

*...projections of proton decay experience...*

from IMB thru Super K

...and onto MegaTonne

*...ideally at Frejus?*

L. R. Sulak

CPPM Marseille and Boston University

## physics motivation...some quotes from theorists

unification theories severely constrained...nucleon decay around corner  
neutrino oscillations exist  $\Rightarrow$  proton decay

synergism with superbeams, neutrino factories

Mauro Mezzetti

search for CP violation in leptons, if mixing angle  $\theta_{13}$  big enough  
sign of  $\Delta m^2$ , using matter interference if sufficient L/E

state of the art...extrapolating IMB and SK to MegaTonne with  $\sim$  no background

“proposed sites”...potential competition...no known  $(100\text{m})^3$  excavation

WIPP, Homestake, San Jacinto, etc...with beams from BNL or FNAL

Frejus, with eventual beams from CERN

Jacques Bouche

depth?...only mass, timing & pixels count...photoelectrons second order  
S-K continues, Hyper-K later? with Japanese Superbeam to both

## potential detector technologies...

water Cherenkov...affordable big mass...especially below  $\pi$  threshold  
balanced scintillation & Cherenkov light, water or oil-based medium

new photodetectors...e.g. pm's half the price of SK?

Pierre Bourgeois

liquid argon...feasible at this scale?

*some quotes from theorists,*

*Murayama, 2003:*

- Baryon Number Violation is naturally expected at some level even without grand unification
- Supersymmetry connects proton decay to Planck-scale physics
- Proton decay suppression may well be due to the same reason why electron is light
- Models suggest rates at “interesting” levels

*Babu, Pati, Wilczek:*

SO(10) models have many more fields at the GUT-scale  
Typically worse than SU(5)

Just above the current limit  $\tau(p \rightarrow K^+ \nu) < 10^{34}$  yrs

*review of the past success...*

## comments on ring-imaging Cherenkov detectors

project prematurely terminated

along with SSC

Dumand '76 -

*...an opportunity seized by Antares*

world's largest calorimeters

IMB '81-

10 kilotons =  $(20\text{m})^3$

salt mine, Ohio

Kamiokande '83- 3 kilotons

heavy metal mine, Japan

Super-K '96-

50 kilotons =  $(40\text{m})^3$

a 2nd site at Kamioka

little to say about the other experiments or the era...

e.g. SNO, Soudan...too small (1 kT) or not totally active (iron absorber)

MB III (1983)

2000 - 8 inch pms and light collectors  
dry suit diver/physicist



Super-K (1996)

half way up first filling

inner detector -

11,000 20 inch pms

outer detector (not visible)

a reconfigured IMB III

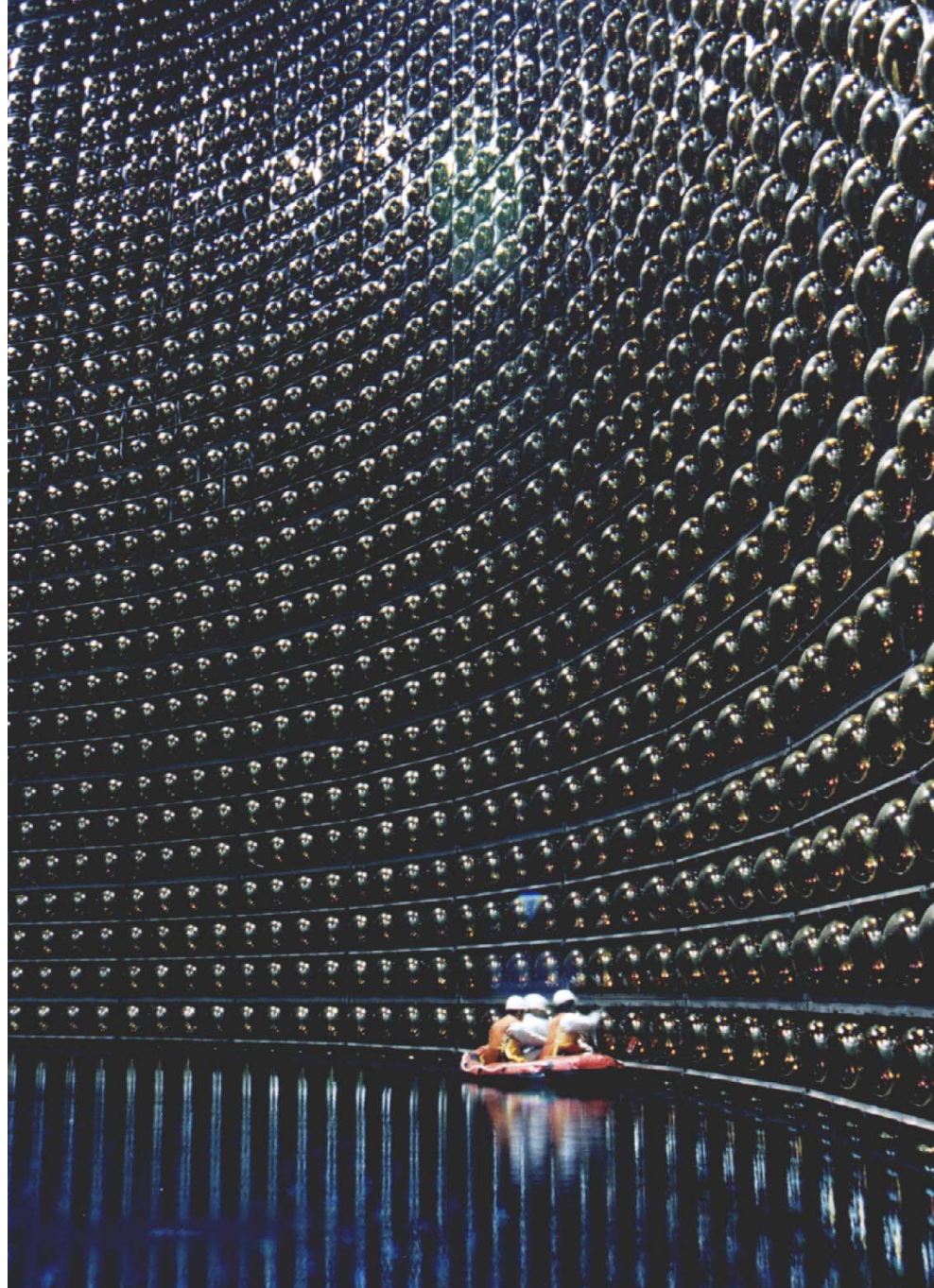
2,000 – 8 inch pms +

wavelength shifting light  
collectors

~\$100 M = ~1/3 pms + DAQ

~1/3 excavation

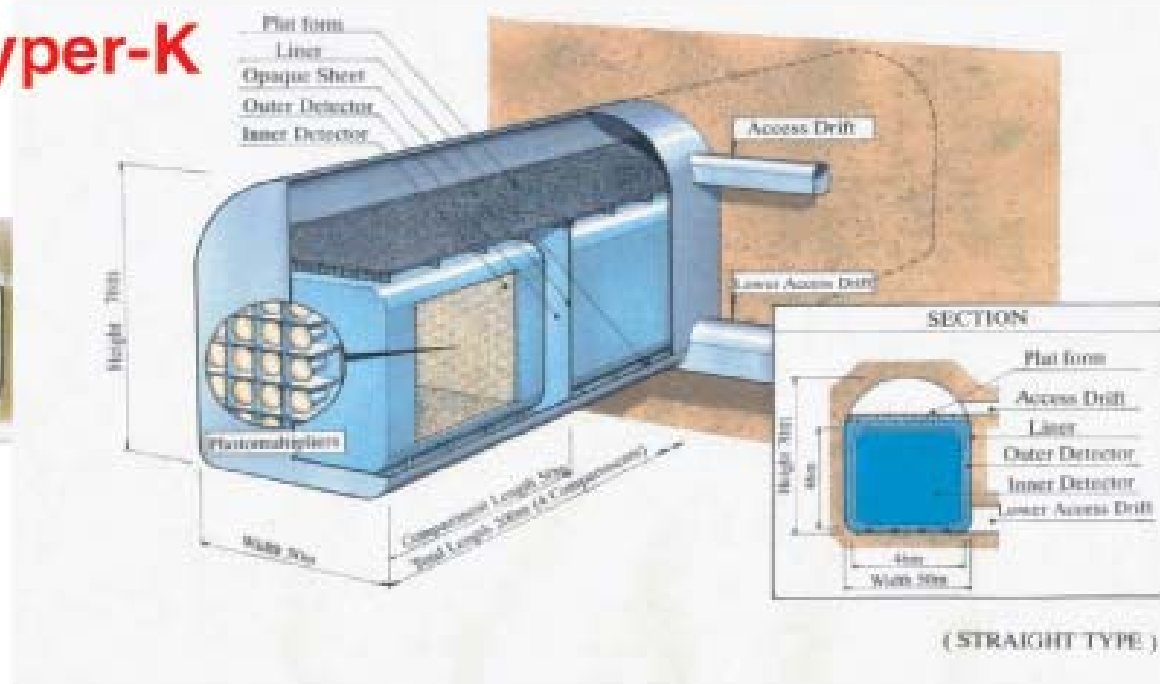
~1/3 tank *et al.*



# Hyper-K



SK

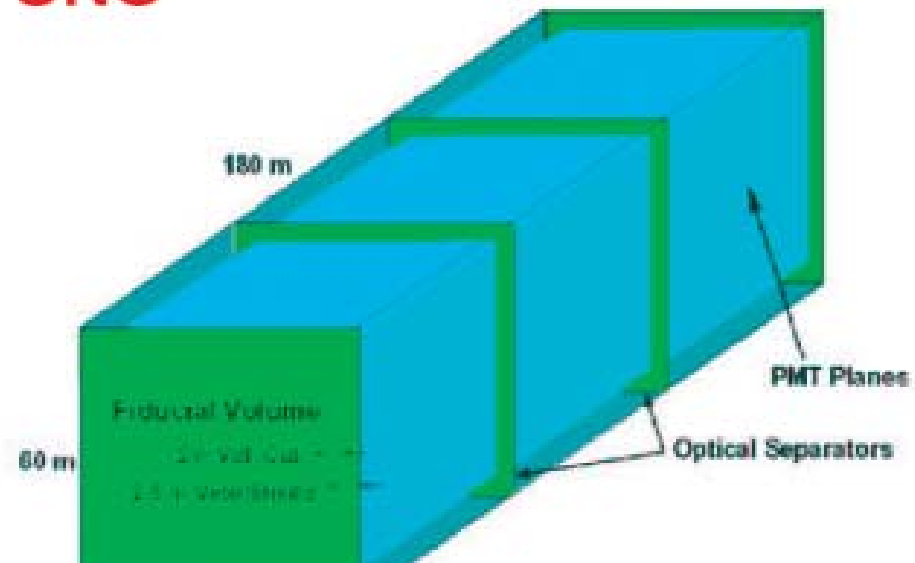


*two proposed*

*future large*

*water detectors...*

# UNO

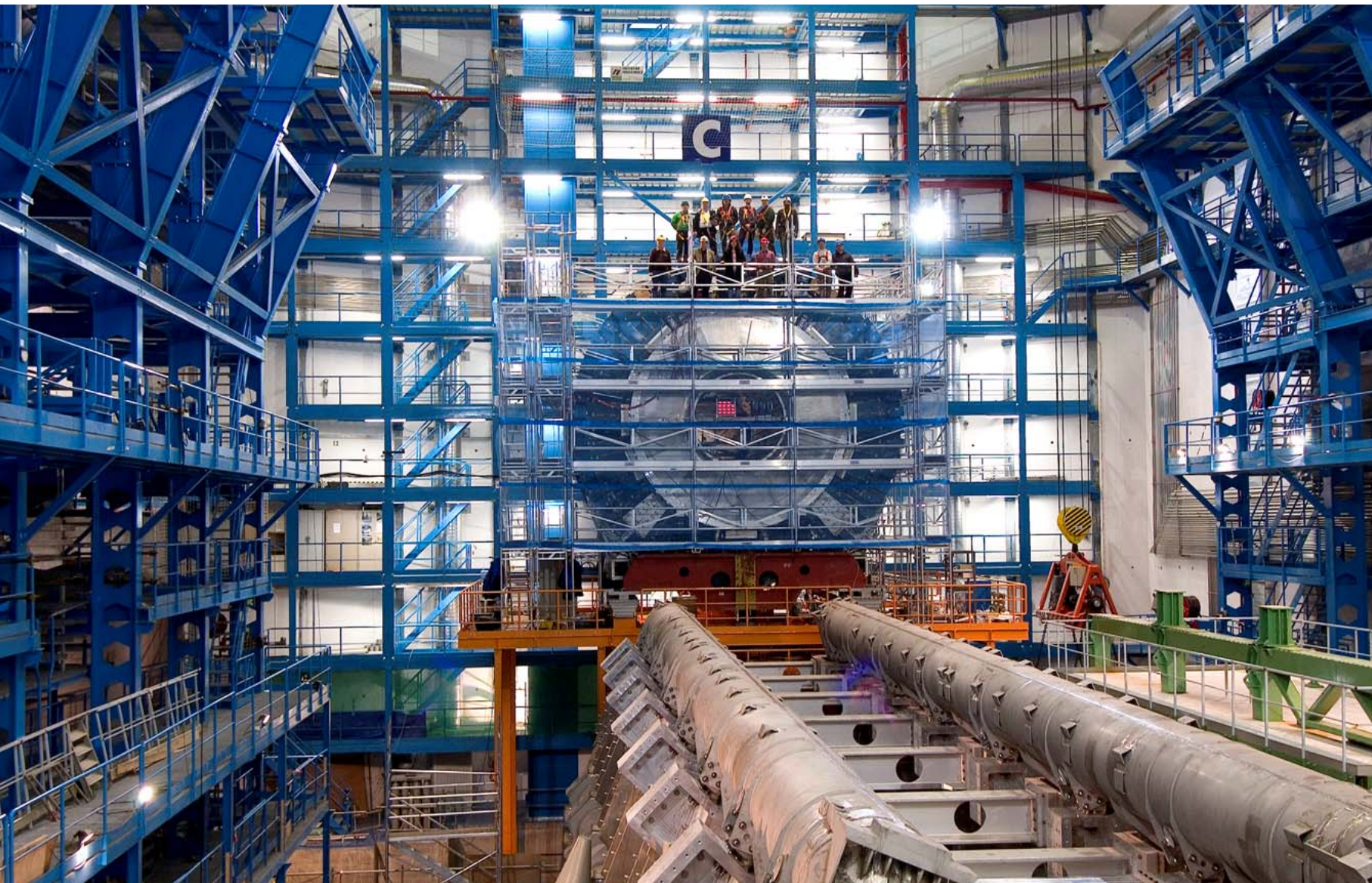




*What is a 50% volume scale up of Super-K cavity?*

~12 Atlas stories...60m high = 200 kT of water...need a few such caverns at depth

...Europe well experienced



*A summary of water detectors: where are we? where going?*

## Proton Decay Search at Super-K I and II

Current State...size, then livetime are everything

IMB: best lifetime limits for most of  $\sim 40$  modes...takes time, 10 years  
blackened spherical volume, surface detectors (22 m/side)  
most economical: minimum # pms, water cheapest medium  
reverse osmosis water attenuation length  $\sim 100$  m  
scales as volume/surface until 100m characterize dimension

