

KamLAND (Anti-Neutrino Status)

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in Astroparticle and Underground Physics

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Itaru Shimizu (Tohoku Univ.)

KamLAND Collaboration

T. Ebihara,¹ S. Enomoto,¹ K. Furuno,¹ Y. Gando,¹ K. Ichimura,¹ H. Ikeda,¹ K. Inoue,¹ Y. Kibe,¹ Y. Kishimoto,¹ M. Koga,¹ Y. Konno,¹ A. Kozlov,¹ Y. Minekawa,¹ T. Mitsui,¹ K. Nakajima,¹, K. Nakajima,¹ K. Nakamura,¹ K. Owada,¹ I. Shimizu,¹ J. Shirai,¹ F. Suekane,¹ A. Suzuki,¹ K. Tamae,¹ S. Yoshida,¹ J. Busenitz,² T. Classen,² C. Grant,² G. Keefer,² D.S. Leonard,² D. McKee,² A. Piepke,² M.P. Decowski,³ S.J. Freedman,³ B.K. Fujikawa,³ F. Gray,³, L. Hsu,³, R. Kadel,³ K.-B. Luk,³ H. Murayama,³ T. O'Donnell,³ H.M. Steiner,³ L.A. Winslow,³ D.A. Dwyer,⁴ C. Jillings,⁴, C. Mauger,⁴ R.D. McKeown,⁴ C. Zhang,⁴ B.E. Berger,⁵ C.E. Lane,⁶ J. Maricic,⁶ T. Miletic,⁶ M. Batygov,⁷ J.G. Learned,⁷ S. Matsuno,⁷ S. Pakvasa,⁷ J. Foster,⁸ G.A. Horton-Smith,⁸ A. Tang,⁸ S. Dazeley,⁹, K. Downum,¹⁰ G. Gratta,¹⁰ K. Tolich,¹⁰ W. Bugg,¹¹ Y. Efremenko,¹¹ Y. Kamyshev,¹¹ O. Perevozchikov,¹¹ H.J. Karwowski,¹² D.M. Markoff,¹² W. Tornow,¹² K. M. Heeger,¹³ F. Piquemal,¹⁴ and J.-S. Ricol¹⁴

(KamLAND Collaboration)



¹Research Center for Neutrino Science, Tohoku University, Sendai 980-8578, Japan

²Department of Physics and Astronomy, University of Alabama, Tuscaloosa, Alabama 35487, USA

³Physics Department, University of California at Berkeley and

Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

⁴W. K. Kellogg Radiation Laboratory, California Institute of Technology, Pasadena, California 91125, USA

⁵Department of Physics, Colorado State University, Fort Collins, Colorado 80523, USA

⁶Physics Department, Drexel University, Philadelphia, Pennsylvania 19104, USA

⁷Department of Physics and Astronomy, University of Hawaii at Manoa, Honolulu, Hawaii 96822, USA

⁸Department of Physics, Kansas State University, Manhattan, Kansas 66506, USA

⁹Department of Physics and Astronomy, Louisiana State University, Baton Rouge, Louisiana 70803, USA

¹⁰Physics Department, Stanford University, Stanford, California 94305, USA

¹¹Department of Physics and Astronomy, University of Tennessee, Knoxville, Tennessee 37996, USA

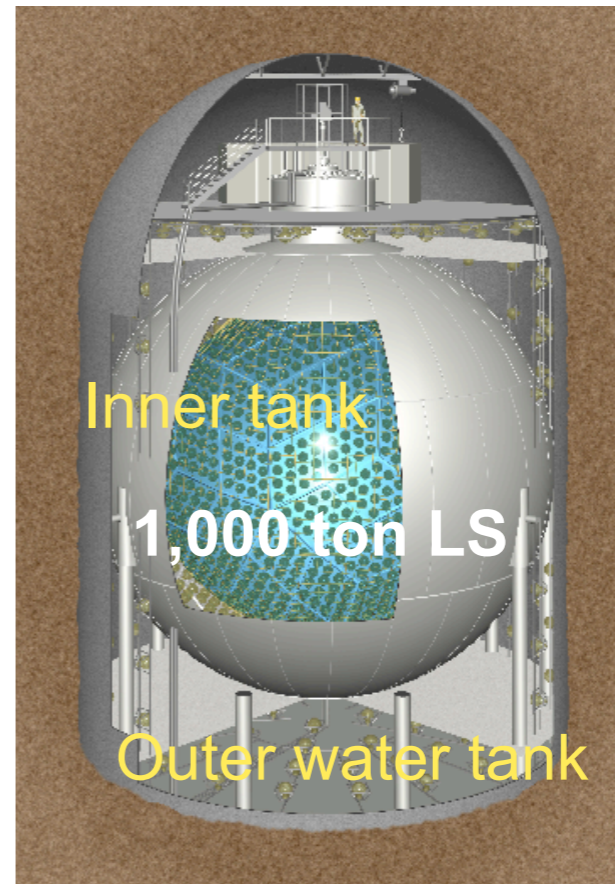
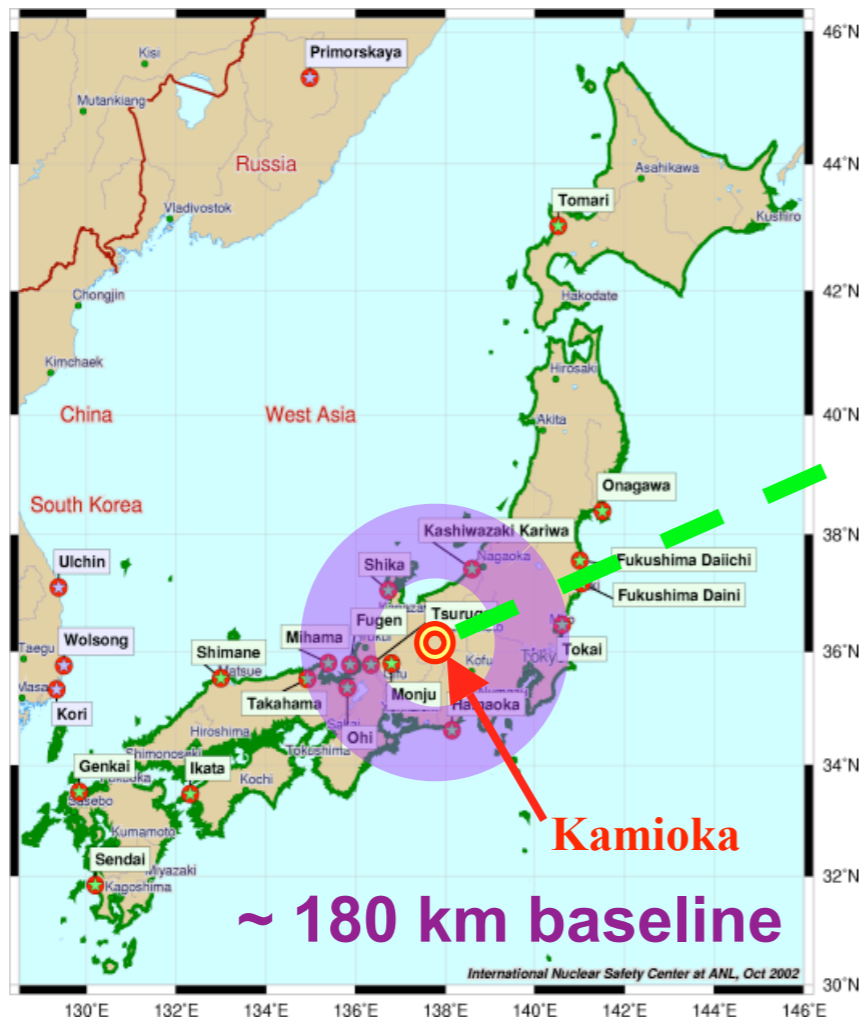
¹²Triangle Universities Nuclear Laboratory, Durham, North Carolina 27708, USA and Physics Departments at Duke University, North Carolina State University, and the University of North Carolina at Chapel Hill

¹³Department of Physics, University of Wisconsin, 1150 University Avenue, Madison, WI 53706, USA

¹⁴CEN Bordeaux-Gradignan, IN2P3-CNRS and University Bordeaux I, F-33175 Gradignan Cedex, France

KamLAND

Kamioka Liquid Scintillator Anti-Neutrino Detector



34% photo-coverage with
1325 17" and 554 20" PMTs

2 flavor neutrino oscillation

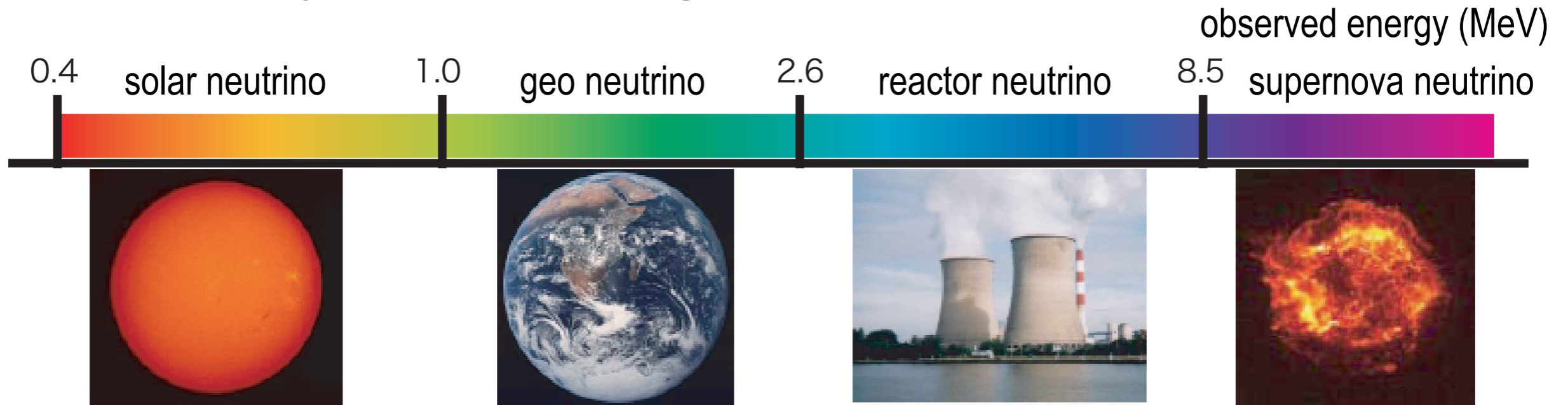
$$P(\nu_e \rightarrow \nu_e) = 1 - \sin^2 2\theta \sin^2\left(\frac{1.27\Delta m^2 [eV^2] l [m]}{E [MeV]}\right)$$

most sensitive region

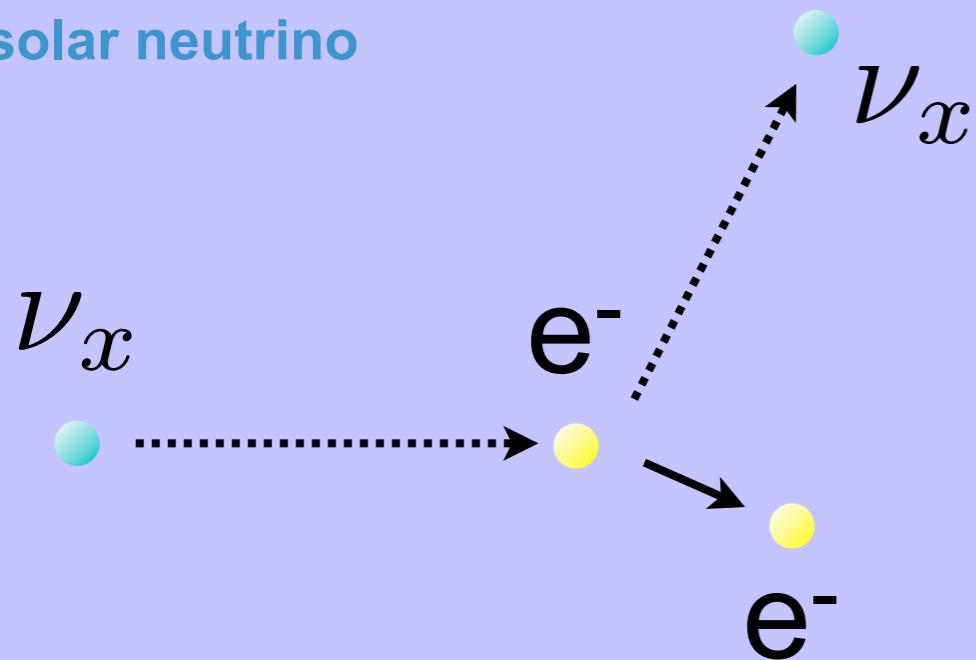
$$\Delta m^2 = (1/1.27) \cdot (E [MeV] / L [m]) \cdot (\pi/2) \\ \sim 3 \times 10^{-5} eV^2$$

