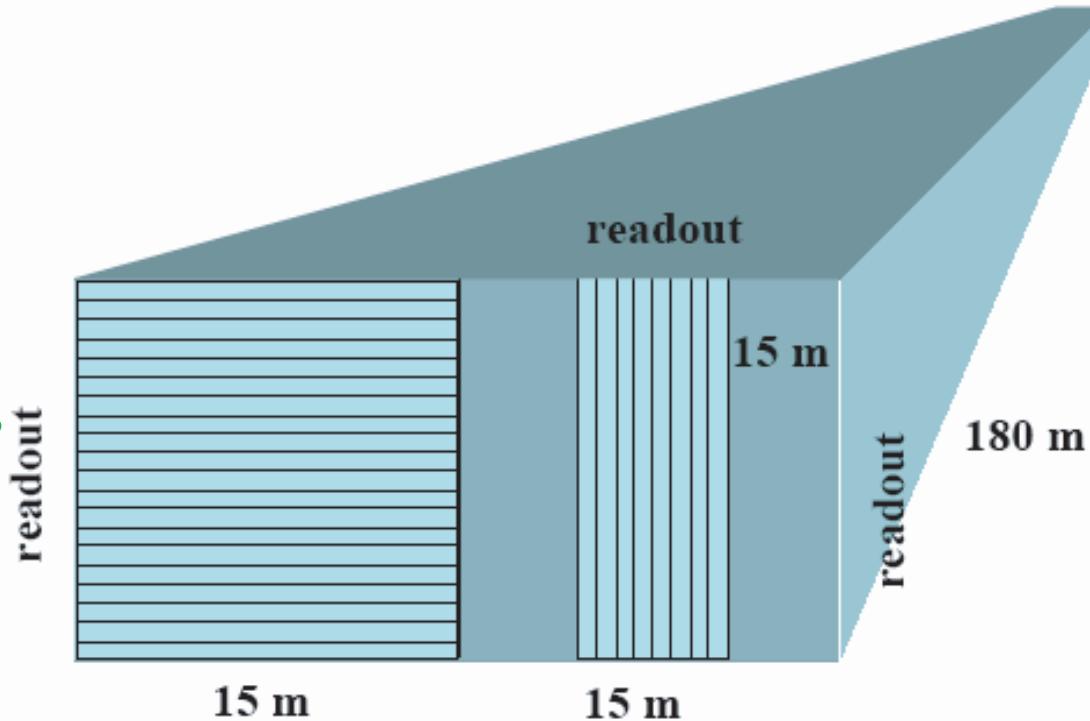
Liquid scintillator contained in
1.2 m x 3 cm x 1.4 m PVC extrusions
30 cells per extrusion
24 extrusions per plane
750 planes (alternating x/y readout)
= 18,000 extrusions
= 540,000 channels

Looped WLS fiber to APD readout

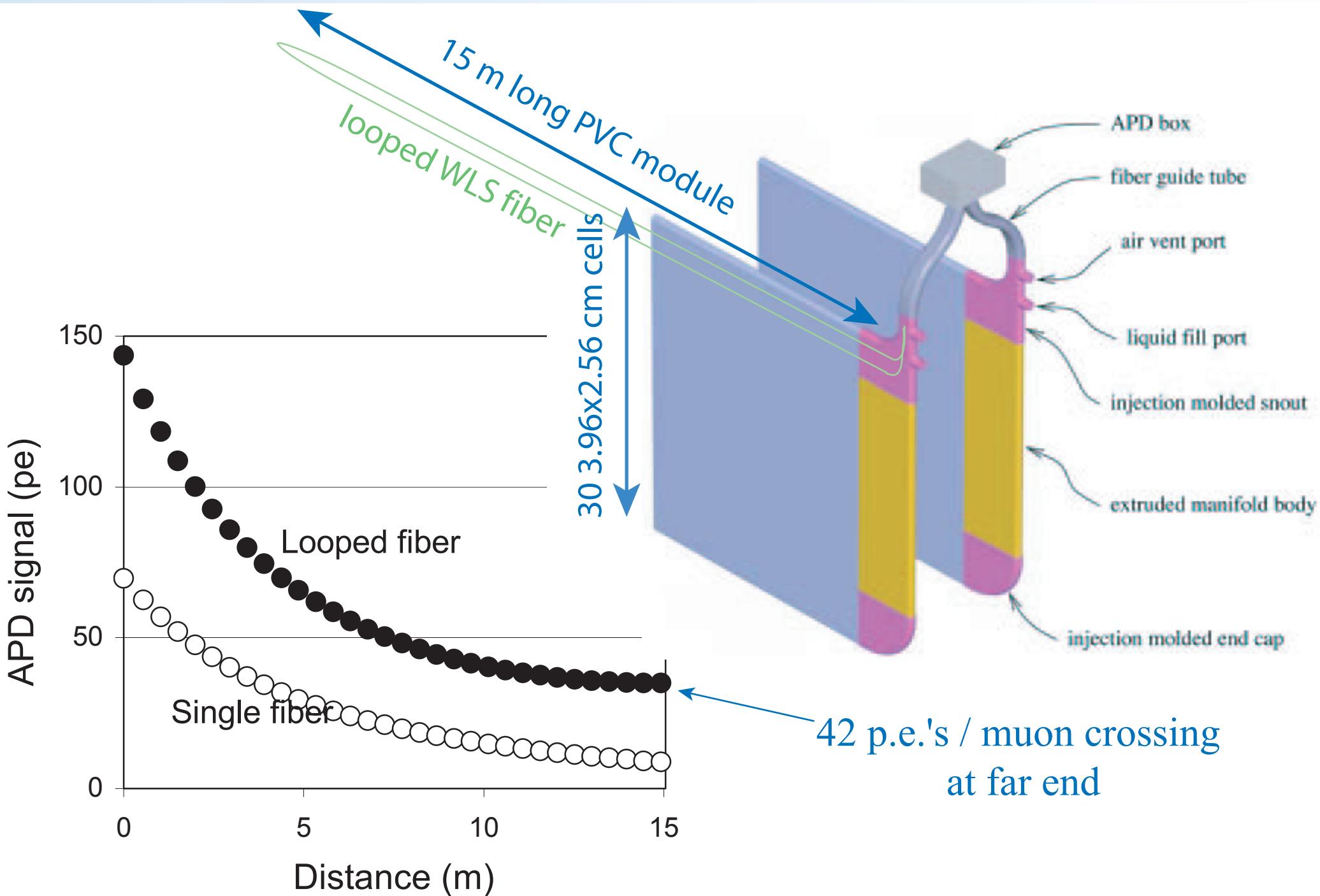
Absorber:

20 cm particle board/plane

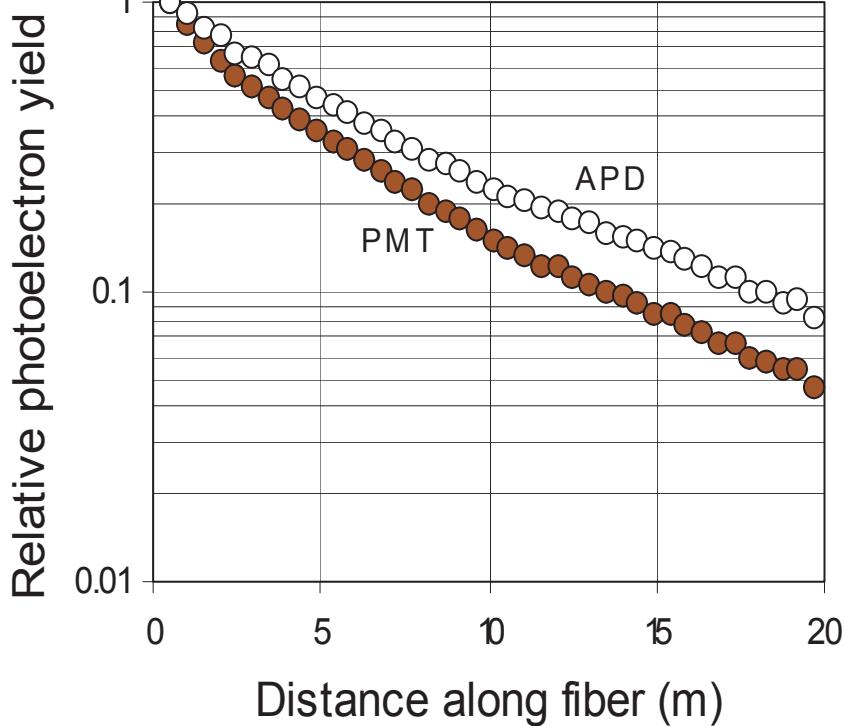
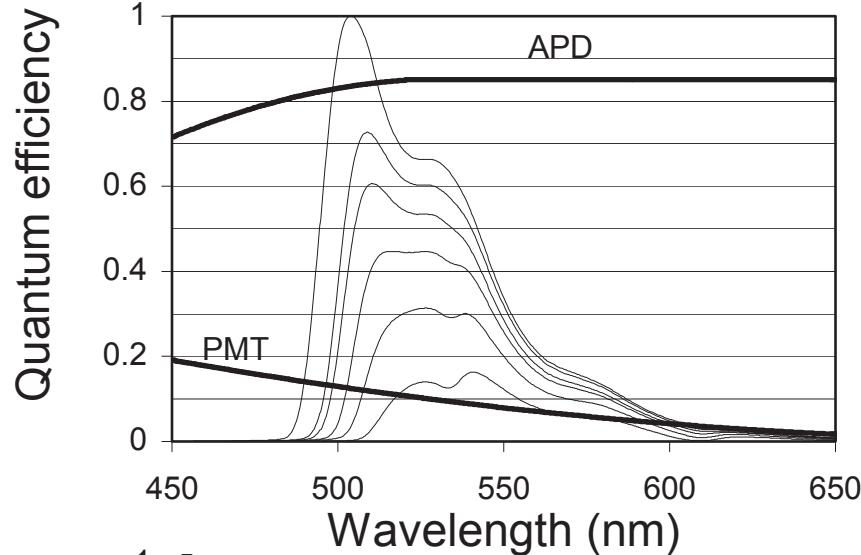
1/3 X_0 per plane