Super-K:  $3.5 \text{ yr} \times 22 \text{ kT} \cong 80 \text{ kT-yr}$

biggest water-filled cavern, 40 m high, 40 m diameter

could probably go to  $60 \text{ m} \times 60 \text{ m}$  in a nearby location

$p \rightarrow e^+ \pi^0 \geq 7 \times 10^{33}$ , background  $\sim 0.2$  events

$p \rightarrow \nu K^+ \geq 1.6 \times 10^{33}$ , background  $\sim 2.2$  events

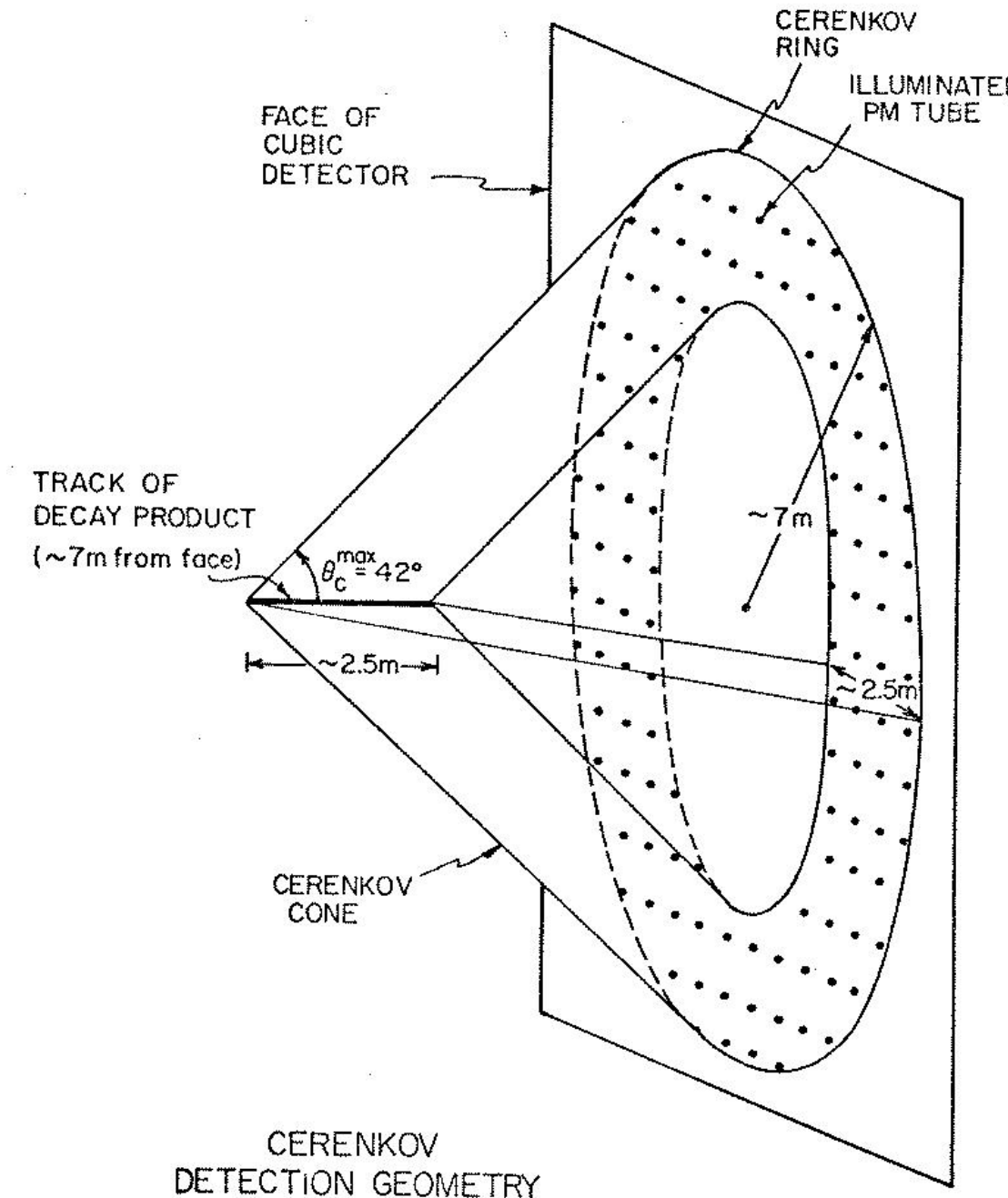
K2K: study of  $\nu$  interactions in water in near detector

with new cuts for SK...eventually be  $\geq 5$  times more restrictive?

appears to be no background for  $10 \text{ yr} \times 0.5 \text{ MegaTon}$  fiducial

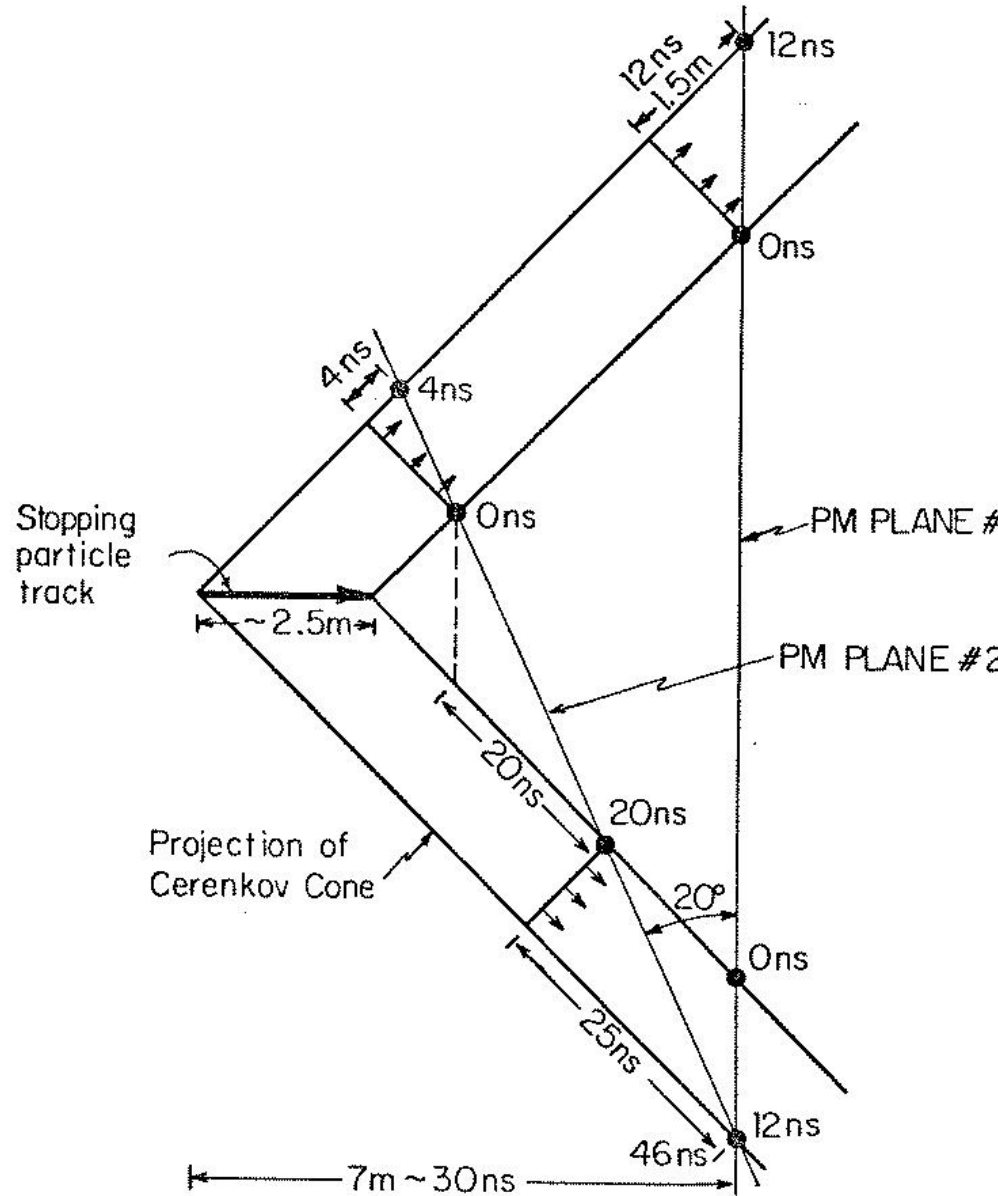
IMB ('81-'90), Kamiokande ('83-'96), Super-K:  $\geq 10$  year livetime realistic

...a charged particle track  
what does it look like?



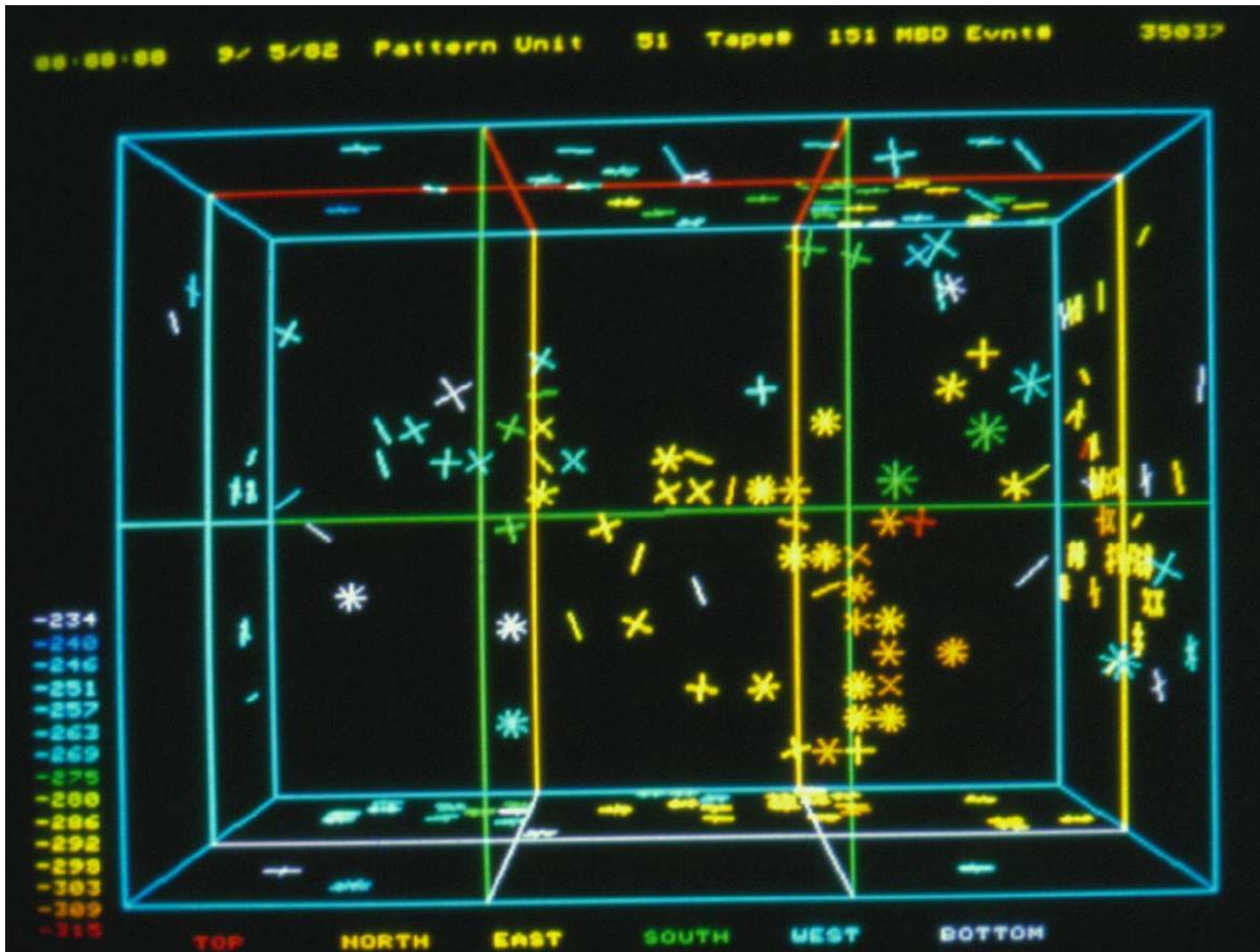
“Neutrino ’79, Bergen” LRS

...why is timing so important?  
 Cherenkov light is directional



“Neutrino ’79, Bergen,” LRS

VIB: best proton decay candidate...pm code = timing in color, 1 pe/slash

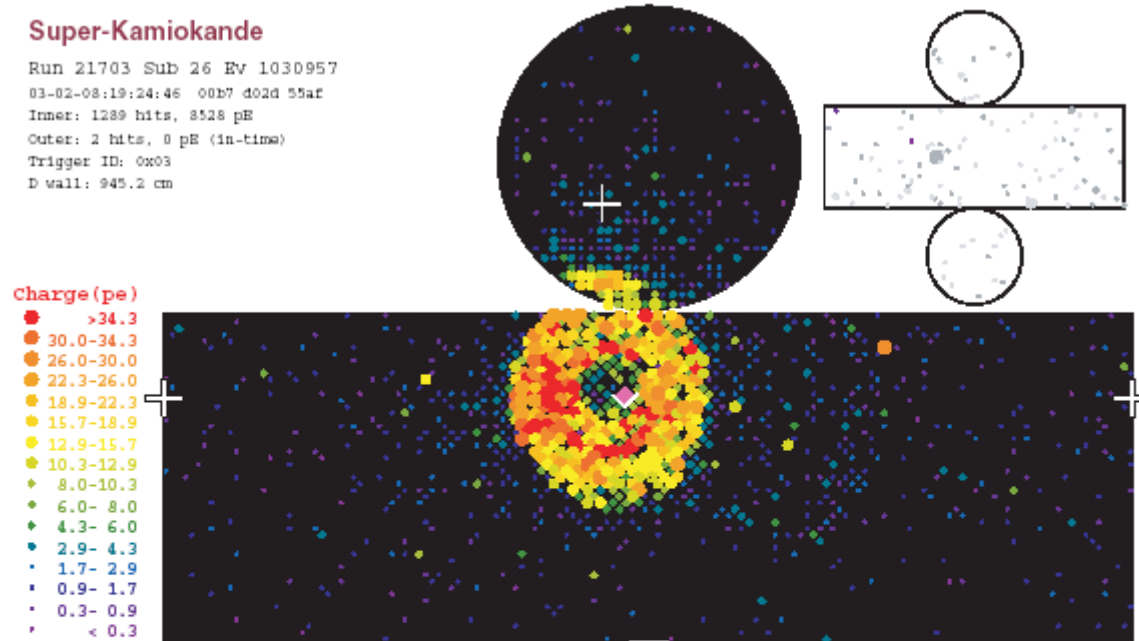


# SK II (2003): muon-neutrino event

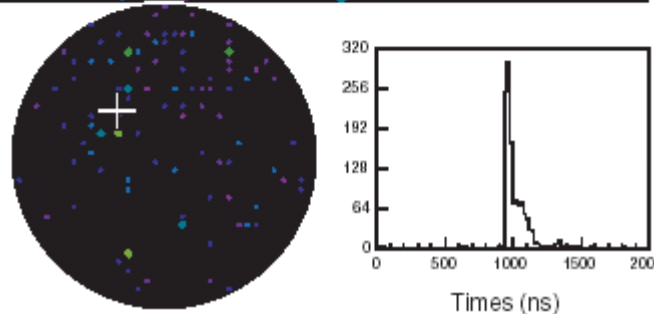
...in time with beam pulse from KEK accelerator 300 km away

## Super-Kamiokande

Run 21703 Sub 26 Ev 1030957  
03-02-08:19:24:46 00b7 d02d 55af  
Inner: 1289 hits, 8528 pE  
Outer: 2 hits, 0 pE (in-time)  
Trigger ID: 0x03  
D wall: 945.2 cm



## Rebuilt Super-K Example Event (from K2K beam)



...sharp ring edges characteristic of a muon track

# typical electron track...fuzzy at edges

Super-Kamiokande

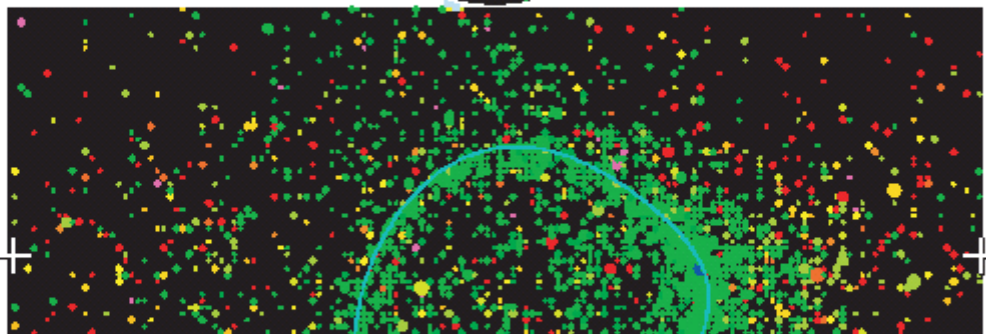
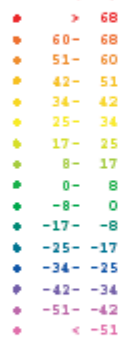
Run 4268 Event 7899421

97-06-23:03:15:57

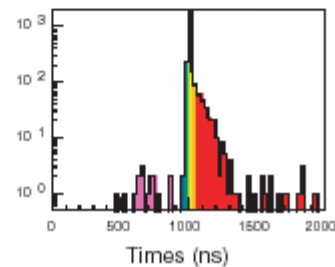
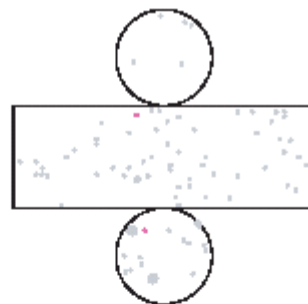
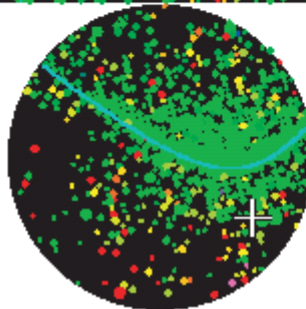
inner: 2652 hits, 5747 px

~620 MeV/c

Resid(ns)



for comparison  
SK-1  
e-like  
PMT time view



*SK summary: where are we going?*

## Proton Decay Search at Super-K II

### Near Term Program:

through September 2005 operate with 1/2 pms

no deterioration of sensitivity for proton decay

SK designed for MeV solar neutrinos...overkill for pdk

complete comparison of atmospheric  $\nu$  background to proton decay

with events in near water detector of K2K

linear gain in limits for “no background” modes

...square root otherwise

develop improved cuts and reconstruction for Hyper-K...detector

invaluable guide to designing post Super-K detector

define size of detector by level of background

set technology

optimize cost



## Future of Kamioka Lab

### Medium Term Goal: Continuation

end 2005 and ~2012

proton decay search with original pm density  
(probably put in gadolinium salt to enhance supernova signal)

≥ 2007 JHF 1

superbeam for neutrino oscillation studies

### Long Term Goal: New Megaton Detector Hyper-K

need significant increase of sensitivity,  $\times 10$ -20

sensitivity = fiducial mass  $\times$  detection efficiency

what mode to focus on?