reactor neutrino : sensitive to LMA solution

Physics Target in KamLAND

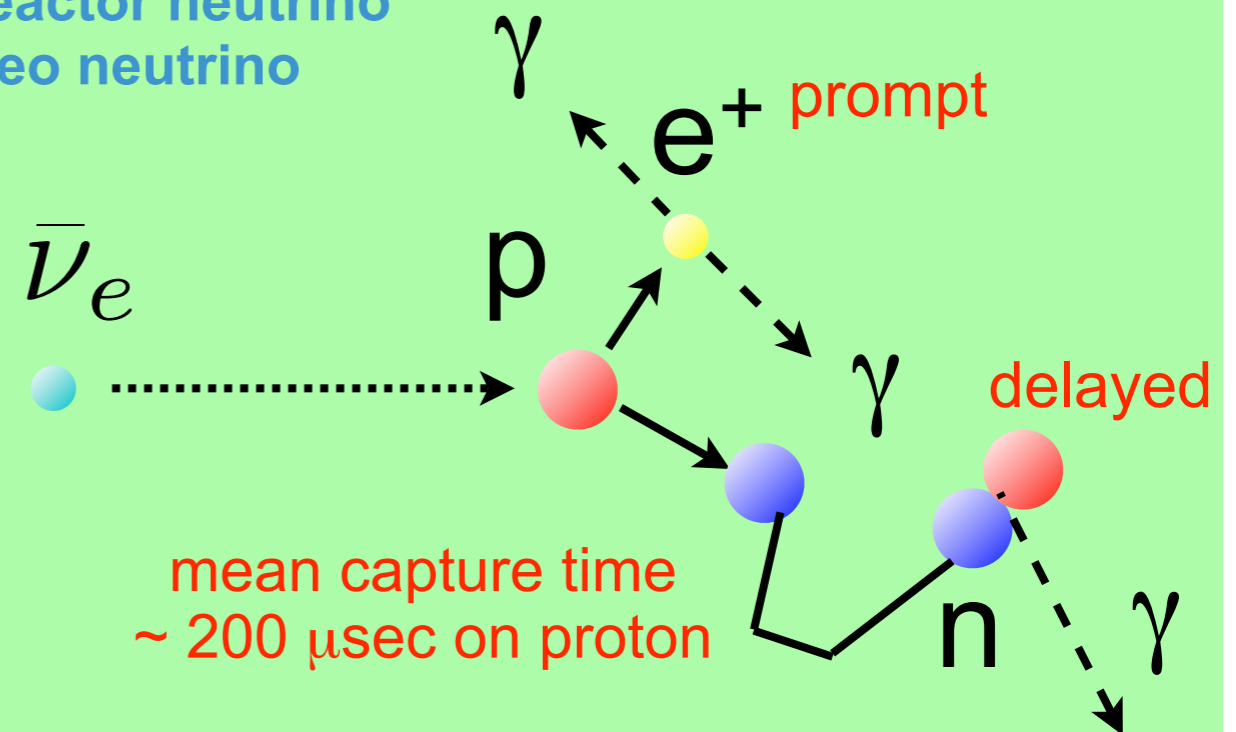


solar neutrino



neutrino detection by electron scattering

reactor neutrino geo neutrino



anti-neutrino detection by inverse beta-decay

Reactor and Geo Neutrino Analysis

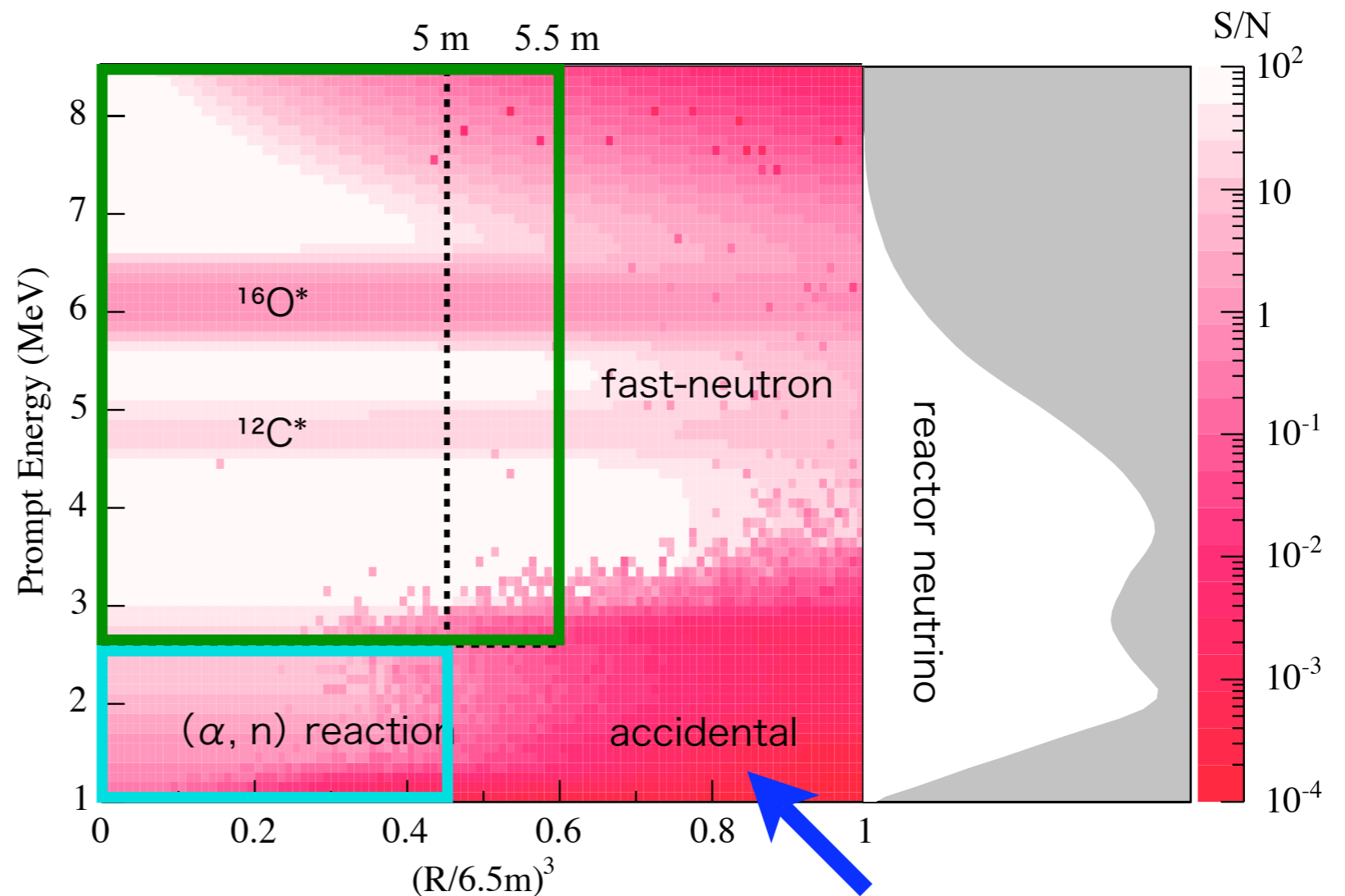
previous result

S / B ratio map (energy v.s. radius)

separated analysis window for reactor and geo neutrinos

reactor neutrino
(2.6 - 8.5 MeV, R 5.5 m)

geo neutrino
(0.9 - 2.6 MeV, R 5.0 m)



large accidental B.G.
caused by external γ -rays

Analysis improvement

- (1) efficient **accidental** background rejection
- (2) combined analysis of **reactor** and **geo** neutrinos

Anti-Neutrino Event Selection

(a) Accidental B.G. discrimination

discriminator based on 5 parameters (E_d , ΔR , ΔT , R_p , R_d)

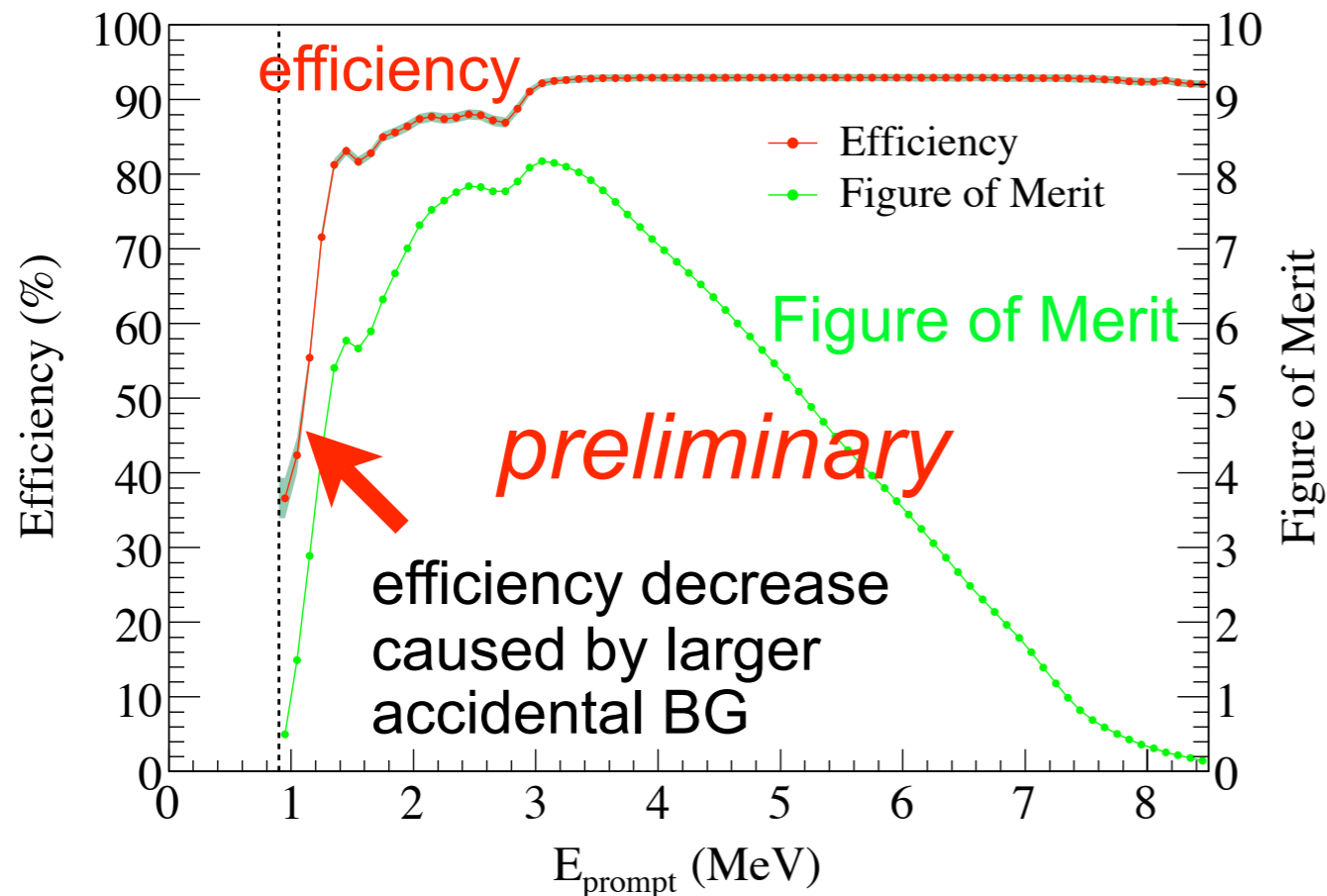
$$L_{\text{ratio}} = \frac{f_{\bar{\nu}}}{f_{\bar{\nu}} + f_{\text{accidental}}} \quad f : \text{PDF}$$

Selection : Maximize "Figure of Merit" $\frac{S}{\sqrt{S + B_{\text{accidental}}}}$

