

NOW 2008 - and Then



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für Kernphysik, Heidelberg**



NOW 2008

Conca Specchiulla (Otranto, Italy), September 6-13, 2008

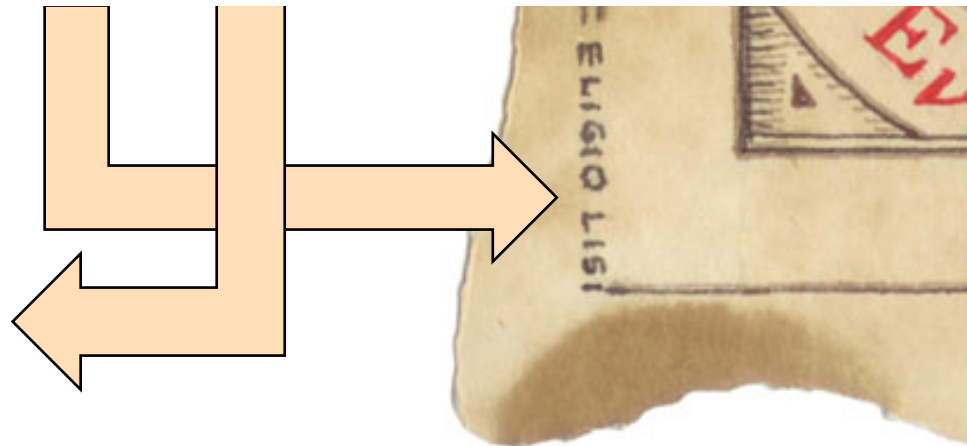
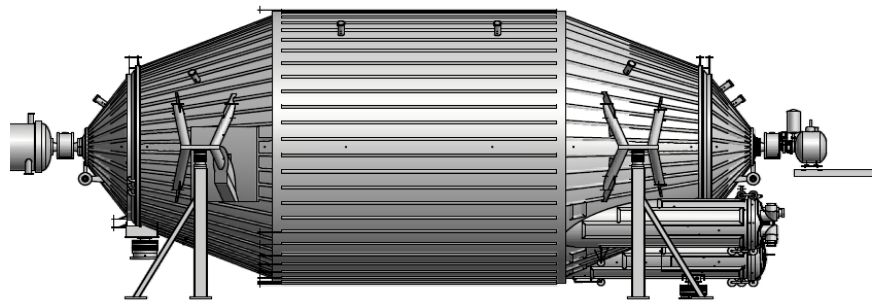
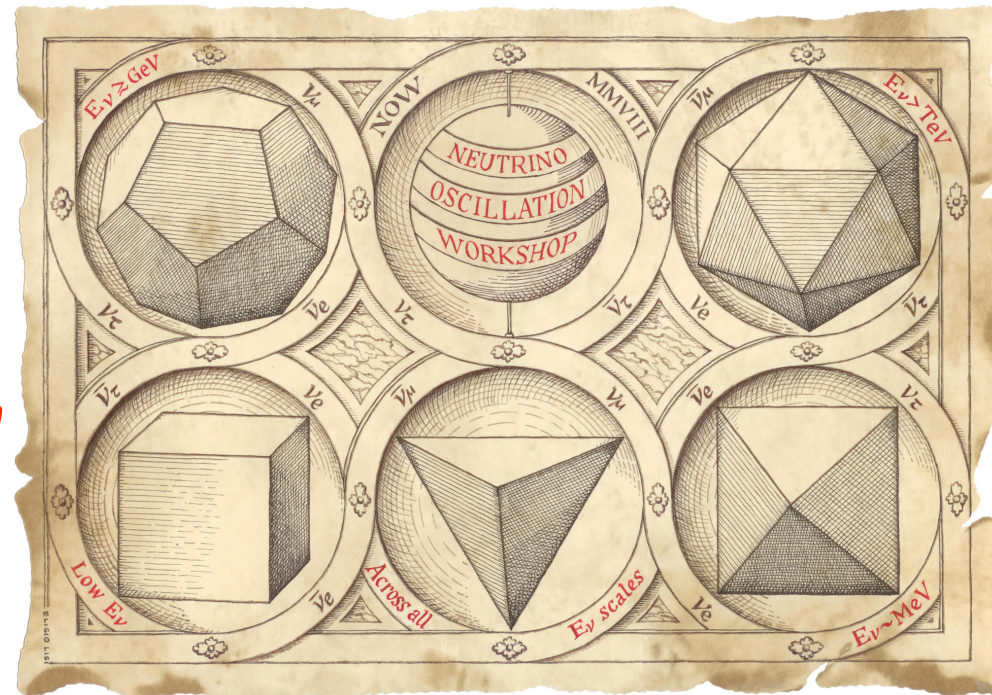


The Task

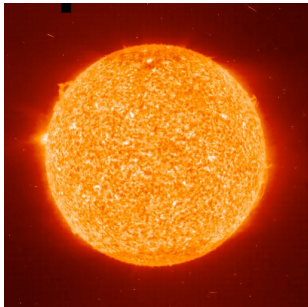
G. Fogli: The Ribbon

At this point it remains to interpret the role of the **ribbon**, which goes around the solids without interruption. (without a break) (uninterruptedly).

- relate the “Platonic solids”
- interpret connections
- the corners and cracks
- the larger picture

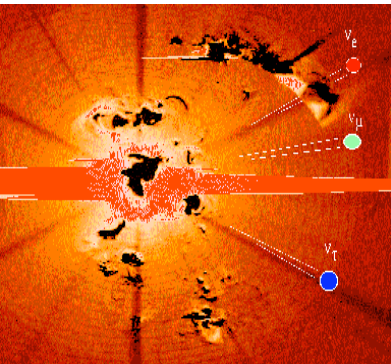
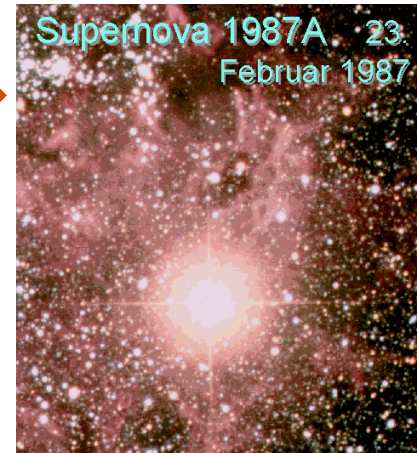


Neutrino Topics



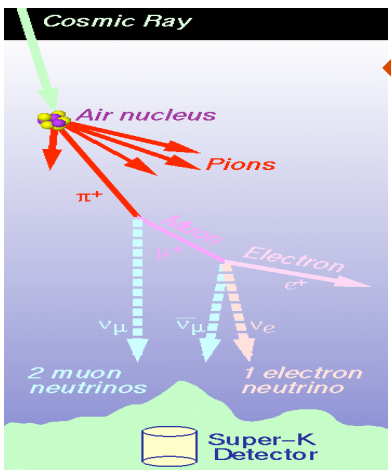
← Sun

Astronomy: →
Supernovae
GRBs
UHE ν 's



← **Cosmology**

Reactors →

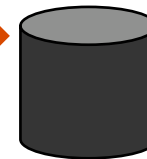


← **Atmosphere**

Accelerators →



β -Sources →



← **Earth**

Four Methods of Mass Determination

- **kinematical**
- **lepton number violation**
 ↔ Majorana nature
- **astrophysics & cosmology**
- **oscillations**

β-decay: energy spectrum

G. Drexlin

measurement of $m(\nu_e)$
based on kinematics & energy conservation

$$m(\nu_e) = \sqrt{\sum_{i=1}^3 |U_{ei}^2| \cdot m_i^2}$$

incoherent sum

$$\frac{d\Gamma_i}{dE} = C \cdot p \cdot (E + m_e) \cdot (E_0 - E) \cdot \sqrt{(E_0 - E)^2 - m_i^2} \cdot F(E, Z) \cdot \theta(E_0 - E - m_i)$$



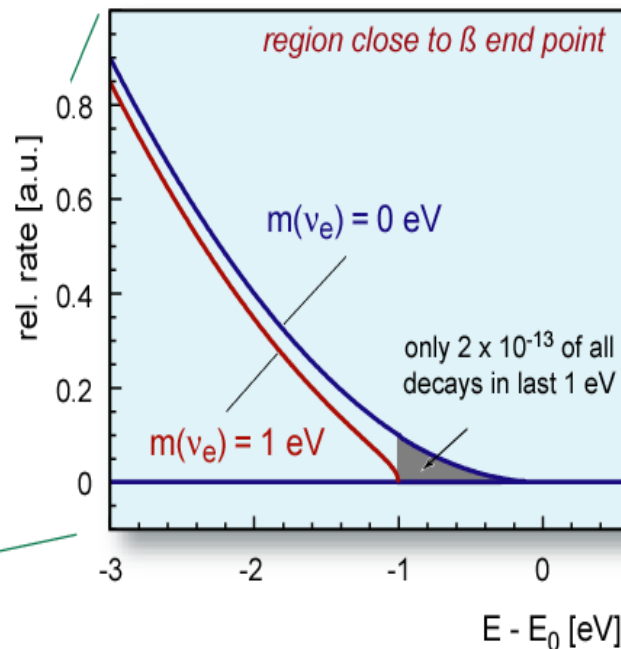
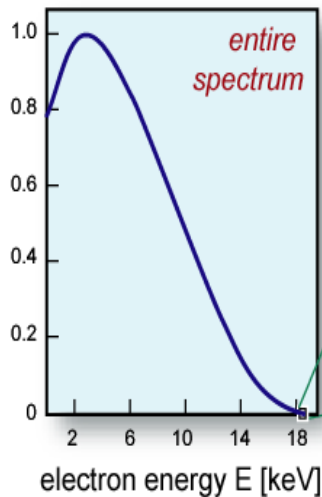
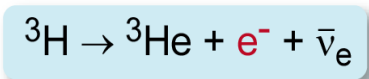
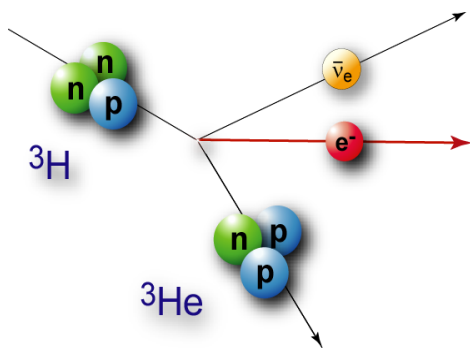
$$G_F^2 \frac{m_e^5}{2\pi^3} \cos^2 \theta_C |M|^2$$



(ν-mass)²



Fermi function



investigation of the kinematics of β -decay:

→ only **model independent** measurement of absolute ν mass scale

↔ cosmology ...

MARE: staged approach based on microcalorimeters ^{187}Re β -decay

MARE-I ~300 detectors with $m(\nu) \sim 2$ eV

MARE-II ~50.000 detectors with $m(\nu) \sim 0.2$ eV

if successful R&D & if funded

KATRIN: designed as 'ultimate' tritium β -decay experiment

initial runs Q4/2010

model-independent

status (Mainz, Troizk): $m_\nu < 2.3$ eV

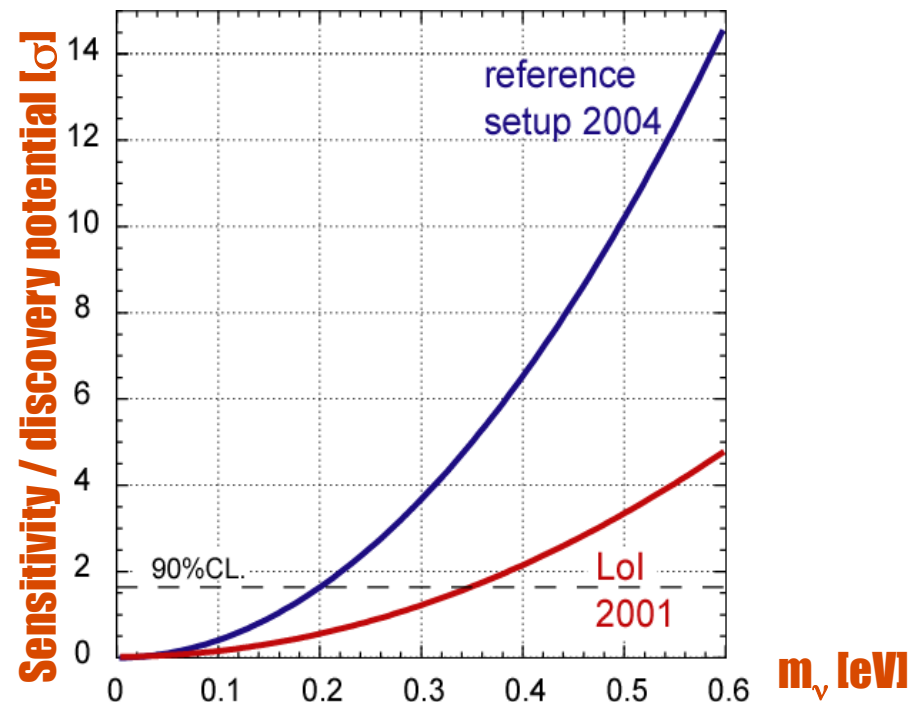
potential: KATRIN, MARE-II

sensitivity (90% CL)

$m(\nu) < 200$ meV

discovery potential

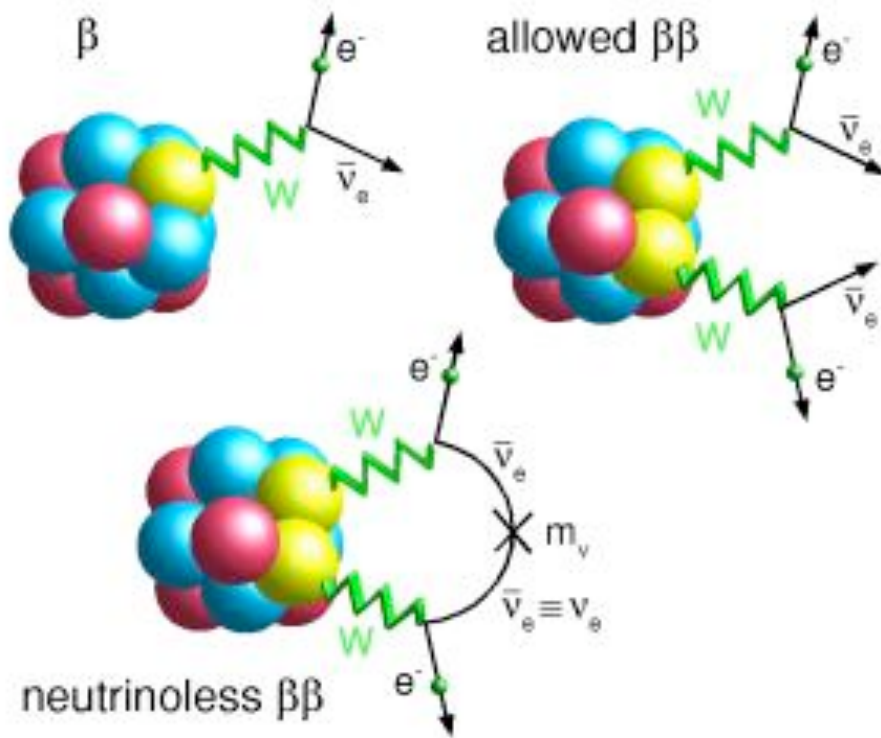
$m(\nu) = 350$ meV (5σ)



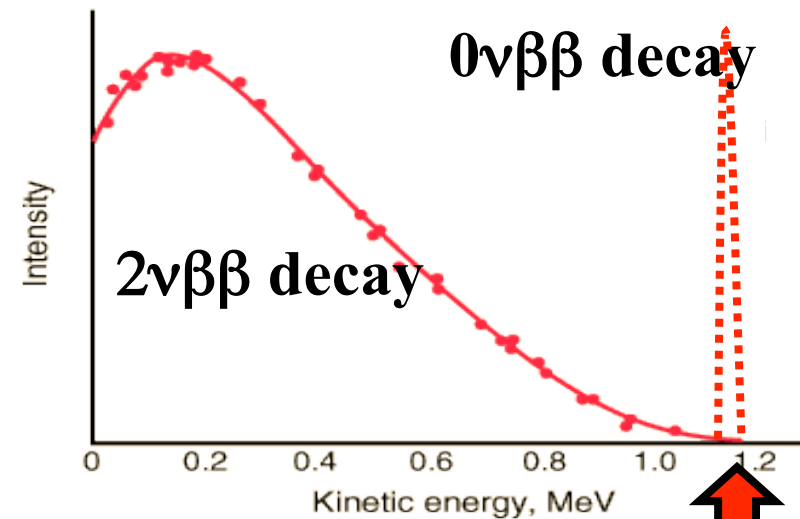
Four Methods of Mass Determination

- **kinematical**
- **lepton number violation**
 ↔ Majorana nature
- **astrophysics & cosmology**
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$0\nu\beta\beta$ Decay Kinematics



$2\nu\beta\beta$ decay of ^{76}Ge observed:
 $\tau = 1.5 \times 10^{21}$ y



Majorana $\nu \rightarrow 0\nu\beta\beta$ decay

warning:

other lepton number violating processes...