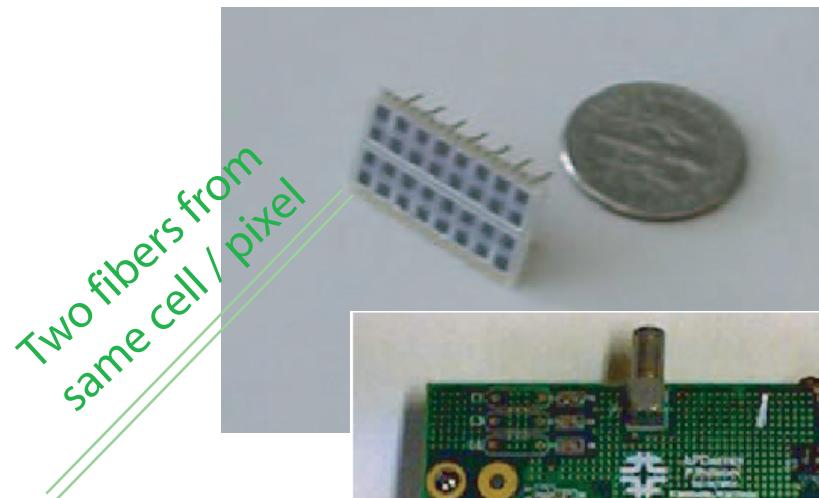
Module Design



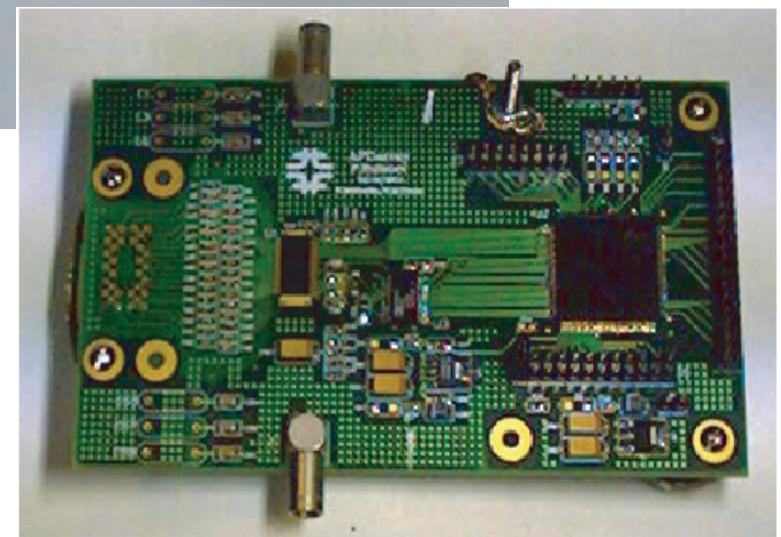
Avalanche Photo Diode (APD) Readout



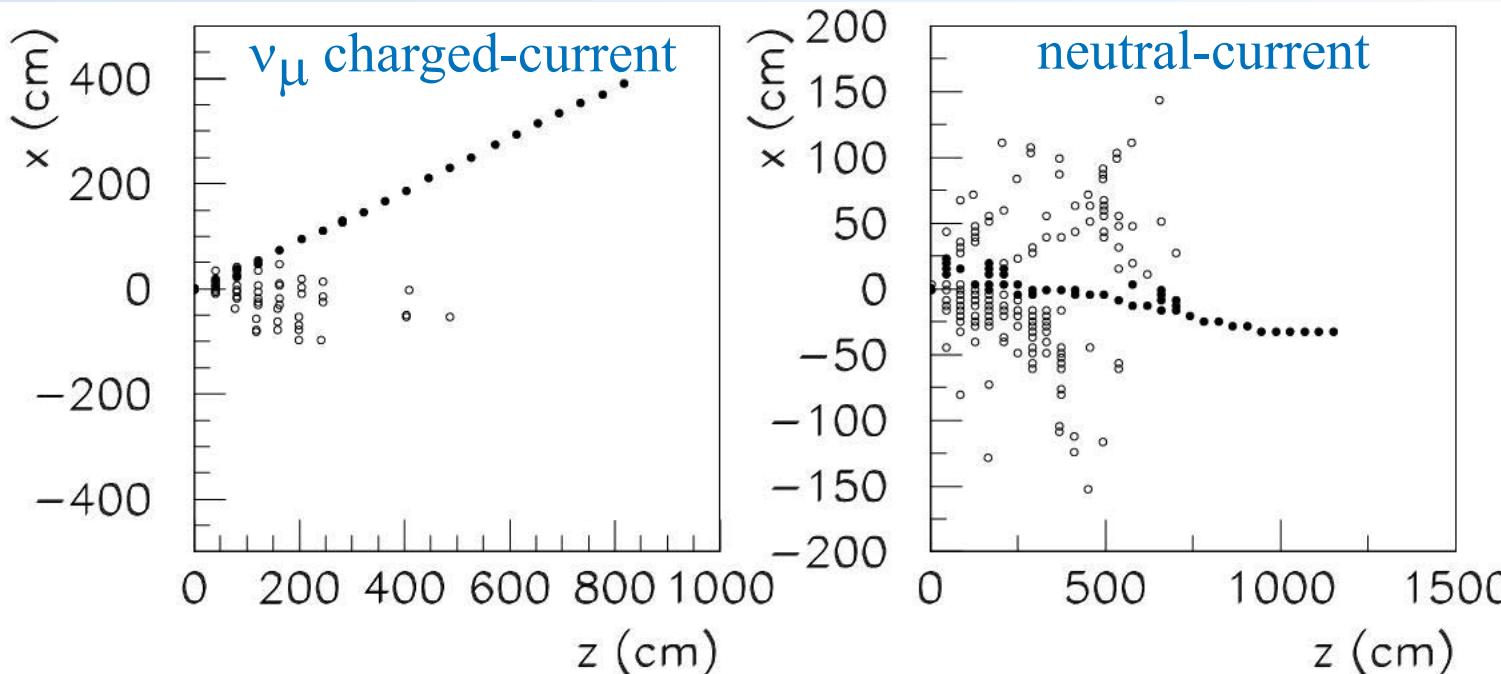
- 85% QE
- x100 gain (operate at -15° C)
- Long wavelength sensitivity is advantage for long fibers (x2 at far end)
- Low cost: \$12/chan FEE+Trigger+DAQ base cost



Prototype
readout board



Baseline detector event displays



(1 of 2 detector views shown)

Event selection

- cuts on #planes and total pulse height
- at least 3 hits on a track
- 75% of hits in track
- hits/plane > 1.5
- $\cos(\theta_{\text{beam}}) > 0.8$
- likelihood analysis of event shape variables

18% signal efficiency

v_μ CC rejection 1600:1

NC rejection 600:1

beam v_e rejection 12:1

~160 signal events

~ 44 background events

FOM=24

$\sin^2 2\theta_{23} = 1$, $\sin^2 2\theta_{13} = 0.1$, $\Delta m^2 = 2.5 \times 10^{-3}$ eV²

50kt x 5 years x 4e20 POT/year

Life on the surface: Active shield

Low duty cycle:

10 us spill every 2 s (1/200,000)

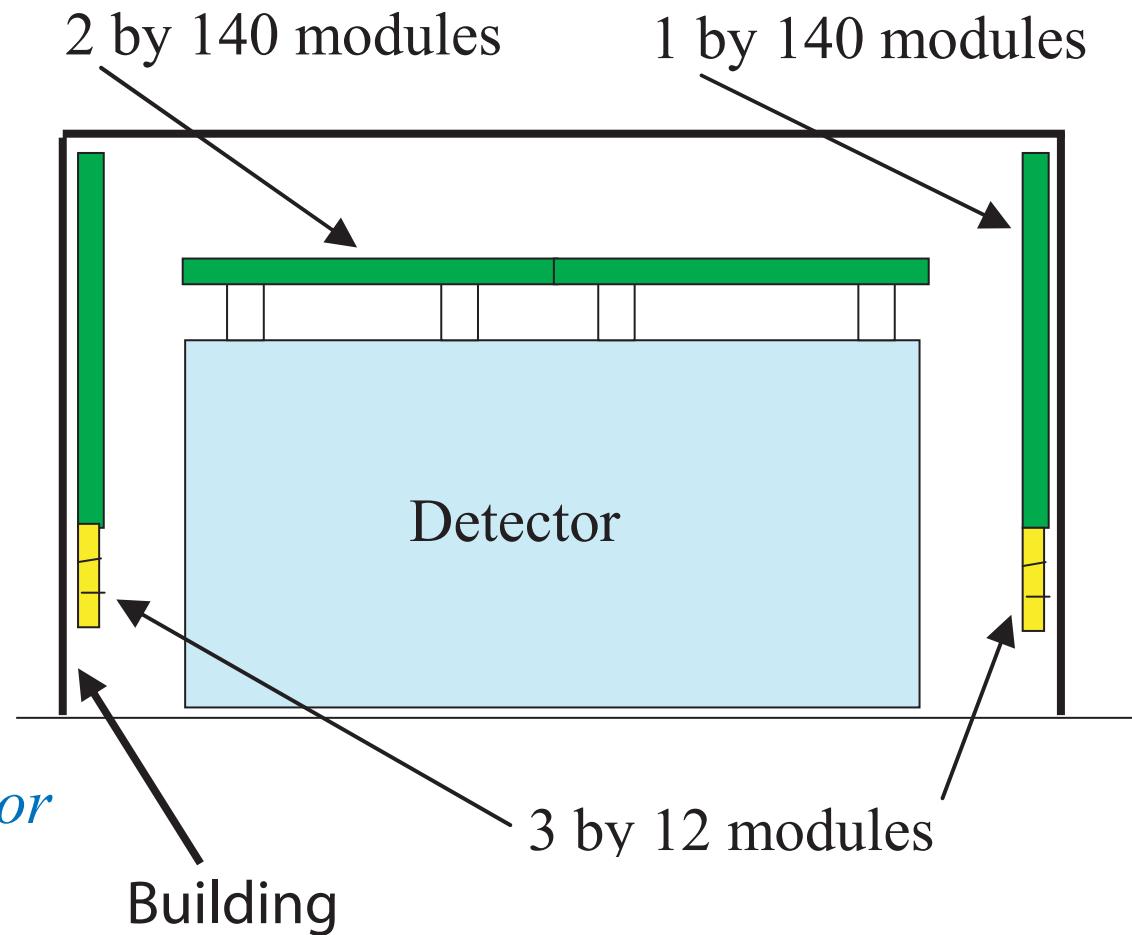
"live" only 100 s per year!