$K^+$  detection could drive technology

Caveat: no Hyper-K start unless

some indication of proton decay, or

$\theta_{13}$  big enough

## Some Post Super-Kamiokande detector ideas

which detector technology?

0.5 - 2 MT water Cherenkov

Titanic - a sunken, water and pm-filled tanker

Y. Suzuki -

detector with balanced scintillator / Cherenkov oil

Svoboda

liquid argon LANNDD

Cline

## Which Detector Technology? water underground

water Cherenkov - lowest cost per megaton  
underground

Super-K 50 kT total, 22 kT fiducial

...due to 2 m veto + 2 m fiducial cut

scale-up volume of Super-K  $\times 2$  to 100 T...rock supportable at new site

...repeat array of 5  $\sim$  double Super-K tanks

maximum cavern size sets ultimate limit...what is possible at Frejus?

excellent for oscillations at neutrino energies below  $\pi$  production

do not want to confuse with  $\nu_\mu$ ,  $\nu_e$  conversion

only solution for massive, economical far detector at 2000-4000 km?

for sign of neutrino mass difference

virtual proposals

Hyper-K, probably for JHF 2,  $\geq 2012$ , with 4MW superbeam

UNO

## Which detector technology: sunken water tank?

water Cherenkov - lowest cost per megaton

### undersea

submerged vessel sunk and anchored on bottom  
e.g. used liquid natural gas tanker

no excavation

avoid dominant construction time and cost

outfit at surface; fill with osmosis water; raise to surface for maintenance

moveable from source for oscillation studies

but, >100m depth...must use pressure-tolerant enclosures

(no bioluminescence, sea currents movement

...e.g. for advocates of MegaTon nestled inside Antares)

“Supersymmetry...

as generated so many thousands of papers it must be correct’

Shelly Glashow



what are the options, continued?

## new detector technology for $\nu K^+$ ?

What if SUSY discovered at LHC ? Focus detector on  $\nu K^+$  ?

What if Super-K gets a candidate for  $\nu K^+$  ?

fill with dilute scintillator to gain a factor of 4 in rate, 16 in signal?

### 2 options to boost sensitivity for $p \rightarrow \nu K^+$

1) replace with high light yield scintillator oil, *e.g.* Kamland, or

2) balanced doping of oil or water: *e.g.* LSND / Miniboone

a) isotropic scintillator light

gives calorimetry and timing signature of  $K^+$

b) but dilute so that Cherenkov signal not overwhelmed

preserves ring-imaging and directionality

$\nu K^+$  detection efficiency increases 10%  $\rightarrow$  40%

potential problem:  $\mu/e$  discrimination degraded? Now under study

technical information to come from Miniboone

*e. g.* electron *vs.*  $\pi^0$  discrimination

MegaTon project: \$1B/MT of oil? No, use “gin & tonic” in water

## *Super-K Lifetime limits...*



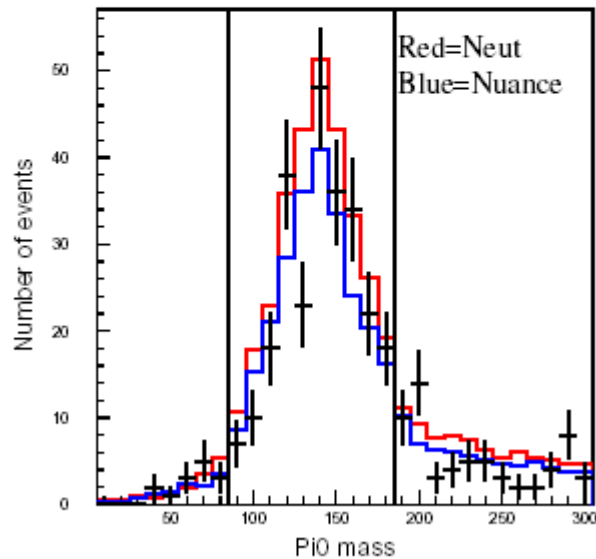
### Summary of cuts

- 2 or 3 rings
- All rings e-like
- $85 \text{ MeV} < m_\pi < 185 \text{ MeV}$ 
  - Applied only to 3-ring events
- 0 decay electrons
- $800 \text{ MeV} < m_p < 1050 \text{ MeV}$
- $P_p < 250 \text{ MeV}$

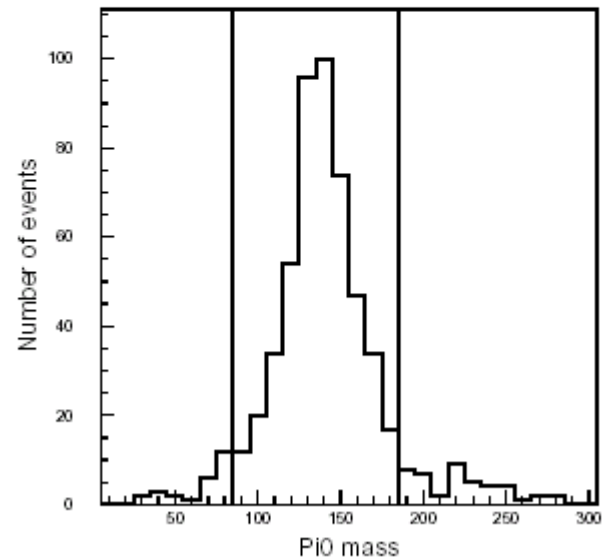
*preliminary results, from Scott Clark*

# Reconstructed $\pi^0$ mass, SK-I

For 3-ring events, reconstruct the  $\pi^0$  mass for the 2 rings coming closest to 135 MeV. This must be between 85 MeV and 185 MeV



Atmospheric data & MC



$e^+\pi^0$  MC



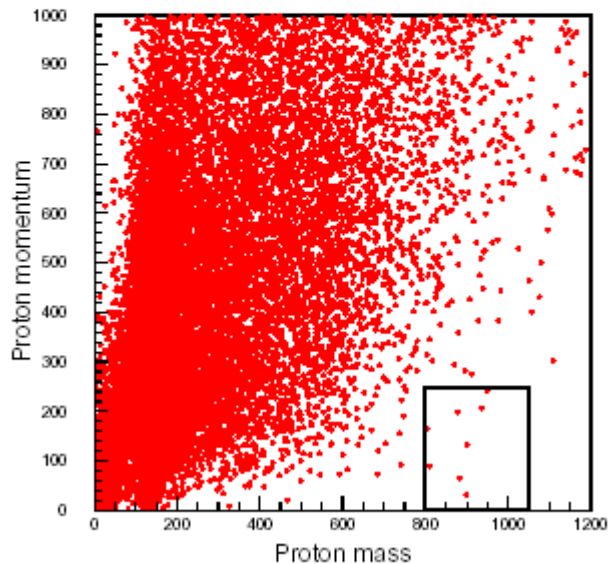
*Two 100 year simulations...*

## Proton mass & momentum, SK-I

Require:  $800 \text{ MeV} < m_p < 1050 \text{ MeV}$

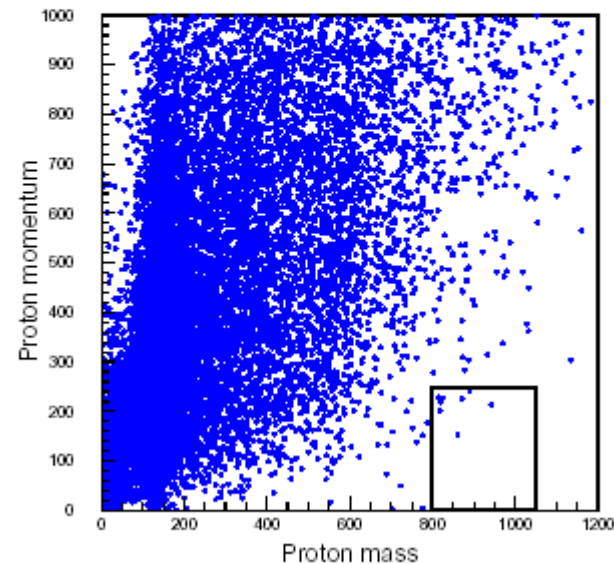
$p_p < 250 \text{ MeV}$

8 BG events (0.3 in 1489 days)



Atmospheric Neut MC

9 BG events (0.4 in 1489 days)



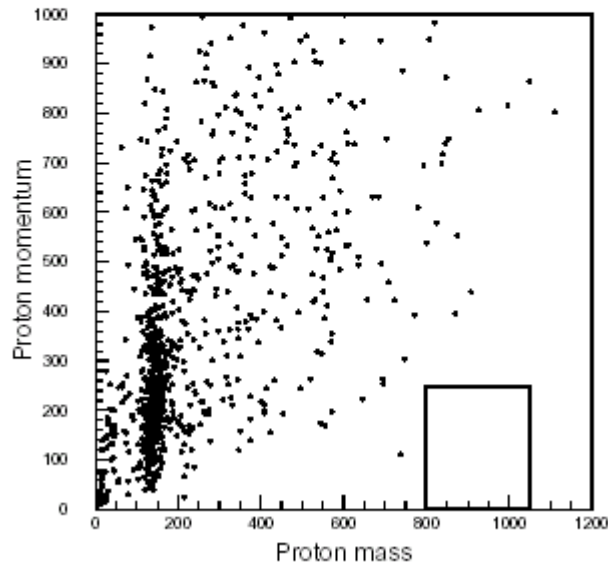
Atmospheric Nuance MC

# Proton mass & momentum, SK-I

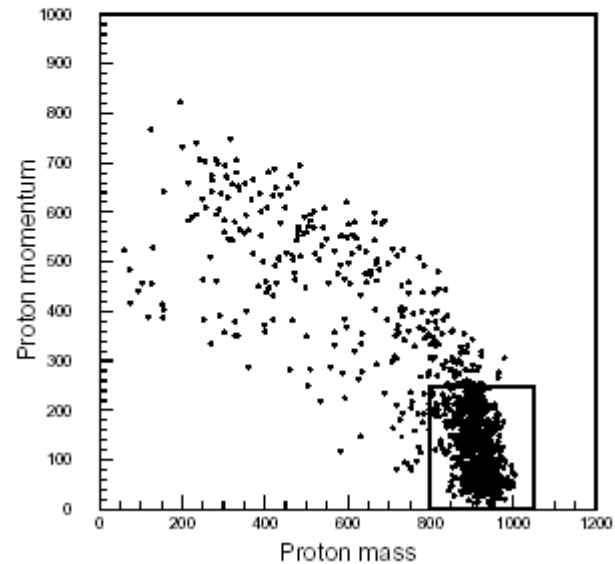
Require:  $800 \text{ MeV} < m_p < 1050 \text{ MeV}$

$p_p < 250 \text{ MeV}$

0 candidates



Atmospheric Data



$e^+\pi^0$  MC

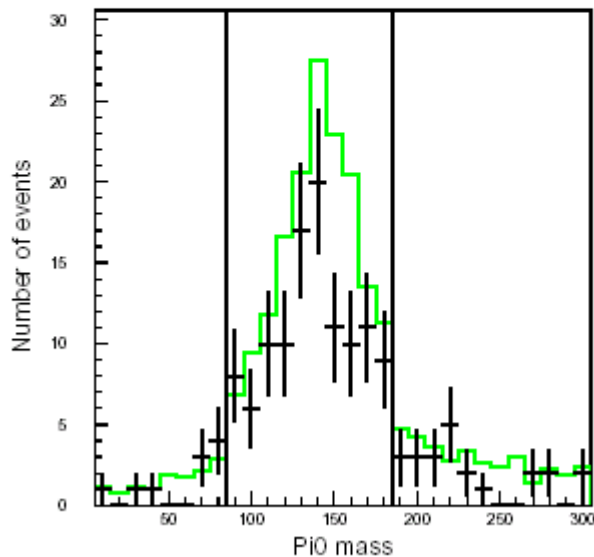
## Summary: SK-I

- 1489 days of data
  - 0 candidates
  - 0.3-0.4 expected background
- Detection efficiency 40.9%
  - 18.6% 2-ring, 22.3% 3-ring
- $\tau/B > 5.4 \times 10^{33}$  years at 90% confidence

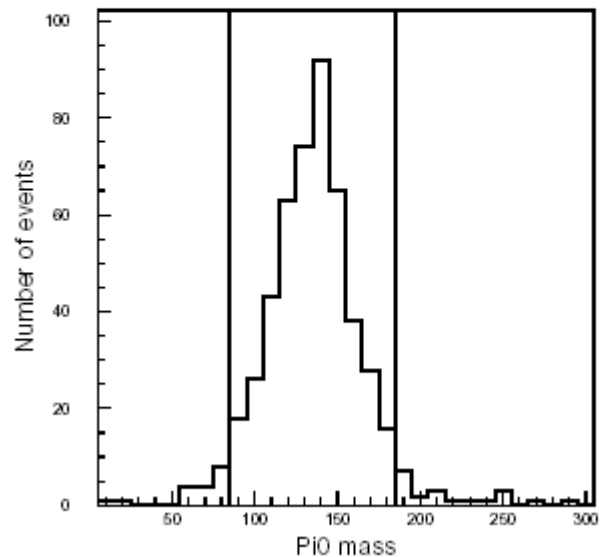
*preliminary results, from Scott Clark*

# Reconstructed $\pi^0$ mass, SK-II

For 3-ring events, reconstruct the  $\pi^0$  mass for the 2 rings coming closest to 135 MeV. This must be between 85 MeV and 185 MeV.



Atmospheric data & MC



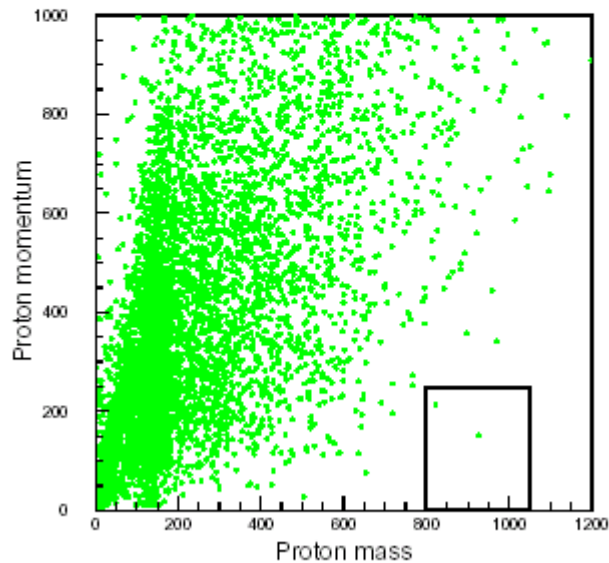
$e^+\pi^0$  MC

# Proton mass & momentum, SK-II

Require:  $800 \text{ MeV} < m_p < 1050 \text{ MeV}$

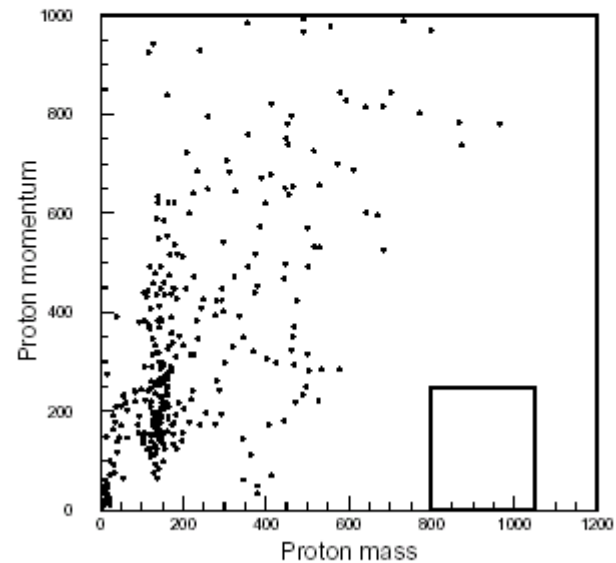
$p_p < 250 \text{ MeV}$

2 BG events = 0.1 in 421 days



Atmospheric MC

0 candidates



Atmospheric Data

## Conclusions

- A analysis (neglecting systematic uncertainties for now) for  $p \rightarrow e^+ \pi^0$  on SK-I and SK-II has been done.
- No candidate events are observed in either the SK-I or the SK-II data.
- Limits are placed on the partial lifetime at 90% confidence:
  - From SK-I:  $\tau/B > 5.4 \times 10^{33}$  years
  - From SK-II:  $\tau/B > 1.5 \times 10^{33}$  years
  - Combined limit:  $\tau/B > \mathbf{6.9 \times 10^{33}}$  years