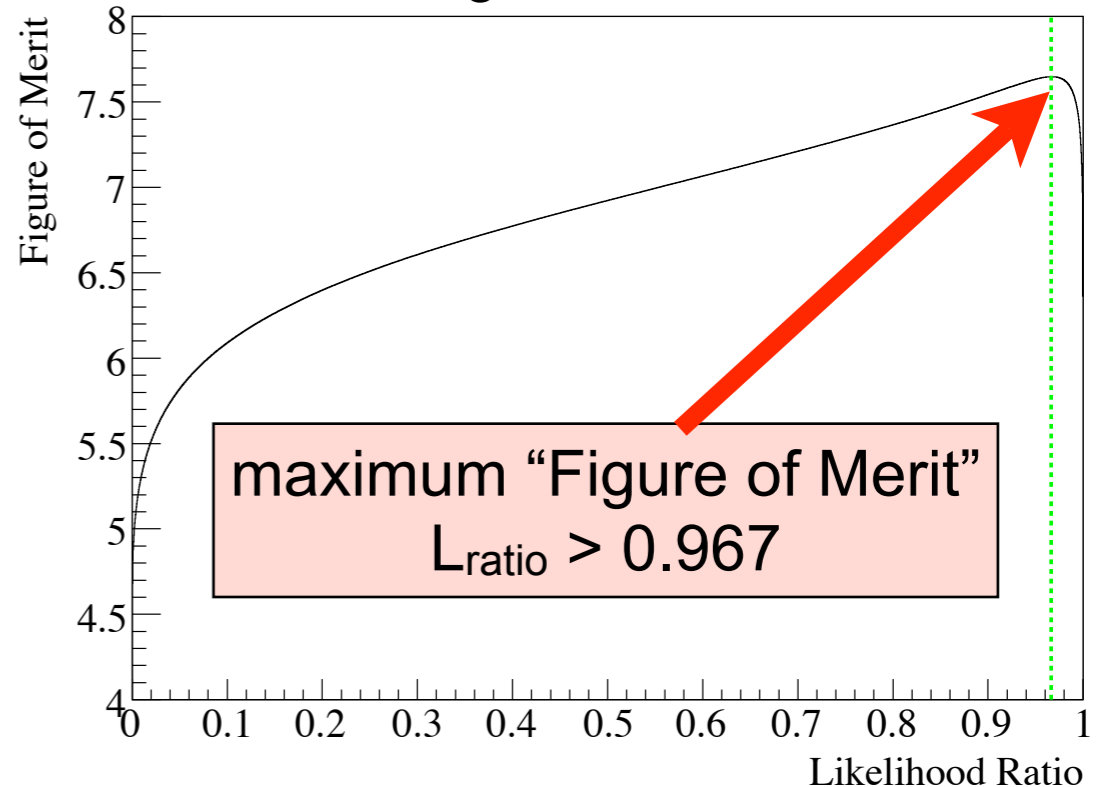
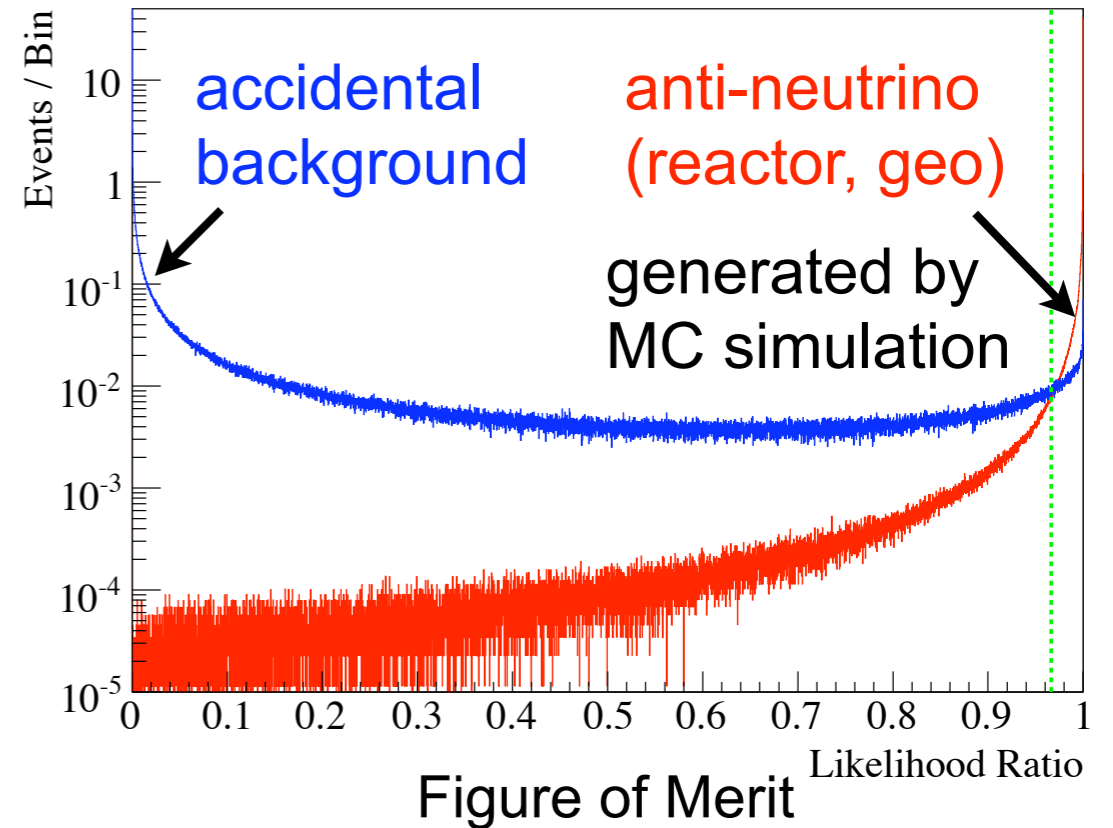
(b) μ spallation cut

- $\Delta T_{\mu} > 2$ s after showing μ ($\Delta Q > 10^6$ p.e.)
- $\Delta T_{\mu} > 2$ s or $\Delta L > 3$ m after non-showering μ

Detection efficiency



$2.2 < E_{\text{prompt}} < 2.3$ MeV



Systematic Uncertainty

“full volume” calibration lowered the fiducial volume error

(4.7% in previous analysis)

preliminary

Detector related

Reactor related

Fiducial volume

1.8%

$\bar{\nu}_e$ spectra

2.4%

Energy scale

1.5%

Reactor power

2.1%

L-selection eff.

0.6%

Fuel composition

1.0%

OD veto

0.2%

Long-lived nuclei

0.3%

Cross section

0.2%

Time lag

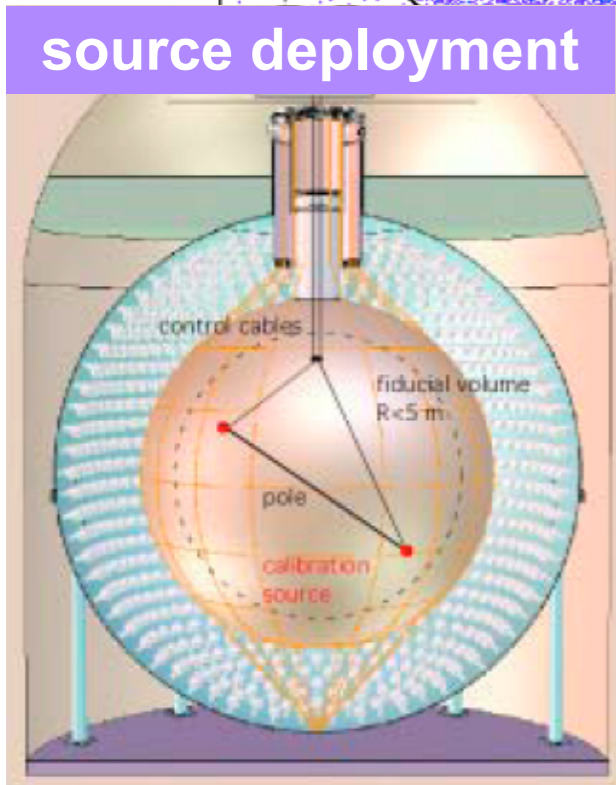
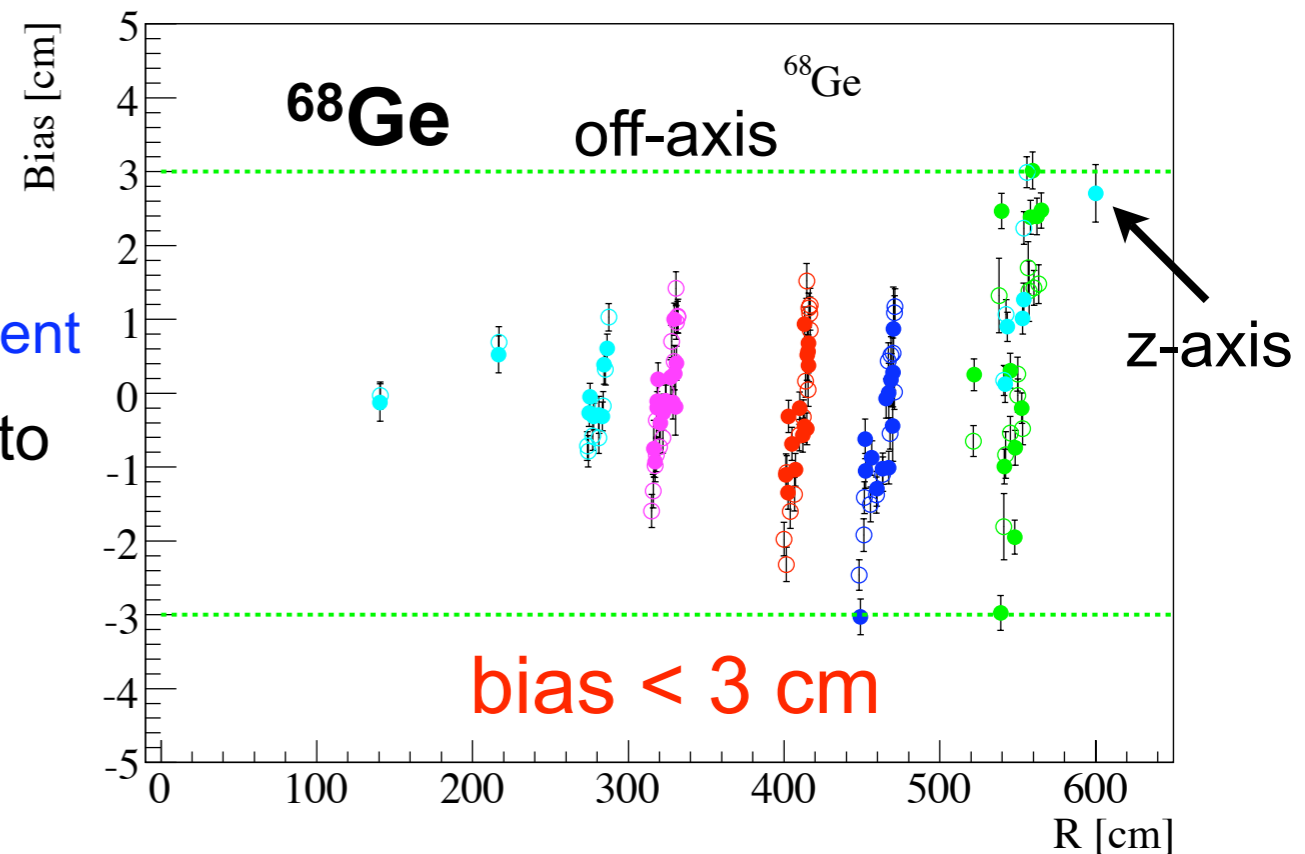
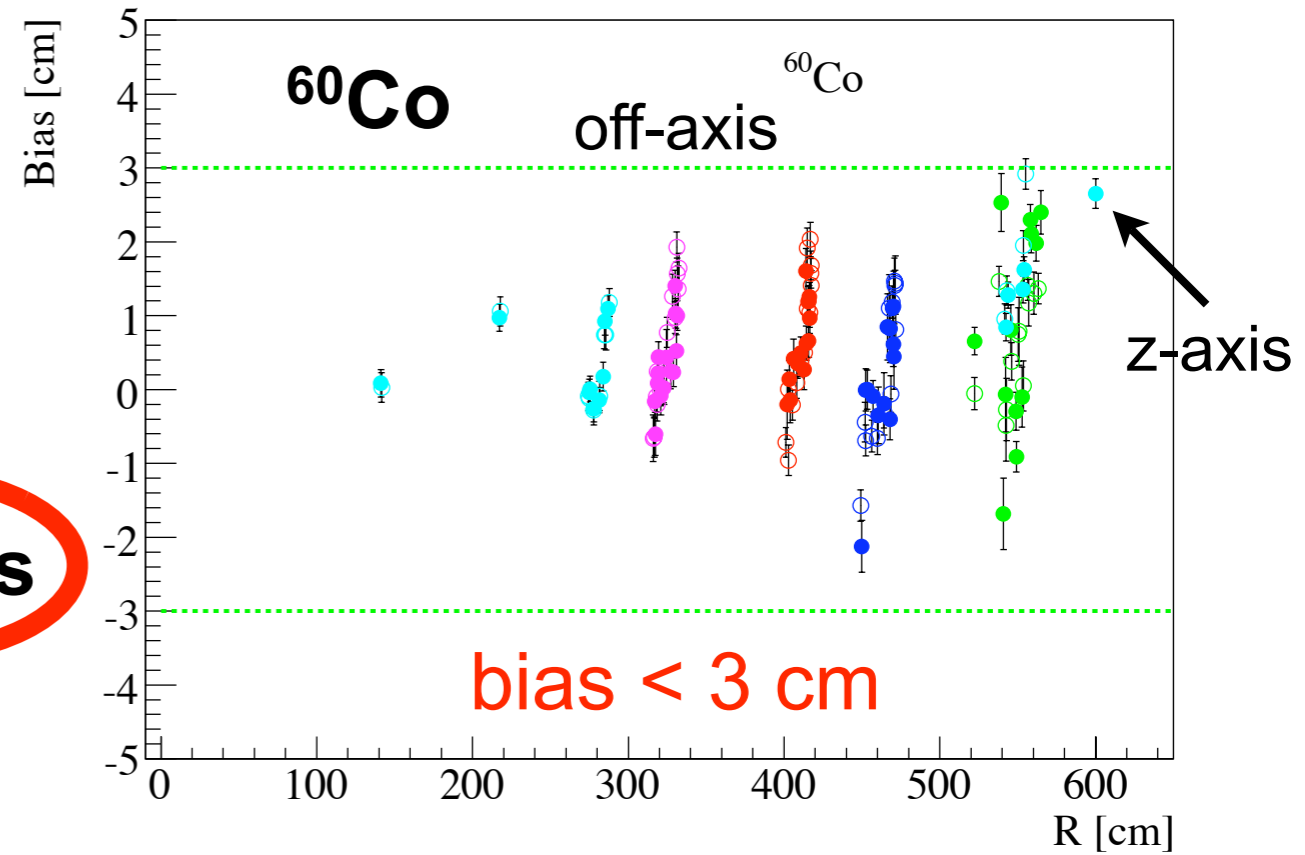
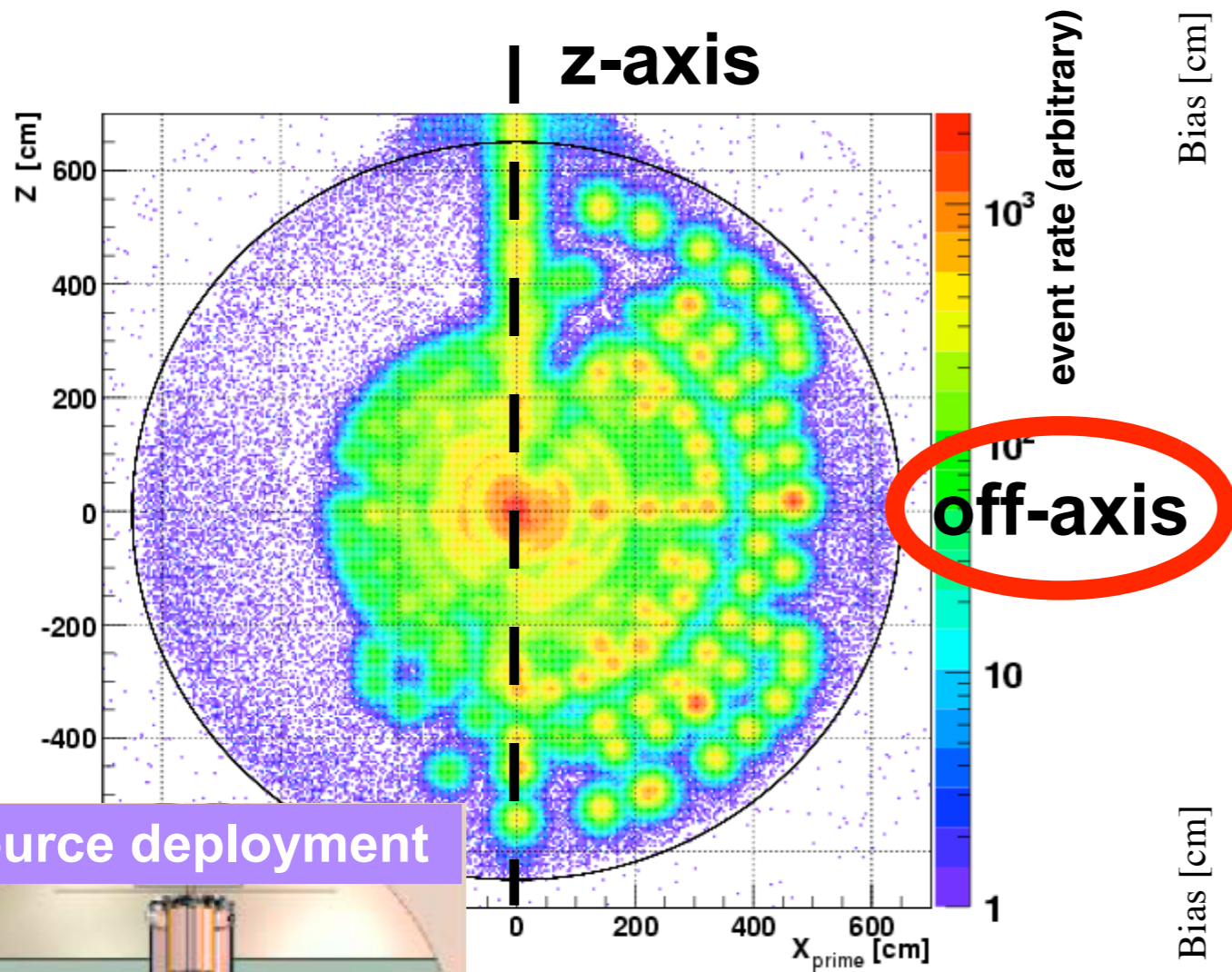
0.01%