1.3×10^8 m/year total in spill

$0.65 \mu/500$ ns DAQ gate over
5000 m² detector area

No passive shield required

Plan to test with prototype detector



We are planning an active veto shield in baseline design

In totally active detector option no shield is required

Totally Active Scintillator Detector (TASD)

25 kilo-ton total mass

 4 kilo-ton passive

 21 kilo-ton active (85%)

Liquid scintillator contained in
1.28 m x 4.9 cm x 17.5 m PVC extrusions

32 cells per extrusion

24 extrusions per plane

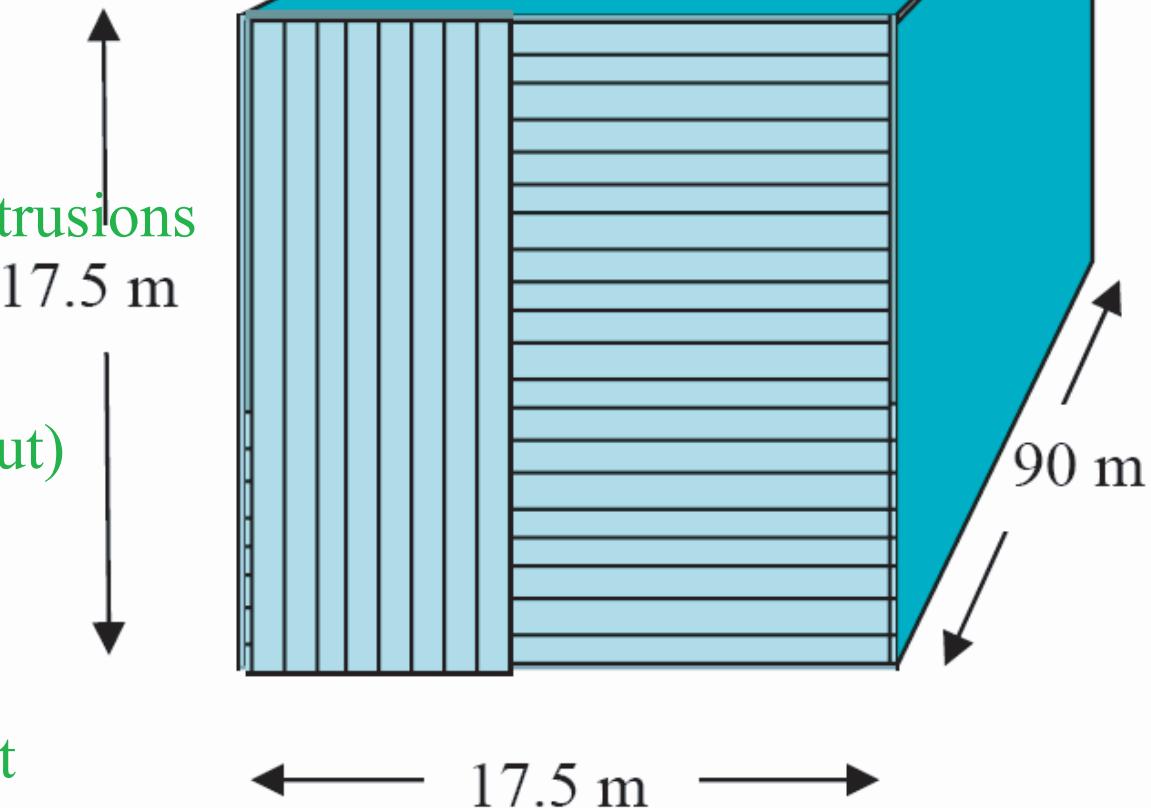
1845 planes (alternating x/y readout)

= 25,830 extrusions

= 826,560 channels

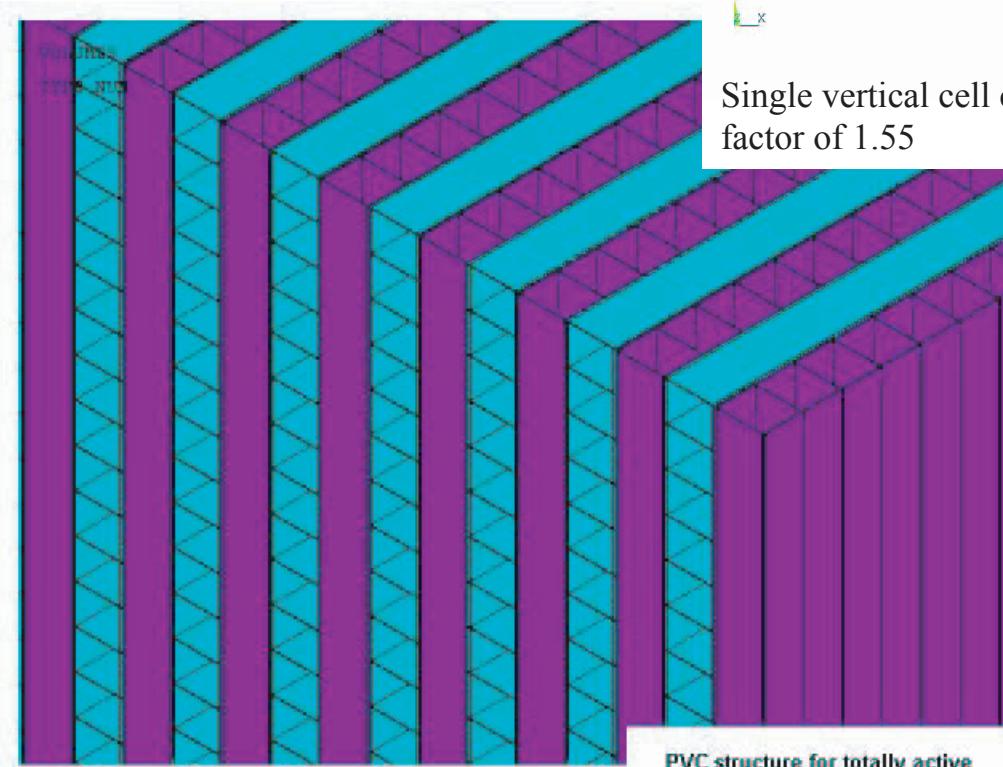
Looped WLS fiber to APD readout

No absorber

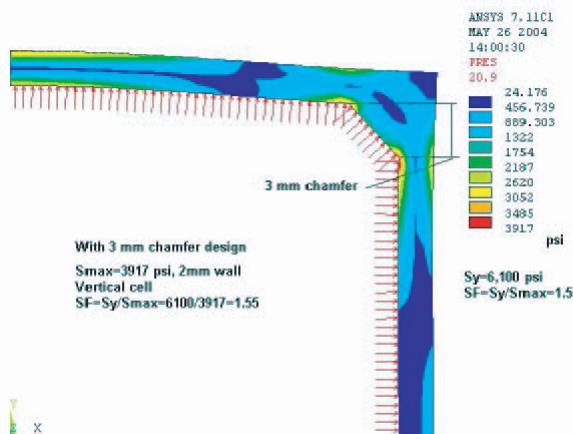


Engineering for TASD

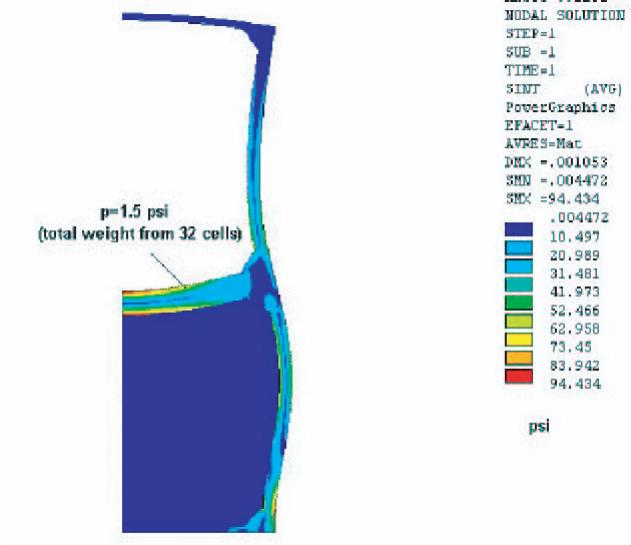
Engineering of TASD is somewhat more challenging than baseline detector



Studies of "free standing" cells



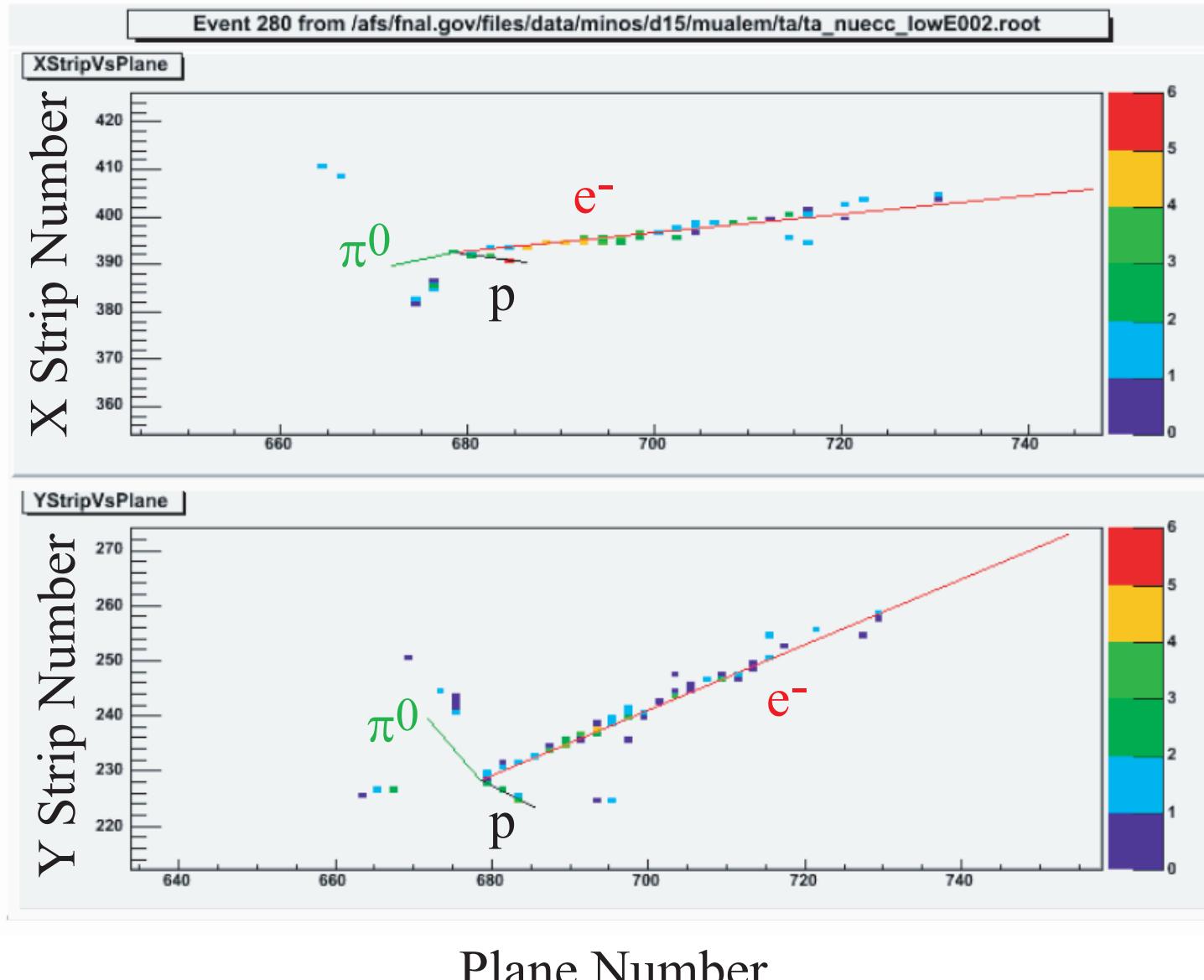
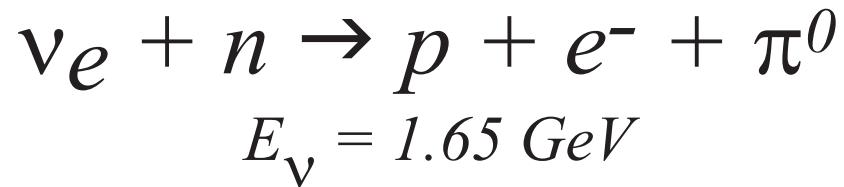
Single vertical cell can hold load with safety factor of 1.55