*preliminary results, from Scott Clark*



## Summary of cuts

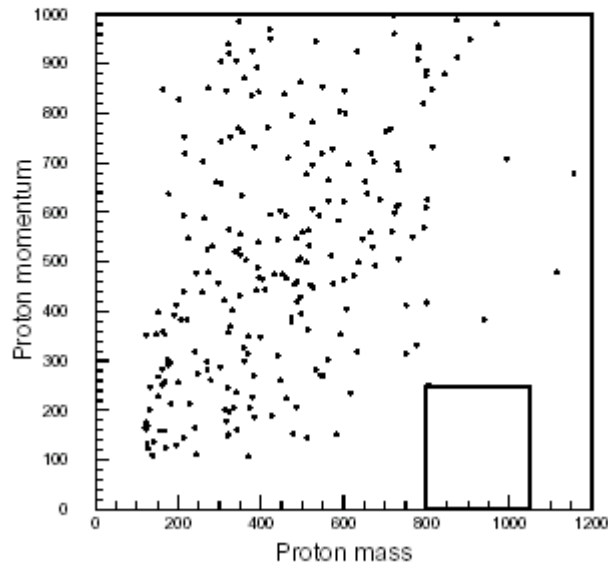
- 2 or 3 rings
- 1 ring m-like, all others e-like
- $85 \text{ MeV} < m_{\pi} < 185 \text{ MeV}$ 
  - Applied only to 3-ring events
- 1 decay electron
- $800 \text{ MeV} < m_p < 1050 \text{ MeV}$
- $P_p < 250 \text{ MeV}$

# Proton mass & momentum, SK-I

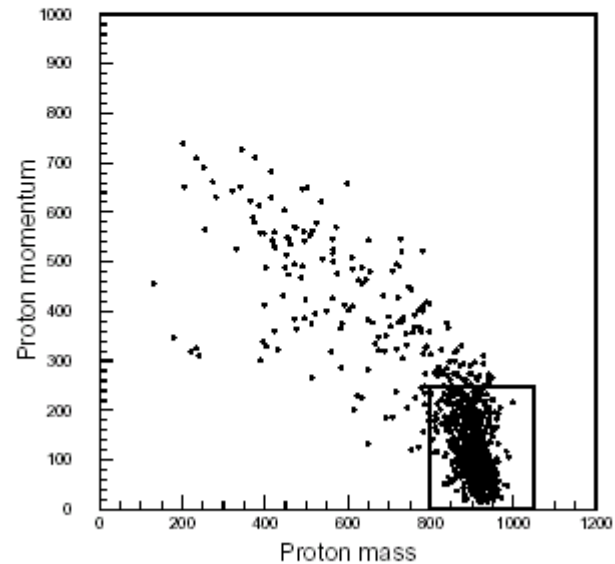
Require:  $800 \text{ MeV} < m_p < 1050 \text{ MeV}$

$p_p < 250 \text{ MeV}$

0 candidates



Atmospheric Data



$\mu^+\pi^0$  MC



# Final Conclusions

$p \rightarrow (e/\mu)^+ \pi^0$  limits for SK-I and SK-II are found.

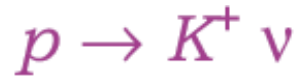
No candidate events are observed in either the SK-I or the SK-II data.

Limits are placed on the partial lifetime at 90% confidence:

- $e^+ \pi^0$  Combined limit:  $\tau/B > 6.9 \times 10^{33}$  years
- $\mu^+ \pi^0$  Combined limit:  $\tau/B > 5.4 \times 10^{33}$  years

*preliminary results, from Scott Clark*

# Super-K lifetime limits...



Favored SUSY decay mode

Note: Also  $p \rightarrow \pi^+ \nu$  in some circumstances

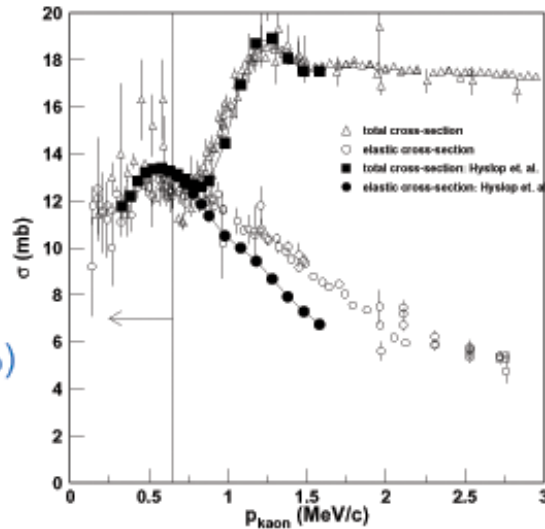
(Strassler and Babu, see also Goto and Nihei).

Also  $\mu^+ K^0$ , Babu, Pati & Wilczek

Momentum of  $K^+$  is 340 MeV/c: below  $\bar{C}$ -threshold

Nuclear Interaction:  
cross section is small  
scattering is elastic

$\Rightarrow K^+$  escapes nucleus  
and decays at rest (90%)



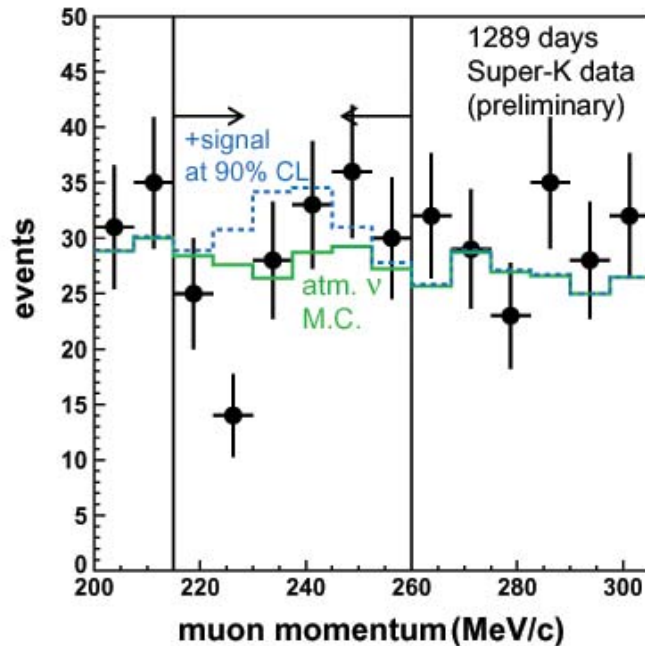
Branching ratios:  $K^+ \rightarrow \mu^+ \nu_\mu$  65%  
 $K^+ \rightarrow \pi^+ \pi^0$  21%

*From Ed Kearns*

# Super-K lifetime limits...

## $K^+ \rightarrow \mu^+(236 \text{ MeV}/c) \nu$ search

1  $\mu$ -like ring  
1 decay electron  
 $215 < p_\mu < 260 \text{ MeV}/c$   
B.R. x efficiency = 33%



from joint fit to background + signal:  
 $\tau/B(p \rightarrow \nu K^+) > 4.4 \times 10^{32} \text{ yr}$

# Super-K lifetime limits...

## Gamma Tag for $p \rightarrow \nu K^+$

Nuclear Shell Model:  
 $^{16}\text{O} (p_{3/2}) \rightarrow ^{15}\text{N}^* + \text{proton hole}$   
 de-excites by 6.3 MeV gamma

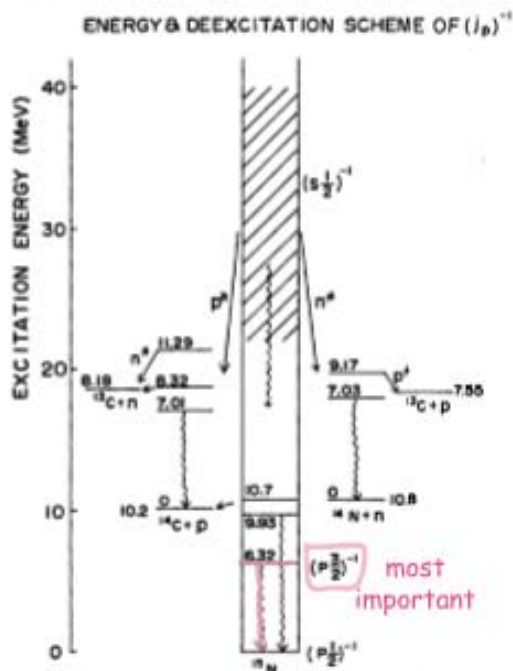
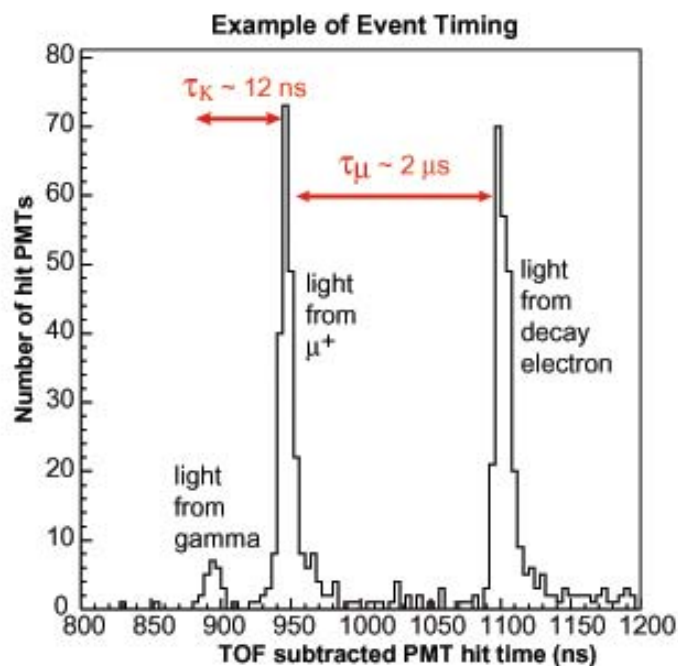


FIG. 2. Level scheme of proton-hole states in  $^{15}\text{N}$  and their deexcitation modes. Energies are given in units of MeV.  $p^*$  and  $n^*$  are the protons and neutrons emitted from the continuum (unbound) region, respectively.

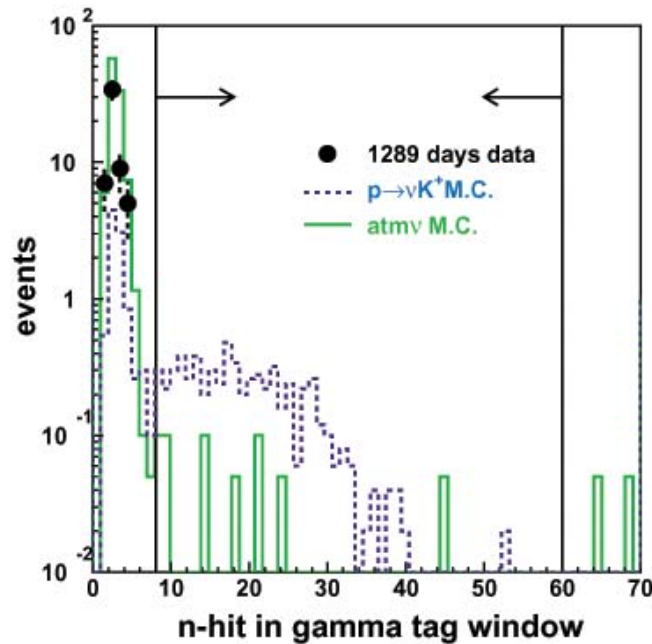
Coincidence signature:  
 proton decay to  $K^+n$  accompanied by prompt  $\gamma$   
 $K^+$  is below Cherenkov threshold: no light  
 followed by  $K^+$  decay to  $\mu^+\nu$  ~12 ns later  
 followed by muon decay to electron ~ 2 ms later



# Super-K Lifetime limits...

## $K^+ \rightarrow \mu^+(236 \text{ MeV}/c) \nu$ search with gamma tag

count PMT hits  
in 12-ns sliding window  
*preceding* light from muon  
B.R. x efficiency = 8.8%



0 events detected, background = 0.5 events:

$$\tau/B(p \rightarrow \nu K^+) > 10 \times 10^{32} \text{ yr}$$

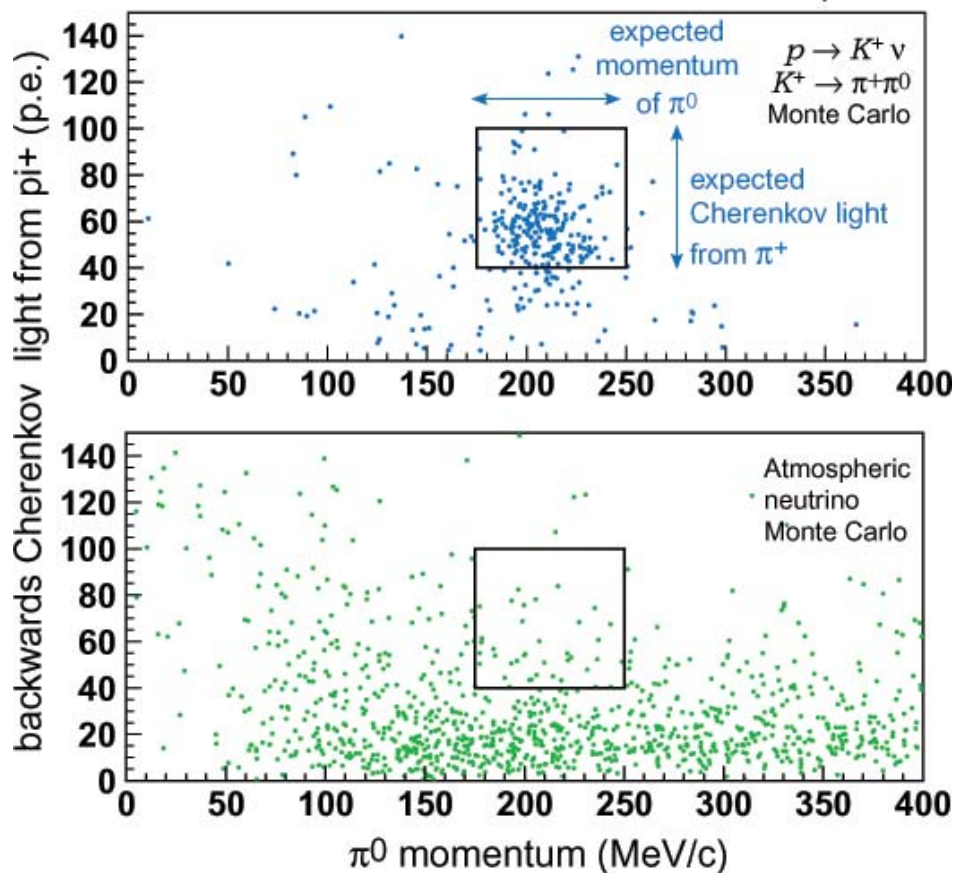
# Super-K lifetime limits...

## $K^+ \rightarrow \pi^+ \pi^0$ search

momentum of  $\pi^+$  is only 205 MeV/c:  
barely above Cherenkov threshold

require 1 decay electron,  $\pi^0$  mass

BR x efficiency = 6.8%

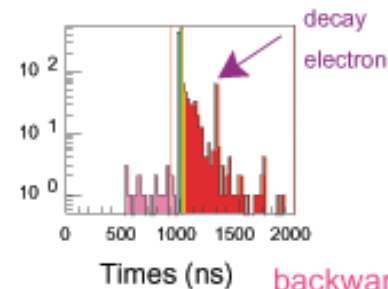


background estimate  
~2.3 events/100 kt-yr

## Super-Kamiokande

Run 1000000 Event 474  
1997-06-25:12:59:29  
Time to prev. event: 0.0us  
Inner: 1395 hits, 2128 pE  
Outer: 16 hits, 9 pE (in-time)  
Trigger ID: 0x03

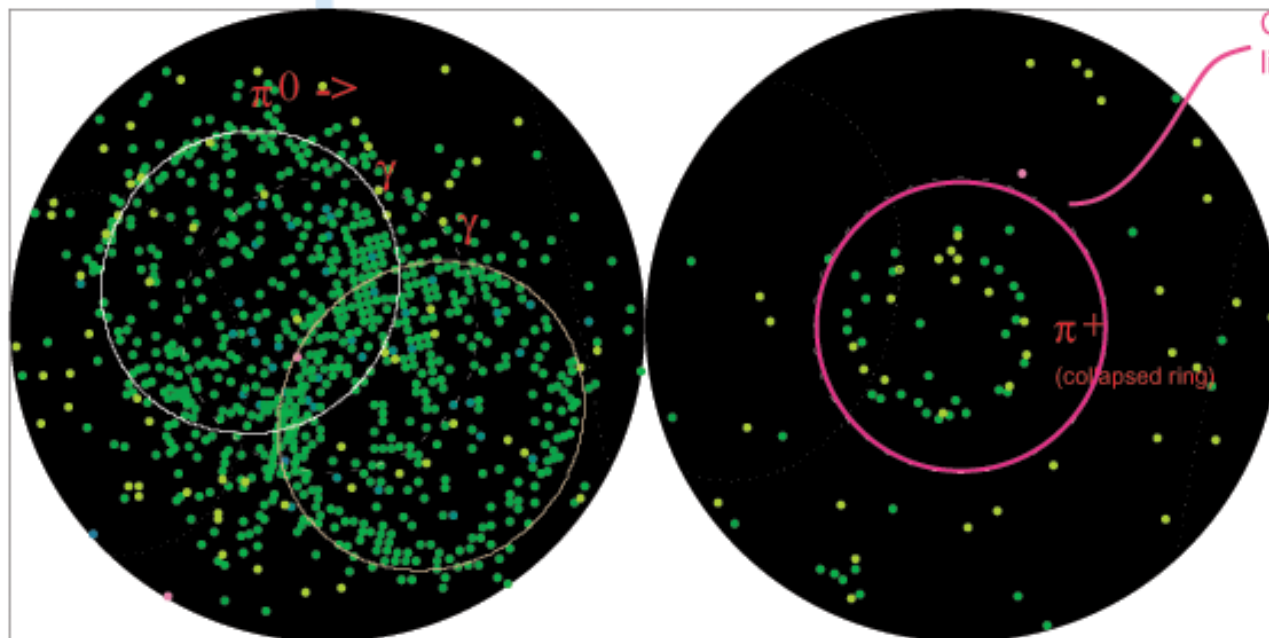
## Forward-backward hemisphere view of Monte Carlo event