2.4%

3.4%

Total systematic uncertainty : 4.1%

Full Volume Calibration

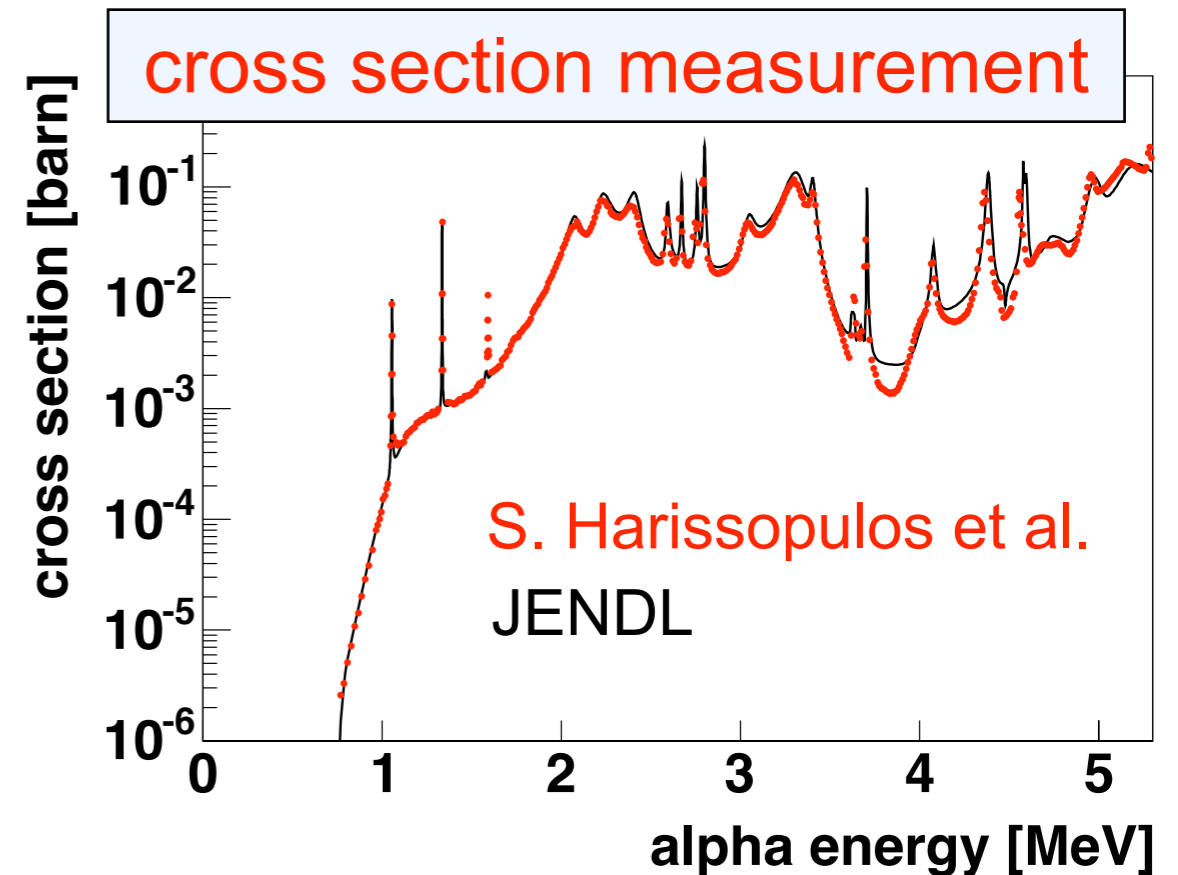
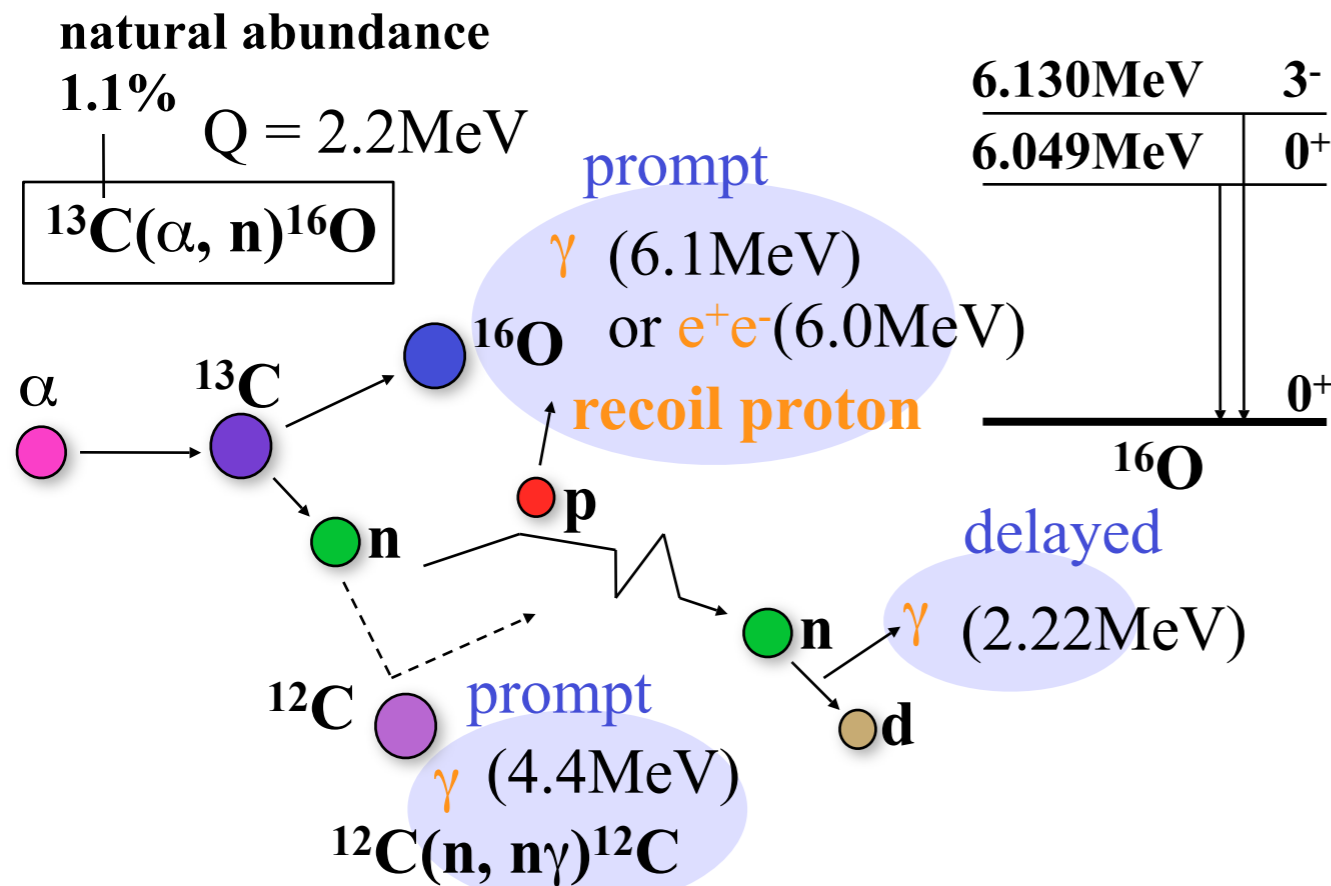