- signal at known Q-value
- $2\nu\beta\beta$ background (resolution)
- nuclear backgrounds
- ➔ use different nuclei

NME's: Relating Lifetimes & Neutrino Masses

rate of $0\nu\beta\beta$

phase space

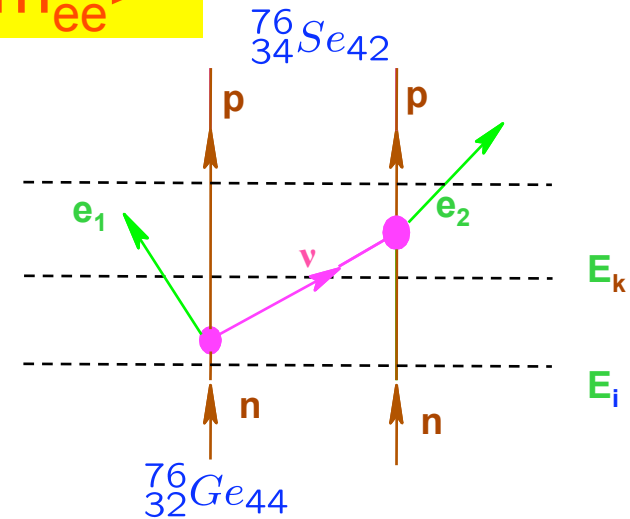
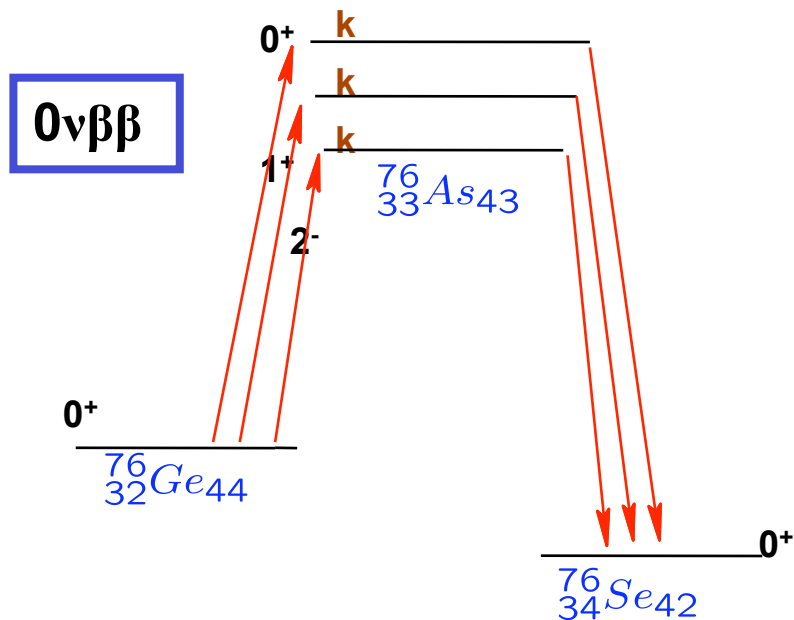
nuclear matrix elements

effective Majorana neutrino mass

$$1/\tau = G(Q,Z) |M_{\text{nucl}}|^2 \langle m_{ee} \rangle^2$$

nuclear matrix elements:

→ virtual excitations of intermediate states



A. Faessler, J. Suhonen, V. Rodin:

good progress in TH errors

→ QRPA & shell model

→ reduced uncertainties

→ what is a 1σ theory error?

Summary: Neutrino Mass from $0\nu\beta\beta$

Theory with R-QRPA and $g_A = 1.25$

Exp. Klapdor et al. Mod. Phys. Lett. A21,1547(2006) ; ^{76}Ge

$T(1/2; 0\nu\beta\beta) = (2.23 +0.44 -0.31) \times 10^{25}$ years; 6σ

- $\langle m(\nu) \rangle = 0.24$ [eV] (exp ± 0.02 ; theor ± 0.01) [eV]
Bonn CD, no short range correlations
- $\langle m(\nu) \rangle = 0.22$ [eV] (exp ± 0.02 ; theor ± 0.01) [eV]
Bonn CD, Consistent Brückner Correlations
- $\langle m(\nu) \rangle = 0.24$ [eV] (exp ± 0.02 ; theor ± 0.01)
Argonne, Consistent Brückner Correlations
- $\langle m(\nu) \rangle = 0.30$ [eV] (exp ± 0.03 ; theor ± 0.01) [eV]
Bonn CD, Fermi Hypernetted Chain (Argonne in nuclei)
- $\langle m(\nu) \rangle = 0.26$ [eV] (exp ± 0.02 ; theor ± 0.01) Bonn CD, UCOM (AV18 in D)
- $\langle m(\nu) \rangle = 0.31$ [eV] (exp ± 0.03 ; theor ± 0.02) [eV] Bonn CD, Jastrow

$0\nu\beta\beta$ Experiments

Future projects[#] (a broad brush, personal view)

→ G. Gratta

Isotope	Experiment	Main principle	Fid mass	Lab
^{76}Ge	Majorana [†]	Eres, 2site tag, Cu shield	30+30kg	SUSEL
	Gerda [†]	Eres, 2site tag, LAr shield	18→40 kg	G Sasso
	MaGe/GeMa	See above	~1ton	DUSEL? G Sasso?
^{150}Nd	SNO+	Size/shielding	56 kg	SNOlab
^{150}Nd or ^{82}Se	SuperNEMO [†]	Tracking	100-200 kg	Canfranc Frejus
$^{130}\text{Te}^*$	CUORE	E Res.	204 kg	G Sasso
^{136}Xe	EXO	Tracking	150 kg	WIPP
		Ba tag, Tracking	1-10ton	DUSEL?

≥ 2 elements

see also talks by
 J. Janisko (GERDA)
 F. Bellini (CUORE)
 L. Vala (SuperNEMO)
 F. Sanchez (NEXT)
 C. Jillings (SNO+)

Exciting time for neutrino-less double beta decay!

Several 100kg-class experiments will start data taking in the next 2-3 years.

R&D for ton-class experiments is on-going.

Neutrino-less Double β -Decay

Beta particle (electron)

Majorana $\nu \rightarrow 0\nu 2\beta$ decay

$\propto |\langle m_{ee} \rangle| = |\sum m_i U_{ei}^2| \leq 0.35 \text{ eV ?}$

Heidelberg-Moscow experiment

$$m_{ee} = |m_{ee}^{(1)}| + |m_{ee}^{(2)}| \cdot e^{i\Phi_2} + |m_{ee}^{(3)}| \cdot e^{i\Phi_3}$$

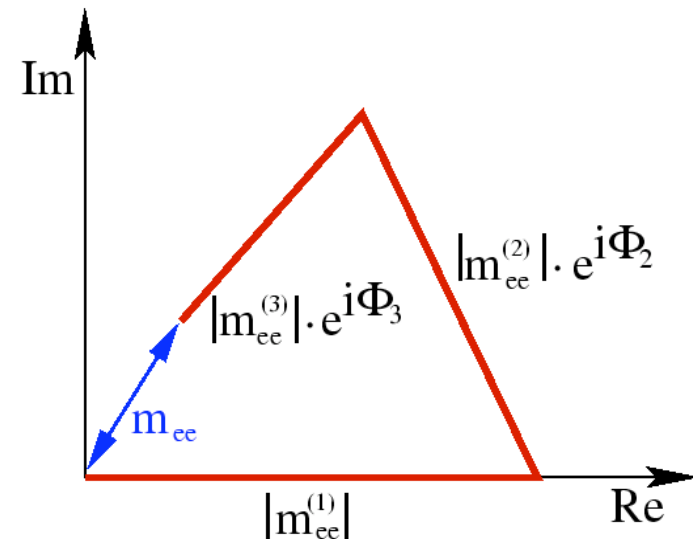
$$|m_{ee}^{(1)}| = |U_{e1}|^2 m_1$$

$$|m_{ee}^{(2)}| = |U_{e2}|^2 \sqrt{m_1^2 + \Delta m_{21}^2}$$

$$|m_{ee}^{(3)}| = |U_{e3}|^2 \sqrt{m_1^2 + \Delta m_{31}^2}$$

solar $\Rightarrow |U_{e1}|^2, |U_{e2}|^2, \Delta m_{21}^2$ atmosph. $\Rightarrow |\Delta m_{31}^2|$ CHOOZ $\Rightarrow |U_{e3}|^2 < 0.05$

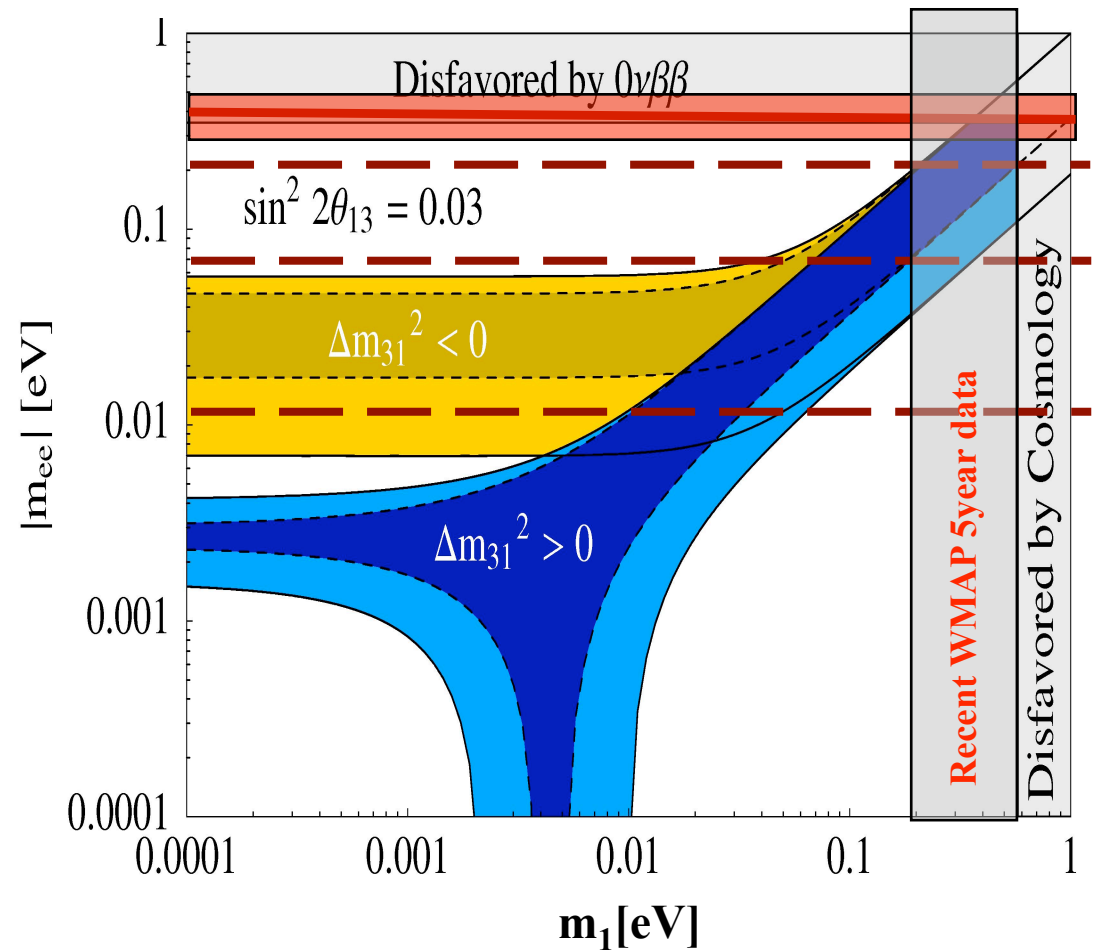
\rightarrow free parameters: $m_1, \text{sign}(\Delta m_{31}^2), \text{CP-phases } \Phi_2, \Phi_3$



Claim of part of the original Heidelberg-Moscow experiment
 \leftrightarrow cosmology \rightarrow ,tension‘

aims of new experiments:

- test HM claim
- $(\Delta m_{31}^2)^{1/2} \simeq 0.05\text{eV} \pm \text{errors}$
 - \rightarrow reach $\sim 0.01\text{eV}$
 - \rightarrow CUORE
 - \rightarrow GERDA phases I, II, (III)
 - \rightarrow EXO, SuperNEMO, ...



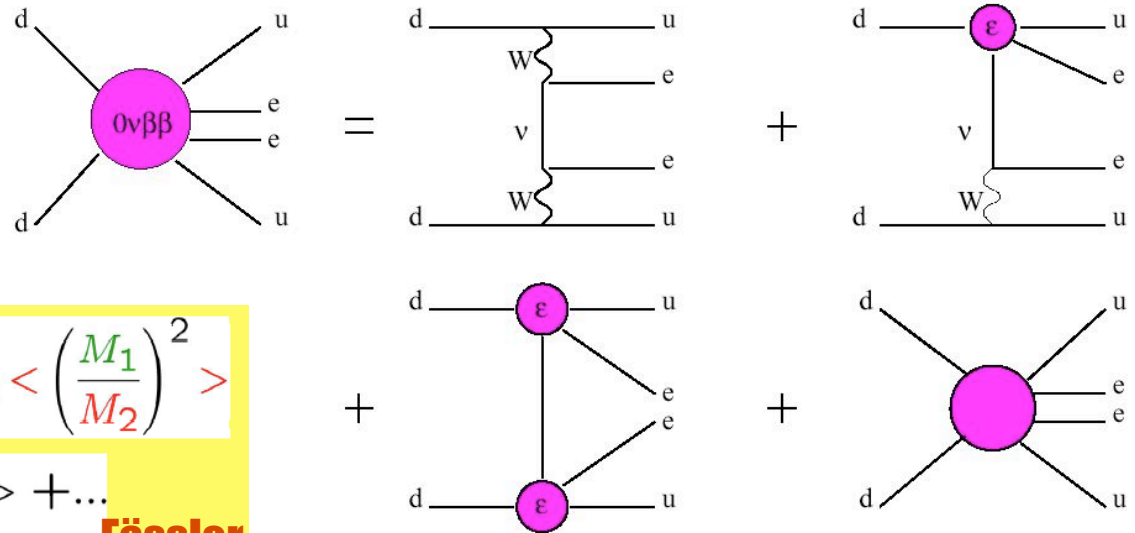
Comments:

- cosmology: limitation by systematical errors \rightarrow another factor ~ 5 ?
- $0\nu\beta\beta$ nuclear matrix elements \rightarrow unavoidable **theory** error in m_{ee}
- $\Delta m^2 > 0$ allows complete cancellation
- $0\nu\beta\beta$ from *other* new BSM lepton number violating operators

... this may not be the full story

LR, RPV-SUSY, ...

→ other operators which violate L ↔ NSI's



$$T = M_m \langle m_\nu \rangle + M_\theta \langle tg\vartheta \rangle + M_{WR} \left\langle \left(\frac{M_1}{M_2} \right)^2 \right\rangle + M_{SUSY} \lambda_{111}^2 + M_{VR} \left\langle \frac{m_p}{M_{VR}} \right\rangle + \dots$$

Fässler

Schechter+Valle: Any L violating operator → radiative mass generation → Majorana nature of ν's

However: Might be a tiny correction to a much larger Dirac mass

→ very promising interplay of neutrino mass determinations, cosmology, LHC, LVF experiments and theory

→ see talks by A. de Gouvea, F. Cei, F. Joaquim, L. Merlo,

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Neutrinos & Cosmology

- Dark Matter ~ 25% & Dark Energy 70%
- mass of all neutrinos: $0.001 \leq \Omega_\nu \leq 0.02$
- baryonic matter $\Omega_B \sim 0.04$

Neutrino mass contribution

possibly as big as all baryonic matter >> visible matter
much more COLD dark matter & dark energy
neutrinos are an important hot dark matter component

Comological impact of neutrinos:

- hot component in structure formation: $330\nu/\text{cm}^3 \times \text{mass} \rightarrow$
structure formation & ν -mass and properties \rightarrow M. Cirelli
- Big Bang Nucleosynthesis \rightarrow G. Miele
- Baryon asymmetry \rightarrow Leptogenesis \rightarrow M. Pluemacher,
G.Branco, S. Petcov, Romanino

Present Day Acceleration

Source: Robert Kirshner

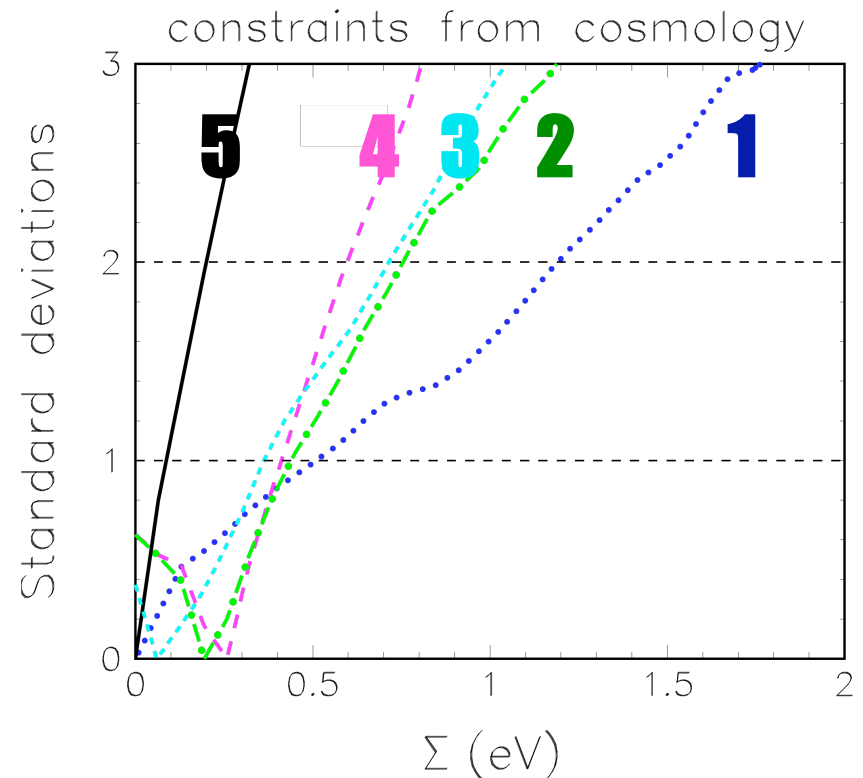
Source: David Aguilar, Harvard-Smithsonian Center for Astrophysics

Cosmological Neutrino Mass Limit(s)

Talk of A. Palazzo:

Fogli et al., Phys. Rev. D 78, 033010 (2008) [arXiv:0805.2517v3]

<i>data set</i>	<i>2σ limit</i>
1- CMB	1.19 eV
2- CMB + HST + SN-Ia	0.75 eV
3- CMB + LSS	0.72 eV
4- CMB + HST + SN + BAO	0.60 eV
5- CMB + HST + SN + BAO + Ly-α	0.19 eV



- reliability of Ly α ?
- systematic limitations for increased precision?