$p \rightarrow \nu K^+ \rightarrow \pi^+ \pi^0$

Resid(ns)

- > 45
- 40- 45
- 34- 40
- 28- 34
- 22- 28
- 17- 22
- 11- 17
- 5- 11
- 0- 5
- -5- 0
- -11- -5
- -17- -11
- -22- -17
- -28- -22
- -34- -28
- < -34



(only hits in time window drawn)

expect only small amount of light outside backwards cone

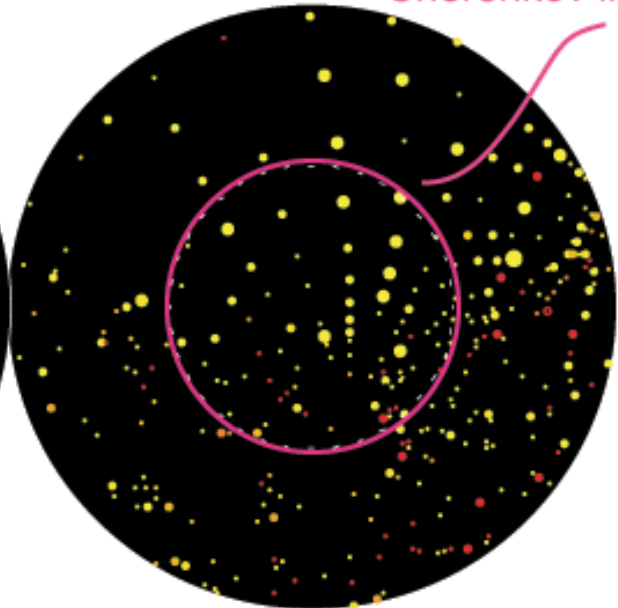
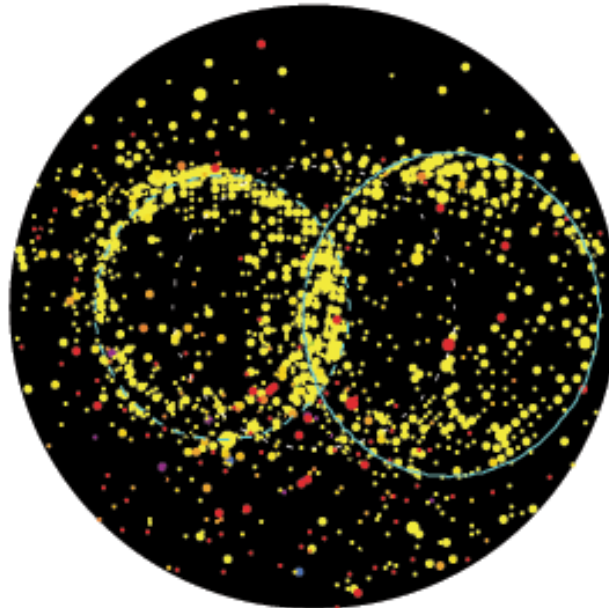
# Super-Kamiokande

Run 7944 Sub 203 Ev 27128713  
99-10-12:23:00:23  
Inner: 1572 hits, 2794 pE  
Outer: 3 hits, 3 pE (in-time)  
Trigger ID: 0x07  
D wall: 200.2 cm  
FC, mass = 141.3 MeV/c<sup>2</sup>

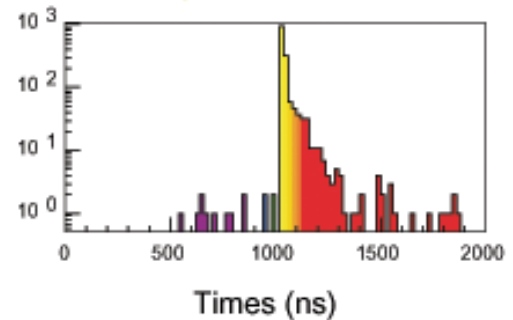
## Forward-backward hemisphere view of PMT hits as seen from reconstructed vertex

### Resid (ns)

- > 114
- 100- 114
- 85- 100
- 71- 85
- 57- 71
- 42- 57
- 28- 42
- 14- 28
- 0- 14
- -14- 0
- -28- -14
- -42- -28
- -57- -42
- -71- -57
- -85- -71
- < -85

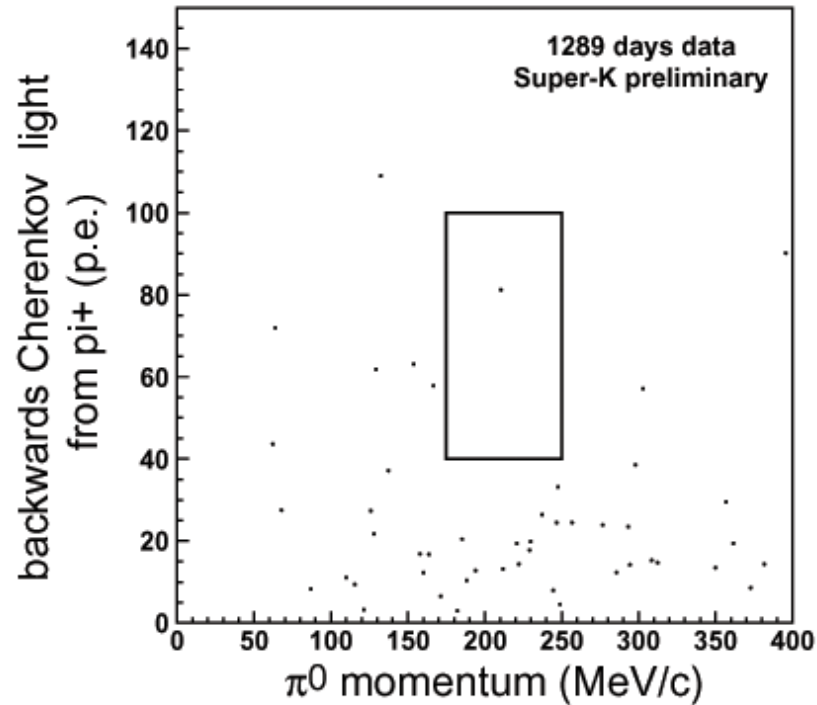


too much light outside of search cone opposite  $\pi^0$  momentum vector





Super-K Data:  $p \rightarrow K^+ \nu$   
 $K^+ \rightarrow \pi^+ \pi^0$



result for 79.3 kt·yr (Super-K preliminary):

1 candidates

6.8% efficiency

1.7 events background

$\tau/B (p \rightarrow K^+ \nu) > 5.9 \times 10^{32} \text{ yr}$  (90% C.L.)

**For final limit: combine all three results (they are independent)**

# Summary of Super-K Limits

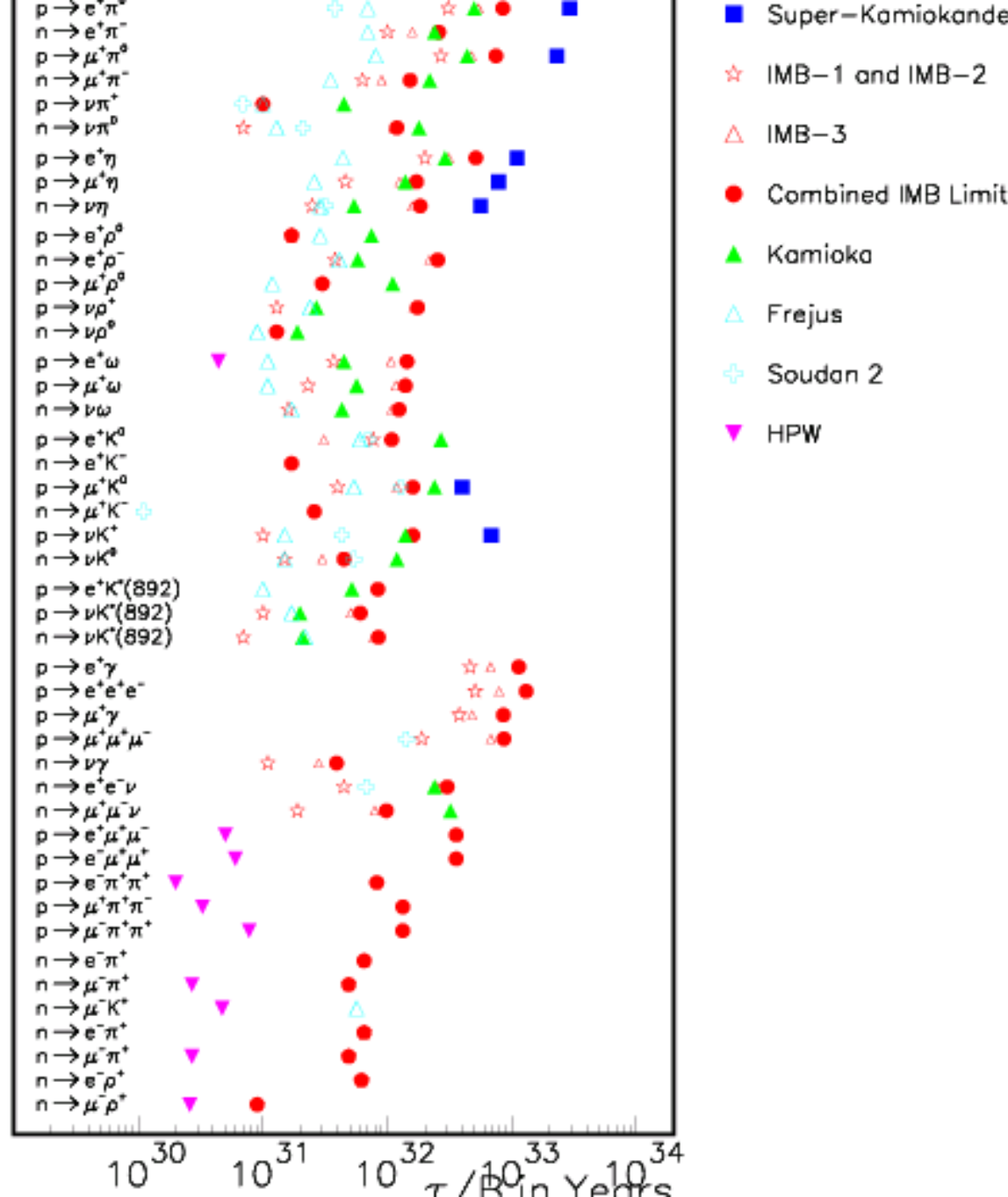
mode	exposure (kt• yr)	$\epsilon B_m$ (%)	observed event	B.G.	$\tau/B$ limit ( $10^{32}$ yrs)
$p \rightarrow e^+ + \pi^0$	79	43	0	0.2	50
$p \rightarrow \mu^+ + \pi^0$	79	32	0	0.4	37
$p \rightarrow e^+ + \eta$	45	17	0	0.3	11
$p \rightarrow \mu^+ + \eta$	45	12	0	0	7.8
$n \rightarrow \bar{\nu} + \eta$	45	21	5	9	5.6
$p \rightarrow e^+ + \rho$	61	6.8	0	0.6	6.1
$p \rightarrow e^+ + \omega$	61	3.3	0	0.3	2.9
$p \rightarrow e^+ + \gamma$	70	71	0	0.1	73
$p \rightarrow \mu^+ + \gamma$	70	60	0	0.2	61
$p \rightarrow \bar{\nu} + K^+$	79				16
$K^+ \rightarrow \nu \mu^+$ (spectrum)		33	--	--	4.4
prompt $\gamma + \mu^+$		8.8	0	0.5	10
$K^+ \rightarrow \pi^+ \pi^0$		6.8	1	1.7	5.9
$n \rightarrow \bar{\nu} + K^0$	79				3.0
$K^0 \rightarrow \pi^0 \pi^0$		9.6	25	33.8	3.2
$K^0 \rightarrow \pi^+ \pi^-$		4.6	10	6.7	1.1
$p \rightarrow e^+ + K^0$	70				5.4
$K^0 \rightarrow \pi^0 \pi^0$		11.8	1	1.4	8.8
$K^0 \rightarrow \pi^+ \pi^-$					
2-ring		6.2	6	1.0	1.5
3-ring		1.4	0	0.2	1.4
$p \rightarrow \mu^+ + K^0$	70				10
$K^0 \rightarrow \pi^0 \pi^0$		6.1	0	1.1	6.2
$K^0 \rightarrow \pi^+ \pi^-$					
2-ring		5.3	0	1.5	5.4
3-ring		2.8	1	0.2	1.8

# Nucleon Lifetime Limits

IMB: 45 decay modes

mass is everything,

MEGATON is needed



## Detector Technology: Liquid Argon

### liquid argon time projection chamber - Icarus

*everything charged visible... $3 \times 3 \times 0.6 \text{ mm}^3$  pixels*

300T half-module studied at surface...now in Gran Sasso

1.5 m drift, 1.8 ms lifetime (vs. 30 ms needed for scaling up to pdk)

2 x 1200T = 3 kT originally proposed for Gran Sasso  
safety a consideration

evaluation *in situ* underway

see muon decay pix

reconstruction of stopping muons and decay:

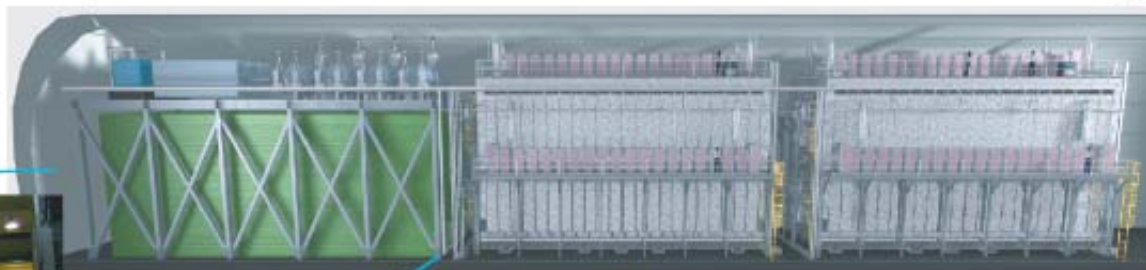
vertices of end of muon and beginning of e

dE/dx vs. range for stopping muons

cross-check with multiple scattering

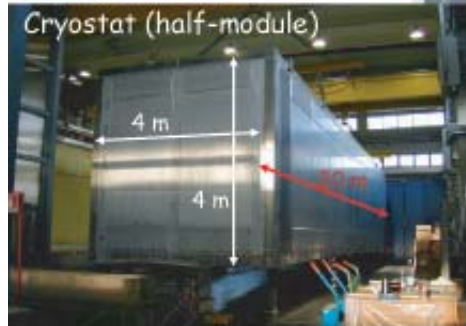
# ICARUS Experiment at Gran Sasso

← ~ 95 m →



A. Eridato  
NOON 2003

Cryostat (half-module)



## Liquid Argon TPC

- mm resolution
- 300 t module tested on surface
- 600 t module to be installed underground this year
- final proposed size 3 kt  
(relatively small mass but...)