“4pi calibration” system for the off-axis source deployment

bias < 3 cm corresponds to 1.8% volume uncertainty

cross-checked by ¹²B/¹²N uniformity

(α, n) Background Estimation



neutron yield difference < 4%

(α, n) background estimation

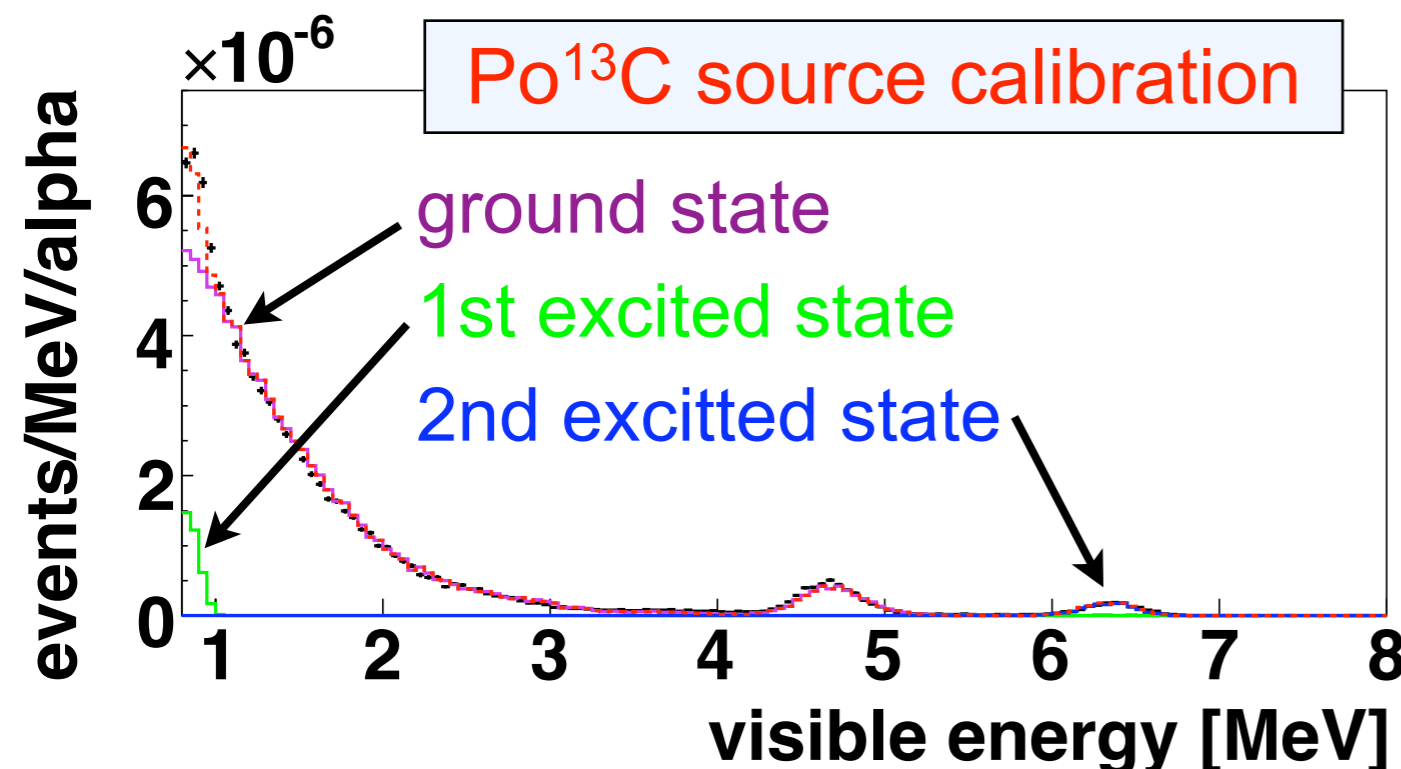
163.3 ± 18.0 events for ground state

18.7 ± 3.7 events for excited state

Estimation uncertainty

11% for ground state

20% for excited state



Rate Analysis above 2.6 MeV

“Reactor” rate analysis
(2.6 MeV threshold)

No osci. expected 1549

Background 63

(see Poster Sessions : Ichimura and Minekawa et al.)

Observed events 985

Ratio = (obs. - B.G.) / No osci.
 $0.594 \pm 0.020(\text{stat}) \pm 0.026(\text{syst})$

8.5 σ disappearance significance

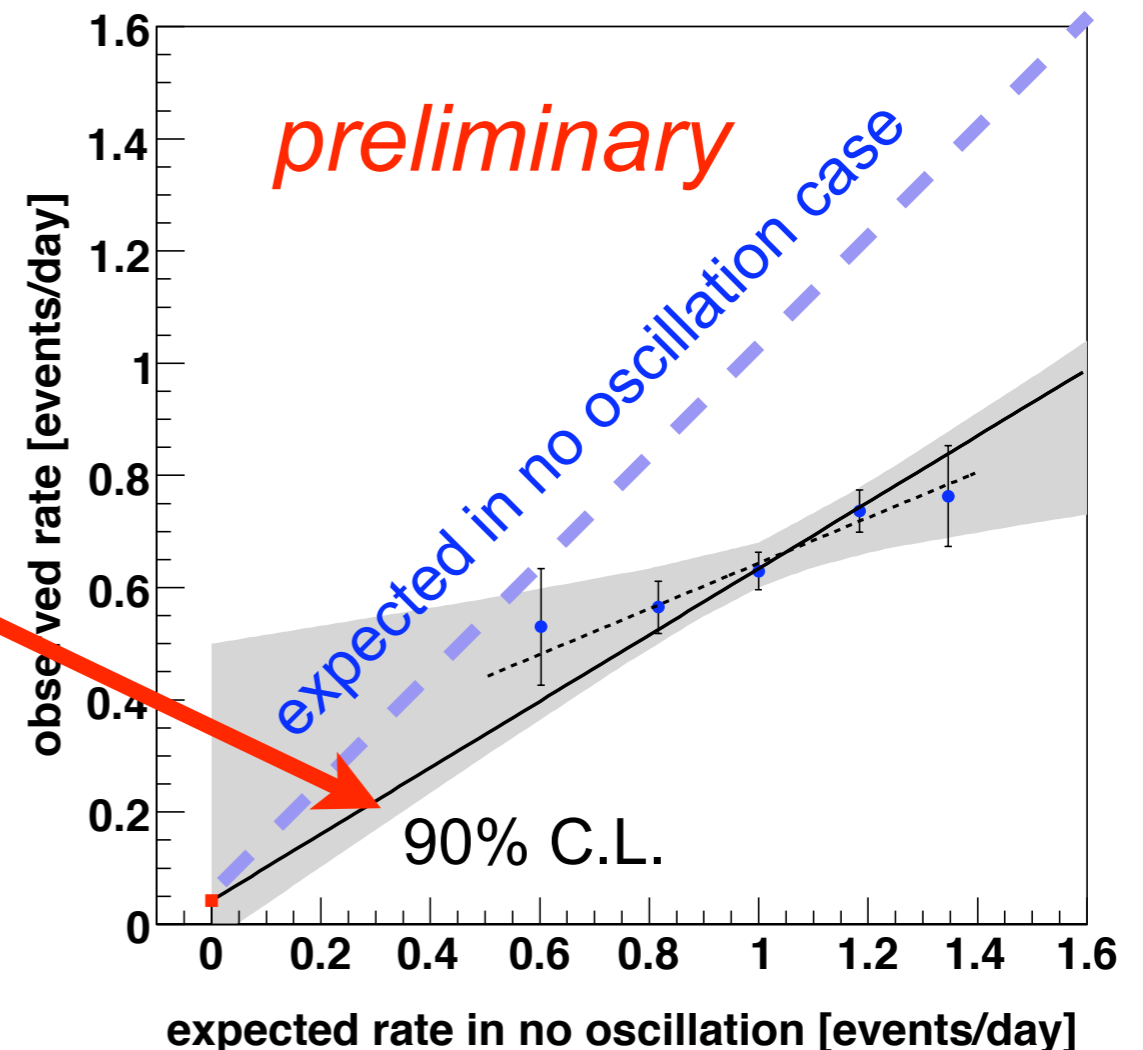
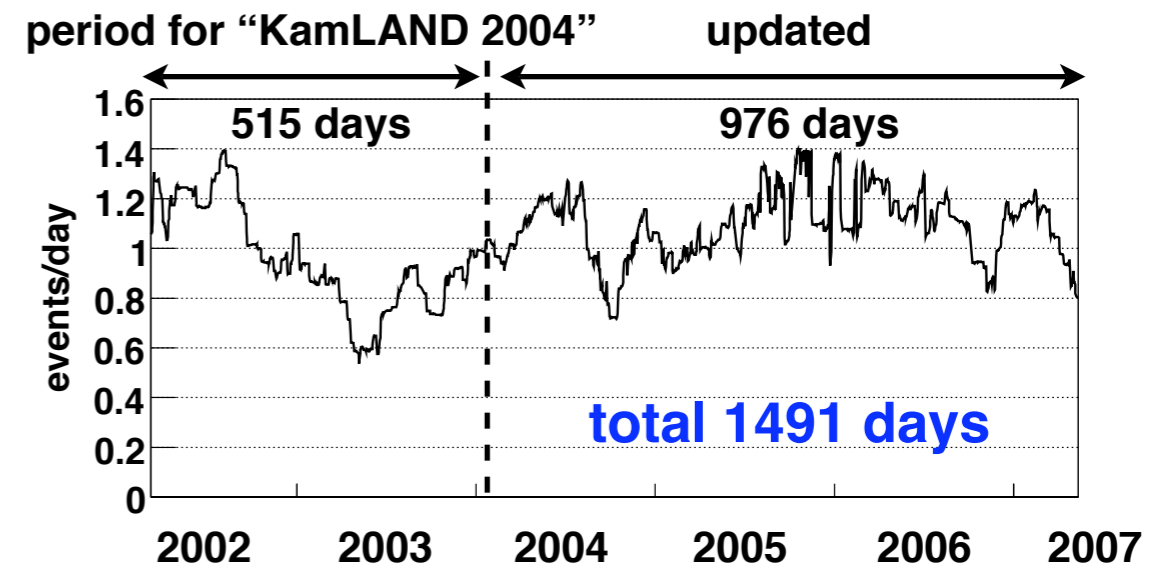
Fit constrained through B.G. expected

$$\chi^2 / \text{ndf} = 3.1 / 4$$

Fit with a horizontal line

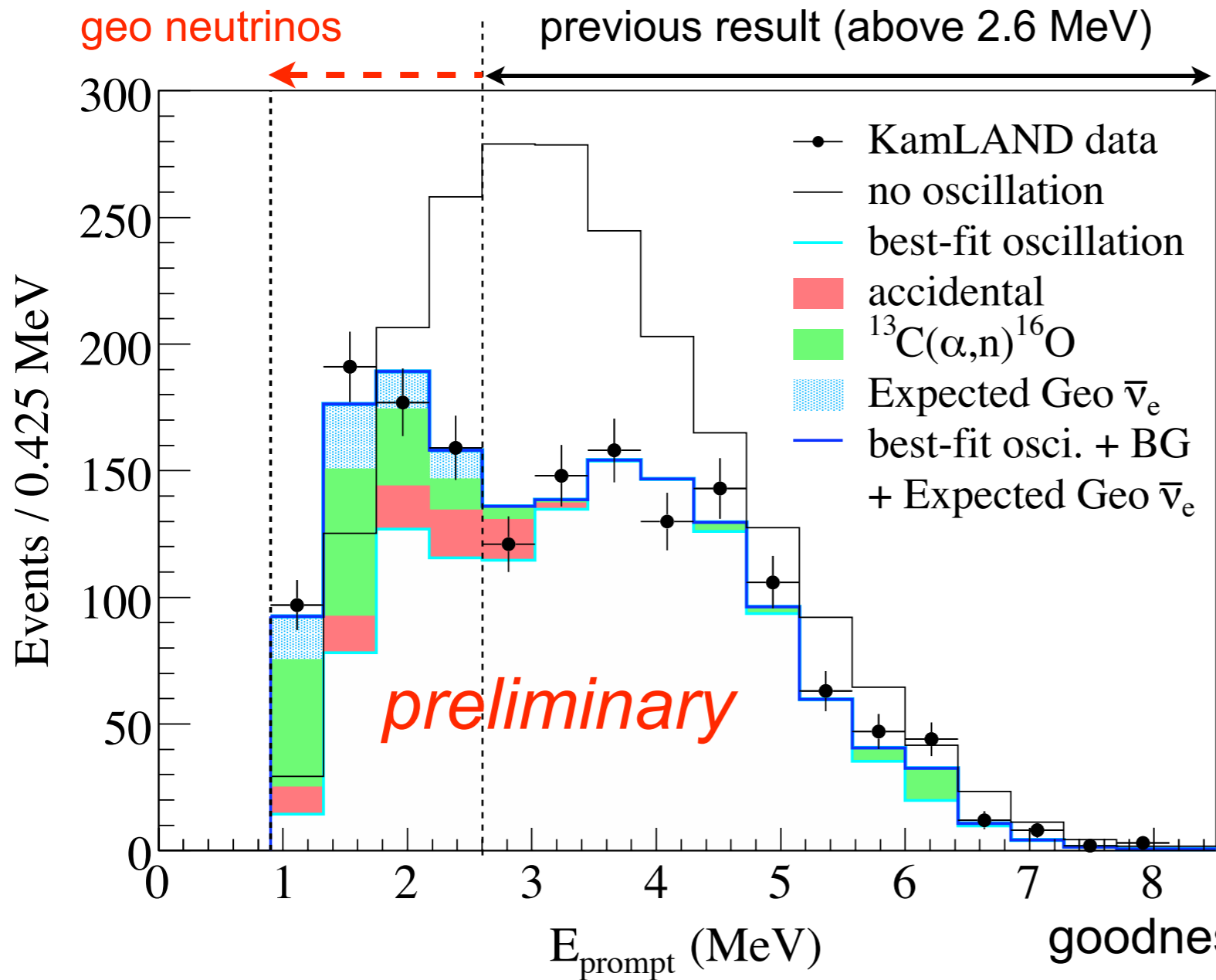
$$\chi^2 / \text{ndf} = 11.8 / 4$$

(1.9% C.L.)



Energy Spectrum above 0.9 MeV

exposure : 2881 ton-year (3.8 × 766 ton-year for “KamLAND 2004”)



“Geo + Reactor”
combined analysis

No osci. expected 2178

Background
(w/o geo neutrino) 276

(Ichimura and Minekawa et al.)