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 \leftrightarrow **Majorana nature**
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I. Shimizu: KamLAND results & update on solar / cleaning

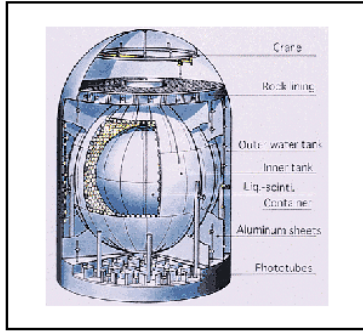
G. Prior: SNO phase III

J. Evans: Latest results from MINOS

Y. Kurimoto: SciBooNE

L. Scotto-Lavina: OPERA

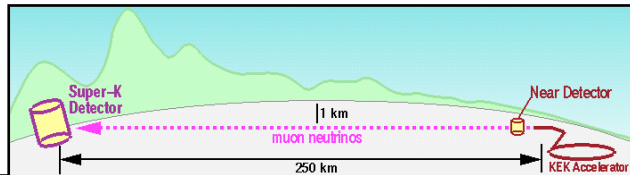
Status of Neutrino Oscillations



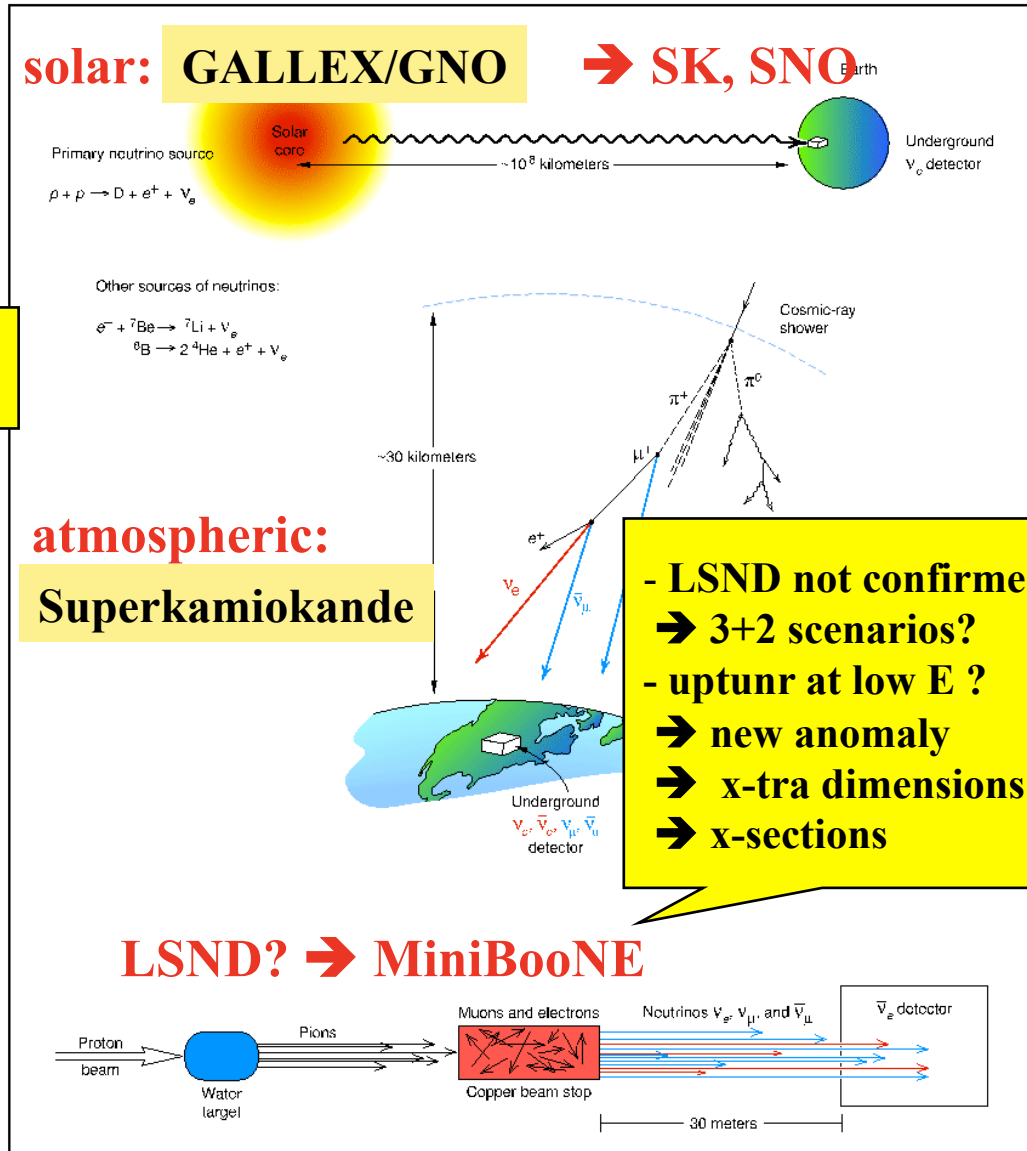
improved results

Reactors: KAMLAND

**Beams: K2K → MINOS
→ OPERA**



$\Delta m_{21}^2 = (7.67 \pm 0.18) * 10^{-5} \text{ eV}^2$
 $\sin^2 \theta_{12} = 0.312 \pm 0.019$
 $\Delta m_{31}^2 = (2.39 \pm 0.1) * 10^{-3} \text{ eV}^2$
 $\sin^2 \theta_{23} = 0.466 \pm 0.073 - 0.058$
 $\sin^2 2\theta_{13} < 0.11$ **Chooz+global fit**
sensitivity to L/E

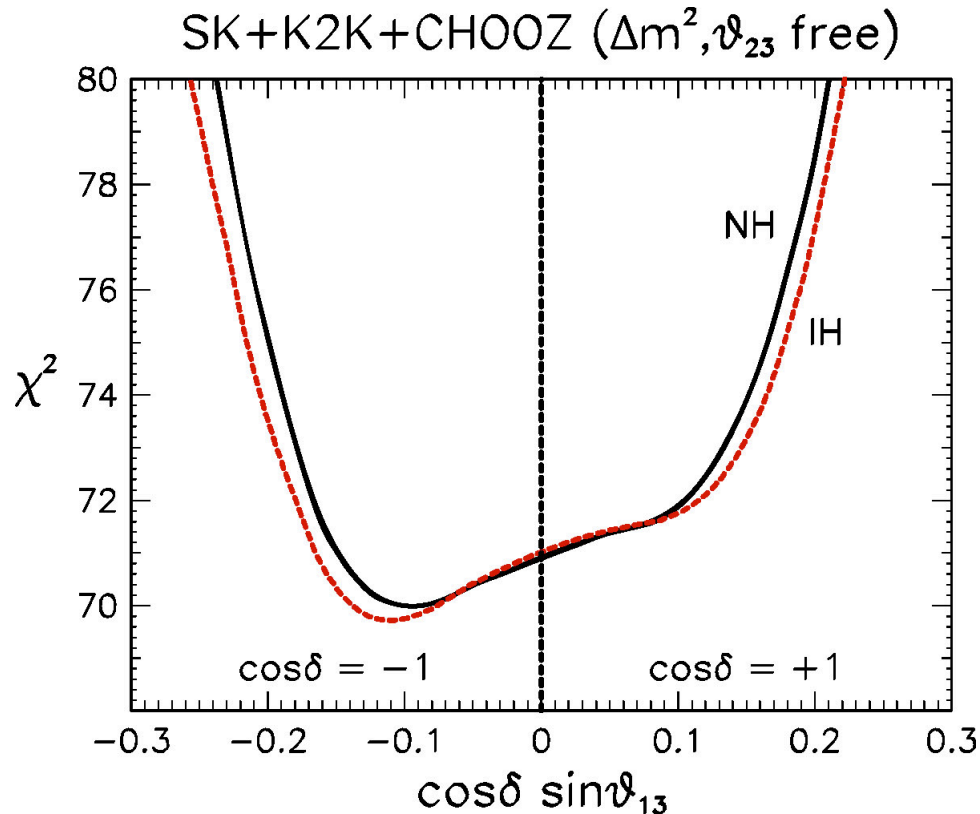


- LSND not confirmed !
→ 3+2 scenarios?
- upturn at low E ?
→ new anomaly
→ x-tra dimensions
→ x-sections

Talk of A. Palazzo:

Fogli et al., Phys. ReV. D 78, 033010 (2008) [arXiv:0805.2517v3]

Parameter	$\delta m^2 / 10^{-5} \text{ eV}^2$	$\sin^2 \theta_{12}$	$\sin^2 \theta_{13}$	$\sin^2 \theta_{23}$	$\Delta m^2 / 10^{-3} \text{ eV}^2$
Best fit	7.67	0.312	0.016	0.466	2.39
1σ range	7.48 – 7.83	0.294 – 0.331	0.006 – 0.026	0.408 – 0.539	2.31 – 2.50
2σ range	7.31 – 8.01	0.278 – 0.352	< 0.036	0.366 – 0.602	2.19 – 2.66
3σ range	7.14 – 8.19	0.263 – 0.375	< 0.046	0.331 – 0.644	2.06 – 2.81



Fogli et al. [arXiv:0806.2649]

Excess of sub-GeV e-like events
and sub-leading 3f effects \longleftrightarrow
“solar” dm^2

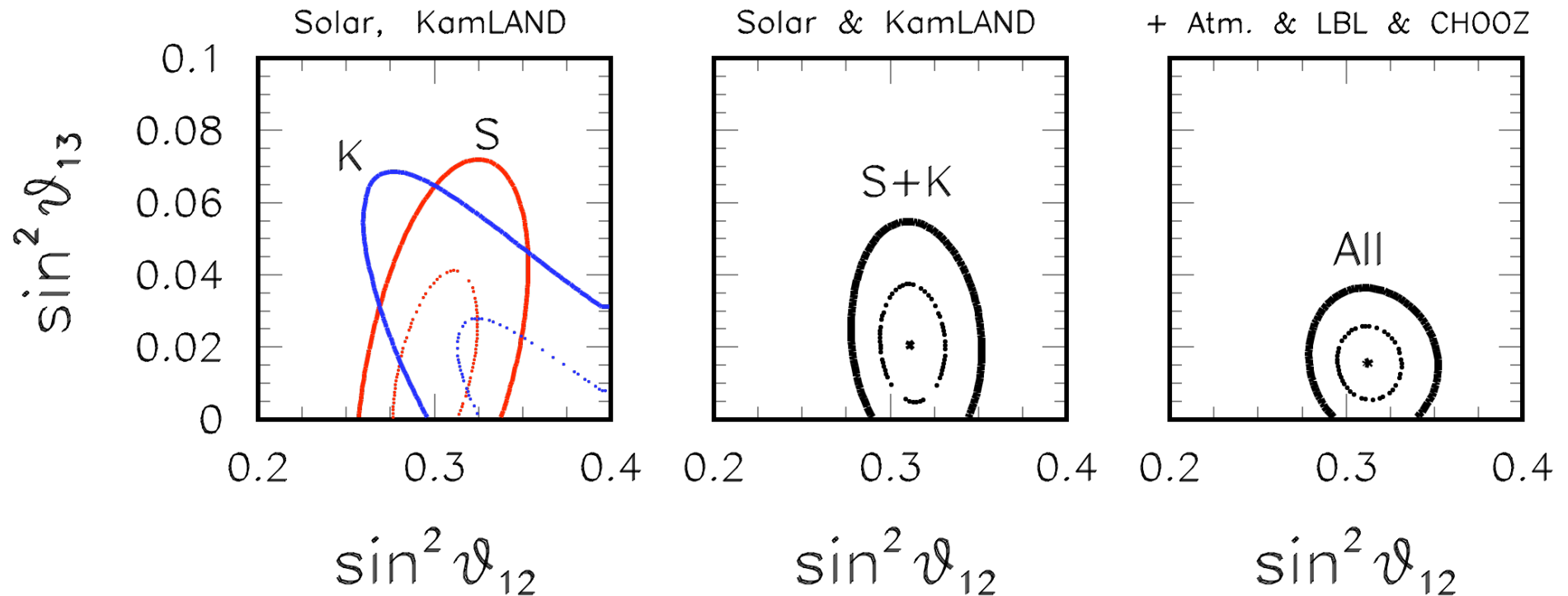
$\rightarrow \sim 1\sigma$ preference for $\sin^2 \theta_{13} > 0$

Effect also seen by other groups:
Schwetz et al.

...

Combining solar and KamLAND

A. Palazzo:



$\sim 1.2\sigma$ preference for $\theta_{13} > 0$?how robust?

Future Precision Oscillation Physics

Precise measurements → 3f oscillation formulae

$$\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \begin{array}{l} \text{x Majorana-} \\ \text{CP-phases} \end{array}$$

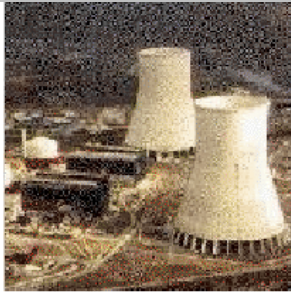
θ_{23} $S_{13} \rightarrow 3 \text{ flavour effects}$ θ_{12}
 $\rightarrow \text{CP phase } \delta$

Aims: → improved precision of the leading 2x2 oscillations
→ detection of generic 3-neutrino effects: θ_{13} , CP violation

Complication: Matter effects → effective parameters in matter

→ expansion in small quantities θ_{13} and $\alpha = \Delta m_{\text{sol}}^2 / \Delta m_{\text{atm}}^2$
Burguet-Castell et al., Akhmedov et al. ...