**Fine grained reconstruction of neutrino interactions  
(atmospheric or CNGS beam)**



~100 evts/yr/600t from atmospheric  $\nu$   
~280 CC  $\nu_\tau$  in 3 kt, CNGS beam (5 yr)

*will detailed events reveal anything new?  
Liquid Argon TPC in near site of beams may  
provide valuable exclusive cross section data*

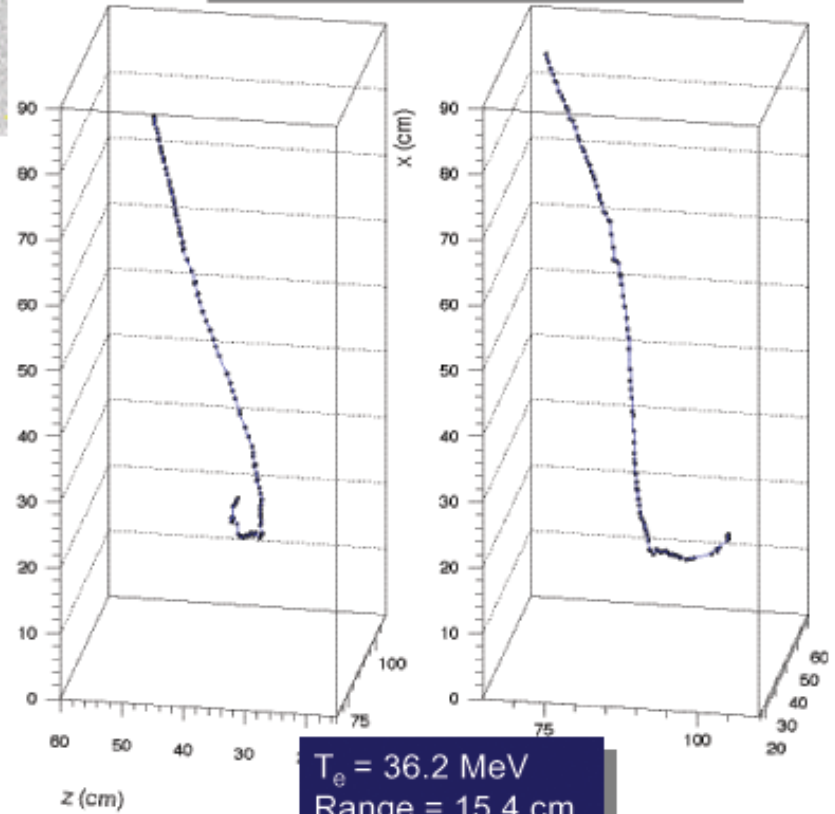
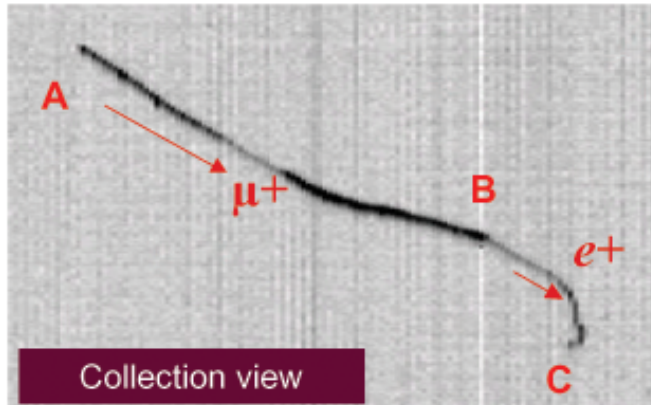
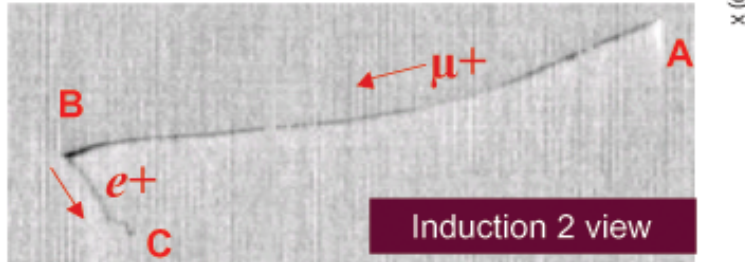
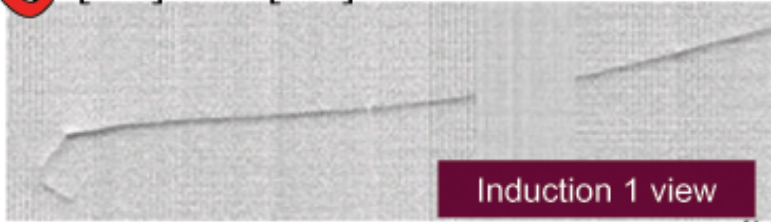
**Charge**

# 3D reconstruction of a stopping muon

$$\mu^+[AB] \rightarrow e^+[BC]$$

(Reconstruction is automatic)

Run 939 Event 95 Right chamber



## Detector Technology: Liquid Argon

scale up Icarus to LANNDD 70 kT for  $K^+$  detector

~ factor 6 higher efficiency than water for  $K^+$

...so cavern could be smaller *if only want to concentrate on  $K^+$*

$\Rightarrow$  420 kT effective mass of water (8 x Super-K total)

but efficiency for  $e^+ \pi^0$  and many other modes same as with water

liquid argon would need  $\times 6$  bigger cavity for a broad search of pdk

cost: \$200 M for the 70kT of liquid argon alone (420 kT too expensive?)

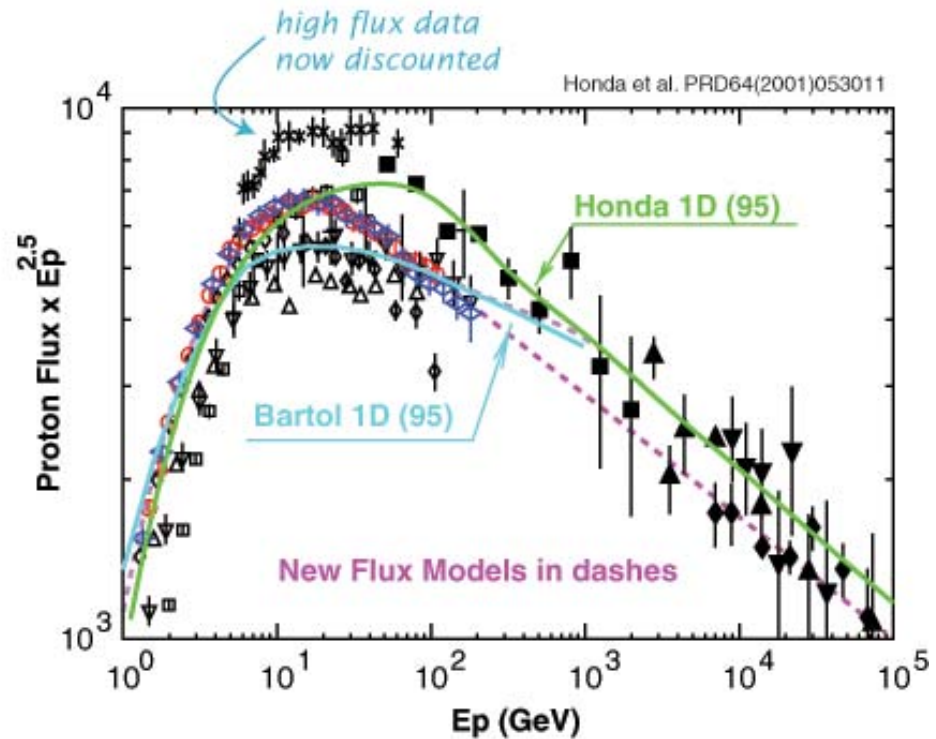
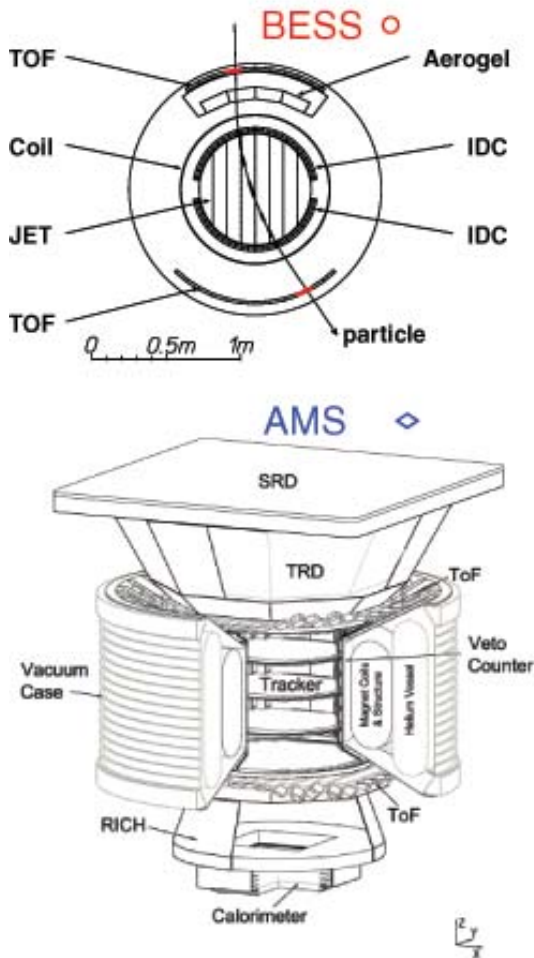
proposed sites: Frejus and WIPP in New Mexico

*LLANND*  
*Large Liquid Argon*  
*Neutrino*  
*and*  
*Nucleon Decay*  
*Detector*  
  
*70 kT magnetized*





# New Precise Data on Primary Cosmic Ray Flux



most important  
for atmospheric  $\nu$

Note: Honda and Bartol neutrino fluxes generally agreed with each other (despite primary flux difference) due to different hadronic models.

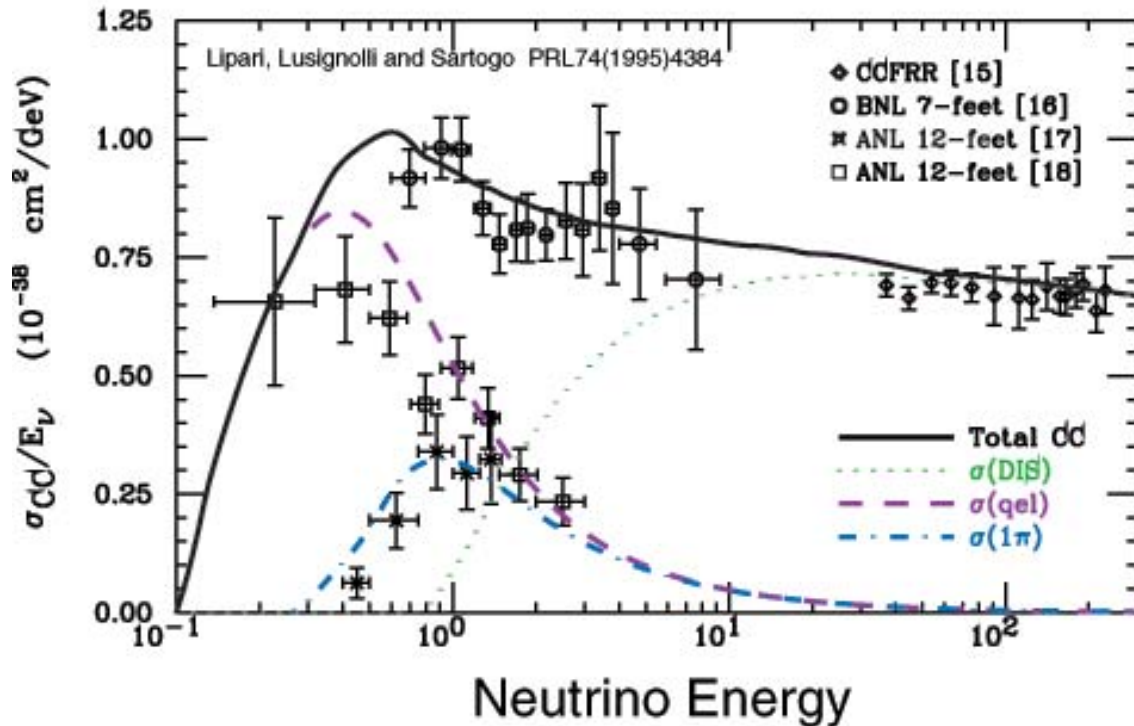
# Neutrino Cross Sections

Long baseline near detector data (K2K) influence atm  $\nu$  Monte Carlo:

$$M_A(QE) = 1.11 \text{ GeV} \quad M_A(1p) = 1.21 \text{ GeV}$$

K2K oscillation analysis insensitive

atmospheric neutrino analysis under study



## Quasi-Elastic

$$\nu_{\mu} N \rightarrow \mu^{-} N'$$

V-A

Llewellyn-Smith 1972

## Single Pion

$$\nu_{\mu} N \rightarrow \mu^{-} N' \pi$$

resonance production

Rein & Seghal 1981

## Coherent Pion (not shown)

$$\nu_{\mu} {}^{16}O \rightarrow \mu^{-} X \pi$$

Marteau et al.

## Deep Inelastic Scattering

$$\nu_{\mu} N \rightarrow \mu^{-} N' \text{ hadrons}$$

GRV94 parton distribution

with Bodek 2001

## Nuclear Effects

Fermi motion

Pauli blocking

Nuclear rescattering

From Ed Kearns

*meanwhile in America?*

## Two Neutrino/Astro Physics Planning Processes

American Physical Society Neutrino Study...soon to be public      Freedman & Kayser

set priorities for experiments: astro, reactor, solar & atmospheric, etc.

highest long term: massive detector for pdk + oscillations, highest

highest medium term: NOVA, off-axis neutrino oscillations  
double beta decay @ 100 kg level

SF Initiative: A “Deep Engineering and Science Underground Laboratory”      Sadoulet

8 potential sites: Anderson, Henderson, Homestake, Icicle Creek,  
Kimburton, San Jacinto, Soudan, WIPP

3 solicitations for proposals:

1) site independent physics justification and requirements, funded 10 days ago

2) site specific proposals, due 28/2/05, 3-5 to be chosen ~7/05

likely at least one massive & one deep (geo- and bio- want more)

3) full proposals due 2/06... Goal: funding in ‘09 presidential budget  
staged, \$1-1.5 B expected for detector + proton driver + superbeam  
likely will include possibility of international site for massive detector

beside geography,

why so many facilities?

for Megaton...

the bigger the better.