Horizontal cells can hold load with safety factor of 3.3

Forming laminate of adjacent x/y planes (using adhesive, ultra-sonic welding, PVC welding, ...) gives safety factor of 20

Typical TASD Event



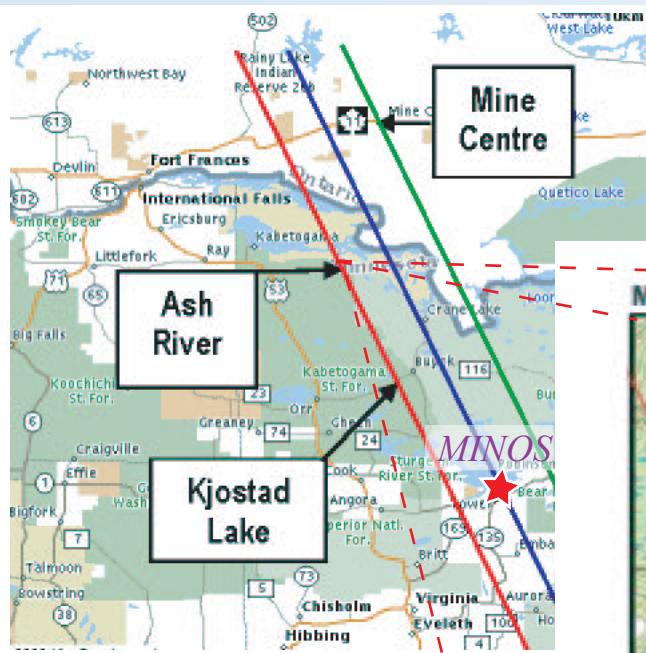
Signal efficiency
32% (18% baseline)

signal/background
7.7 (4.6 baseline)

signal/sqrt(bg.)
26 (24.5 baseline)

color prop. to pulse height

Possible sites for detector



Ash River (*baseline site*)

L = 810-812 km, T = 12-15 km

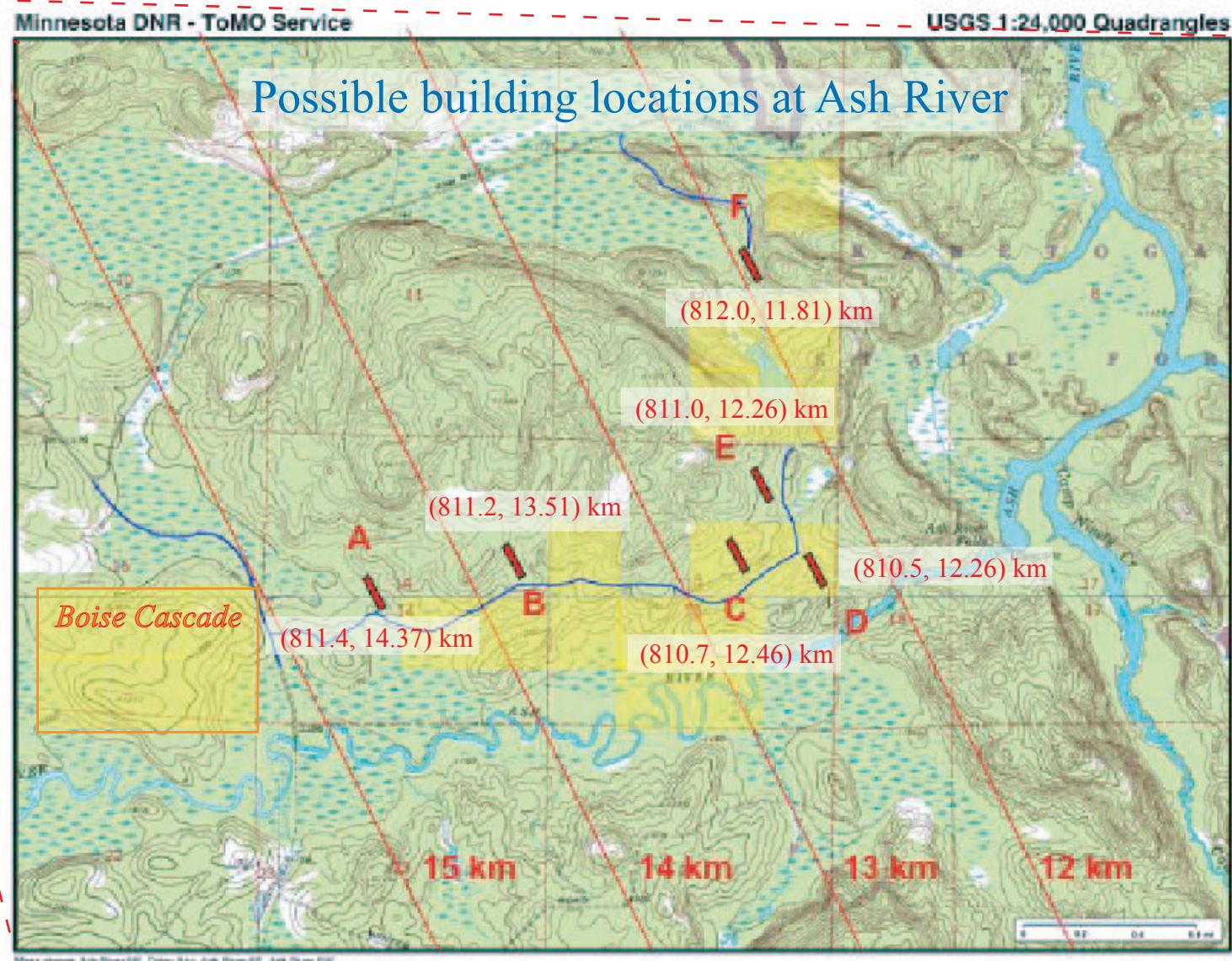
Kjostad Lake

L = 775 km, T = 11-15 km

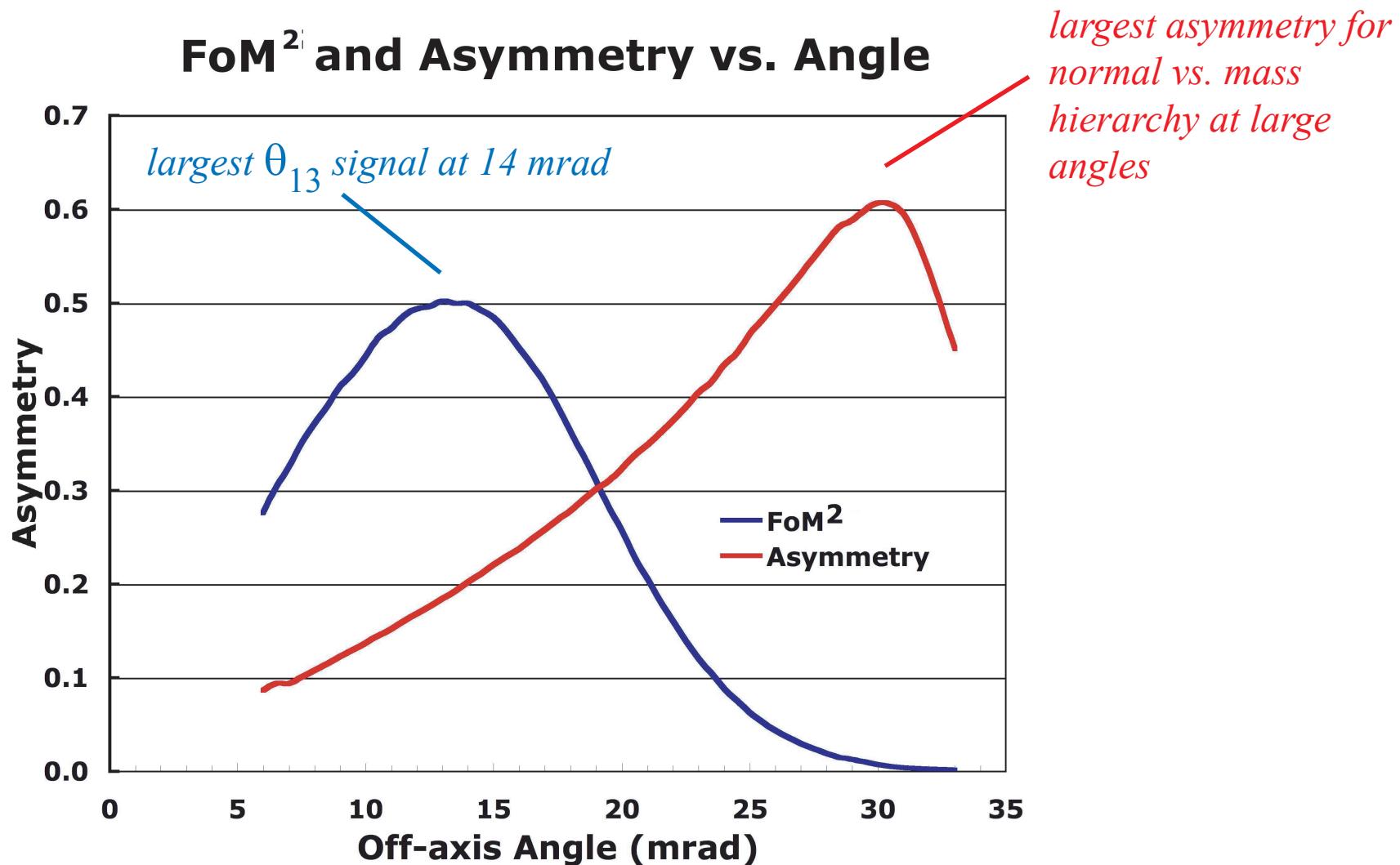
Mine Center

L = 845 km, T = 5 km

Several possible sites along US-Canada border
All have year-round truck access, power, and nearby towns