Future Precision with Reactor Experiments



$\bar{\nu}_e$

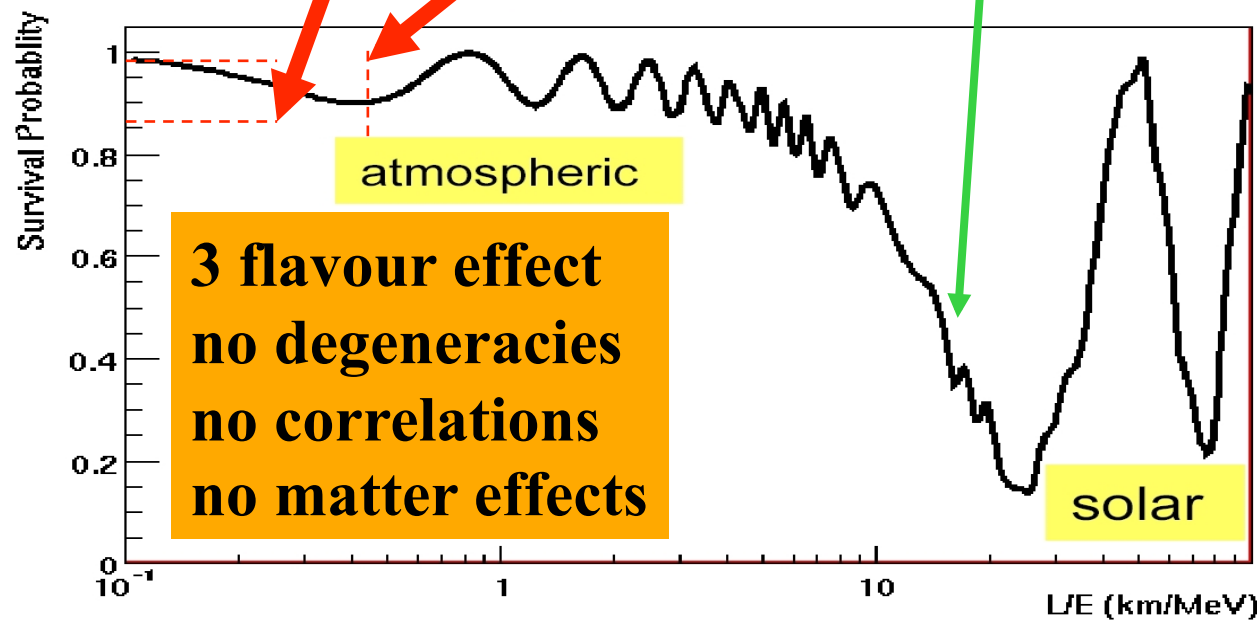
near detector (170m)

$\bar{\nu}_e$

far detector (1700m)

identical detectors → many errors cancel

$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{31}^2 L}{4E_\nu} - \left(\frac{\Delta m_{21}^2 L}{4E_\nu} \right)^2 \cos^4 \theta_{13} \sin^2 2\theta_{12}$$



- Double Chooz
- Daya Bay
- Reno
- Angra

clean & precise θ_{13} measurements

E=4MeV → 2km 4km 40km 80km

New Reactor Experiments

S. Peeters

(overview & Double Chooz)

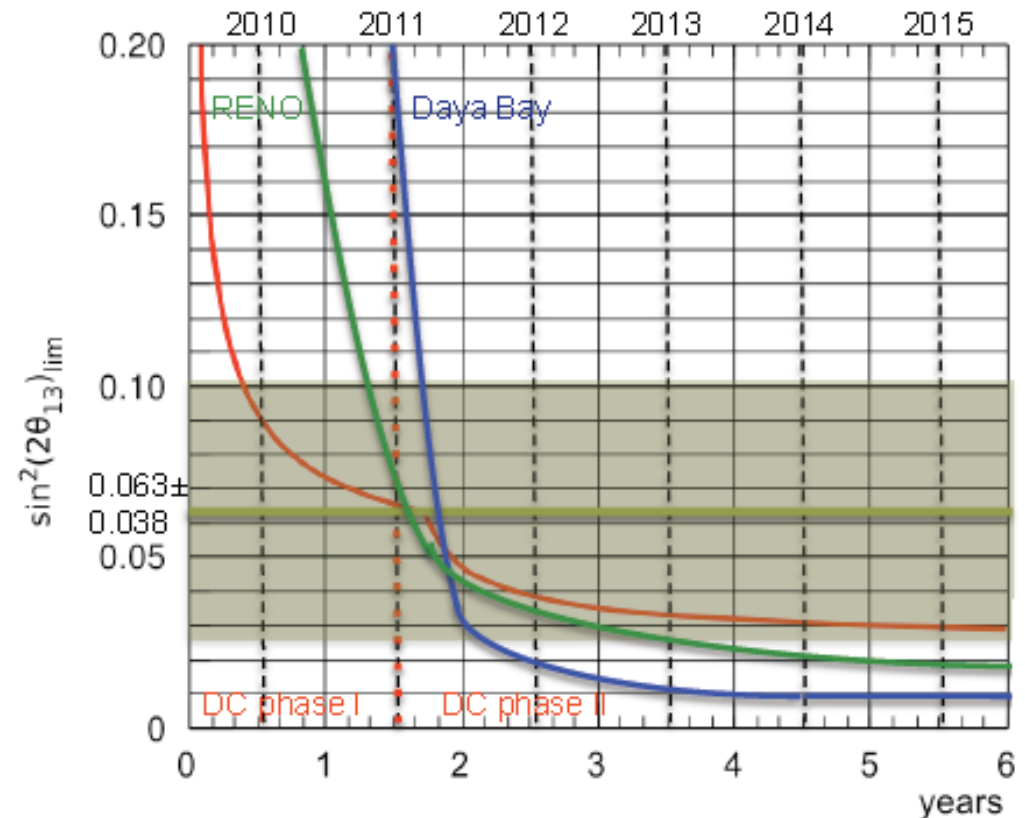
Y. Oh (RENO)

D. Lhuillier (monitoring)

Experiments	Location	Thermal Power (GW)	Distances Near/Far (m)	Depth Near/Far (mwe)	Target Mass (tons)
Double-CHOOZ	France	8.7	280/1050	60/300	10/10
RENO	Korea	16.4	290/1380	120/450	15/15
Daya Bay	China	11.6	360(500)/1985(1613)	260/910	40×2/80

- promising experiments
- precision is very demanding:
 - stable scintillator (c.f. Chooz!)
 - technological challenges
 - backgrounds
 - systematics
 - different optimism on schedules

→ comparison possible, but they depend on non-trivial assumptions...



Future Precision with New Neutrino Beams

- conventional beams, superbeams
→ MINOS, CNGS, T2K, NOvA, T2H,...
- β -beams
→ pure ν_e and $\bar{\nu}_e$ beams from radioactive decays; $\gamma \simeq 100$
- neutrino factories
→ clean neutrino beams from decay of stored μ 's

$$P(\nu_e \rightarrow \nu_\mu) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2((1-\hat{A})\Delta)}{(1-\hat{A})^2}$$
$$\pm \sin \delta_{\text{CP}} \alpha \sin 2\theta_{12} \cos \theta_{13} \sin 2\theta_{13} \sin 2\theta_{23} \sin(\Delta) \frac{\sin(\hat{A}\Delta) \sin((1-\hat{A})\Delta)}{\hat{A}(1-\hat{A})}$$
$$+ \sin \delta_{\text{CP}} \alpha \sin 2\theta_{12} \cos \theta_{13} \sin 2\theta_{13} \sin 2\theta_{23} \cos(\Delta) \frac{\sin(\hat{A}\Delta) \sin((1-\hat{A})\Delta)}{\hat{A}(1-\hat{A})}$$
$$+ \alpha^2 \sin^2 2\theta_{12} \cos^2 \theta_{23} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}$$

↳ correlations & degeneracies, matter effects

Channels of interest

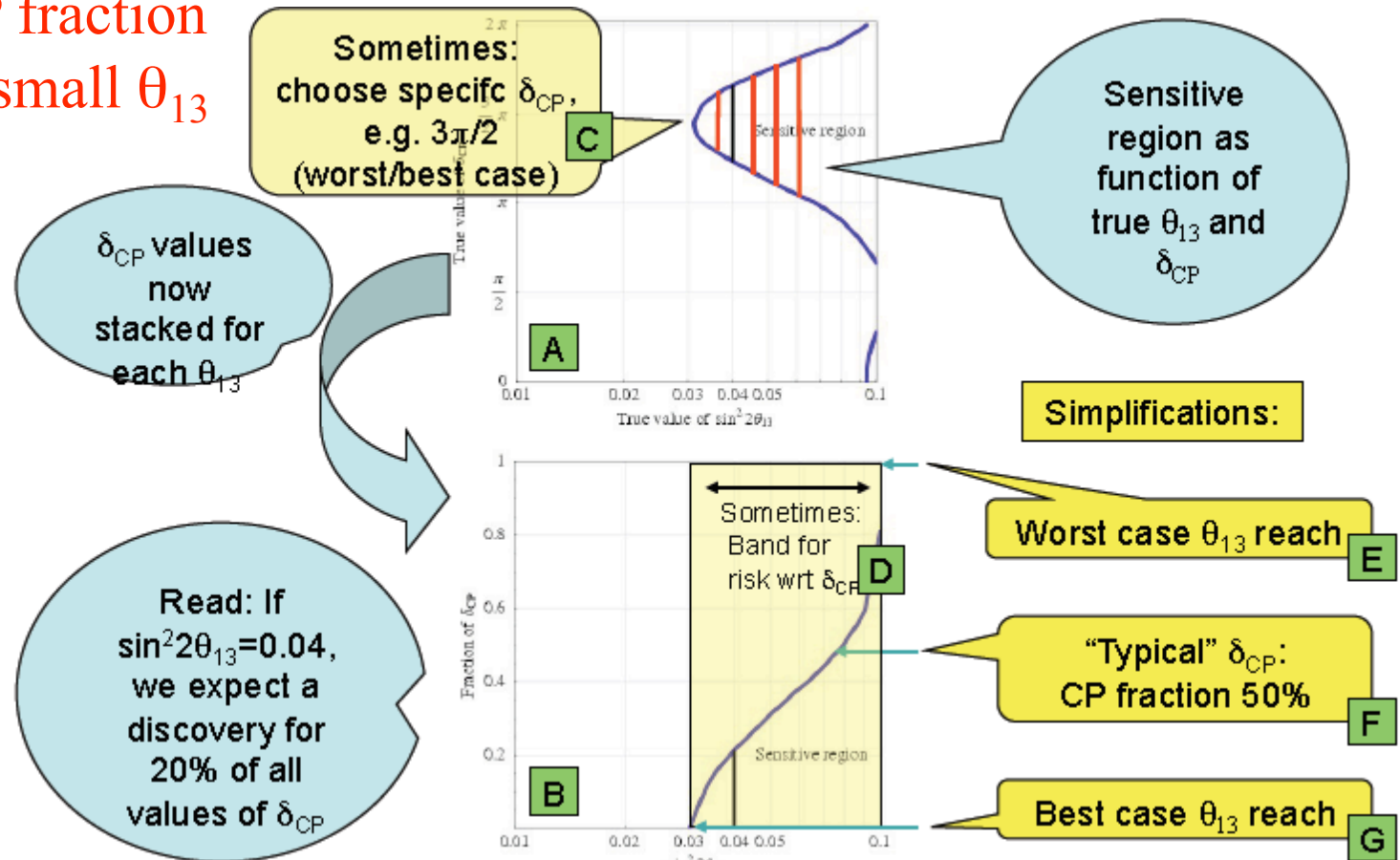
- Disappearance for Δm_{31}^2 , θ_{23} : $\nu_\mu \Rightarrow \nu_\mu$
 $1 - P_{\mu\mu} = \sin^2 2\theta_{23} \sin^2 \Delta_{31} + \text{h.o.t.}$ $\Delta_{31} = \Delta m_{31}^2 L/(4E)$
NB: We expand in $\sin 2\theta_{13}$ and $\alpha \equiv \Delta m_{21}^2/\Delta m_{31}^2$ ($|\alpha| \sim 0.03$)
- Appearance for θ_{13} , CPV, MH:
 - Golden: $\nu_e \Rightarrow \nu_\mu$ (NF/BB) or $\nu_\mu \Rightarrow \nu_e$ (SB)
(e.g., De Rujula, Gavela, Hernandez, 1999; Cervera et al, 2000)
 - Silver: $\nu_e \Rightarrow \nu_\tau$ (NF – low statistics!?)
(Donini, Meloni, Migliozzi, 2002; Autiero et al, 2004)
 - Platinum: $\nu_\mu \Rightarrow \nu_e$ (NF: difficult!)
(see e.g. ISS physics working group report)
- Other appearance: $\nu_\mu \Rightarrow \nu_\tau$ (OPERA, NF?)
- Neutral currents for new physics
(e.g., Barger, Geer, Whisnant, 2004; MINOS, 2008)

Resolving degeneracies: information beyond rates @ fixed L/E

→ different L/E, different channels & beams, spectral info., ...

Optimization:

- more than one quantity (θ_{13} , δ , θ_{ij} , hierarchy, BSM,...),
- definitions: CP fraction
- physics: large/small θ_{13}



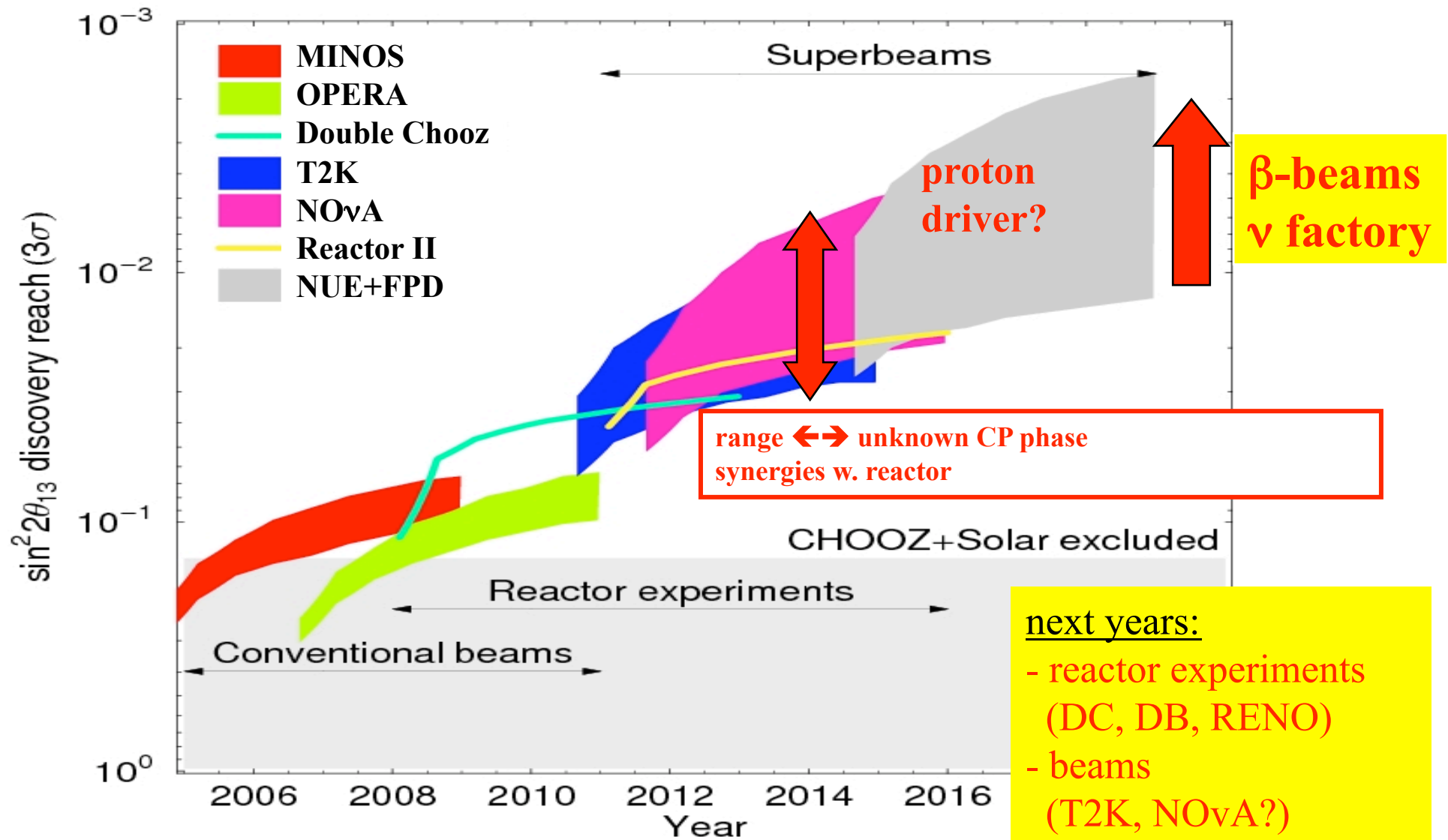
H. Kakuno (T2K → impressive progress)
M. Goodman (MINOS, NOvA & US perspective)
F. Ronga (LBL@ LNGS & EU perspective)
P. Sala (LNGS ν beam)
M. Mezzetto (beta beams)
S. Goswami (INO)
F. Dufour (T2KK)
A. Guglielmi (MODULAR)
T. Schwetz (future atmospheric ν 's)
M. Bonesini (R&D towards a neutrino factory)
K. Long (International design study)



- 1) build and exploit approved projects (T2K, ..., reactor)**
- 2) be aware of technological, physics (ν , LHC, LFV) and political uncertainties / problems**
 - now: get R&D for new beams, detectors done ...!**
 - aim at decisions depending on results in a few years from now (2012?)**

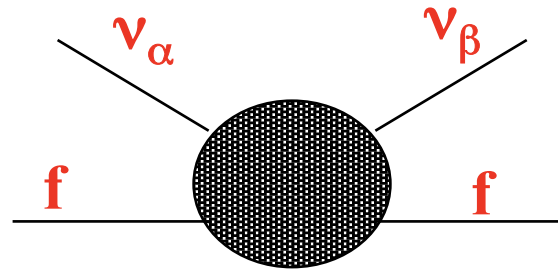
see e.g. LAGUNA – talk by L. Oberauer

θ_{13} – Now and in the Future



New Physics, NSIs & ν -Oscillations

See talks by
T. Rashba and T. Ota



$$|\epsilon| \simeq \frac{M_W^2}{M_{NSI}^2}$$

Future precision oscillation experiments:

Source	⊗	Oscillation	⊗	Detector
<ul style="list-style-type: none"> - neutrino energy E - flux and spectrum - flavour composition - contamination - symmetric $\nu/\bar{\nu}$ operation 		<ul style="list-style-type: none"> - oscillation channels - realistic baselines - MSW matter profile - degeneracies - correlations 		<ul style="list-style-type: none"> - effective mass, material - threshold, resolution - particle ID (flavour, charge, event reconstruction, ...) - backgrounds - x-sections (at low E)

precision experiments might see new effects beyond oscillations!