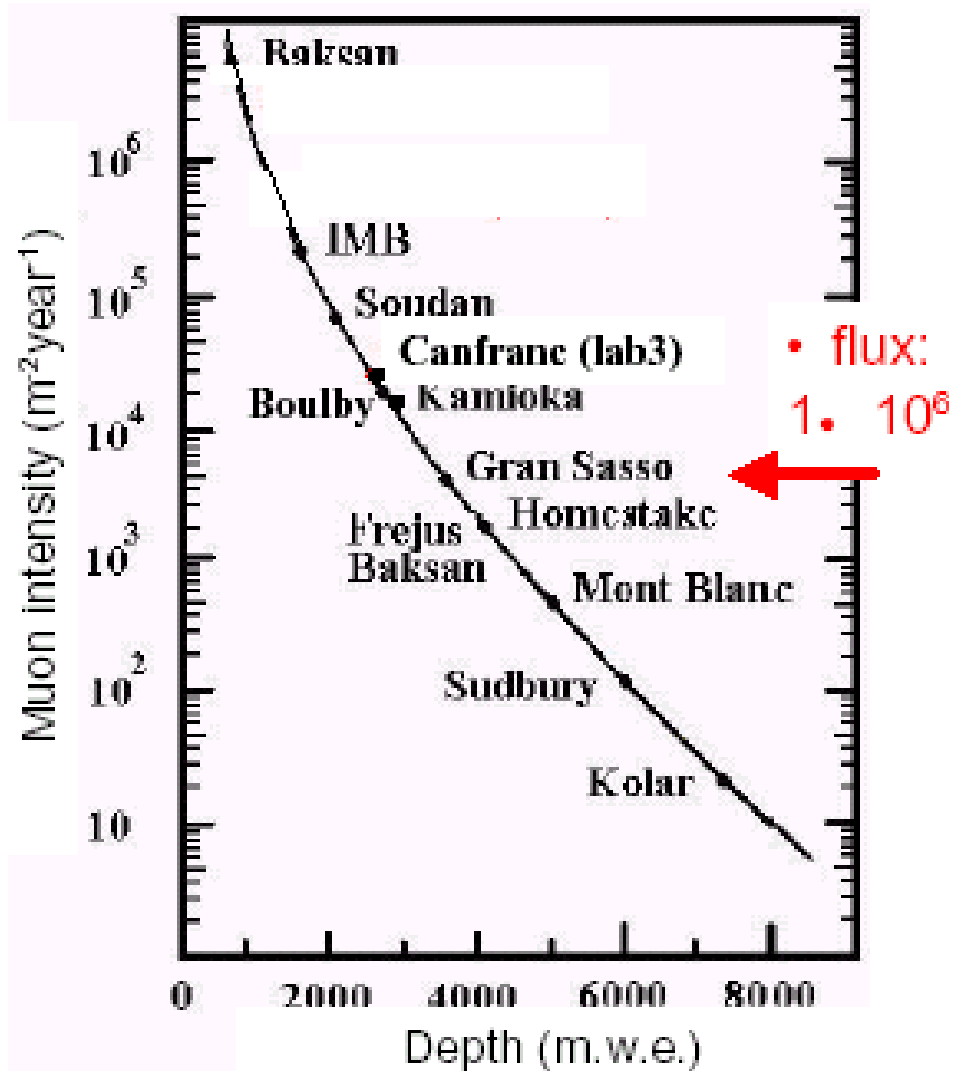
for solar and

background sensitive exp'ts

the deeper, the better

e.g. impossible for IMB

consider SNO-Lab at Sudbury



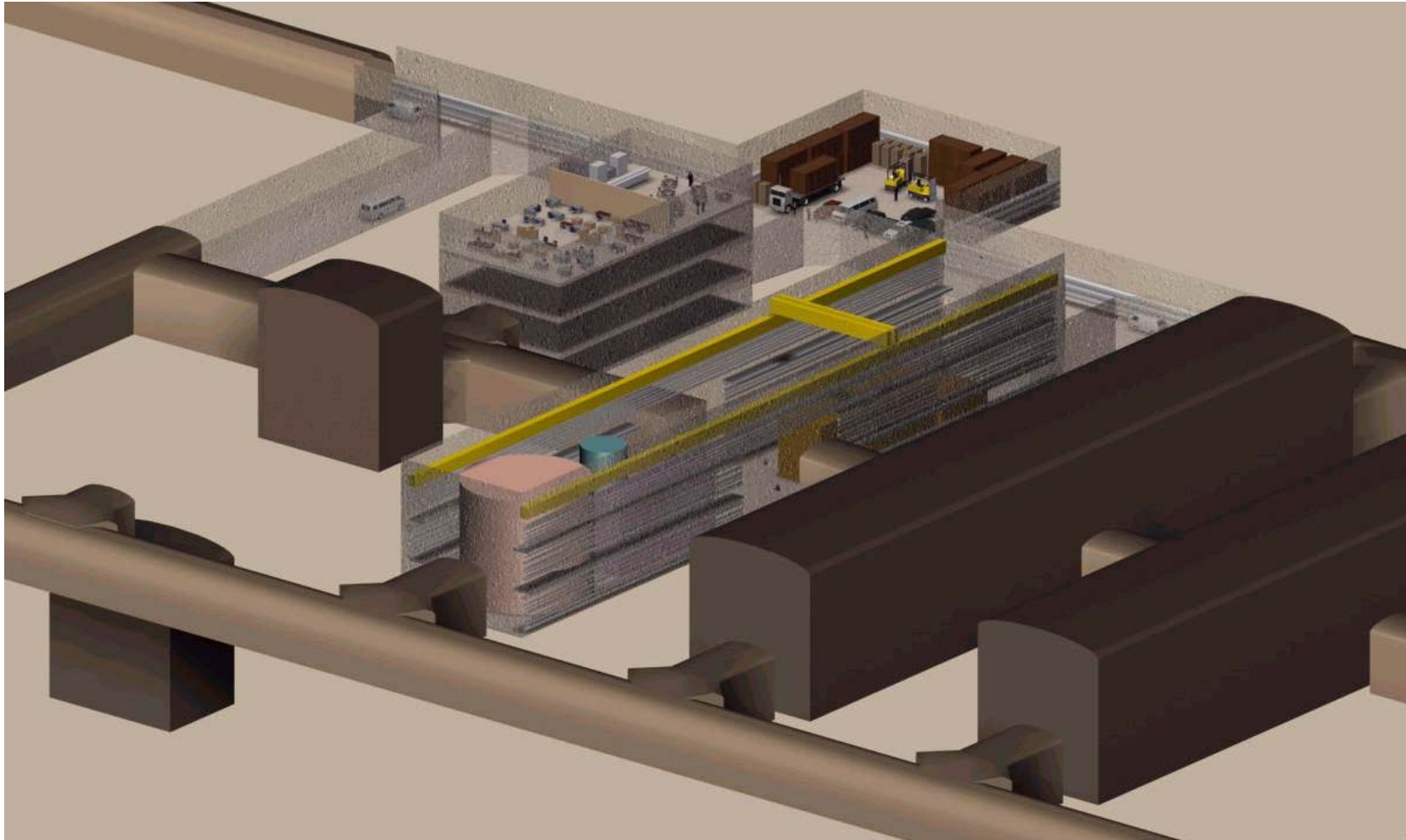
# Possible long baseline proposals in US...



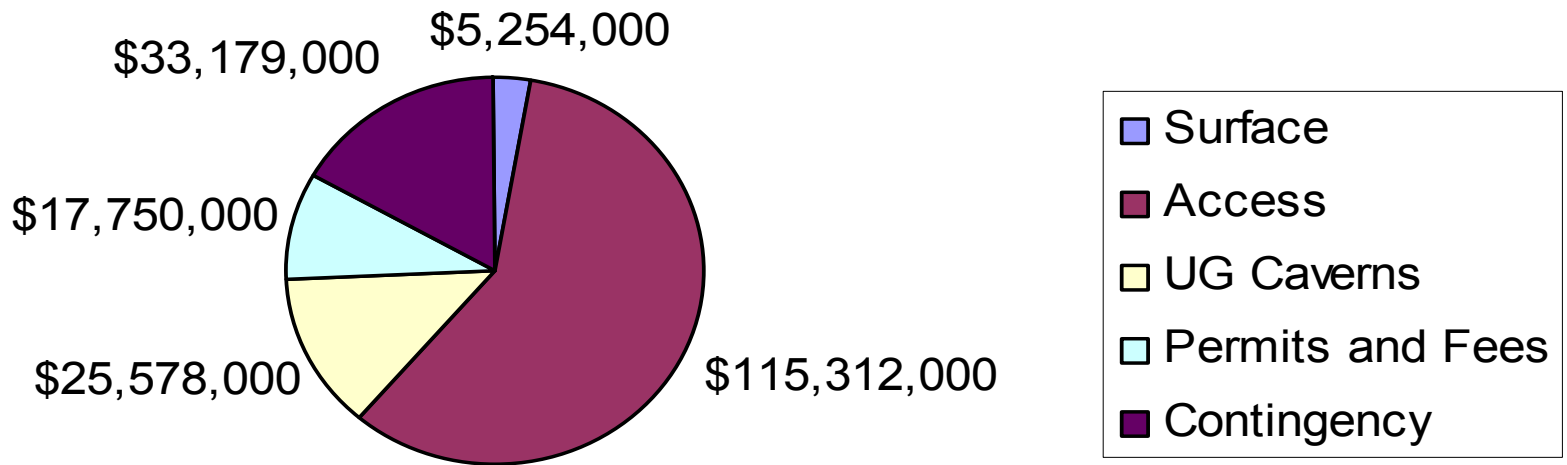
*San Jacinto,  
California  
proposal...*



## Underground Laboratory Complex



## Sample Cost Breakdown (8,000 feet overburden)



Estimated total cost \$200 million



## *San Jacinto proposal...*

### Estimated Operating Costs

WBS 1-Fees	\$100,000
WBS 2-Utility Costs	\$2,604,000
WBS 3-Maintenance	\$560,000
WBS 4-Equipment & Transportation	\$225,000
WBS 5-Staff	\$1,080,000
WBS 6- Programs	\$160,000
WBS 7-Outside Costs & Subcontracts	\$638,000
<b>Total</b>	<b>\$5,367,000</b>

\*

# San Jacinto proposal...

## Project Schedule

Task	Year 1	Year 2	Year 3	Year 4	Year 5	Year 6
Project Start	█					
Preliminary Env. Assessment	█					
Schematic Design	█					
Phase 1 Site Investigation	█					
Draft EIR/EIS / Schematic Design Rpt		█				
Design Development		█				
Phase 2 Site Investigation		█				
Environmental Review & Response Period		█				
EIR/EIS Record of Decision			█			
Const. Documents & Bidding			█			
Construction Authorization				█		
Mobilization & Site Work				█		
TBM Procurement				█		
Tunnel 1 Construction				█		
Cavern A					█	
Cavern B					█	
Support Caverns					█	
Outfitting						█
Beneficial Occupancy						█

*San Jacinto proposal...*

*California funding for*

*detailed engineer costing...*

	1 Tun	2 Tun	1 Tun	2 Tun	1 Tun	2 Tun
<b>1 Land Acquisition, Easements &amp; Usage Fee</b>	<b>\$0</b>	<b>\$0</b>	<b>\$0</b>	<b>\$0</b>	<b>\$0</b>	<b>\$0</b>
1.1. Surface Land Costs						
1.2. Underground Rights Costs						
1.3. Road Easements						
1.4. Utility Easements						
1.5. Public/Private Road Fees						
<b>2 Surface</b>	<b>\$5,254,356</b>	<b>\$5,254,356</b>	<b>\$5,254,356</b>	<b>\$5,254,356</b>	<b>\$5,254,356</b>	<b>\$5,254,356</b>
2.1 Access roads	\$29,356	\$29,356	\$29,356	\$29,356	\$29,356	\$29,356
2.2 Surface Infrastructure						
2.2.1 Electrical and substation	\$105,000	\$105,000	\$105,000	\$105,000	\$105,000	\$105,000
2.2.3 Water	\$50,000	\$50,000	\$50,000	\$50,000	\$50,000	\$50,000
2.2.4 Sewer	\$50,000	\$50,000	\$50,000	\$50,000	\$50,000	\$50,000
2.2.5 Communications	\$100,000	\$100,000	\$100,000	\$100,000	\$100,000	\$100,000
2.3 Buildings						
2.3.1 Building 1-Visitor's Center & Administration	\$2,550,000	\$2,550,000	\$2,550,000	\$2,550,000	\$2,550,000	\$2,550,000
2.3.3 Building 3-Warehouse & Assembly	\$1,350,000	\$1,350,000	\$1,350,000	\$1,350,000	\$1,350,000	\$1,350,000
2.3.4 Building 4-Laboratories	\$1,020,000	\$1,020,000	\$1,020,000	\$1,020,000	\$1,020,000	\$1,020,000
<b>3 Underground Access</b>	<b>\$115,312,744</b>	<b>\$162,217,558</b>	<b>\$118,421,321</b>	<b>\$172,773,234</b>	<b>\$115,272,839</b>	<b>\$166,605,446</b>
3.1 Portal(s)						
3.1.1 Access tunnel portal	\$3,252,000	\$4,407,000	\$3,252,000	\$4,407,000	\$3,252,000	\$4,407,000
3.2 Tunnel(s)						
3.2.1 Access Tunnel	\$56,413,837	\$51,127,005	\$58,732,214	\$53,228,114	\$56,413,837	\$51,127,005
3.2.2 Egress Tunnel	\$0	\$51,461,023	\$0	\$53,228,114	\$0	\$51,461,023
3.2.3 Cross Cuts or Turnouts	\$1,950,000	\$1,950,000	\$1,950,000	\$2,100,000	\$1,950,000	\$1,950,000
3.2.4 Mechanical	\$44,404,290	\$47,835,290	\$45,194,490	\$49,921,490	\$44,404,290	\$47,835,290
3.2.5 Electrical	\$4,451,275	\$0	\$4,451,275	\$4,451,275	\$4,451,275	\$4,451,275
3.2.6 Fire Protection	\$3,830,400	\$3,830,400	\$3,830,400	\$3,830,400	\$3,830,400	\$3,830,400
3.3 Surface haulage	\$1,010,942	\$1,606,840	\$1,010,942	\$1,606,840	\$971,037	\$1,543,440
<b>4 Underground Facilities</b>	<b>\$25,578,546</b>	<b>\$25,578,546</b>	<b>\$30,453,600</b>	<b>\$30,453,600</b>	<b>\$16,987,289</b>	<b>\$16,987,289</b>
4.1 Caverns						
4.1.1 Common Area Cavern	\$1,506,221	\$1,506,221	\$1,506,221	\$1,506,221	\$1,506,221	\$1,506,221
4.1.2 Utility Cavern	\$853,444	\$853,444	\$853,444	\$853,444	\$853,444	\$853,444
4.1.3 Experimental Cavern A	\$4,190,846	\$4,190,846	\$4,190,846	\$4,190,846	\$2,109,741	\$2,109,741
4.1.4 Experimental Cavern B	\$4,989,999	\$4,989,999	\$4,989,999	\$4,989,999	\$0	\$0
4.1.5 Experimental Cavern C	\$0	\$0	\$4,190,846	\$4,190,846	\$0	\$0
4.1.7 Refuge Cavern	\$156,320	\$156,320	\$156,320	\$156,320	\$156,320	\$156,320
4.1.8 Sump	\$156,320	\$156,320	\$156,320	\$156,320	\$156,320	\$156,320
4.2 Tunnels						
4.2.1 Main Street Tunnel	\$1,074,918	\$1,074,918	\$1,330,156	\$1,330,156	\$819,680	\$819,680
4.2.2 Connecting Tunnels	\$2,177,548	\$2,177,548	\$2,504,253	\$2,504,253	\$1,074,918	\$1,074,918
4.3 Underground Infrastructure						
4.3.1 Groundwater Drainage	\$500,000	\$500,000	\$500,000	\$500,000	\$500,000	\$500,000
4.3.2 Mechanical	\$2,717,033	\$2,717,033	\$2,717,033	\$2,717,033	\$2,717,033	\$2,717,033
4.3.3 Electrical	\$632,100	\$632,100	\$632,100	\$632,100	\$632,100	\$632,100
4.3.4 Fire protection	\$1,213,038	\$1,213,038	\$1,213,038	\$1,213,038	\$1,213,038	\$1,213,038
4.3.5 Security	\$0	\$0	\$0	\$0	\$0	\$0
4.3.6 Assembly Areas	\$116,880	\$116,880	\$116,880	\$116,880	\$116,880	\$116,880
4.3.7 Steel & concrete structures	\$5,000,000	\$5,000,000	\$5,000,000	\$5,000,000	\$5,000,000	\$5,000,000
4.4 Surface Haulage	\$293,880	\$293,880	\$396,145	\$396,145	\$131,595	\$131,595
<b>5 Permits, Fees and Professional Services</b>	<b>\$17,750,811</b>	<b>\$23,370,946</b>	<b>\$18,707,410</b>	<b>\$25,219,856</b>	<b>\$16,716,625</b>	<b>\$22,867,289</b>
5.1 Professional Services						
5.1.1 Site Investigation	\$2,432,552	\$3,202,729	\$2,563,643	\$3,456,101	\$2,290,828	\$3,133,770
5.1.2 Schematic Design	\$1,185,165	\$1,560,404	\$1,249,034	\$1,683,850	\$1,116,116	\$1,526,770
5.1.3 Design Development	\$1,925,893	\$2,535,656	\$2,029,681	\$2,736,255	\$1,813,688	\$2,481,000
5.1.4 Construction Documents	\$3,407,350	\$4,486,161	\$3,590,973	\$4,841,067	\$3,208,833	\$4,389,400
5.1.5 Construction Engineering Services	\$5,481,389	\$7,216,867	\$5,776,783	\$7,787,804	\$5,162,036	\$7,061,300
5.1.6 Site Characterization During Construction	\$740,728	\$975,252	\$780,646	\$1,052,406	\$697,572	\$954,200
5.1.7 Environmental Studies	\$740,728	\$975,252	\$780,646	\$1,052,406	\$697,572	\$954,200
5.1.8 Cultural Studies	\$207,404	\$273,071	\$218,581	\$294,674	\$195,320	\$267,100
5.1.9 Public Affairs	\$148,146	\$195,050	\$156,129	\$210,481	\$139,514	\$190,800
5.2 In-House Services, Permits, Owner's Rep.	\$1,481,456	\$1,950,505	\$1,561,293	\$2,104,812	\$1,395,145	\$1,908,400
<b>6 Environmental Mitigation</b>	<b>\$2,000,000</b>	<b>\$2,000,000</b>	<b>\$2,000,000</b>	<b>\$2,000,000</b>	<b>\$2,000,000</b>	<b>\$2,000,000</b>
6.1 Environmental Mitigation	\$2,000,000	\$2,000,000	\$2,000,000	\$2,000,000	\$2,000,000	\$2,000,000
<b>Subtotal</b>	<b>\$165,896,458</b>	<b>\$218,421,406</b>	<b>\$174,836,687</b>	<b>\$235,701,046</b>	<b>\$156,231,109</b>	<b>\$213,714,333</b>
Contingency 20%	\$33,179,292	\$43,684,281	\$34,967,337	\$47,140,209	\$31,246,222	\$42,742,800
<b>Total</b>	<b>\$199,075,749</b>	<b>\$262,105,687</b>	<b>\$209,804,025</b>	<b>\$282,841,255</b>	<b>\$187,477,331</b>	<b>\$256,457,133</b>

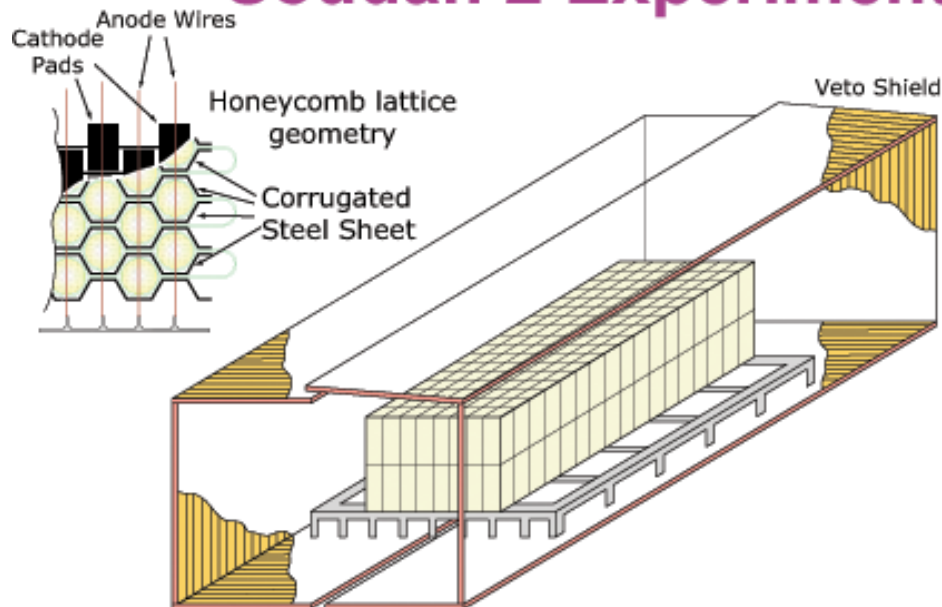


# Proton Decay Detector at Soudan

---

- Water Cerenkov detector
- Believe key to progress is improved photodetectors
- Solid state photodetectors address two significant problems of PMT's: pressure limitation and radioactivity of glass
- Also yield unclear advantage in greater pixelation
- Goal is  $\sim \$1/\text{cm}^2$  with 70% QE (CMOS process)
- Conceptual design is vertical cylinder 50 m in diameter with heights up to 100 m ( $\sim 0.2$  MT)
- Multiple volumes provide required mass