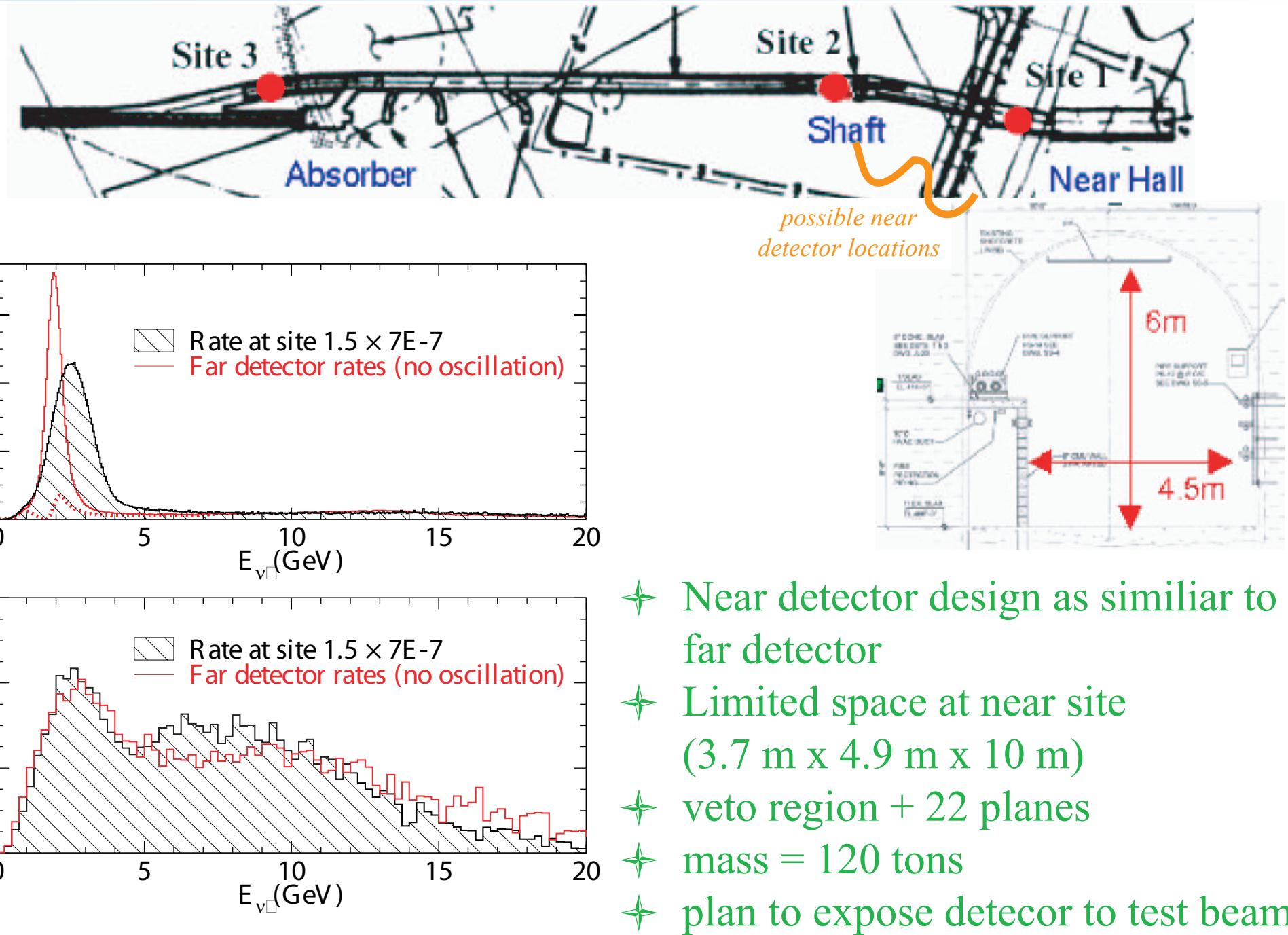


Optimizing location for physics



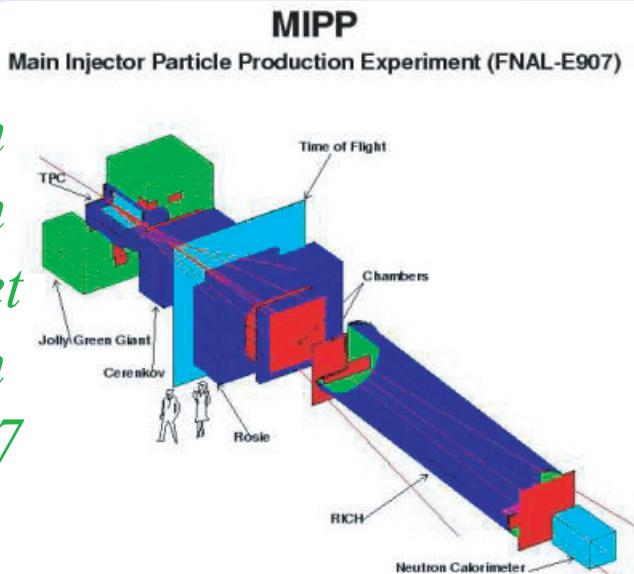
To optimize for θ_{13} sensitivity run slightly above oscillation maximum
To optimize for mass hierarchy run closer to oscillations maximum

Near Detector



Expect NuMI beam to be very well studied

Hadron production on NuMI target from FNAL E907

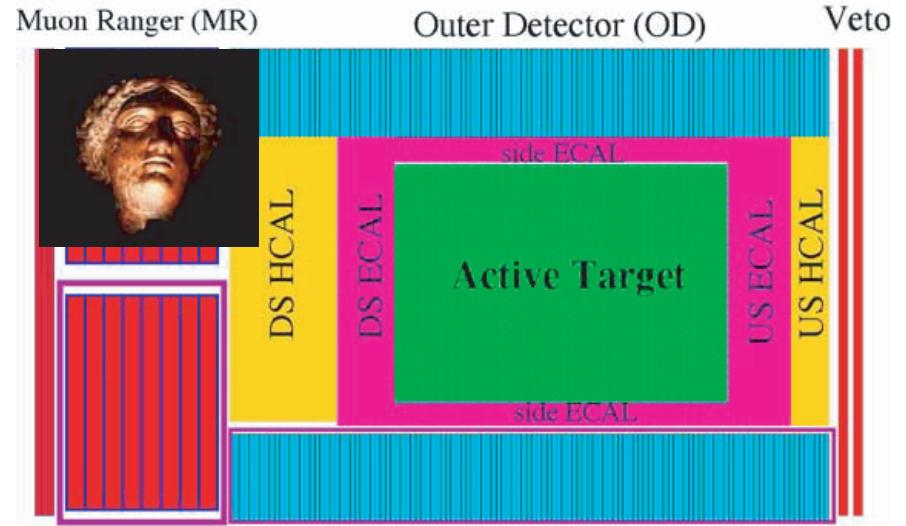


On axis oscillated ν_μ rates from MINOS far detector

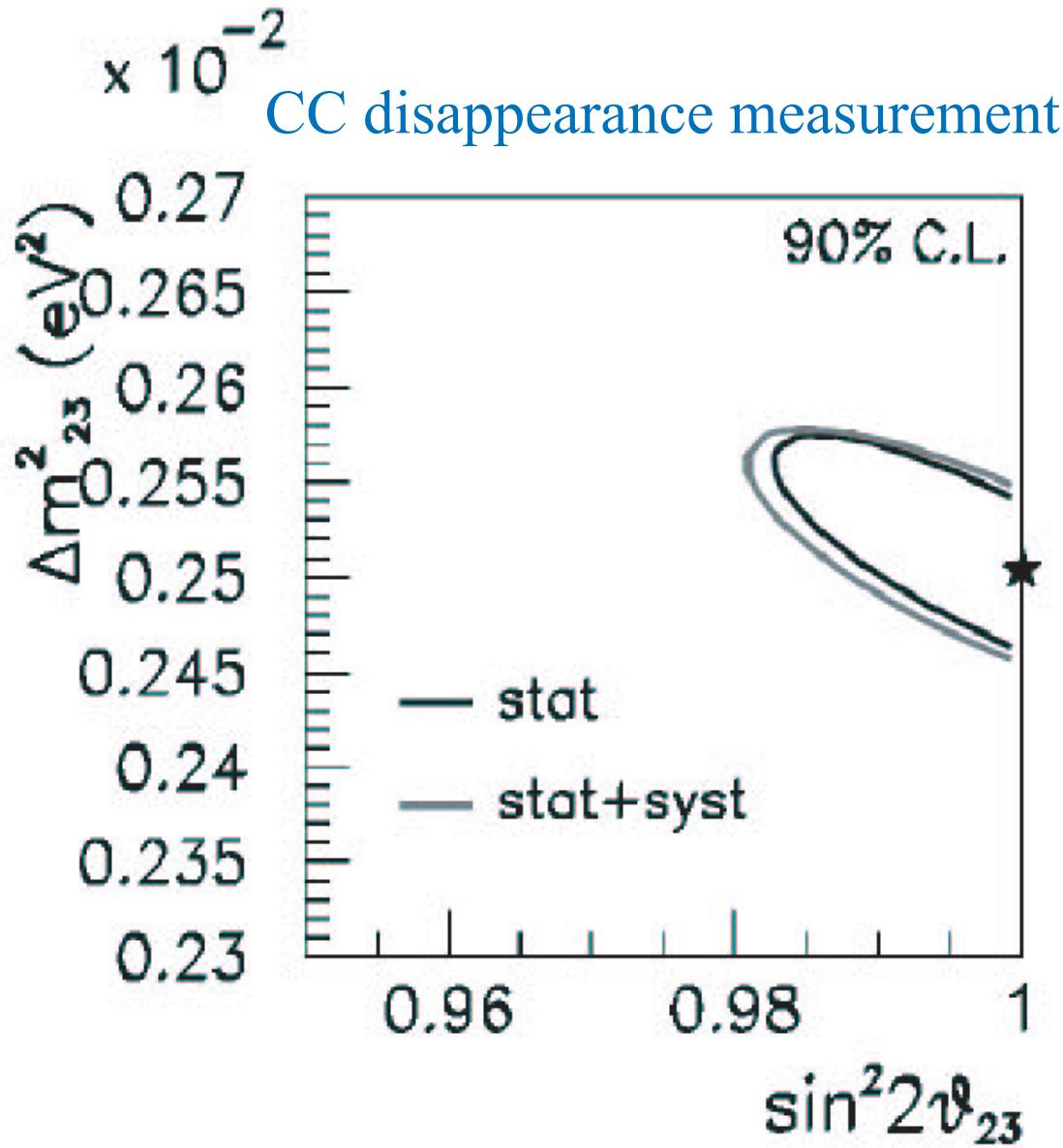


On axis near ν_μ rates from MINOS near detector

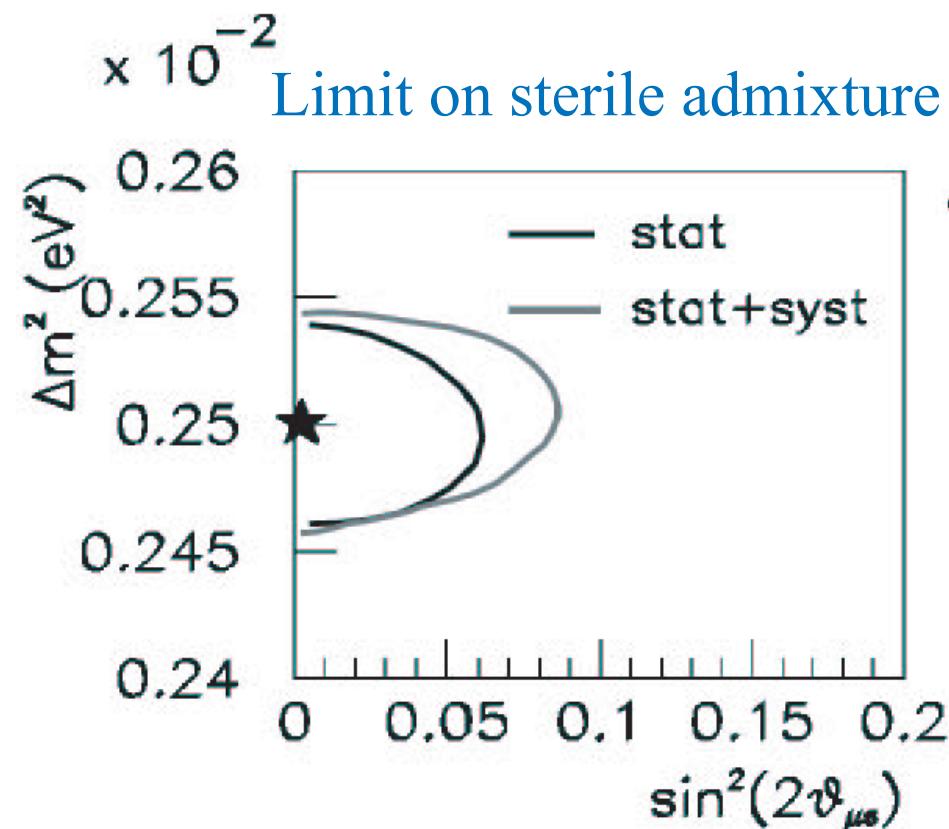
ν_μ and ν_e flux and ν -cross sections from MINERvA