Observed events 1609

best-fit

$(\tan^2\theta, \Delta m^2)$
= $(0.56, 7.58 \times 10^{-5} \text{ eV}^2)$

free parameter : geo neutrinos
(U, Th) = (39.3, 29.4) events

goodness of fit using equal probability bins

best-fit $\chi^2 / \text{ndf} = 21.0 / 16$ (18.0% C.L.)

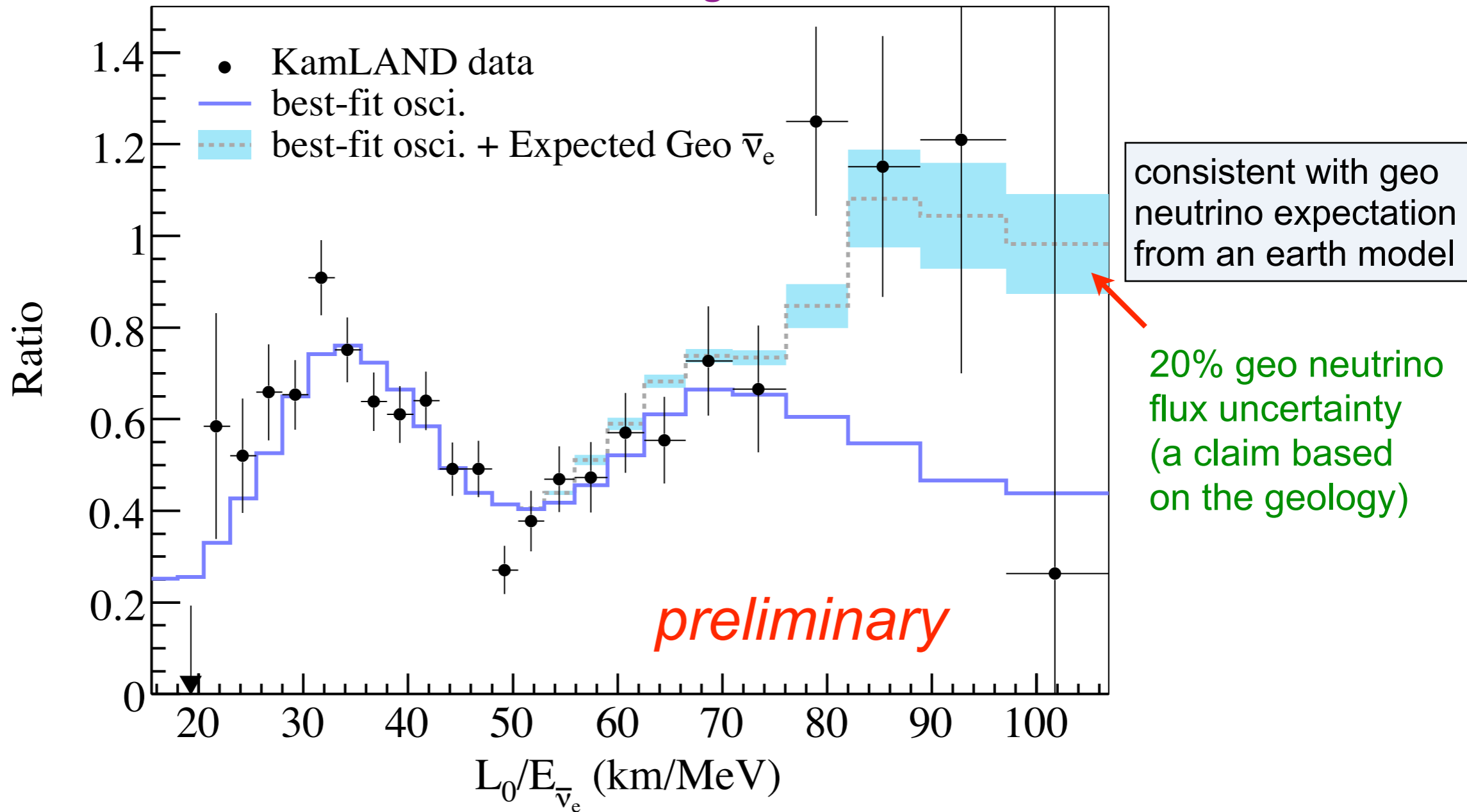
no osci. $\chi^2 / \text{ndf} = 63.9 / 17$

Scaled no oscillation spectrum is excluded at 5.2σ

L/E plot

$$\text{Ratio} = (\text{observed} - \text{B.G.}) / (\text{no osci. expected})$$

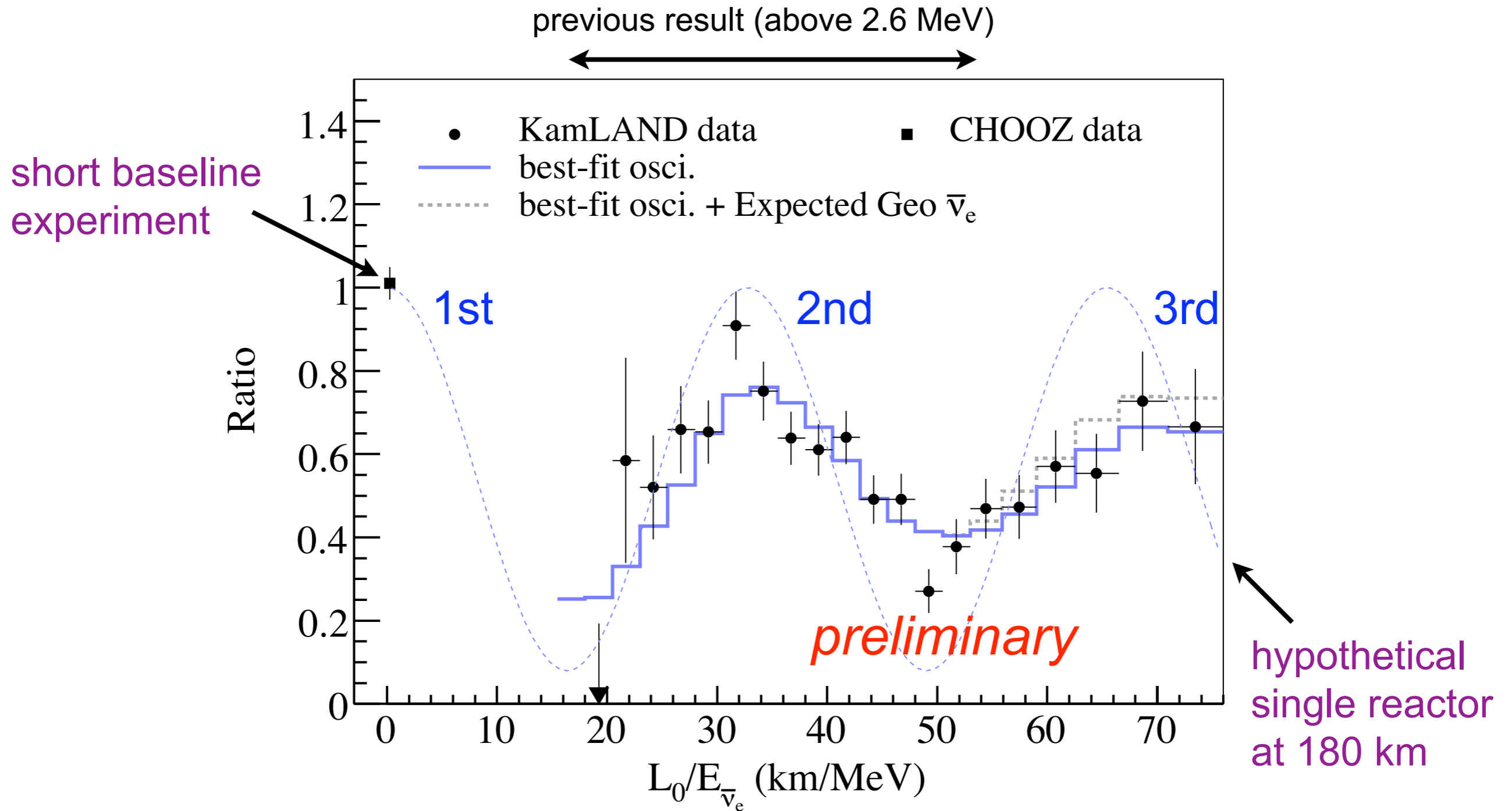
↑ w/o geo neutrino



L_0 : a fixed baseline (180 km)