# Soudan 2 Experiment

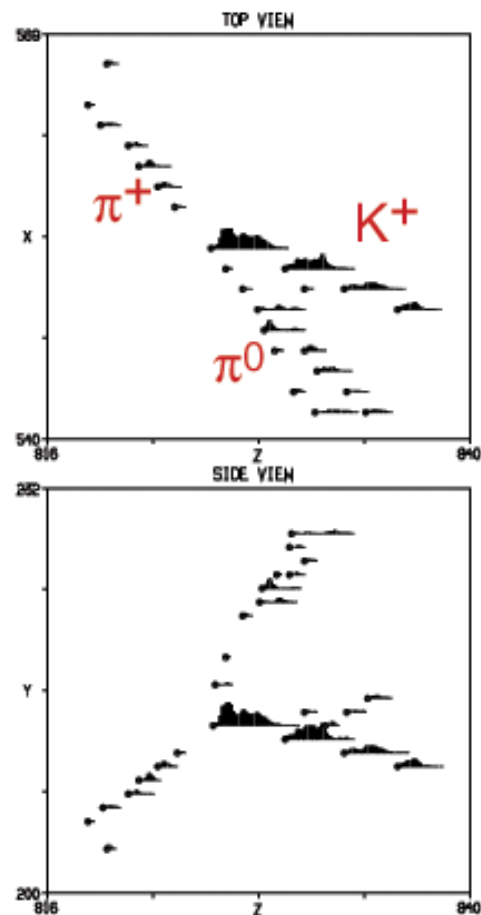


## tracking iron tracking calorimeter

Soudan Mine (Minnesota), 2100 m.w.e., 770 ton (fiducial)  
1 cm spatial resolution with dE/dx sampling

suitable for: non-relativistic particles ( $K^+$ )  
high final state multiplicities

however: greater intranuclear scattering than water  
smaller in size due to cost and complexity



Soudan 2 proton decay M.C.



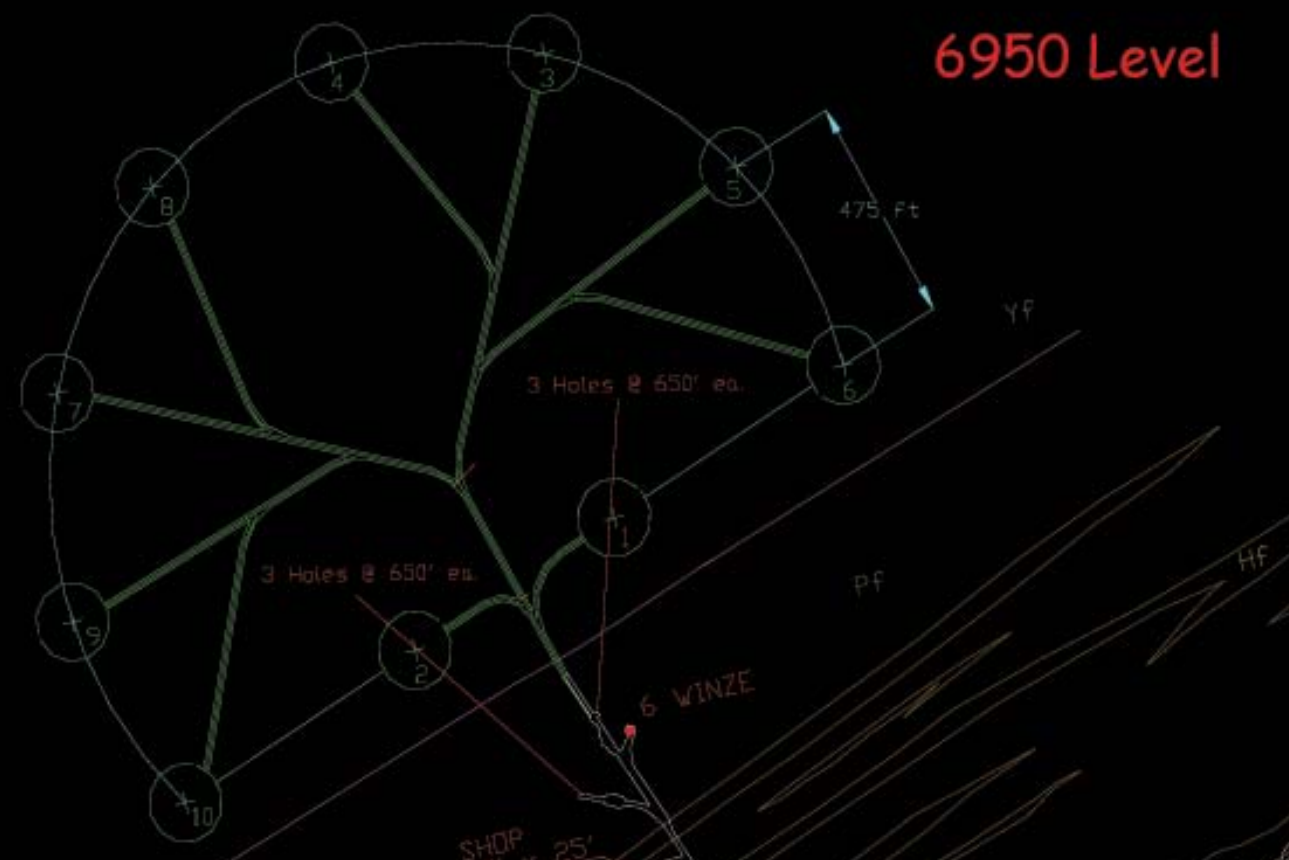
# Proton Decay Detector at Soudan

---

- Build Phase 1 detector 50 m diameter, 50 m high (2 Super-K's) at 700 m depth
- Use Phase 1 detector to develop techniques, software etc.
- Use Phase 1 detector to do physics in coincidence with beam (depth not a problem)
- Extend lab to 2,500 m (8,200 feet)
- Build 5 Phase 2 Detectors at 1,500 to 2,500 m

# MEGATON MODULAR MULTI-PURPOSE DETECTOR

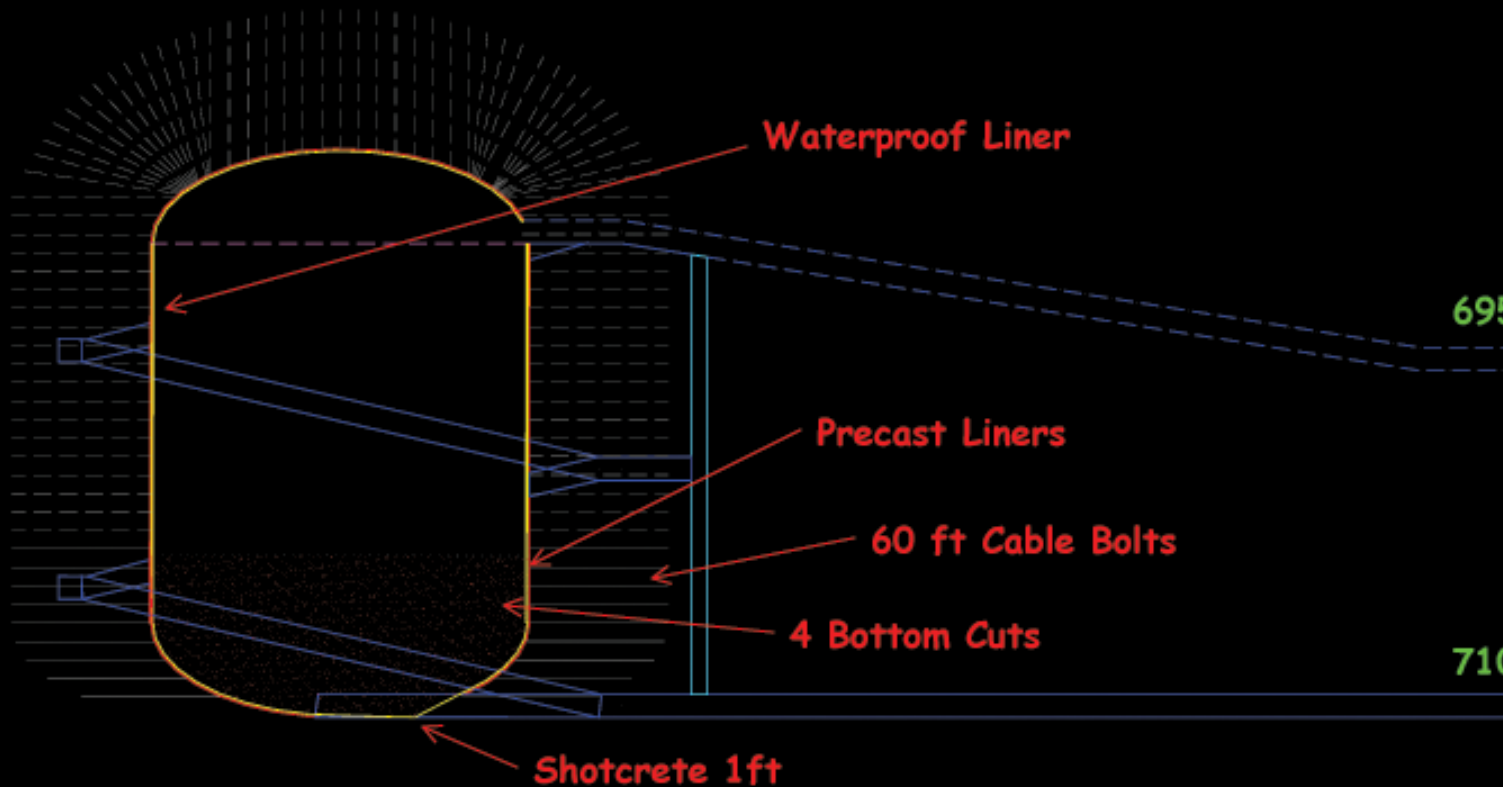
## ✓ Modular Configuration



# MEGATON MODULAR MULTI-PURPOSE DETECTOR

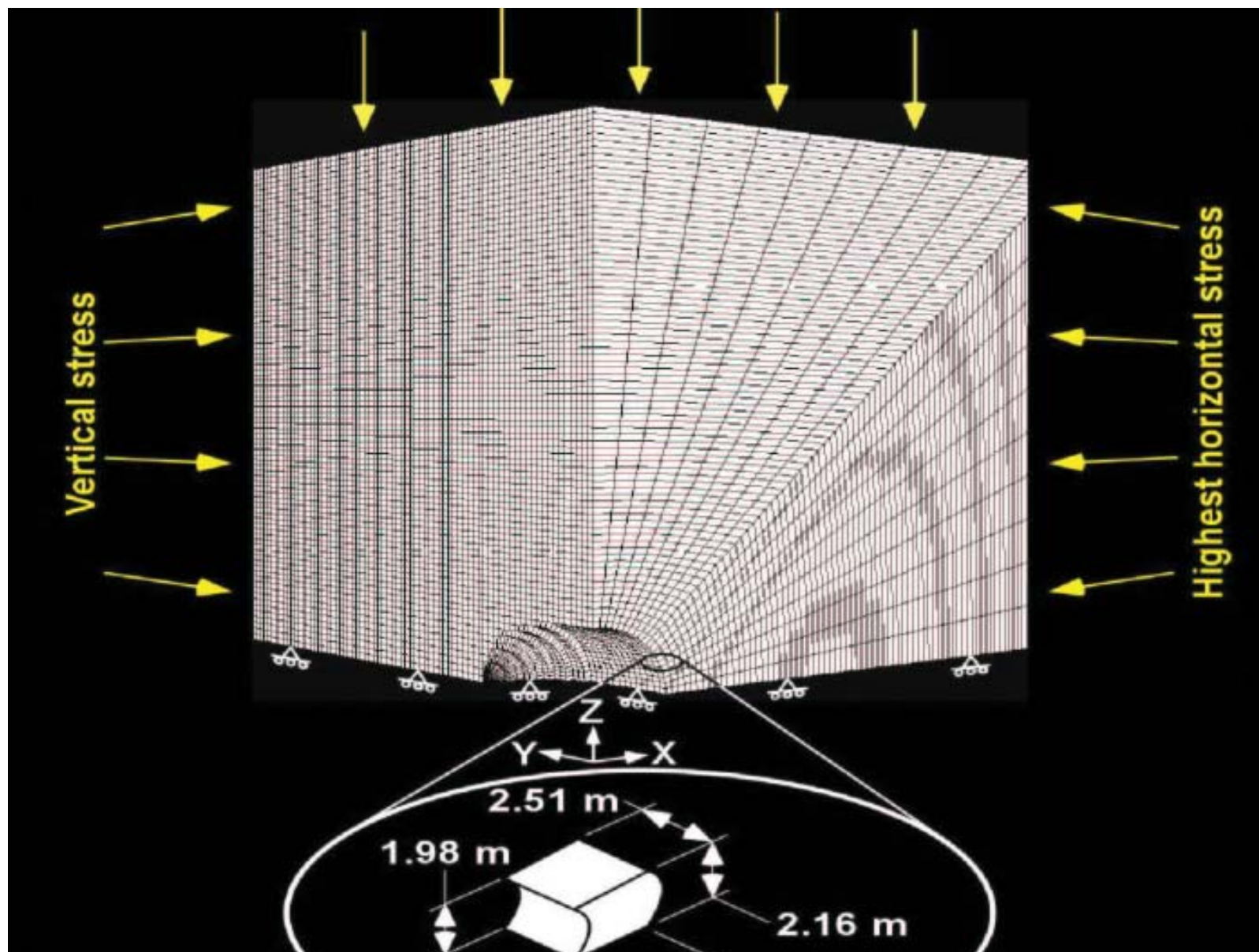
## ✓ Estimated Timeline

Year Four





*Homestake proposal...detail geological engineering in deep site ...*



## Homestake proposal...

### MEGATON MODULAR MULTI-PURPOSE DETECTOR

#### ✓ Estimated Excavation Costs (\$MM)

	# Of Chambers	1	3
⇒ Labor & Benefits		\$ 5.51	\$ 10.94
⇒ Mining & Construction			
★ Equipment Operation		\$ 1.30	\$ 3.89
★ Supplies		\$ 4.51	\$ 13.35
★ Precast Concrete Liner		\$ 3.25	\$ 9.75
⇒ Other (Outside Contractor)		\$ 0.12	\$ 0.37
⇒ 15% Contingency		\$ 2.20	\$ 5.74
<b>TOTAL</b>		<b>\$ 16.89</b>	<b>\$ 44.04</b>

**Features of the Design:**

- 1) Module dimensions 50 m diameter x 50 m high  
- 100 kilotons water
- 2) Parallel construction of modules – one “crew” can build 3 modules in parallel- completion in 4-5 years
- 3) Multiple crews can work at same time – , 2 crews can build 600 kilotons of detector in 4-5 years.
- 4) Concrete liner in each module provides strength and stability, smooth surface for water tight plastic inner liner, surface for photomultiplier installation guides.
- 5) Low construction cost

**physics motivation strong for PDK search**

unification theories severely constrained

synergism with neutrino superbeams,  $\nu$  factories

**preparatory work in progress**

K2K (data now finished) and other precision studies of  $\nu$  interactions coming  
detailed understanding of cross sections and neutrino background,

especially from neutrino-induced single pion modes

vigorous r & d for detector options...phototubes, liquid argon at Gran Sasso

if S-K gets one candidate for PDK, first put scintillator in Super-K?

then make big water detector?

if SUSY, found at LHC, make detector dedicated to  $K^+$  ?

**next generation detector**

build water Cherenkov to largest size limited by atmospheric  $\nu$  background

consistent with geology

## why CPPM? why build Megaton at Frejus?

potential CERN neutrino beam

2 hours from infrastructure of CERN, as well as

12 HEP labs of CEA/CNRS, INFN, and CH with couple of hours

CPPM one of closest in France

For low energy neutrino beam...water detector most economical, ideal L/E

CPPM experience: extensive water Cherenkov : Antares, Km<sup>3</sup>

high numbers of pm's, optimizing electronics, underwater cabling, etc.

Limiting investment: the cavern dominates

in size, in cost ...moving from 20 m, to 40 m, to 60 m

Italy: drilling tunnels...a national pastime

long Gran Sasso experience

France: world renown for grand civil engineering challenges

...highest bridge, biggest caverns for LHC

CERN: 150 M € / year available for new initiatives starting '09?

EU framework projects...MegaTonne a unique, continent wide project

~20 year program, 1 – 1.5 B €, should involve most of Europe