Improvements on other MINOS measurements

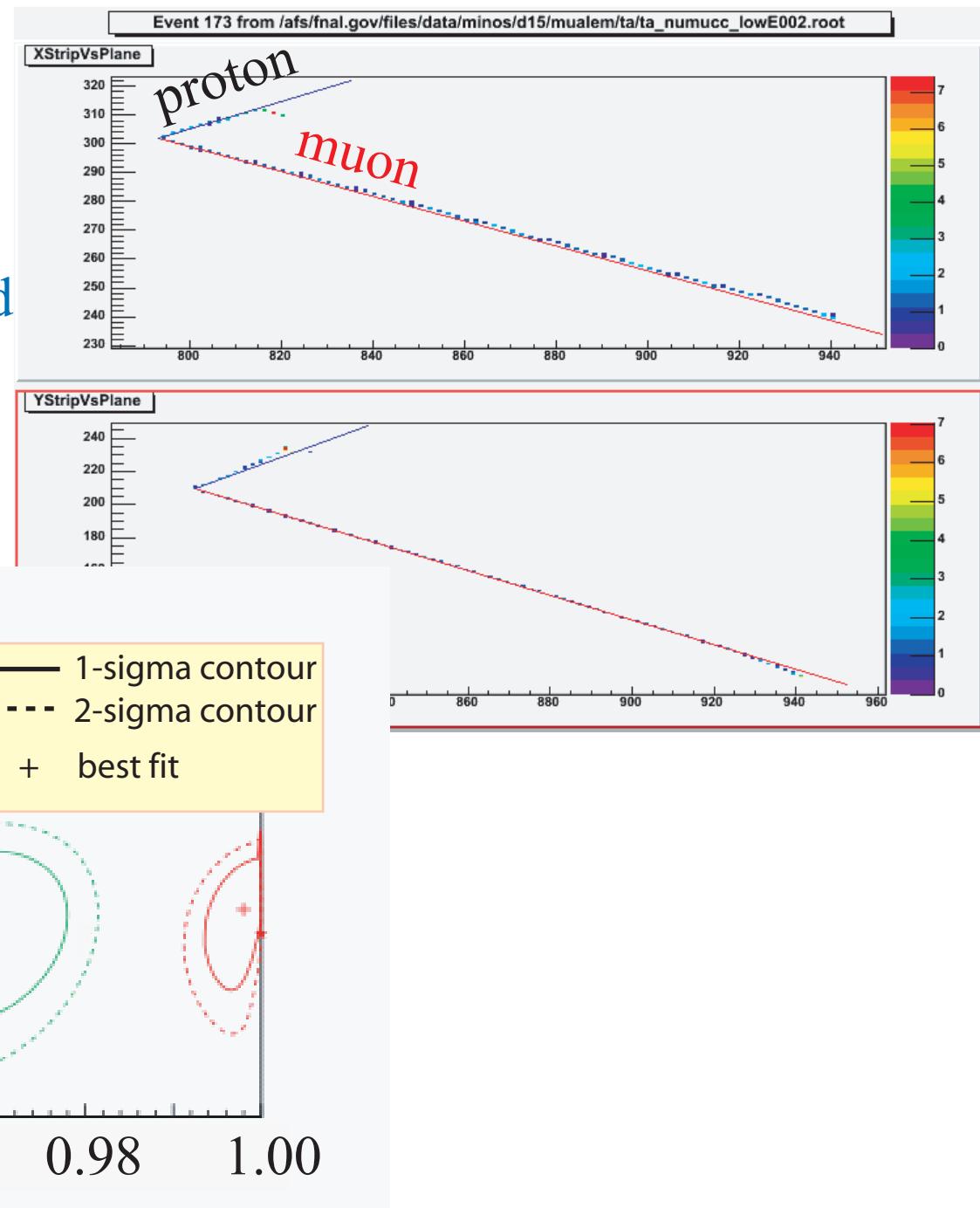


Baseline detector
design

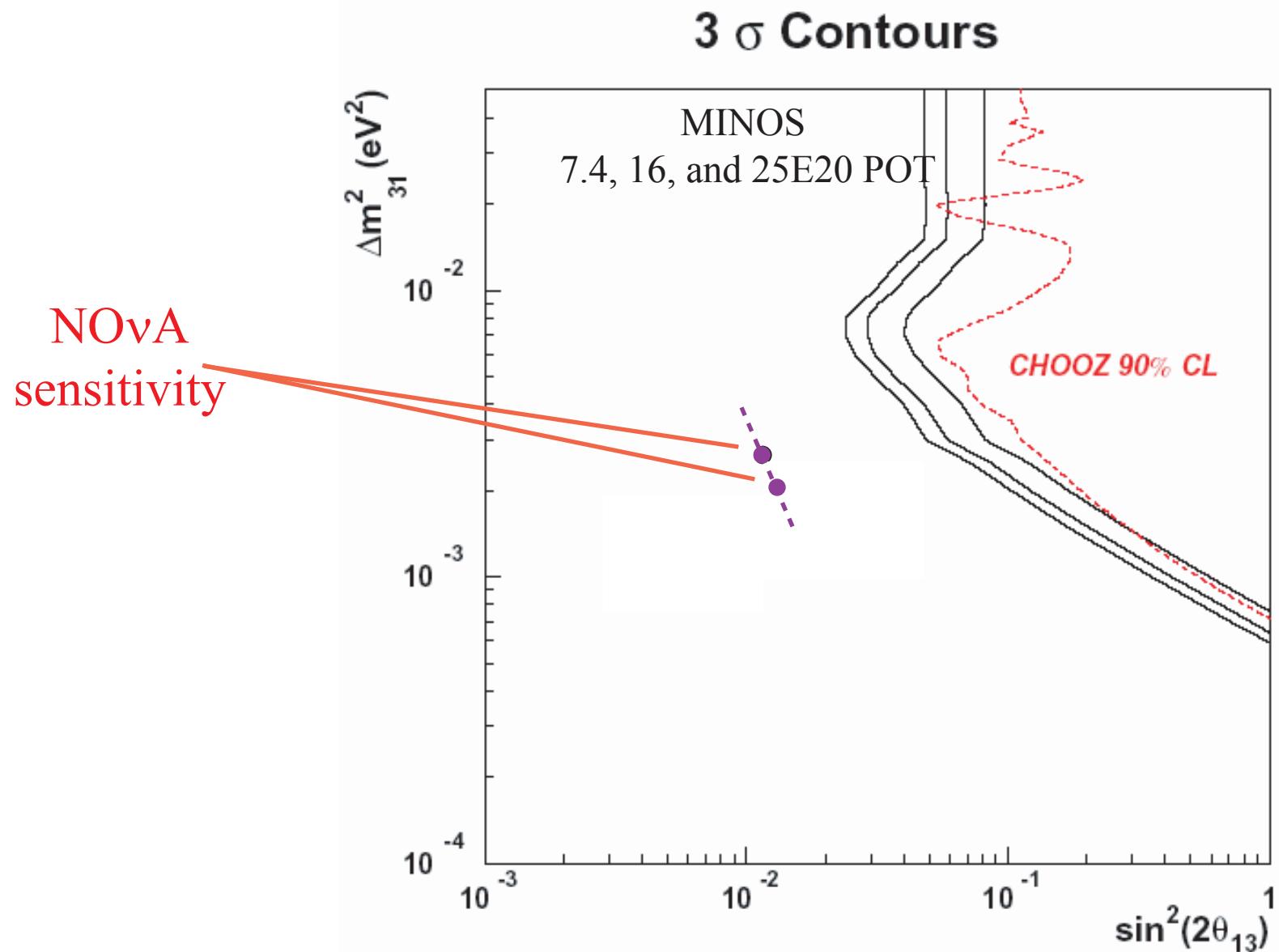


TAS Detector $\sin^2 2\theta_{23}$ measurement

- ◆ Quasi-elastic events are very clean in TASD
- ◆ Excellent energy resolution
- ◆ Essentially zero NC background
- ◆ Allow for clean measurement of $\sin^2 2\theta_{23}$ to roughly 1-2% level



NOvA Sensitivity Compared to CHOOZ/MINOS



Reminder of the problem

Experiments measure oscillation probabilities [this case $P(v_e) = 0.02$ and $P(\bar{v}_e)$]

Ambiguities in $\sin^2(2\theta_{13})$ due to:

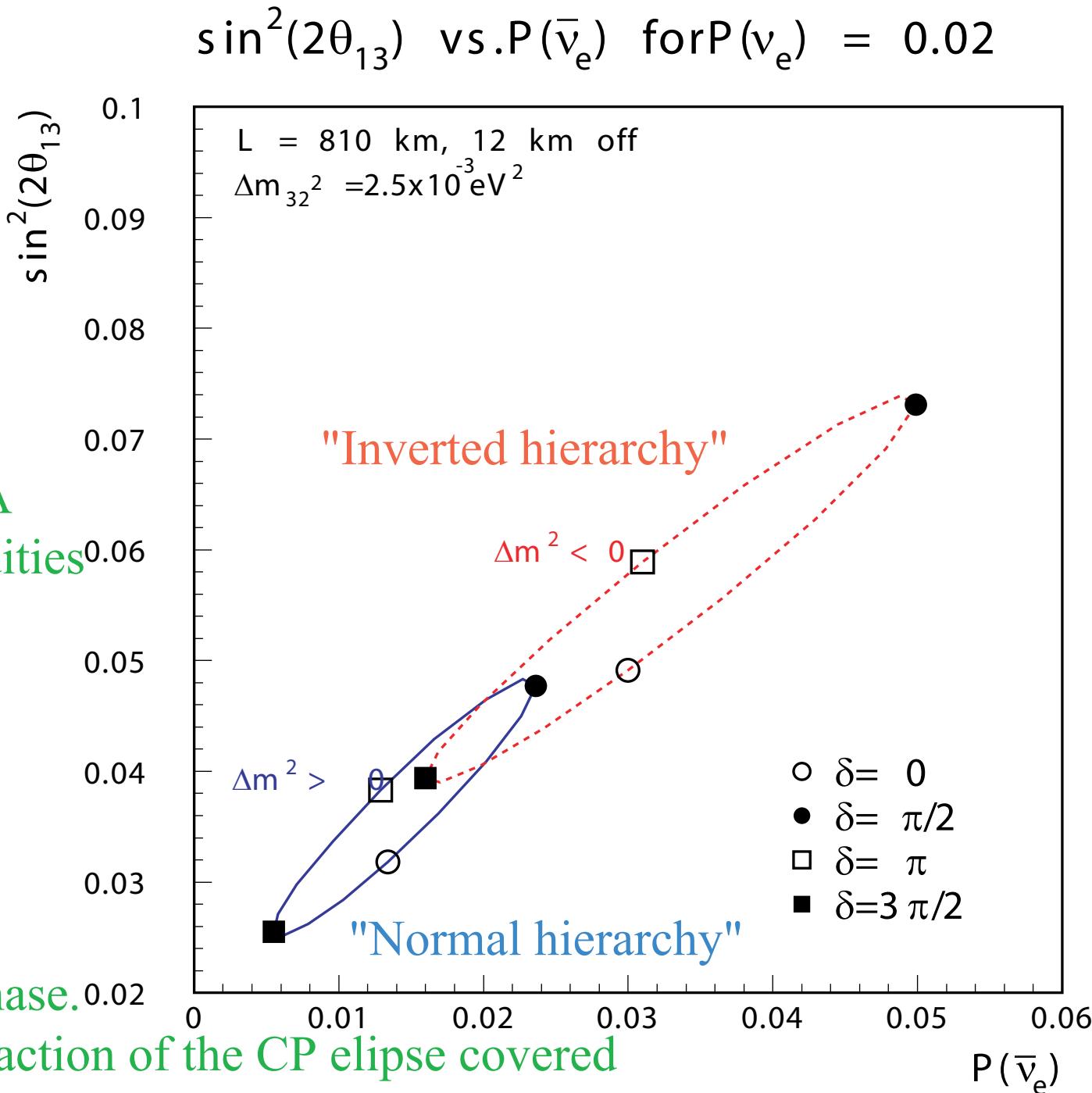
- CP phase δ ,
- mass hierarchy

Use comparison of NOvA and T2K to break ambiguities

Possibly use 2nd detector located a 2nd oscillation maximum

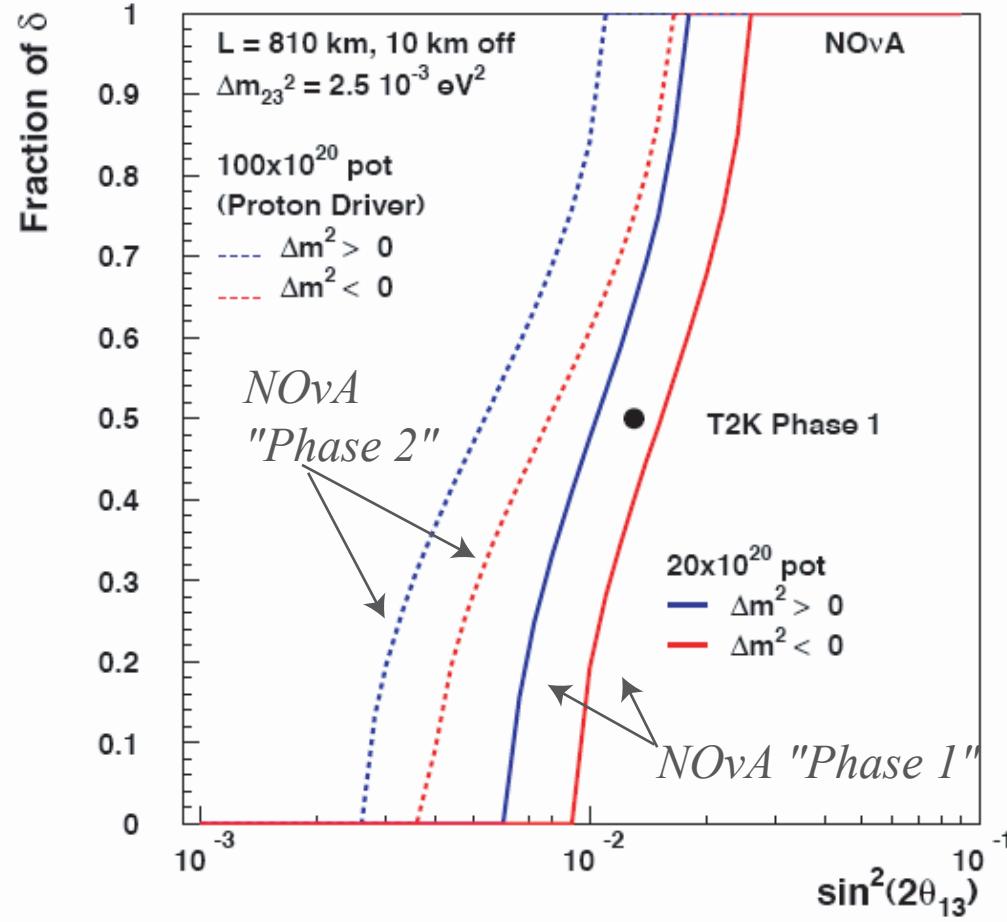
Sensitivity of experiment varies according to CP phase.

Quote sensitivities as a fraction of the CP ellipse covered

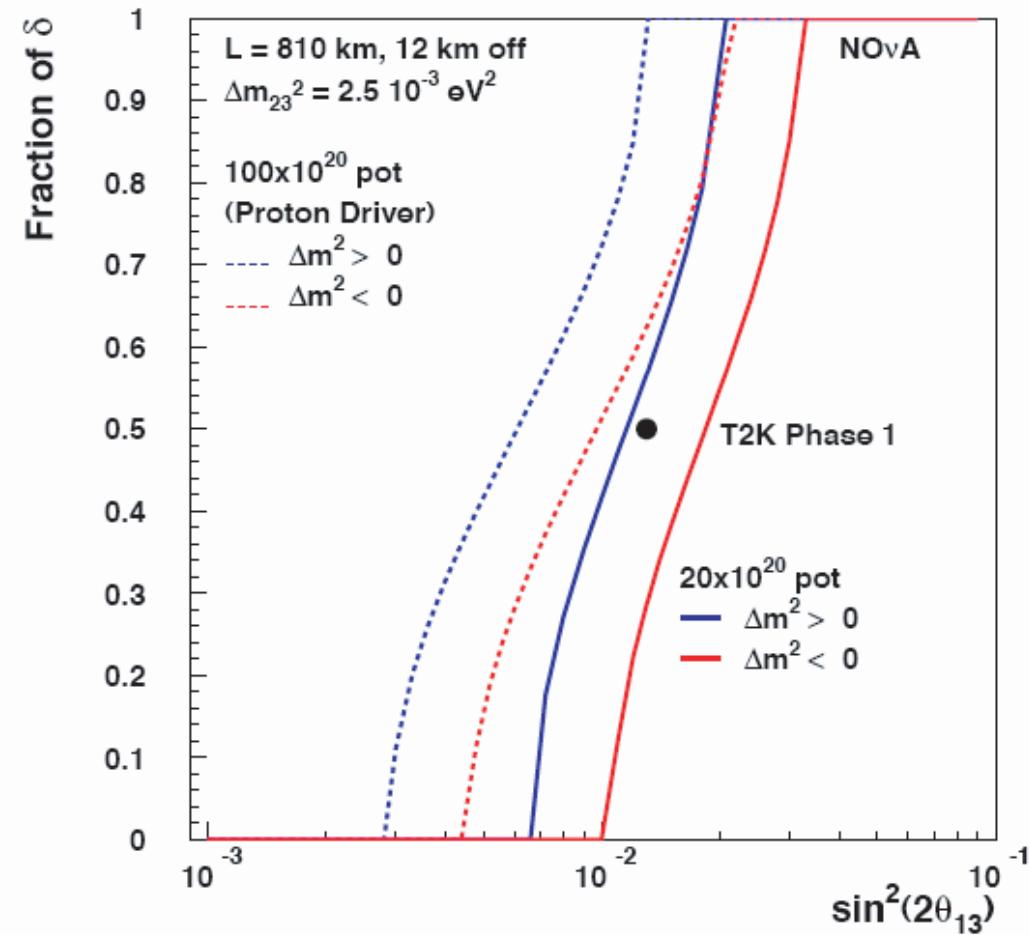


$\sin^2(2\theta_{13})$ Sensitivity

3 σ Sensitivity to $\sin^2(2\theta_{13})$



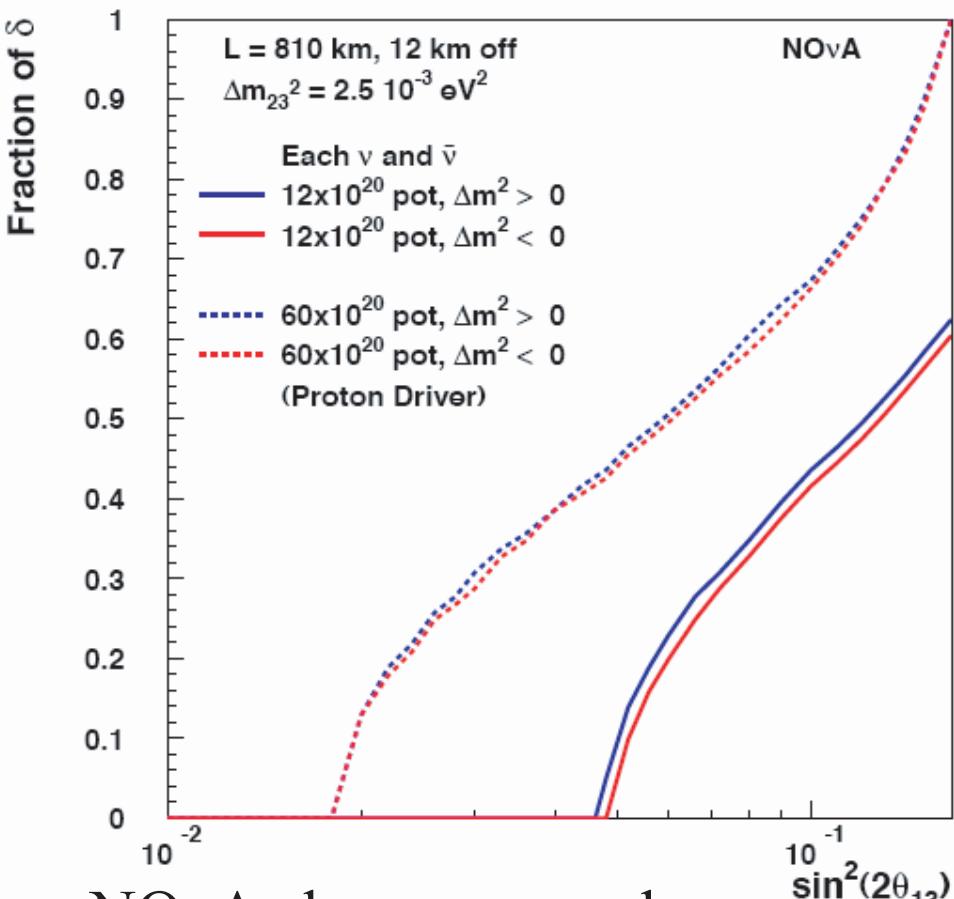
3 σ Sensitivity to $\sin^2(2\theta_{13})$