Distortion effect is clearly illustrated by L/E plot

Neutrino Oscillation



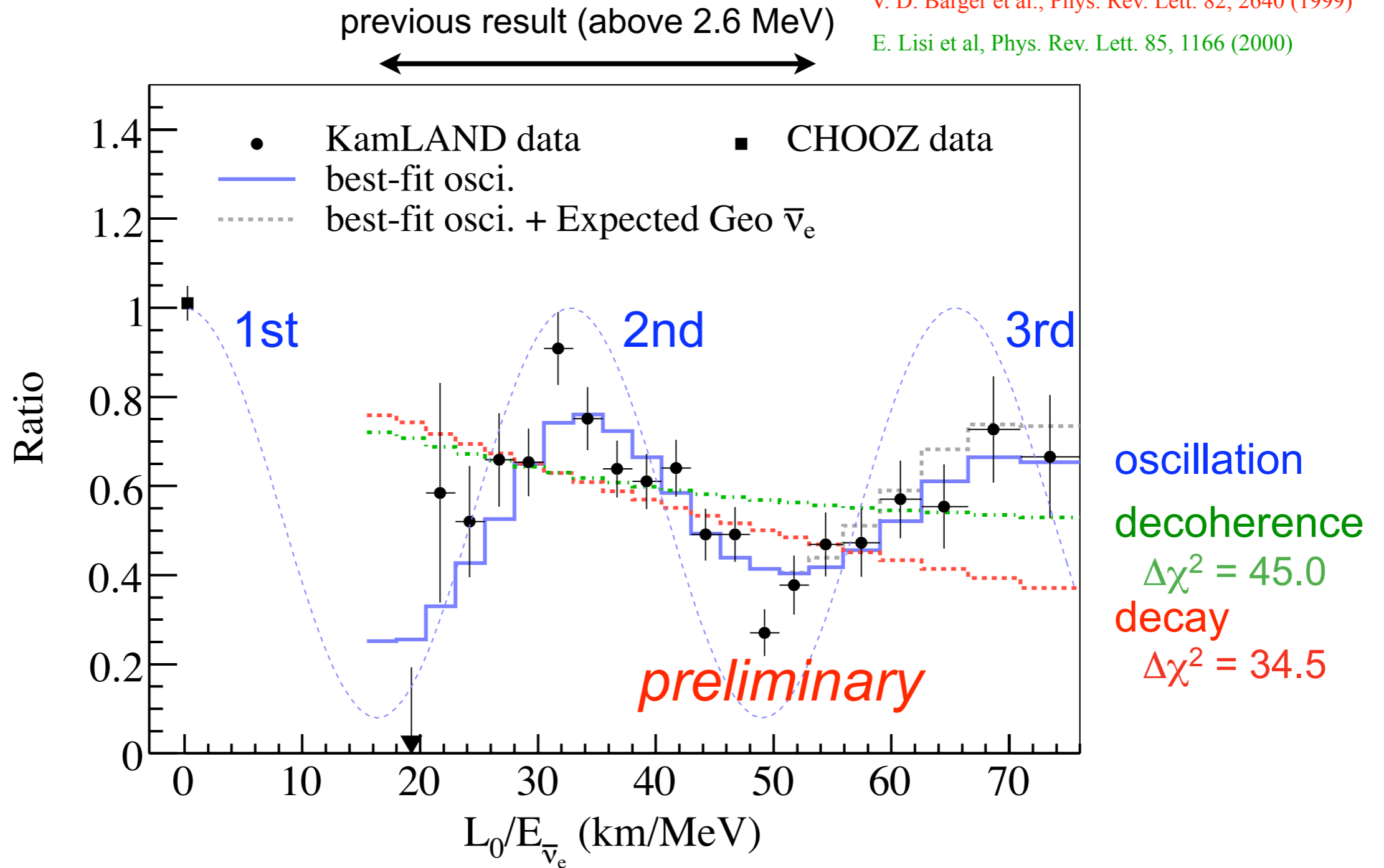
KamLAND covers the 2nd and 3rd maximum

→ **characteristic of neutrino oscillation**

Alternate Hypothesis

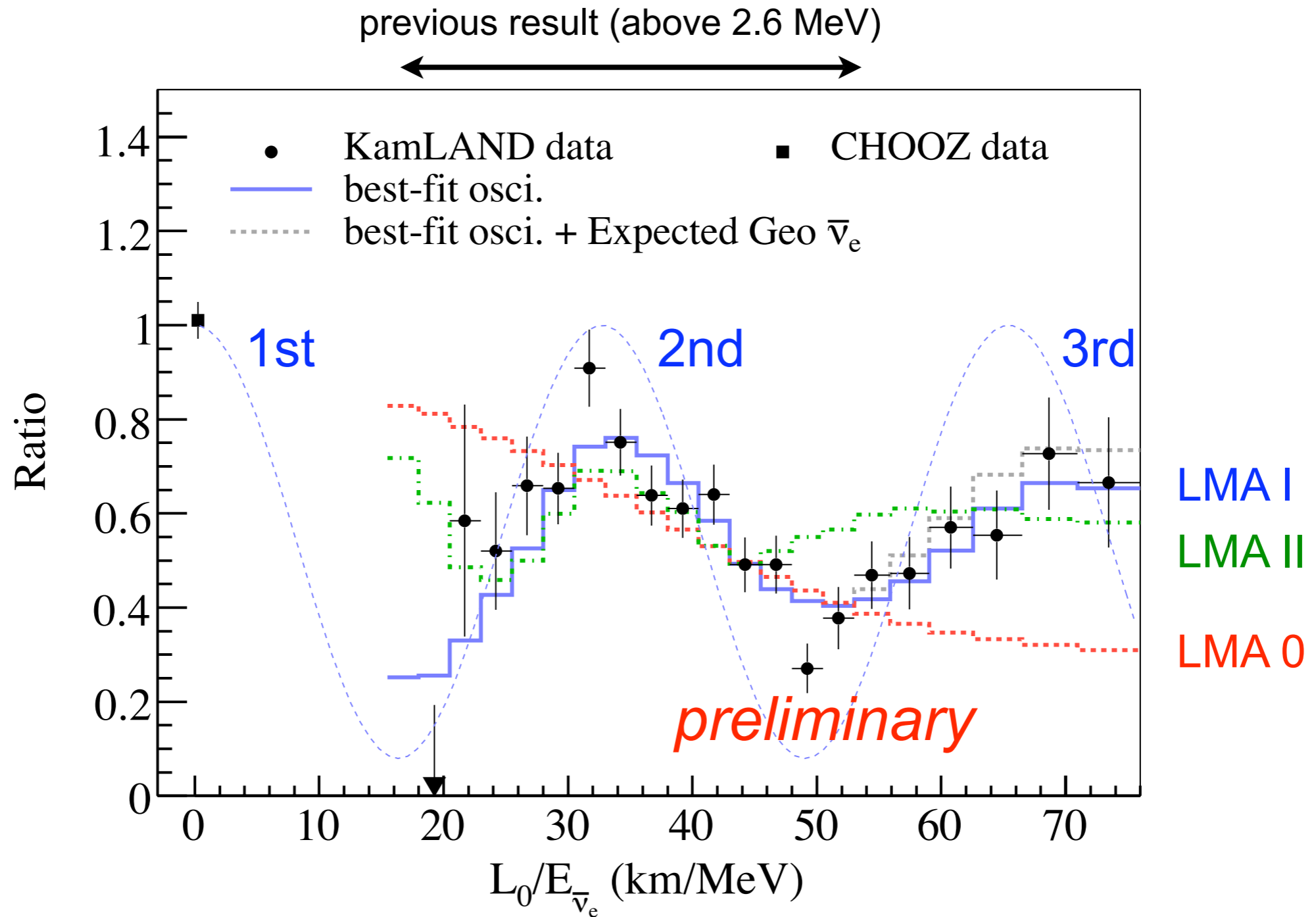
V. D. Barger et al., Phys. Rev. Lett. 82, 2640 (1999)

E. Lisi et al, Phys. Rev. Lett. 85, 1166 (2000)



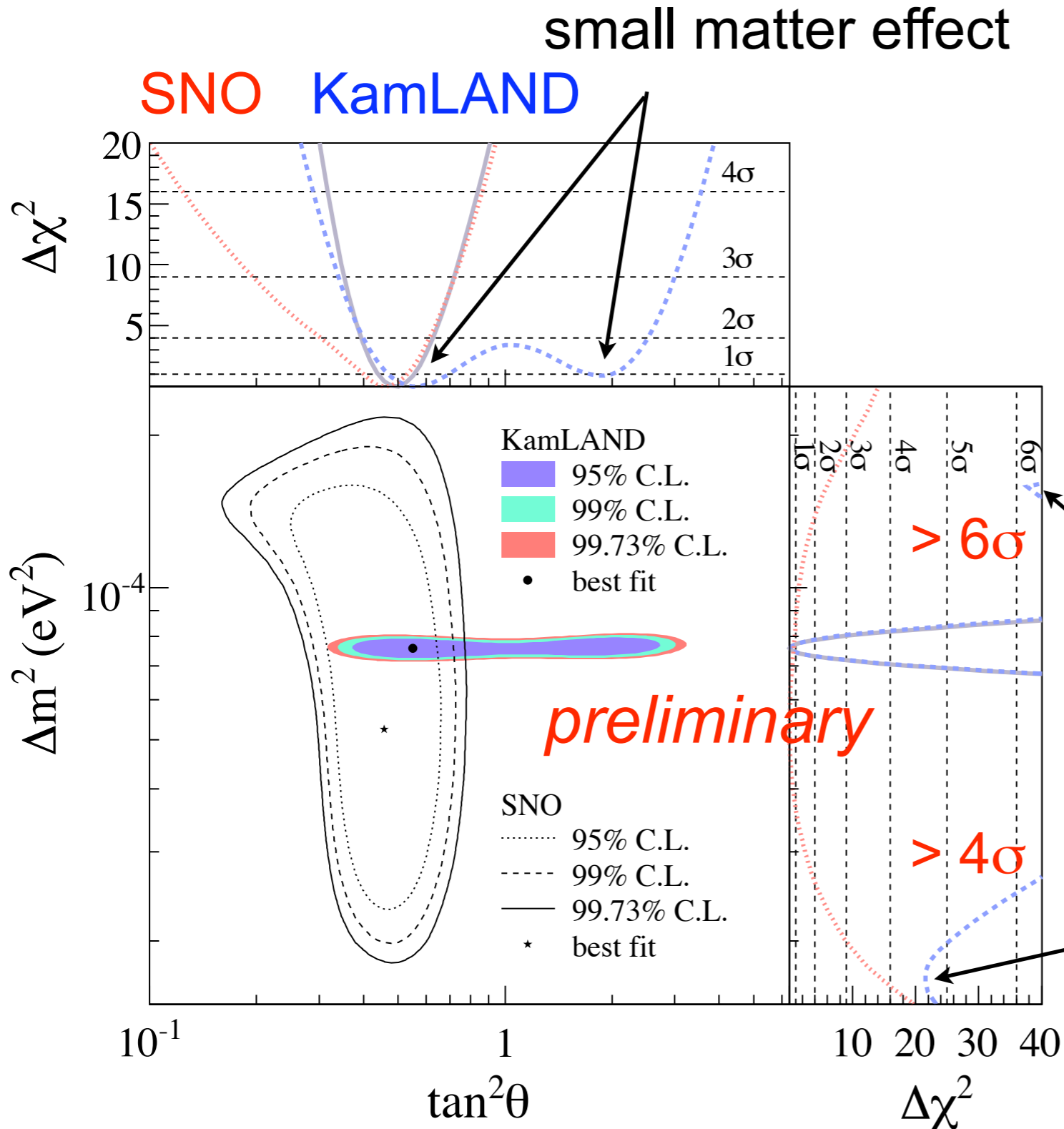
best model is neutrino oscillation

Alternate Wavelength



LMA 0 and LMA II are disfavored at more than 4σ

Oscillation Parameters

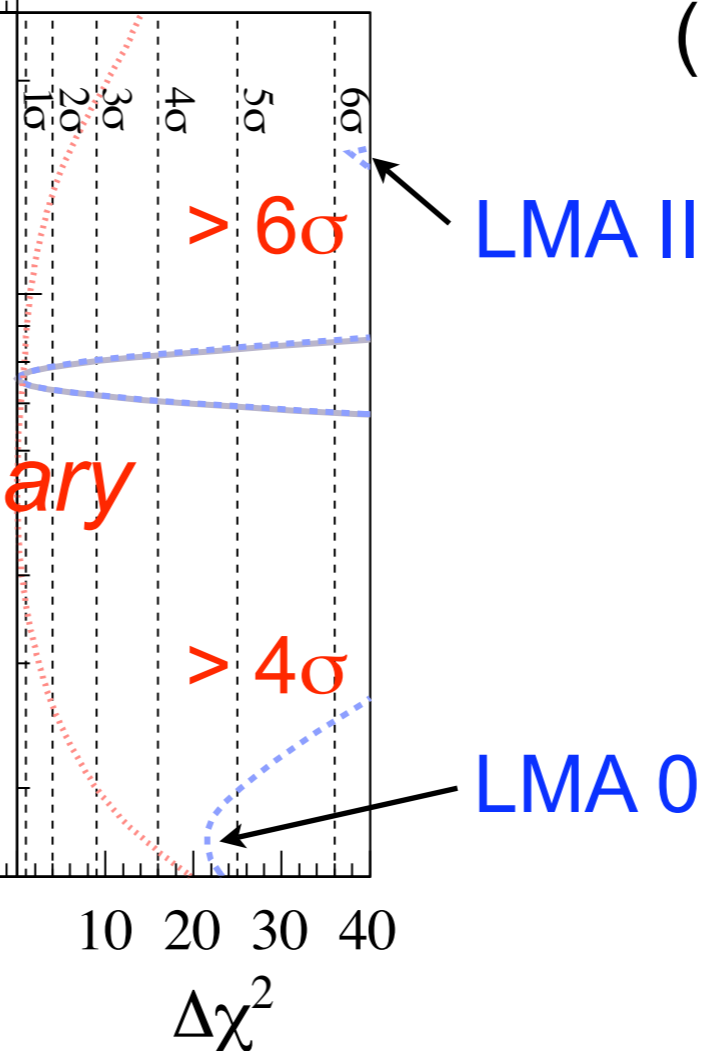


KamLAND only

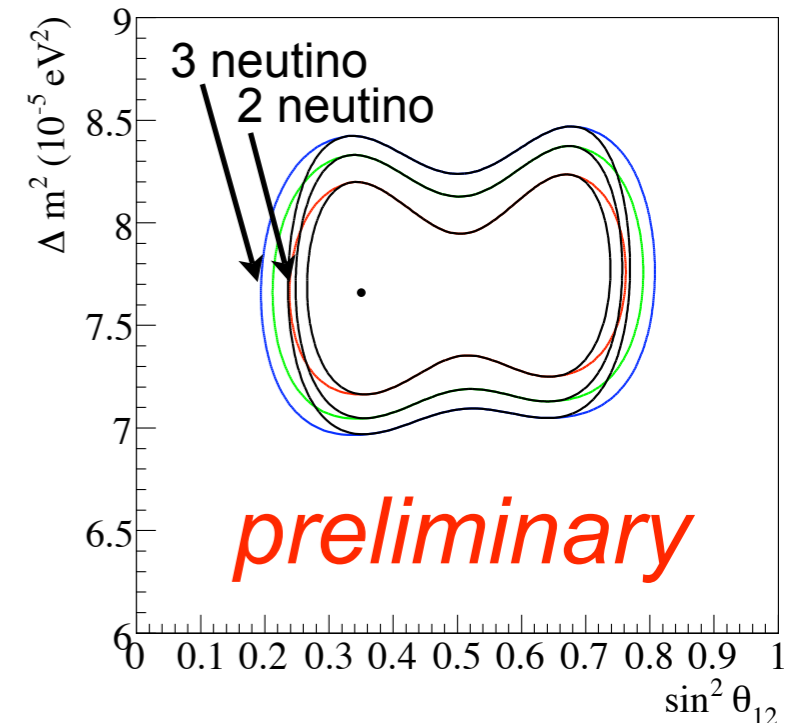
$$\tan^2\theta = 0.56^{+0.14}_{-0.09}$$

$$\Delta m^2 = 7.58^{+0.21}_{-0.20} \times 10^{-5} \text{ eV}^2$$

(marginalized error)