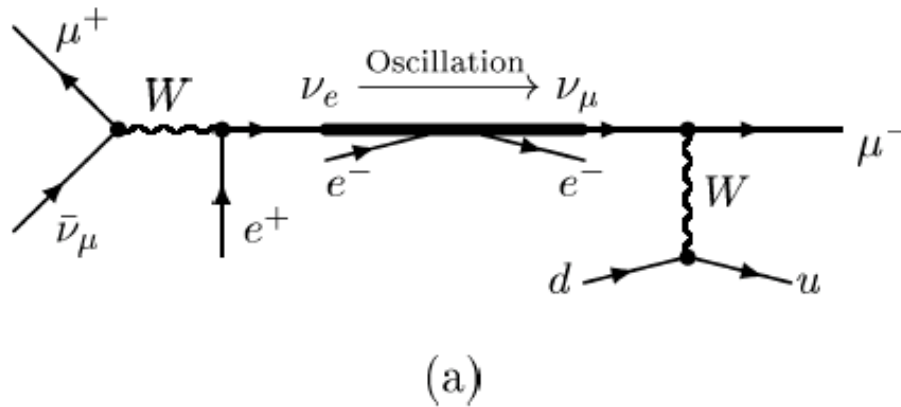
➔ **modifications of 3f oscillation formulae, different L/E**

➔ **small event rates: offset in oscillation parameters**

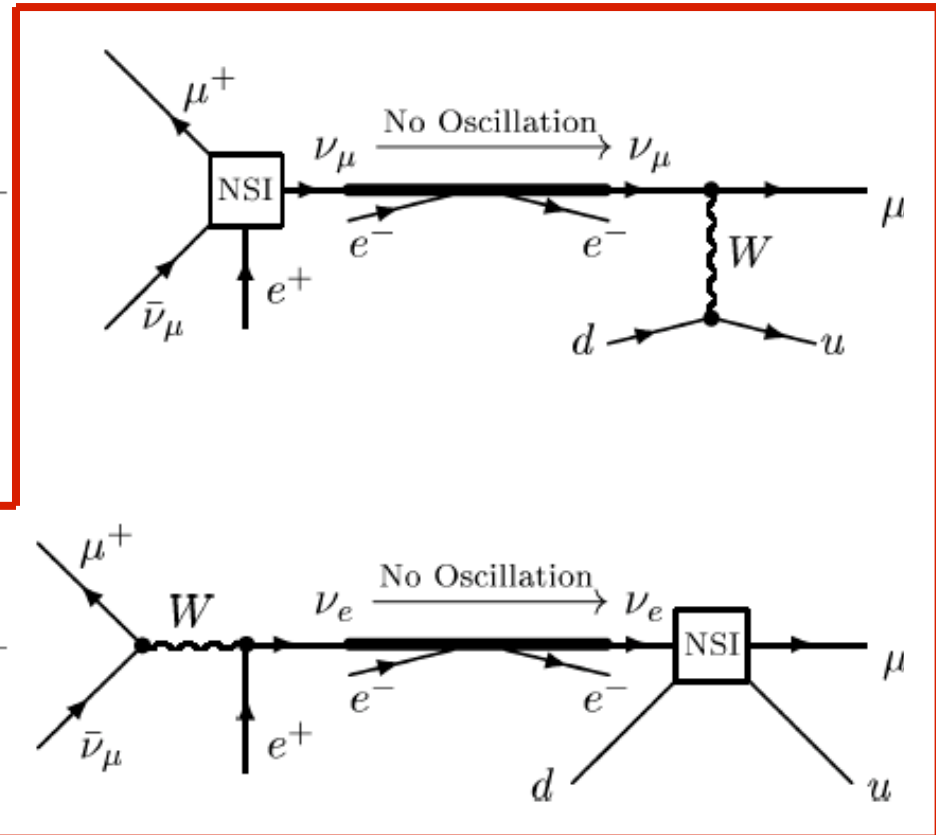
➔ **Non Standard Interactions = NSI's**

NSIs interfere with Oscillations

the “golden” oscillation channel

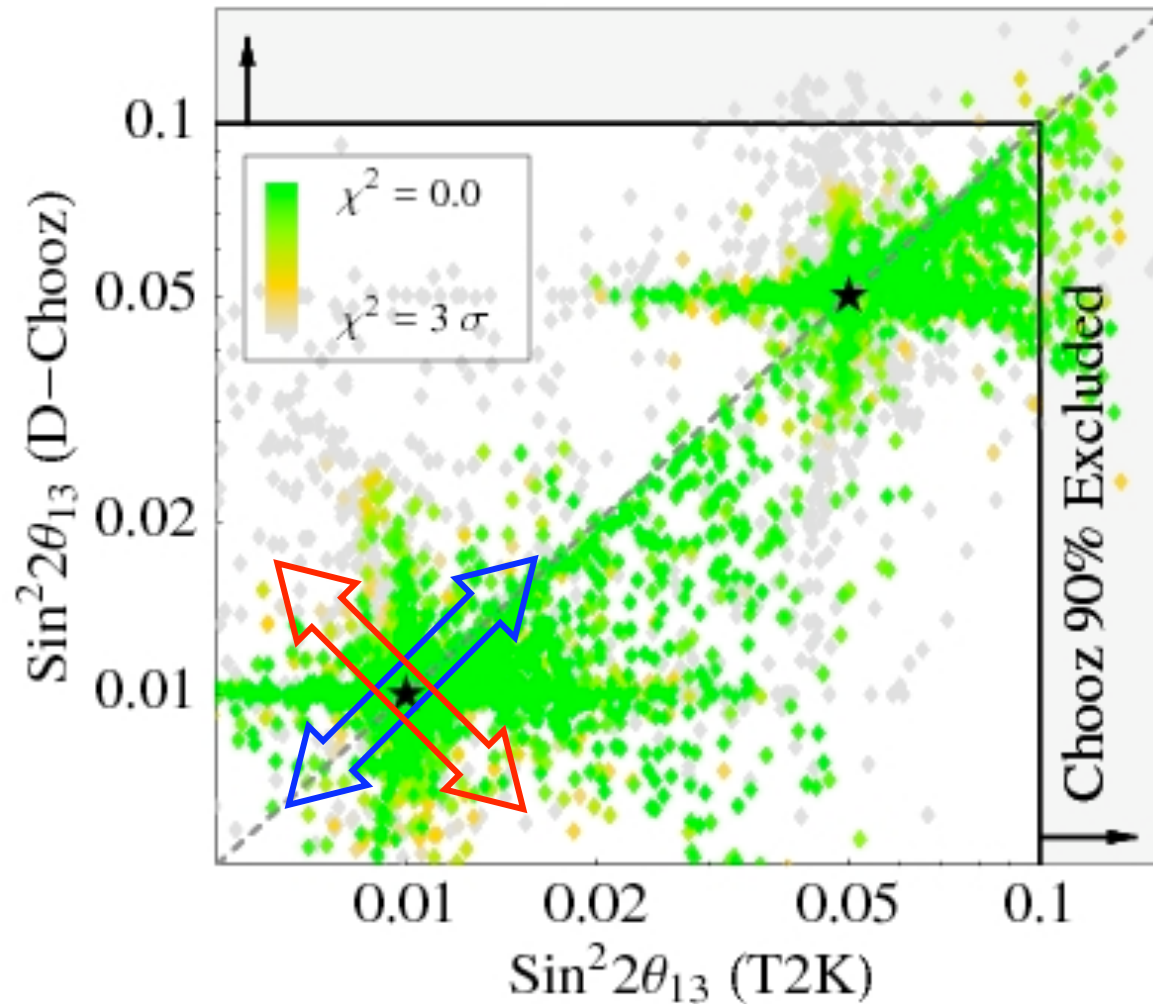


NSI contributions to the “golden” channel



note: interference in oscillations $\sim \epsilon$ \leftrightarrow FCNC effects $\sim \epsilon^2$

NSI: Offset and Mismatch in θ_{13}



redundant measurement of θ_{13}

Double Chooz + T2K

***=assumed 'true' values of θ_{13}**

scatter-plot:

- ϵ values random
- below existing bounds
- random phases

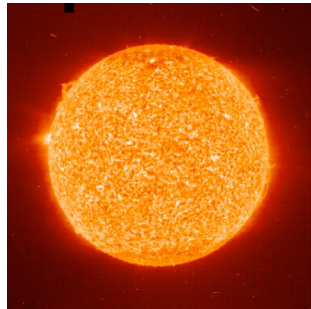
NSIs can lead to:

- **offset**
- **mismatch**

- ➔ **redundancy**
- ➔ **interesting potential**

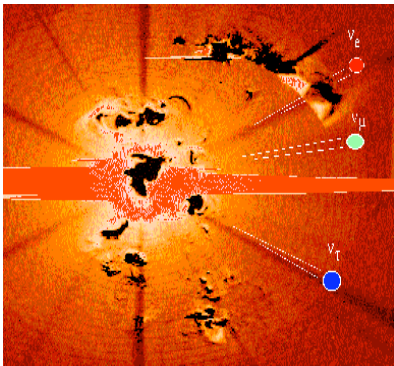
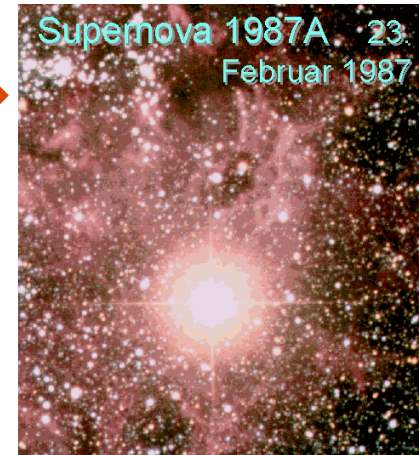
Kopp, ML, Ota, Sato

New Physics: Neutrino Sources



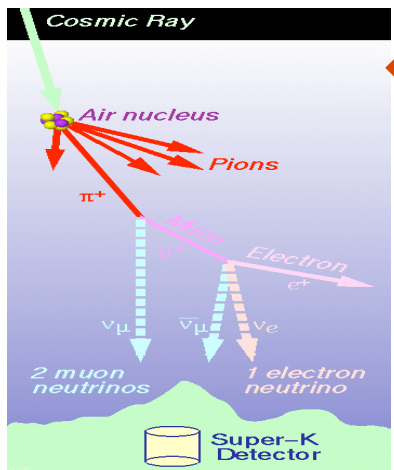
← Sun

Astronomy: →
Supernovae
GRBs
UHE ν 's



← **Cosmology**

Reactors →

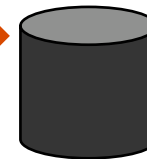


← **Atmosphere**

Accelerators →



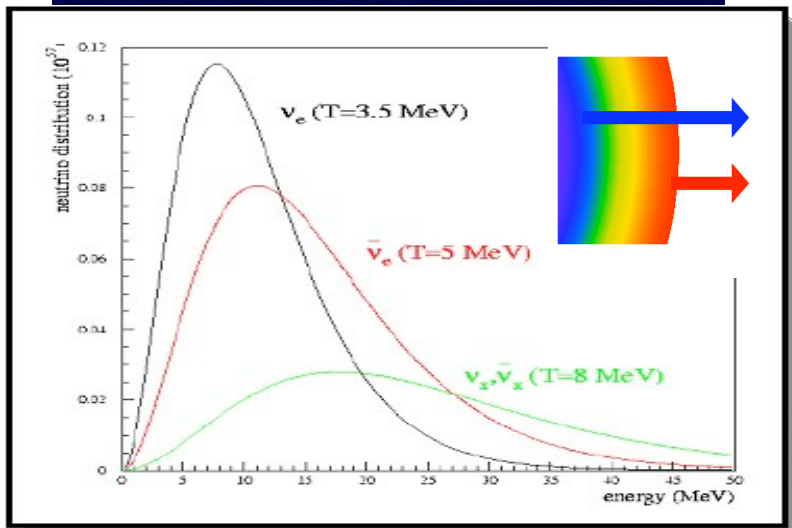
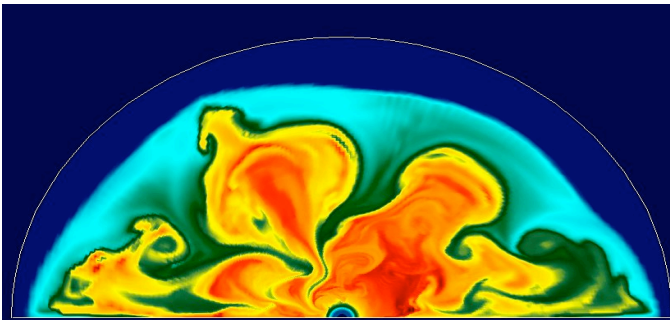
β -Sources →



← **Earth**

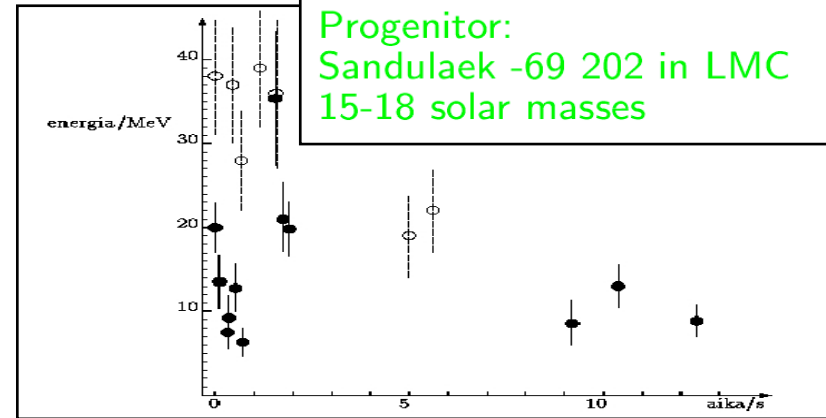
Supernova Neutrinos

- Collaps of a typical star $\rightarrow \sim 10^{57}$ ν 's
- $\sim 99\%$ of the energy in ν 's
- ν 's essential for explosion
- **3d simulations do not explode**
(so far... 2d \rightarrow 3d, \rightarrow convection? ...)



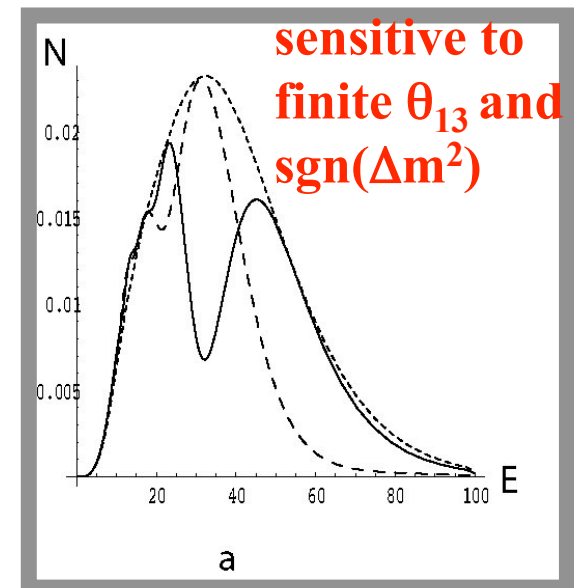
SN1987A neutrino burst

Progenitor:
Sandulaek -69 202 in LMC
15-18 solar masses



MSW: SN & Earth
non-linearity, ...