Resolving the mass hierarchy

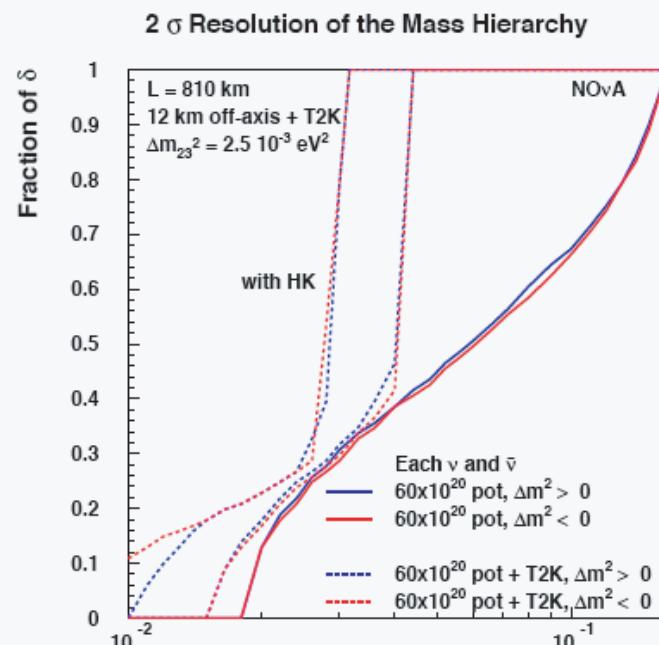
NOvA Alone

2 σ Resolution of the Mass Hierarchy

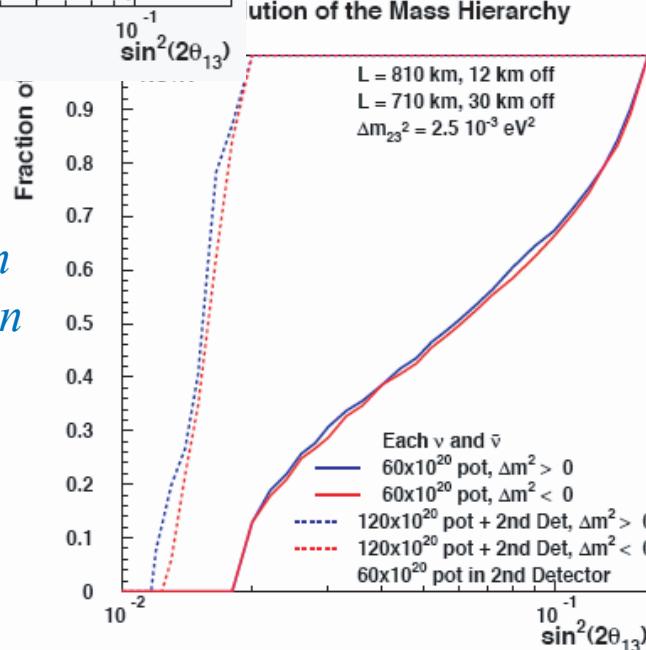


NOvA alone can resolve
hierarchy for large $\sin^2(2\theta_{13})$
over 30-40% of δ phase space. Proton Driver
extends this reach by a factor of 2

in combination with T2K

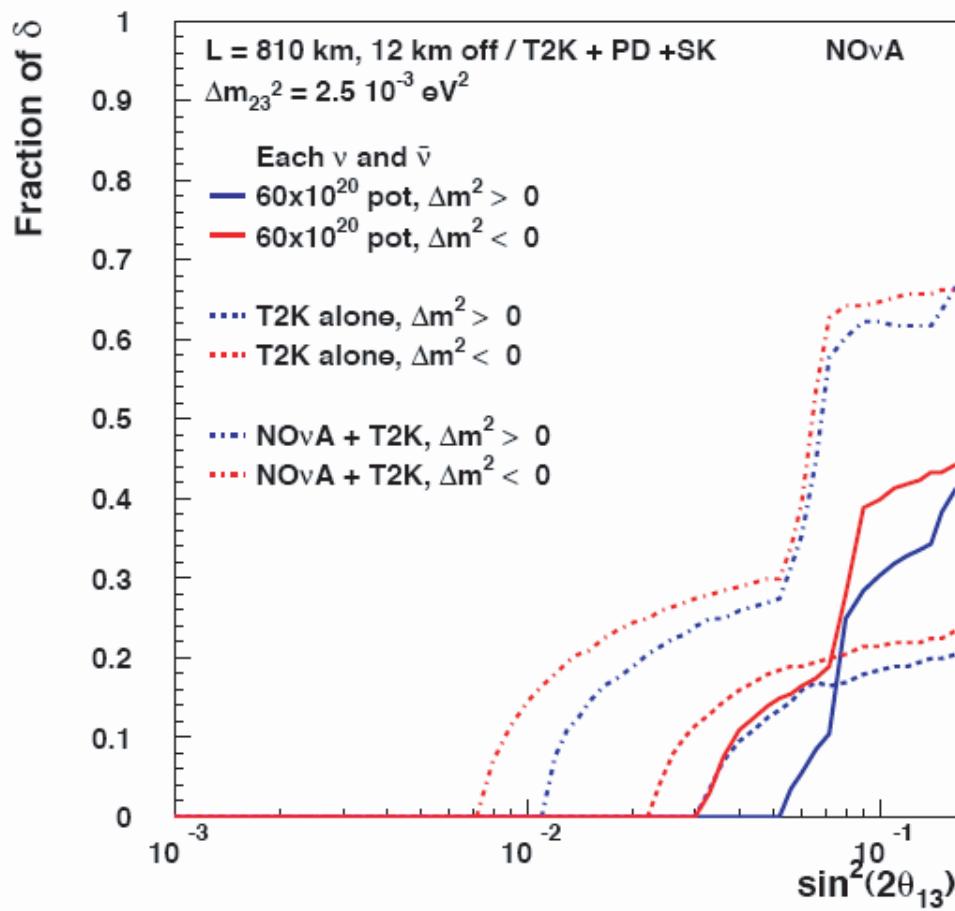


in combination
with detector in
NuMI beam at
2nd maximum

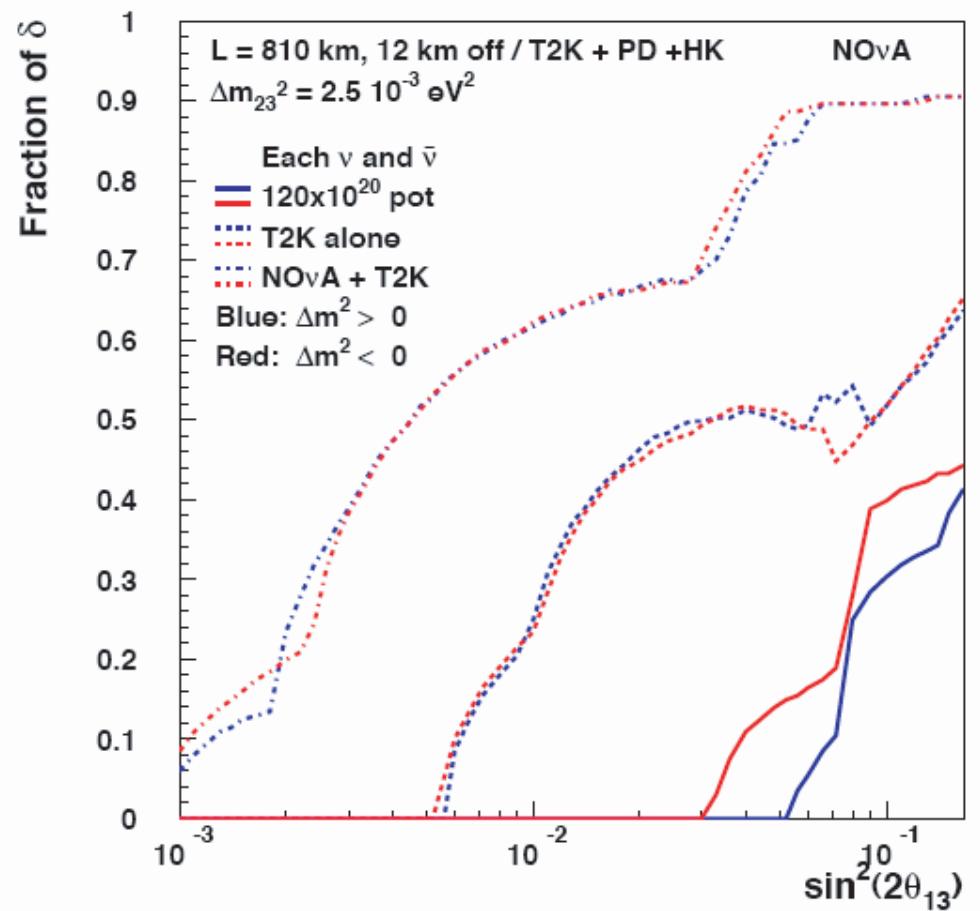


Demonstrating CP violation

3 σ Determination of CP Violation



3 σ Determination of CP Violation



NOvA extends reach of CPV search in $\sin^2(2\theta_{13})$

\$\$\$\$'s

	<u>Baseline Detector</u> <i>(cost in millions of \$'s)</i>	<u>TAS Detector</u> <i>(cost in millions of \$'s)</i>
Far Detector		
<i>Absorber</i>	17	0
<i>Active Detector</i>	39	84
<i>Electronics</i>	11	14
<i>Shipping and Installation</i>	28	17
Far Detector Sub-total	95	115
Building	37	29
Project Management	6	6
Near Detector	5	9
Active Shield	4	0
Total	147	159

- Costing model based on recently completed MINOS far detector
- Includes overhead and contingency

Technically limited schedule

	2004	2005	2006	2007	2008	2009	2010	2011	2012+	future...
Stage 1 Approval	👍									
Final technology decision	◆									
Final Approval		👍								
Start Construction			🏗							
Baseline Detector										
Start Data Collection					😊					
Finish Construction								🏗		
TAS Detector										
Start Data Collection				😊						
Finish Construction						🏗				
Protons/year (goal)		2.5E20	3.0E20	3.5E20	4.0E20					20.E20??
										Proton Driver?

- TASD detector has fewer parts and can be assembled in roughly half the time of the baseline detector
- Proposal calls for plan to reach proton intensity goals building on existing MINOS collaboration with beams divisions

Conclusions

[1] *The NOvA detector options:*

15% active 50 kt liquid scintillator + particle board absorber

85% active ("TASD") 25 kt liquid scintillator + PVC containers

TASD looking very attractive

Technically limited schedule has detector on the air in 2007 or '08

[2] *Physics reach:*

$\sin^2 2\theta_{13}$ down to 0.01-0.2

1-2% $\sin^2 2\theta_{23}$ measurement

possibility to resolve mass-hierarchy via matter effect

extend reach of CPV searches

[3] *NOvA provides a flexible program to study remaining unknowns in neutrino mass and mixing. An experiment with sensitivity to matter effects is crucial.*

[4] *Proposal to be reviewed by FNAL Physics Advisory Committee this month*