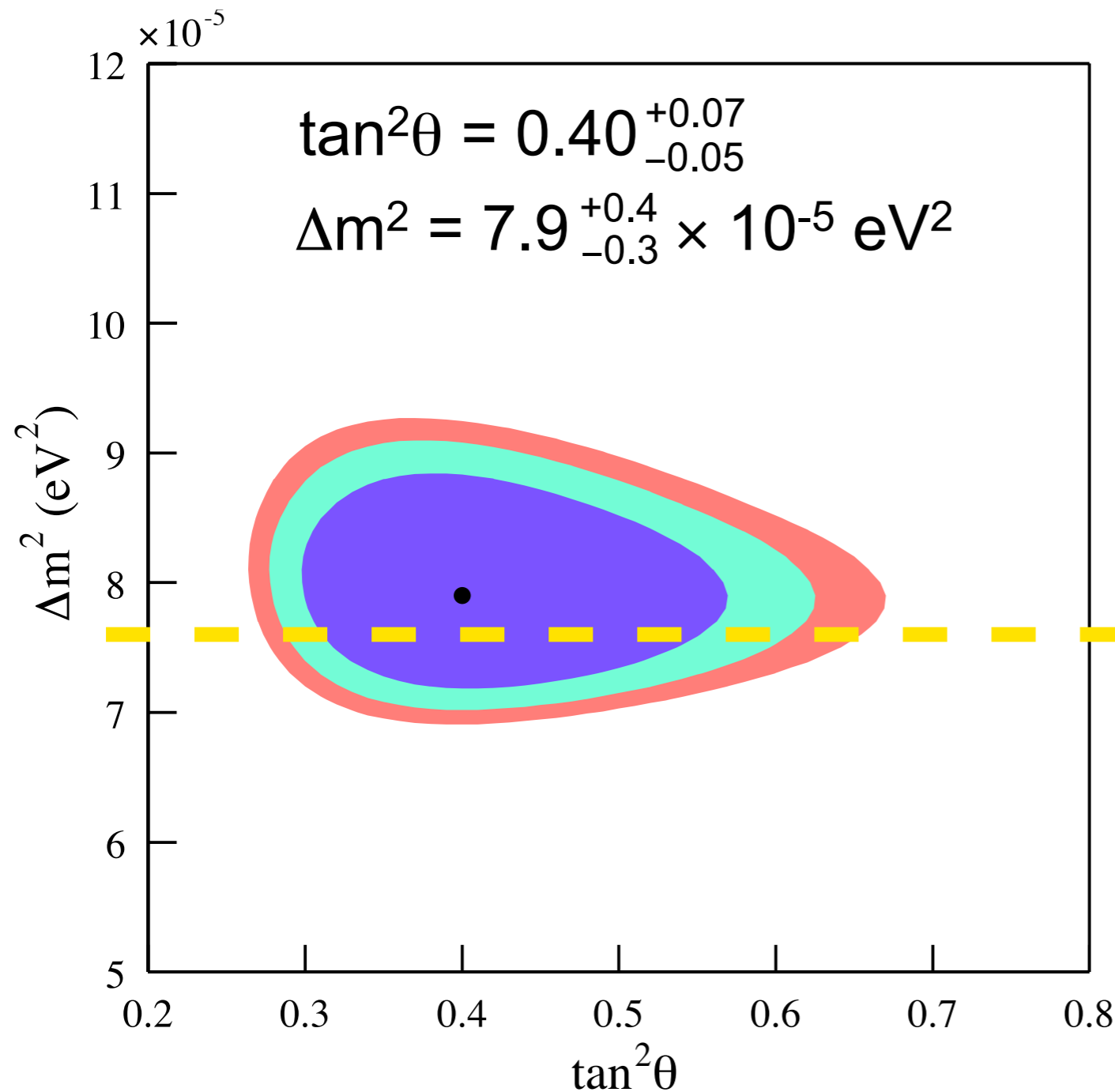
3 neutrino effect



same result for Δm^2

Precise measurement of Δm^2

KamLAND 2004

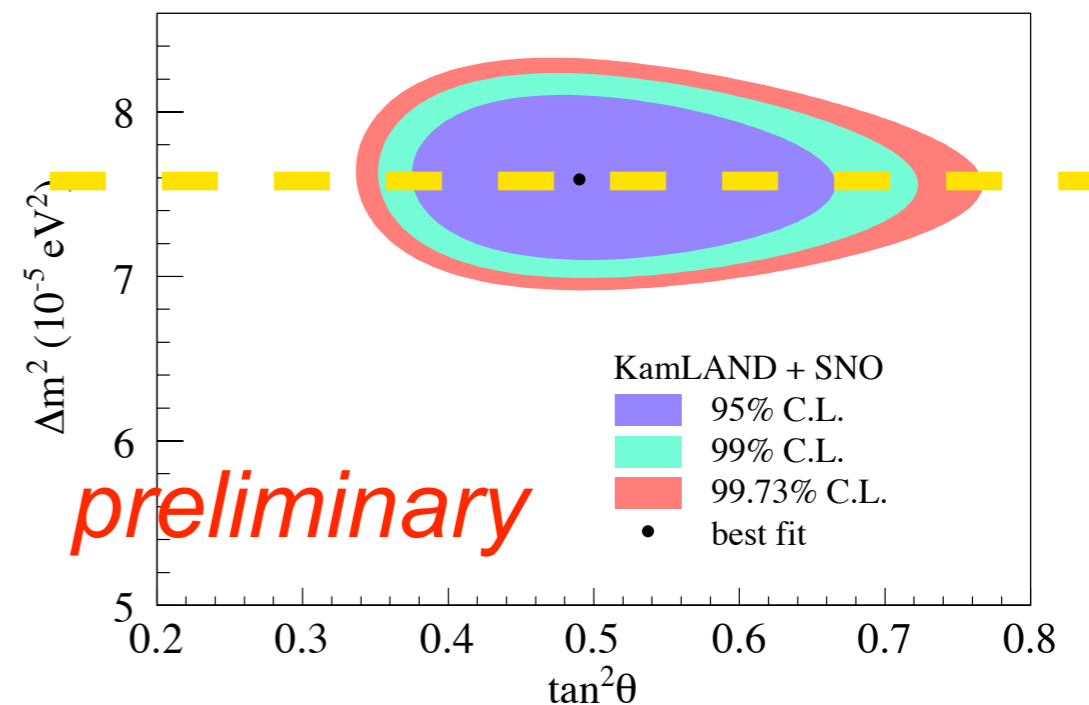


This result

KamLAND + SNO

$$\tan^2\theta = 0.49^{+0.07}_{-0.05}$$
$$\Delta m^2 = 7.59^{+0.20}_{-0.21} \times 10^{-5} \text{ eV}^2$$

Δm^2 : systematic uncertainty 2.0%
dominated by linear energy scale uncertainty

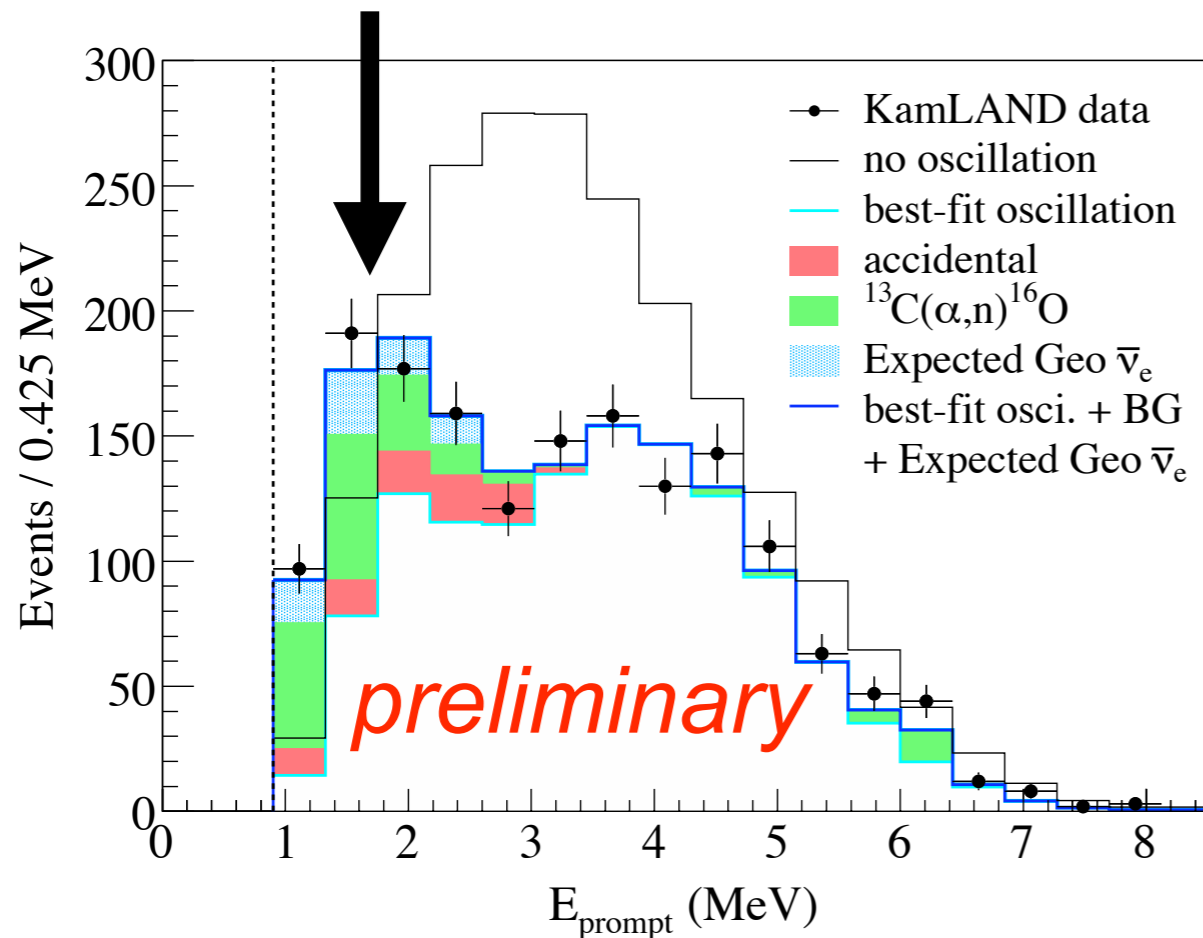


Δm^2 is measured at 2.8% precision by KamLAND

Geo Neutrino Estimation

Analysis : KamLAND (rate + shape + time) + SNO

geo neutrinos (U, Th)



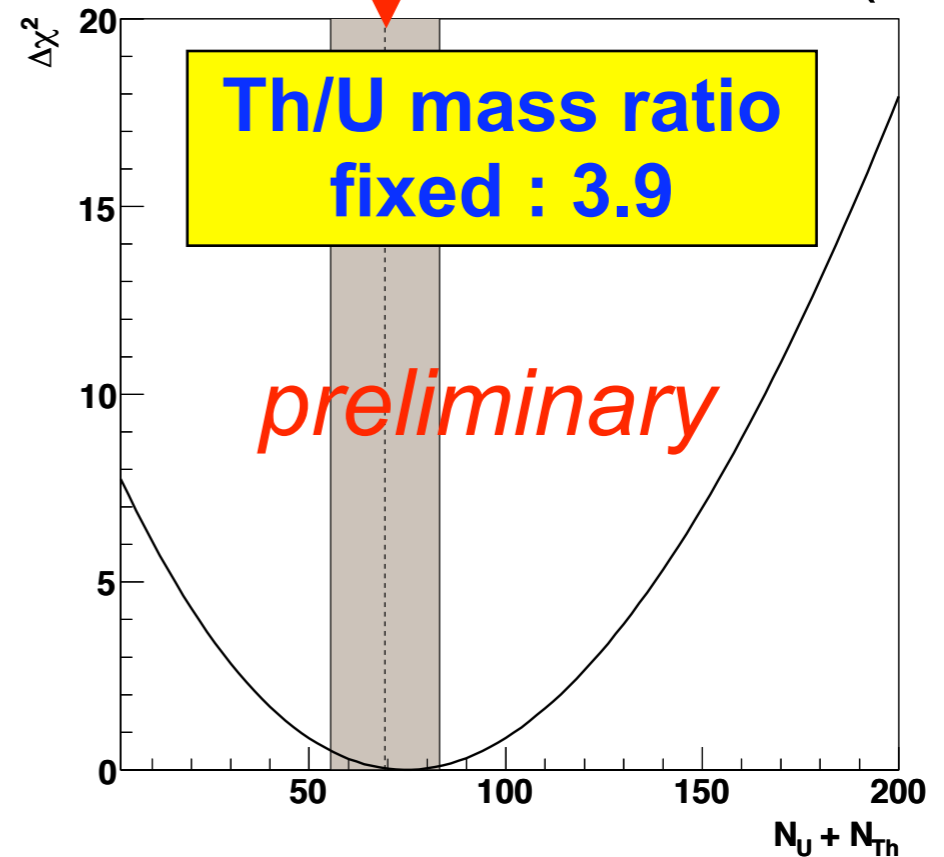
Reference model (16 TW)

U : 56.2 event (28.9 TNU)

Th : 13.1 event (7.6 TNU)

model expected

↓ 69.3 events (36.5 TNU)



U+Th = $74.9^{+27.3}_{-27.2}$ event

$39.4^{+14.4}_{-14.3}$ TNU

(previous result : $57.4^{+32.0}_{-30.0}$ TNU)

TNU (Terrestrial Neutrino Unit) = events/ 10^{32} target-proton/year

Summary

- KamLAND improved sensitivity to $\bar{\nu}_e$ observation.

data-set : 766 ton-yr \rightarrow 2881 ton-yr (α, n) B.G. uncertainty :
E threshold : 2.6 MeV \rightarrow 0.9 MeV 32% \rightarrow 10% (ground state)
syst. uncertainty : 6.5% \rightarrow 4.1% 100% \rightarrow 20% (excited state)

- In the reactor neutrino analyses, we showed

- Oscillatory shape including 2nd and 3rd maximum
- Exclusion of LMA II and O at more than 4σ C.L.
- Precise measurement of oscillation parameters.

KamLAND only $\tan^2\theta = 0.56^{+0.14}_{-0.09}$ $\Delta m^2 = 7.58^{+0.21}_{-0.20} \times 10^{-5} \text{ eV}^2$

KamLAND + SNO $\tan^2\theta = 0.49^{+0.07}_{-0.05}$ $\Delta m^2 = 7.59^{+0.20}_{-0.21} \times 10^{-5} \text{ eV}^2$

- Geo neutrino flux is measured with better precision.