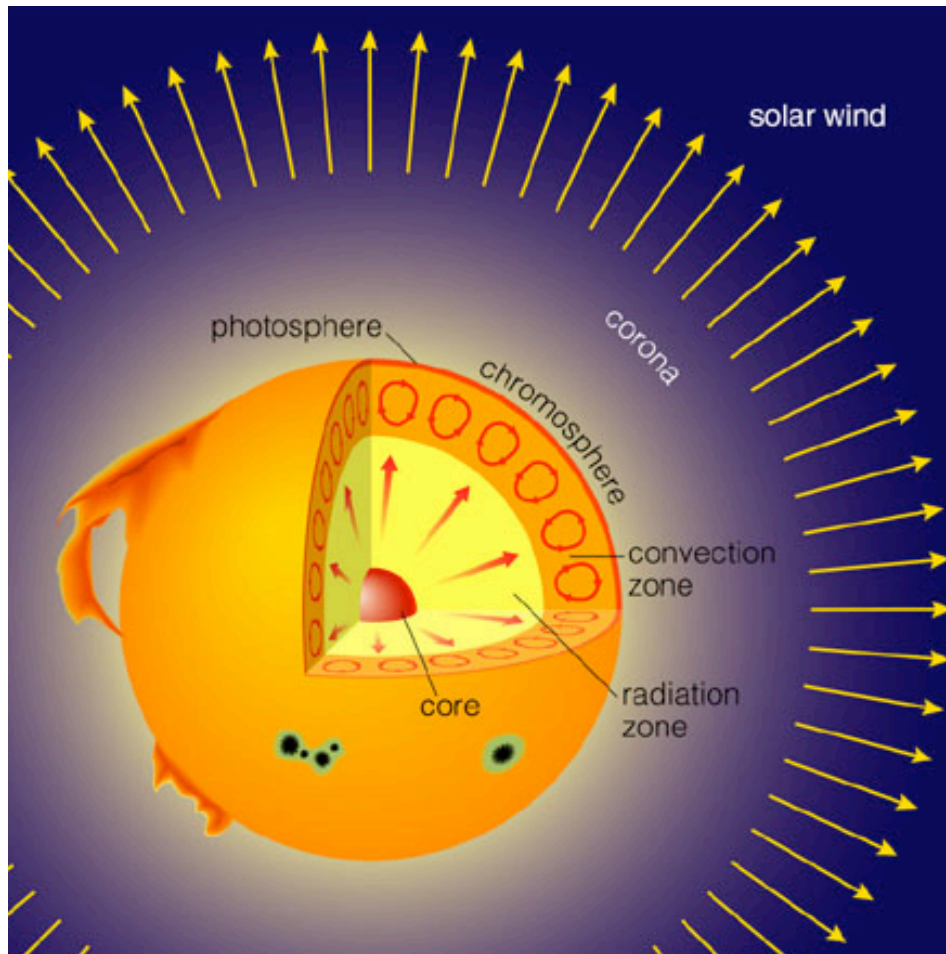
G. Sigl
C. Cardall
B. Dasgupta
A. Esteban



Solar Neutrinos: Learning About the Sun

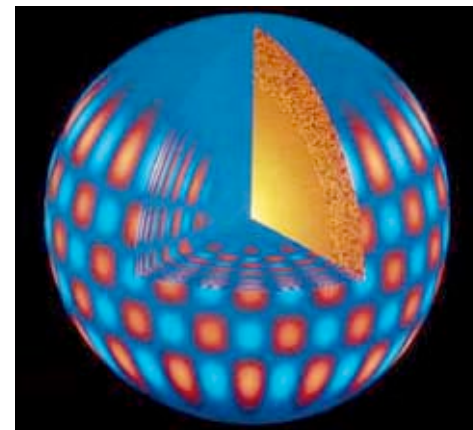
Observables:

- **optical** (total energy, surface dynamics, sun-spots, historical records, B, ...)
- **neutrinos** (rates, spectrum, ...)

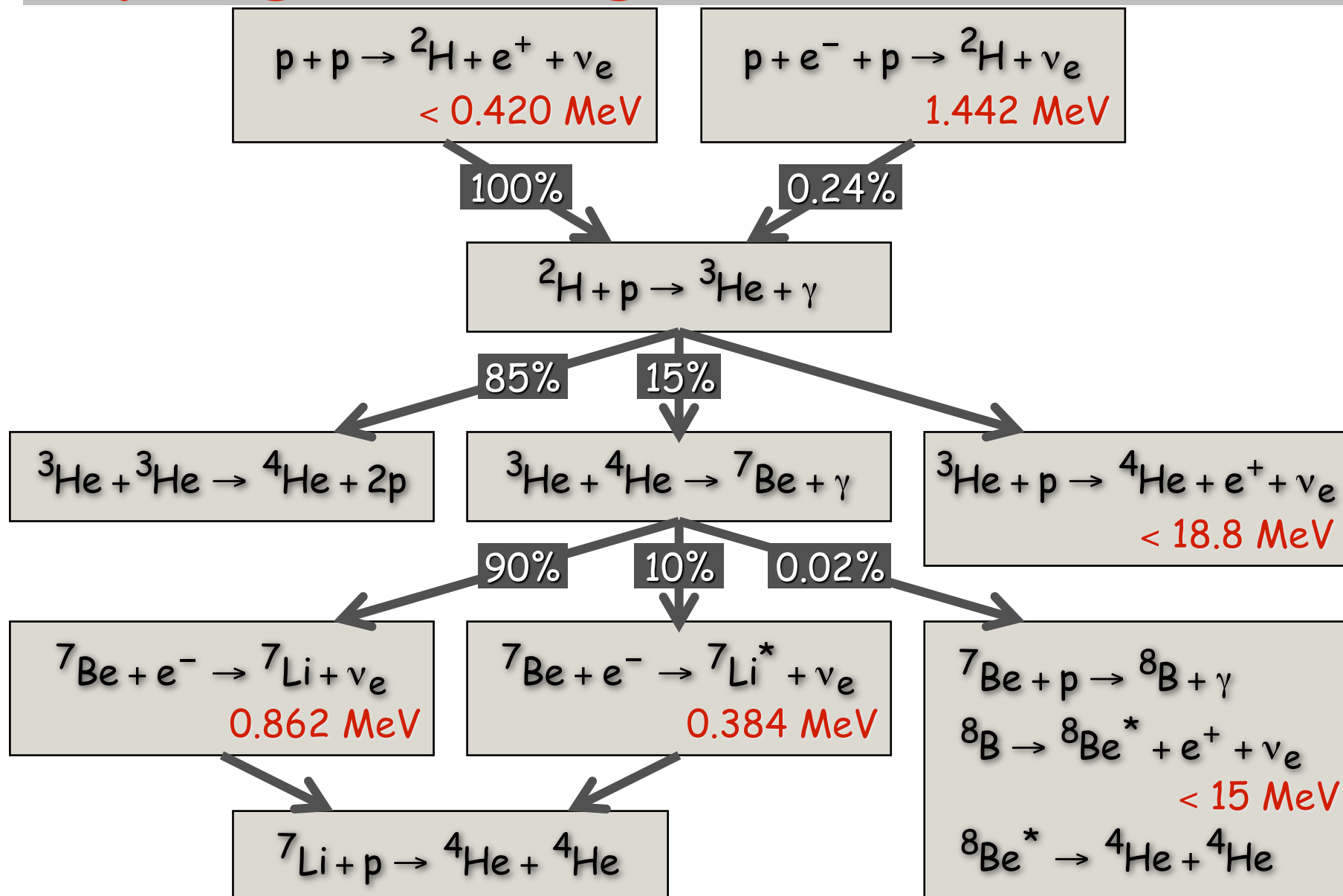


Topics:

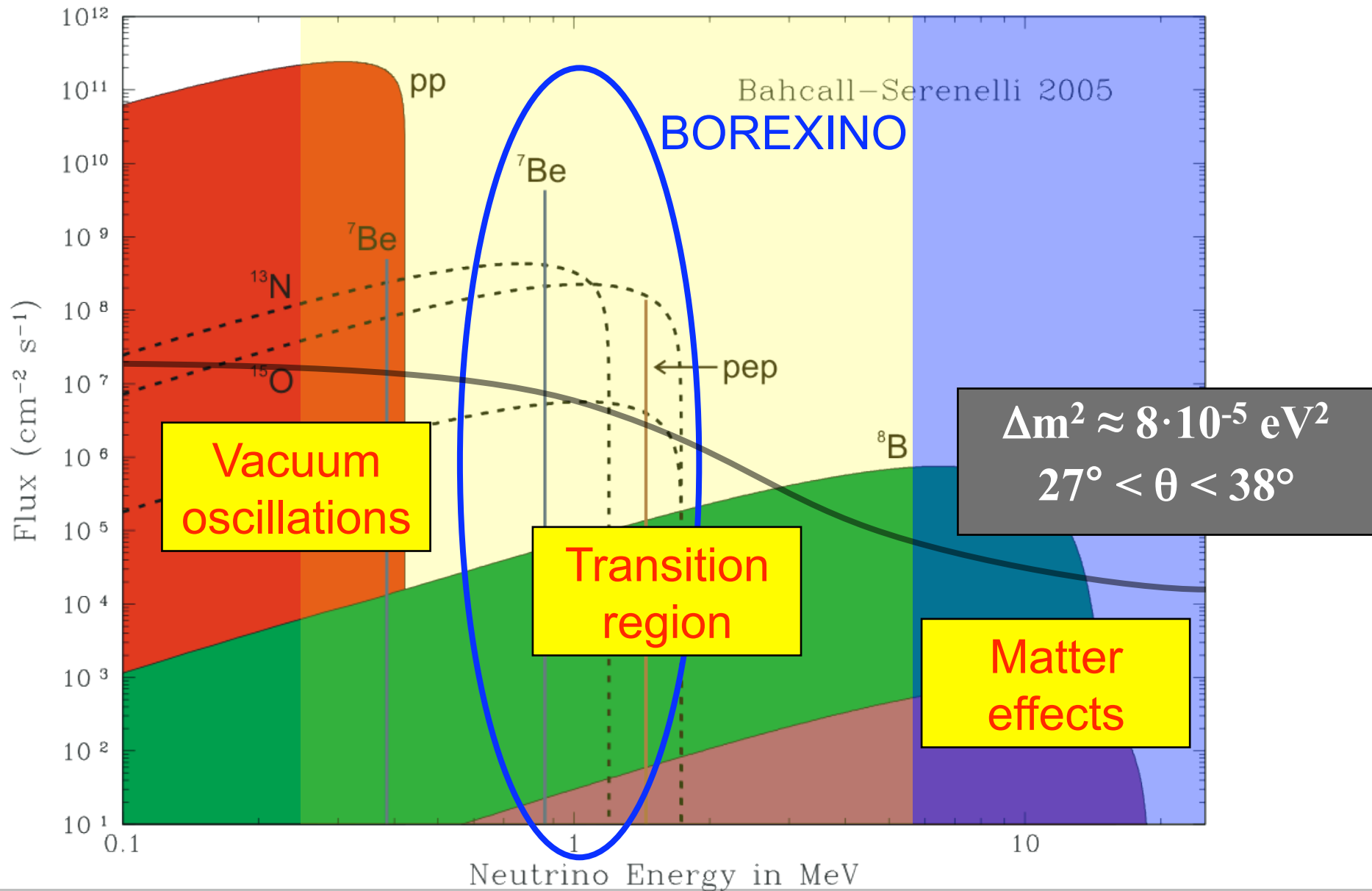
- **nuclear cross sections**
(at finite $T \sim \text{few MeV}$)
- **solar dynamics**
- **helio-seismology**
- **variability**
- **composition**
- ➔ **recent debate about metallicity**



Hydrogen Burning: Proton-Proton Chains

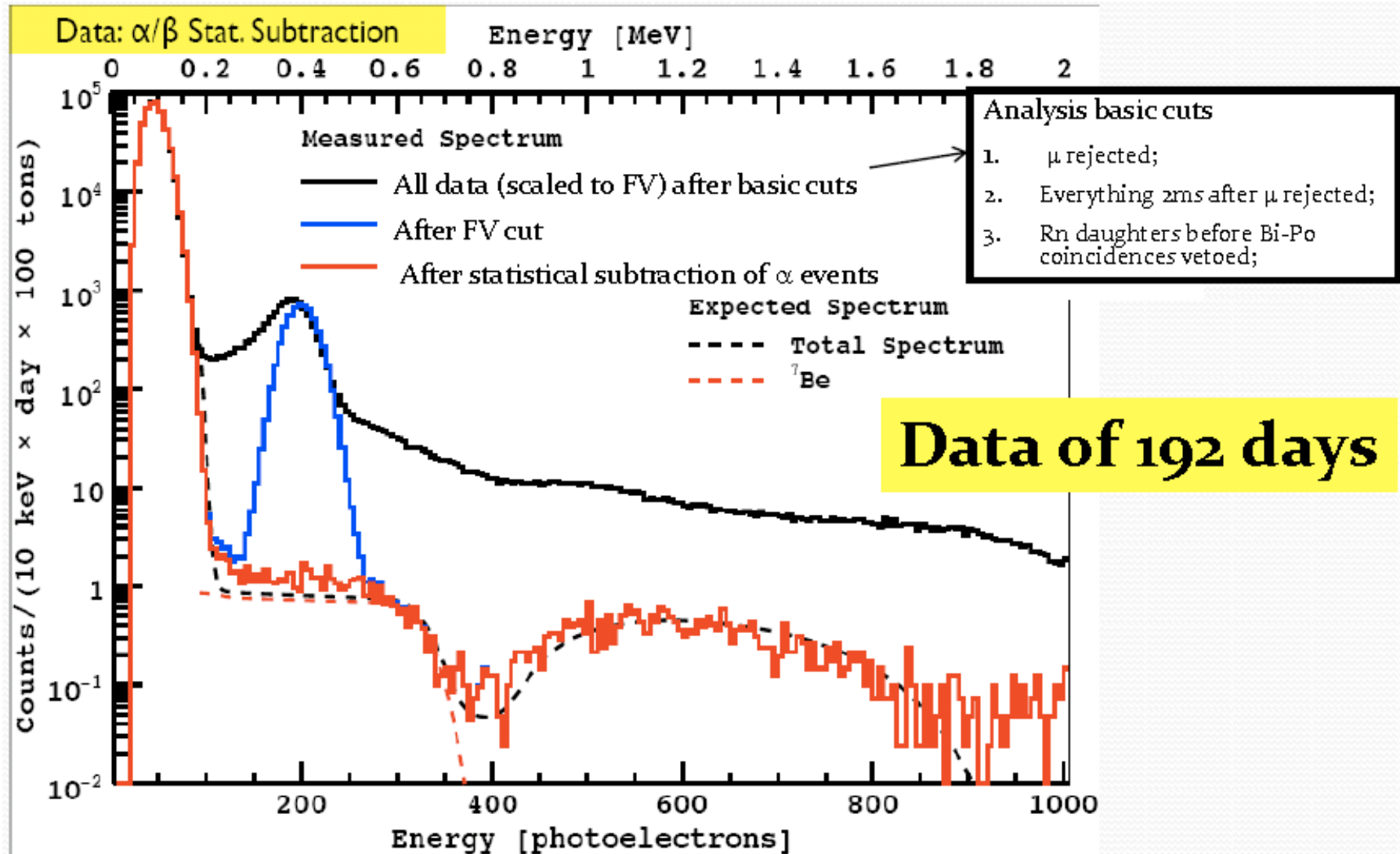


Solar Neutrino Spectroscopy

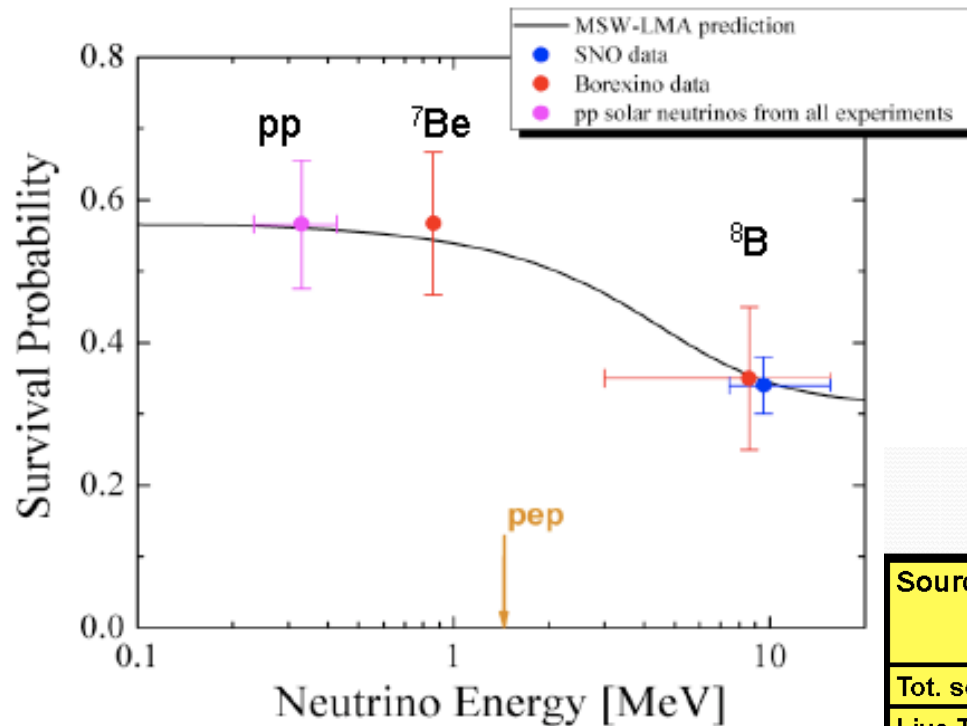


Borexino Results

talks by L. Ludhova and D. Franco



Survival probability after Borexino



First simultaneous measurement in both vacuum-dominated and matter-enhanced regions

Systematic uncertainties

Source	Syst.error (1 σ)
Tot. scint. mass	0.2%
Live Time	0.1%
Efficiency of Cuts	0.3%
Detector Resp.Function	6%
Fiducial Mass	6%
TOT	8.5%

$49 \pm 3_{\text{stat}} \pm 4_{\text{sys}}$ cpd/100 tons

	Expected rate (cpd/100 t)
No oscillation	75 ± 4
BPS07(GS98) HighZ	48 ± 4
BPS07(AGS05) LowZ	44 ± 4

No-oscillation hypothesis rejected at 4σ level

Assuming high-Z SSM (BPS 07) the ^8B rate measurement corresponds to

$P_{ee} (^8\text{Be}) = 0.35 \text{ } 0.10 \text{ @ } 8.6 \text{ MeV mean energy}$

Neutrino magnetic moment

SM with $m_\nu = 0$: $\mu_\nu = 0$

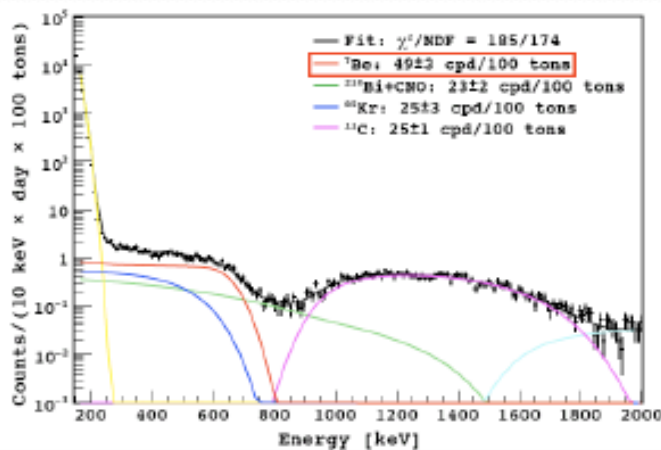
T - kinetic energy of scattered e
 E_ν - neutrino energy

$$\left(\frac{d\sigma}{dT}\right)_W = \frac{2G_F^2 m_e}{\pi} \left[g_L^2 + g_R^2 + \left(1 - \frac{T}{E_\nu}\right)^2 - g_L g_R \frac{m_e T}{E_\nu^2} \right]$$

SM with $m_\nu > 0$: $\mu_\nu > 0$,
 additional EM term
 influencing the cross section
 and thus the spectral shape

$$\left(\frac{d\sigma}{dT}\right)_{EM} = \mu_\nu^2 \frac{\pi \alpha_{em}^2}{m_e^2} \left(\frac{1}{T} - \frac{T}{E_\nu} \right)$$

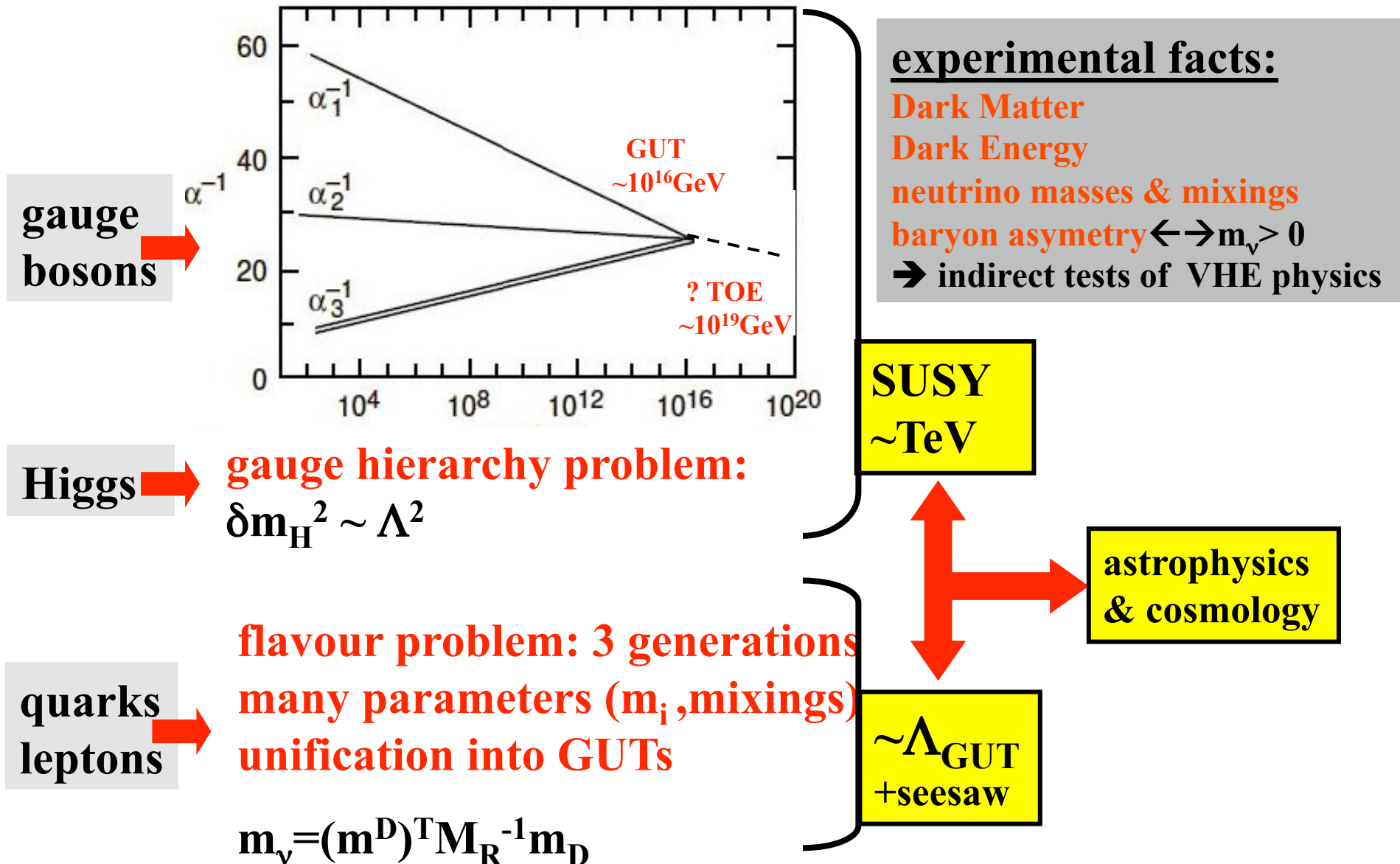
Sensitivity
 enhanced
 @ low energies



Estimate8	Method	90% C.L. $10^{-11} \mu_B$
SuperK	^8B above 5 MeV	< 11
Montanino <i>et al.</i>	^7Be (Borexino data)	< 8.4
GEMMA	Reactor anti- ν	< 5.8
Borexino	^7Be	< 5.4

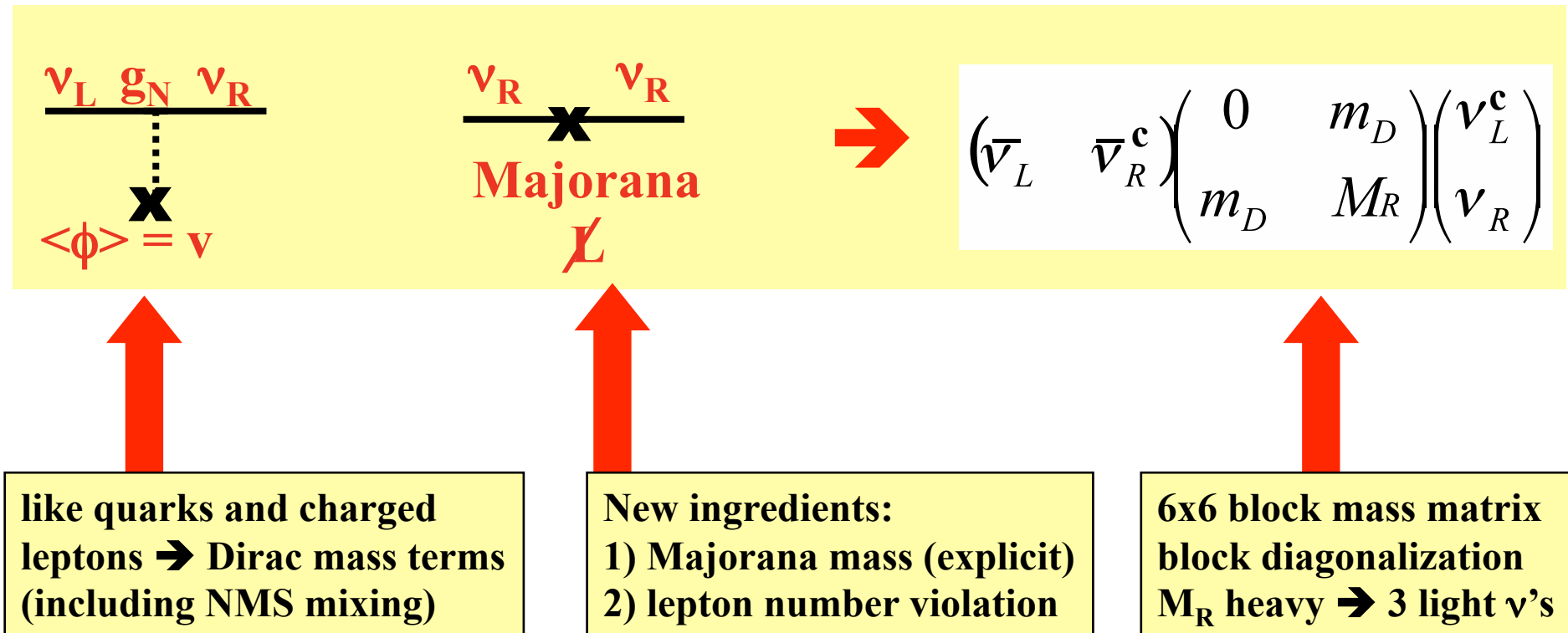
Currently the best experimental limit!

Theory: Different Routes Beyond the SM



Adding Neutrino Mass Terms

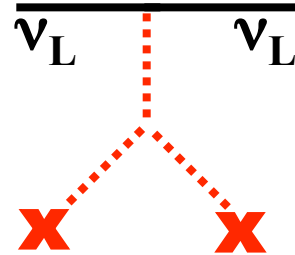
1) Simplest possibility: add 3 right handed neutrino fields



NEW ingredients, 9 parameters \rightarrow SM+

Other Neutrino Mass Operators

2) new Higgs triplets Δ_L :



→ left-handed Majorana mass term:

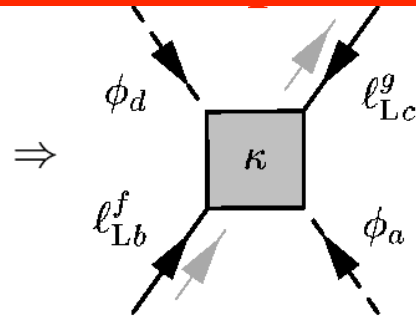
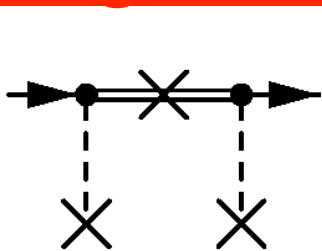
→ $M_L \bar{L} L^c$

3) Both ν_R and new Higgs triplets Δ_L :

→ see-saw type II

$m_\nu = M_L - m_D M_R^{-1} m_D^T$

4) Higher dimensional operators: $d=5, \dots$



$\Leftrightarrow \mathcal{L}_{mass} = \kappa \cdot \bar{\nu}_L^C \nu_L \Phi^T \Phi$

→ $M_L \bar{L} L^c$

5-N) ...

Other effective Operators Beyond the SM

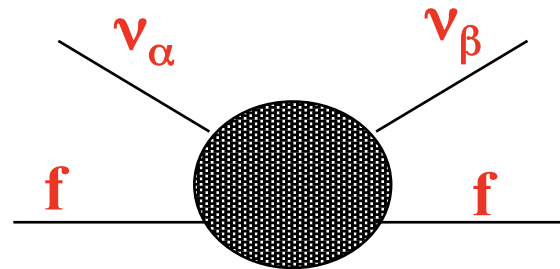
→ effects beyond 3 flavours

→ **Non Standard Interactions = NSIs** → effective 4f operators

$$\mathcal{L}_{NSI} \simeq \epsilon_{\alpha\beta} 2\sqrt{2}G_F (\bar{\nu}_{L\beta} \gamma^\rho \nu_{L\alpha}) (\bar{f}_L \gamma_\rho f_L)$$

- **integrating out heavy physics (c.f. $G_F \leftrightarrow M_W$)**

$$|\epsilon| \simeq \frac{M_W^2}{M_{NSI}^2}$$



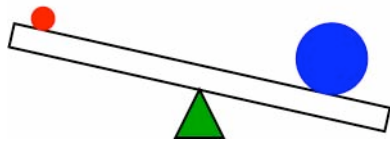
Suggestive Seesaw Features

QFT: natural value of mass operators \leftrightarrow scale of symmetry

$m_D \sim$ electro-weak scale

$M_R \sim$ L violation scale $\leftarrow? \rightarrow$ embedding (GUTs, ...)

See-saw mechanism (type I)



$$m_\nu = m_D M_R^{-1} m_D^T$$

$$m_h = M_R$$

Numerical hints:

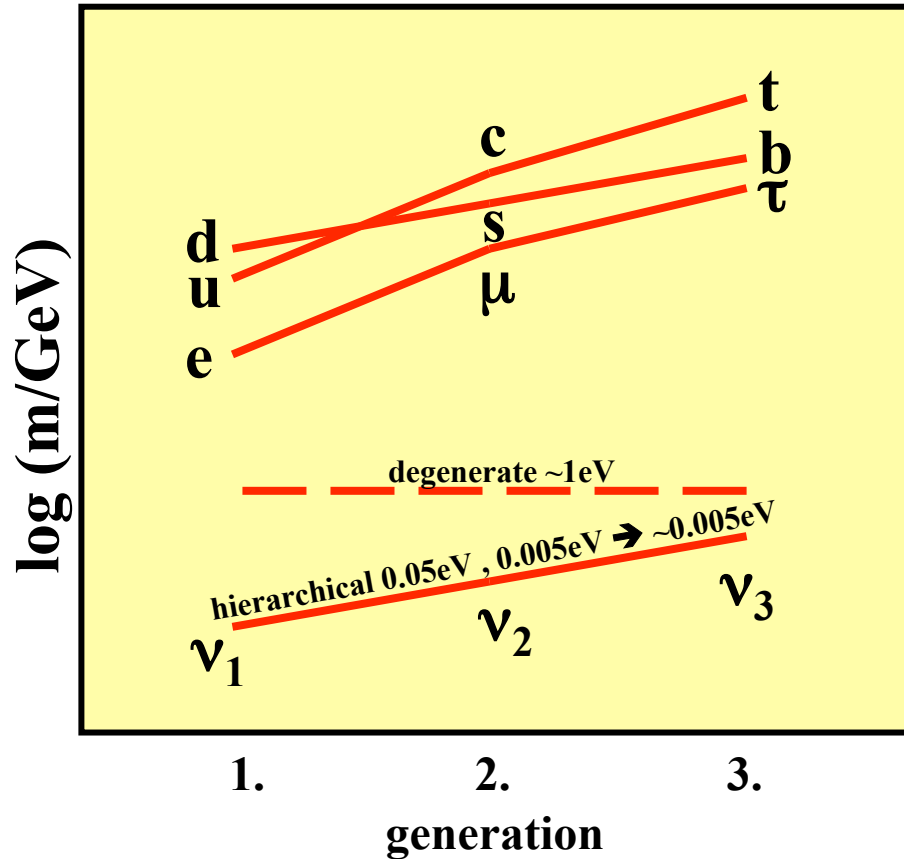
For $m_3 \sim (\Delta m_{\text{atm}}^2)^{1/2}$, $m_D \sim$ leptons $\rightarrow M_R \sim 10^{11} - 10^{16} \text{ GeV}$

$\rightarrow \nu$'s are **Majorana particles**, m_ν probes \sim GUT scale physics!

\rightarrow smallness of $m_\nu \leftrightarrow$ high scale of L , symmetries of m_D, M_R

2nd Look Questions

Quarks & charged leptons → hierarchical masses → neutrinos?



Quarks and charged leptons:

$$m_D \sim H^n ; n = 0, 1, 2 \rightarrow H \geq 20 \dots 200$$

Neutrinos:

$$m_\nu \sim H^n \rightarrow H \leq \sim 10$$

See-saw:

$$m_\nu = -m_D^T M_R^{-1} m_D$$



H	~ 10	≥ 20	?	≥ 20
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- less hierarchy in m_D or correlated hierarchy in M_R ? → theoretically connected!
- mixing patterns: not generically large, why almost maximal, θ_{13} small?

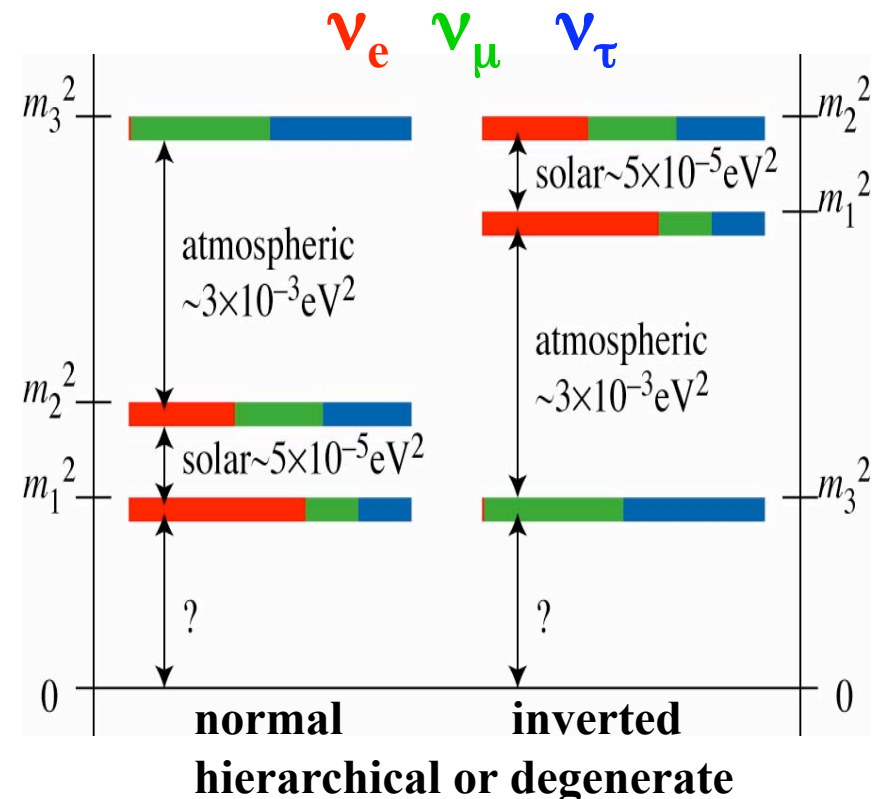
Parameters for 3 Light Neutrinos

mass & mixing parameters: m_1 , Δm^2_{21} , $|\Delta m^2_{31}|$, $\text{sign}(\Delta m^2_{31})$

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \text{diag}(e^{i\alpha}, e^{i\beta}, 1)$$

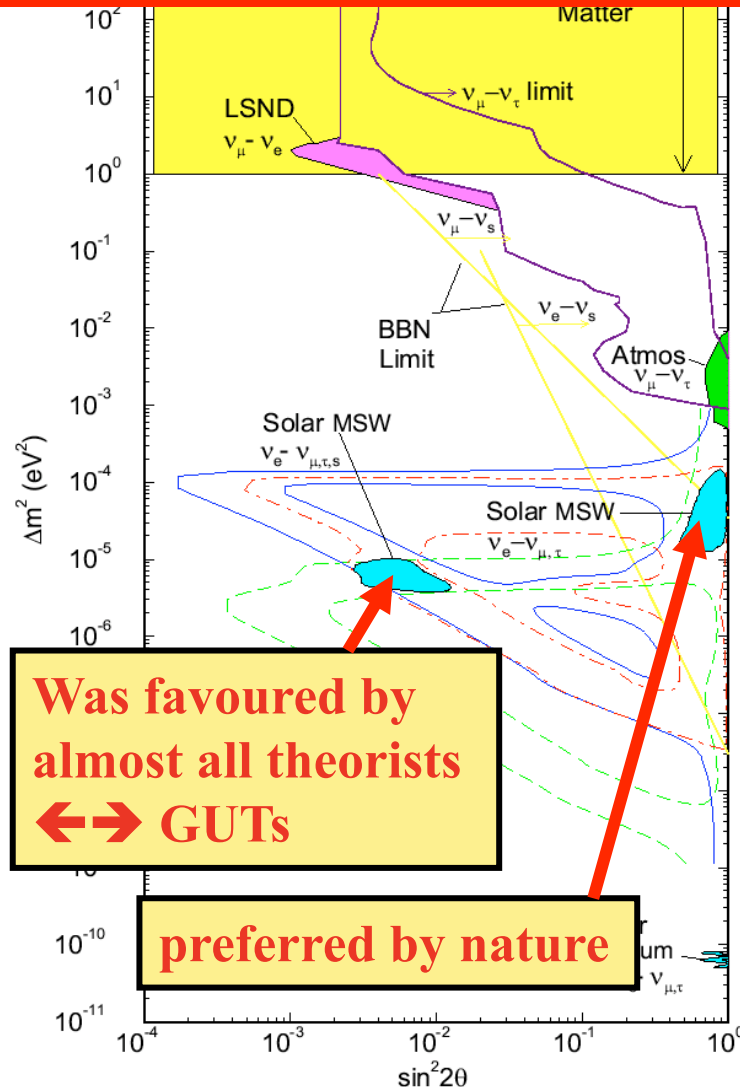
questions:

- Dirac / Majorana
- mass scale: m_1
- mass ordering: $\text{sgn}(\Delta m^2_{31})$
- how small is θ_{13} , θ_{23} maximal?
- leptonic CP violation
- 3 flavour unitarity?
- why 3 generations, $d=4$, gauge group, ...



Learning about Flavour

History: Elimination of SMA



Was favoured by almost all theorists
 ↔ GUTs

preferred by nature

Next: Smallness of θ_{13} , θ_{23} maximal

- models for masses & mixings
- input: known masses & mixings
 - distribution of θ_{13} predictions
 - θ_{13} expected close to ex. bound
 - well motivated experiments

what if θ_{13} is very tiny?
 or if θ_{23} is very close to maximal?

- numerical coincidence unlikely
- special reasons (symmetry, ...)
- answered by coming precision

The larger Picture: GUTs

Gauge unification suggests that some GUT exists

Requirements:

gauge unification

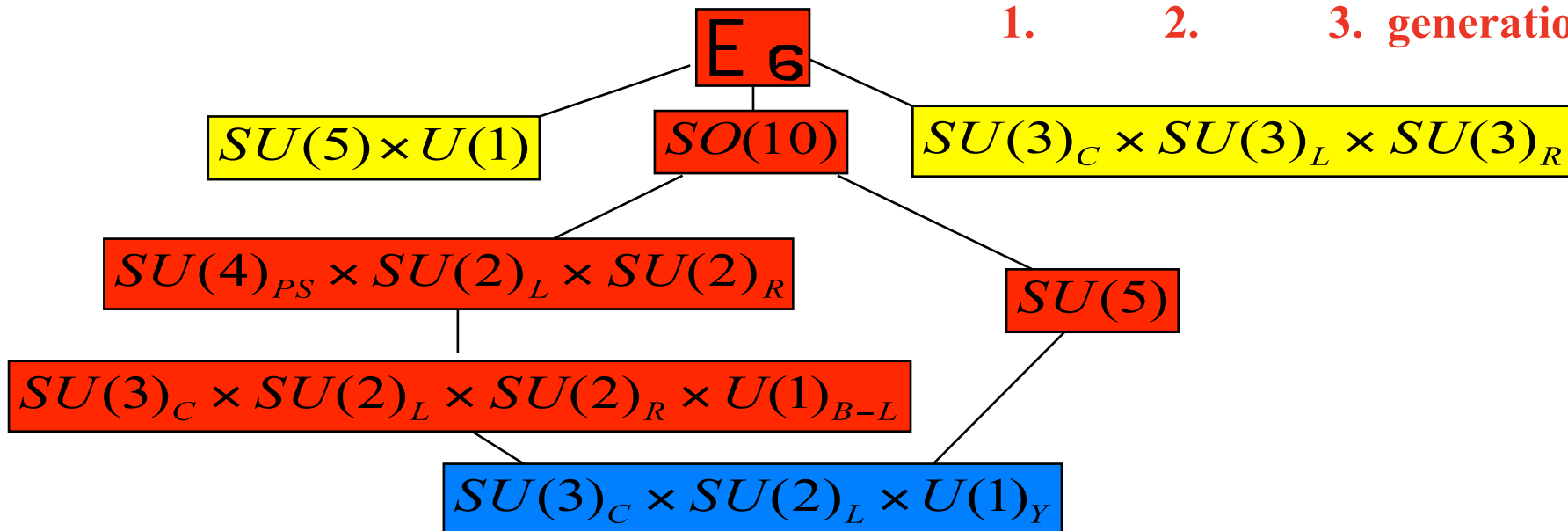
particle multiplets $\leftrightarrow \nu_R$

proton decay

...

Quarks	$2/3$	$2/3$	$2/3$
	u ~5	c ~1350	t 175000
	$-1/3$	$-1/3$	$-1/3$
	d ~9	s ~175	b ~4500
Leptons	$0?$	$0?$	$0?$
	ν_1	ν_2	ν_3
	0.511	105.66	1777.2
	e	μ	τ

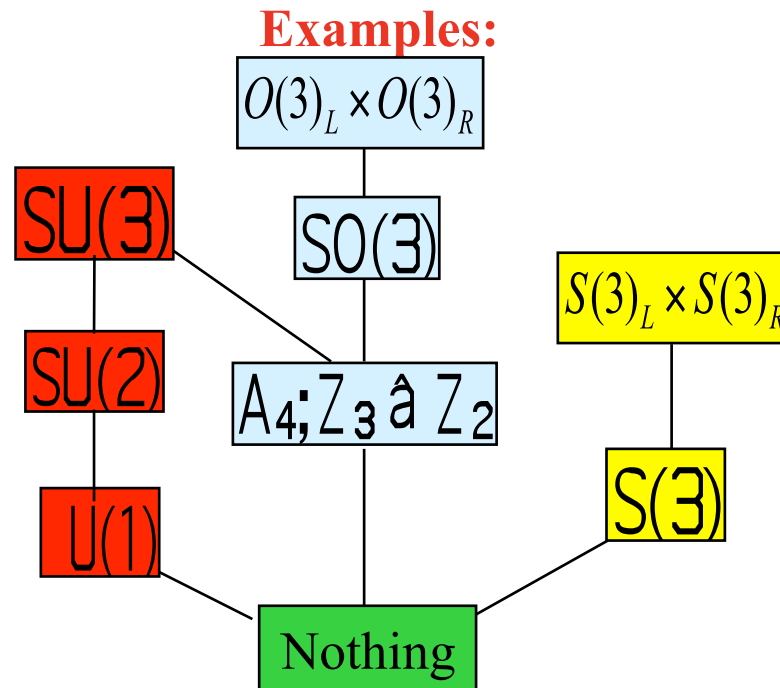
1. 2. 3. generation



Flavour Unification

- so far **no understanding of flavour, 3 generations**
- apparant regularities in quark and lepton parameters
- ➔ flavour symmetries (finite number for limited rank)
- ➔ **symmetry** not texture zeros

Quarks	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$
	u	c	t
	~ 5	~ 1350	175000
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$
	d	s	b
	~ 9	~ 175	~ 4500
Leptons	$0?$	$0?$	$0?$
	ν_1	ν_2	ν_3
	e	μ	τ
	0.511	105.66	1777.2
	1.	2.	3.
	generation		



See talks by:
M.-C. Chen
A. Romanino
W. Rodejohann
L. Merlo

GUT x Flavour Unification

$SO(10)$	Quarks	$\frac{2}{3}$ u ~5	$\frac{2}{3}$ c ~1350	$\frac{2}{3}$ t 175000
		$-\frac{1}{3}$ d ~9	$-\frac{1}{3}$ s ~175	$-\frac{1}{3}$ b ~4500
	Leptons	$0?$ ν_1	$0?$ ν_2	$0?$ ν_3
		0.511 e	105.66 μ	1777.2 τ
		1.	2.	3.
	$SO(3)_F$	generation		

→ GUT group X flavour group

example: $SO(10) \times SU(3)_F$

- SSB of $SU(3)_F$ between Λ_{GUT} and Λ_{Planck}

- all flavour Goldstone Bosons eaten

- discrete sub-groups survive \leftrightarrow SSB

e.g. Z_2, S_3, D_5, A_4

→ structures in flavour space

→ compare with data

GUT x flavour is rather restricted

\leftrightarrow small quark mixings *AND* large leptonic mixings ; quantum numbers

→ so far only a few viable models (without supersymmetry)

rather limited number of possibilities; phenomenological success non-trivial

→ aim: distinguish models further by future precision

Further Implications of Precision

Precision allows to identify / exclude:

- special angles: $\theta_{13} = 0^\circ$, $\theta_{23} = 45^\circ$, ... \leftrightarrow discrete f. symmetries?
- special relations: $\theta_{12} + \theta_C = 45^\circ$? \leftrightarrow quark-lepton relation?
- quantum corrections \leftrightarrow renormalization group evolution

Provides also measurements / tests of:

- MSW effect & coherent scattering
- cross sections (G. Co)
- 3 neutrino unitarity & sterile neutrinos (D. Meloni)
- neutrino decay (admixture...)
- Geophysics via oscillograms (E. Akhmedov)
- electromagnetic properties (A. Studenikin)
- NSI (T. Ota, T. Rashba, M. Cirelli)
- MaVaN scenarios (M. Lattanzi), unparticles (R. Zukanovich)
- \rightarrow synergies with LHC and LFV (A. de Gouvea, F. Joaquiun)

Guessing the Future

Status quo: neutrino revolution → consolidation

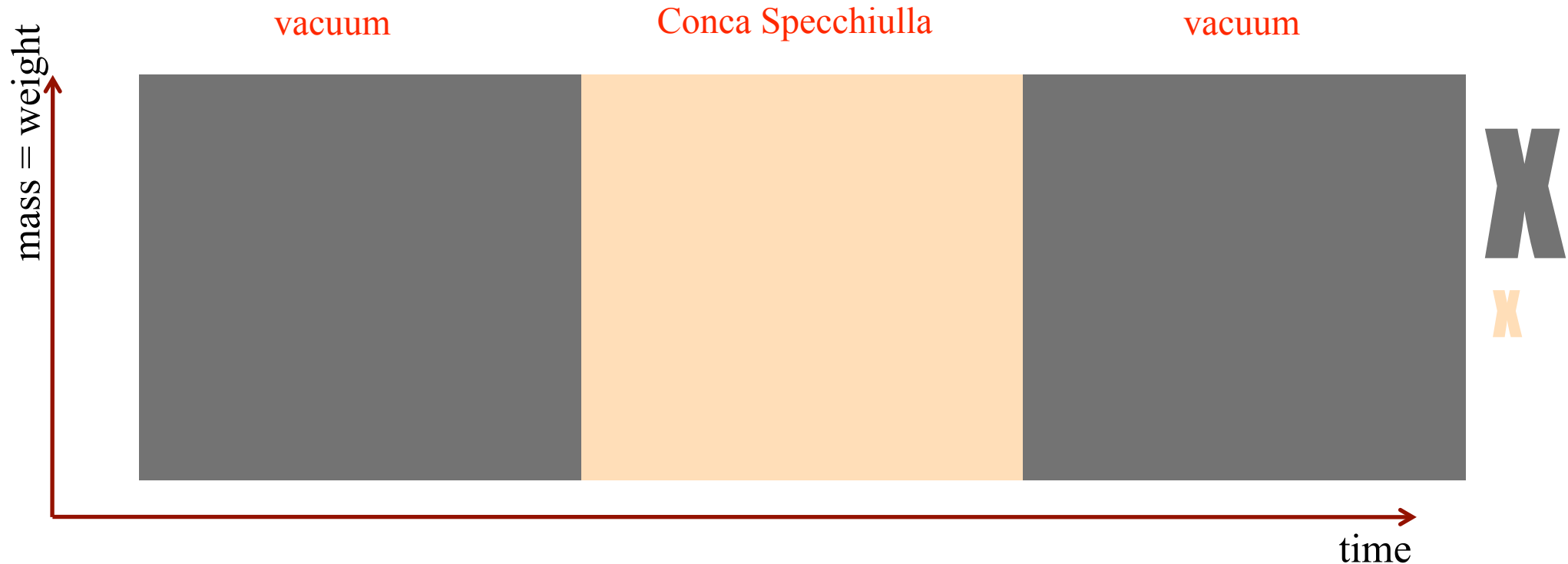
- check and improve knowledge ↔ potential for surprises
- **slowdown** due to larger / more complex experiments
 - dinosaurs & mutations
 - **scale known technologies (with required R&D)**
 - look for **new ideas**
 - **GSI anomaly** → not due to ν -mixing (C. Giunti) (if real...)
 - **Moessbauer neutrinos** (S. Parke) (very difficult but...)
 - **relic neutrino** detection (A. Cocco)
 - ν -masses from **cosmological 21cm observation** (J. Pritchard)
- additional promising areas: **LHC, LFV, astroparticle/cosmology**
 - interesting by themselves & important theoretical interplay

Other Topics (apologies)

- **Neutrino telescopes** (T. Montaruli, U. Katz, J.L. Bazo, E. Presani)
- **Geoneutrinos** (S. Dye)
- **CMB in the light of Planck** (P. de Bernardis)
- **Dark Matter** (N. Fornengo, F. Cafagna)
- **HECR and UHECR** (V. Berezhinsky, L.Nellen, F. Villante, N. Busca, J. Matthews, E. Carmona, P. Serpico, R. Tomas, K.H. Kampert, D. Allard, N. Giglietto)
- **Axions** (A. Mirizzi)

The N.O.W. Effect

(invented by G. Fogli; assumes only an average physicist “X”)



- effect was seen in all runs (2000, 2004, 2006, evidence in 2008)
- high statistics → 6.3 sigma effect
- safe (no black holes or other damages) → can (should be) be repeated

Thanks to

- Gianluigi Fogli**
- Eligio Lisi and all other co-organizers**
- session convenors**
- E. Forini for his nice special talk**
- hotel & staff**

We look forward to NOW 2010!

ELIGIO LISI