

SPL-Fréjus

- Some performances
- Fréjus site
- French Photodetector R&D

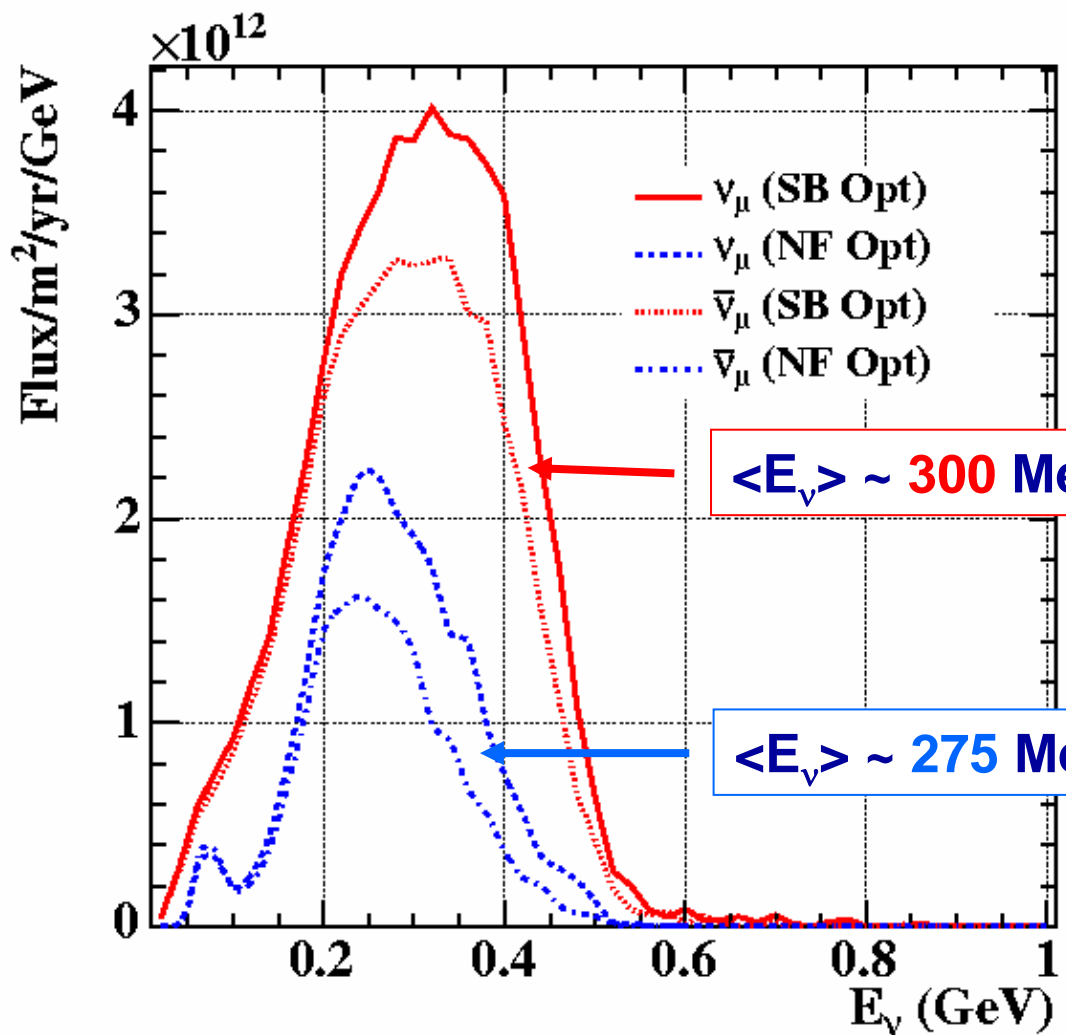
Thanks: A. Cazes, M. Mezzetto, L. Mosca, Th. Schwetz
and IPNO & LAL engeneers.

Some ingredients for physics analysis

- 440kT Water Č located 130km from CERN (see site later)
- Essentially SK analysis with tighter cuts for e/μ id (cf. hep-ph/0105297)
- Use energy resolution dominated by Fermi motion* (200MeV bins)
- 2% systematics on signal & bkgd
- Optimized machine versions: βB (M. Mezzetto) and SB (A.C + J.E.C)
- Use Atmospheric Neutrinos
- GLOBES & NUANCE

*: migration matrix for βB

Fluxes comparison @ 130km



$\sim 95 \nu_{\mu}^{CC}/kT/yr^*$



$\langle E_{\nu} \rangle \sim 300 \text{ MeV}, 1.2 \cdot 10^{12}/m^2/yr$

3.5GeV SPL optimum

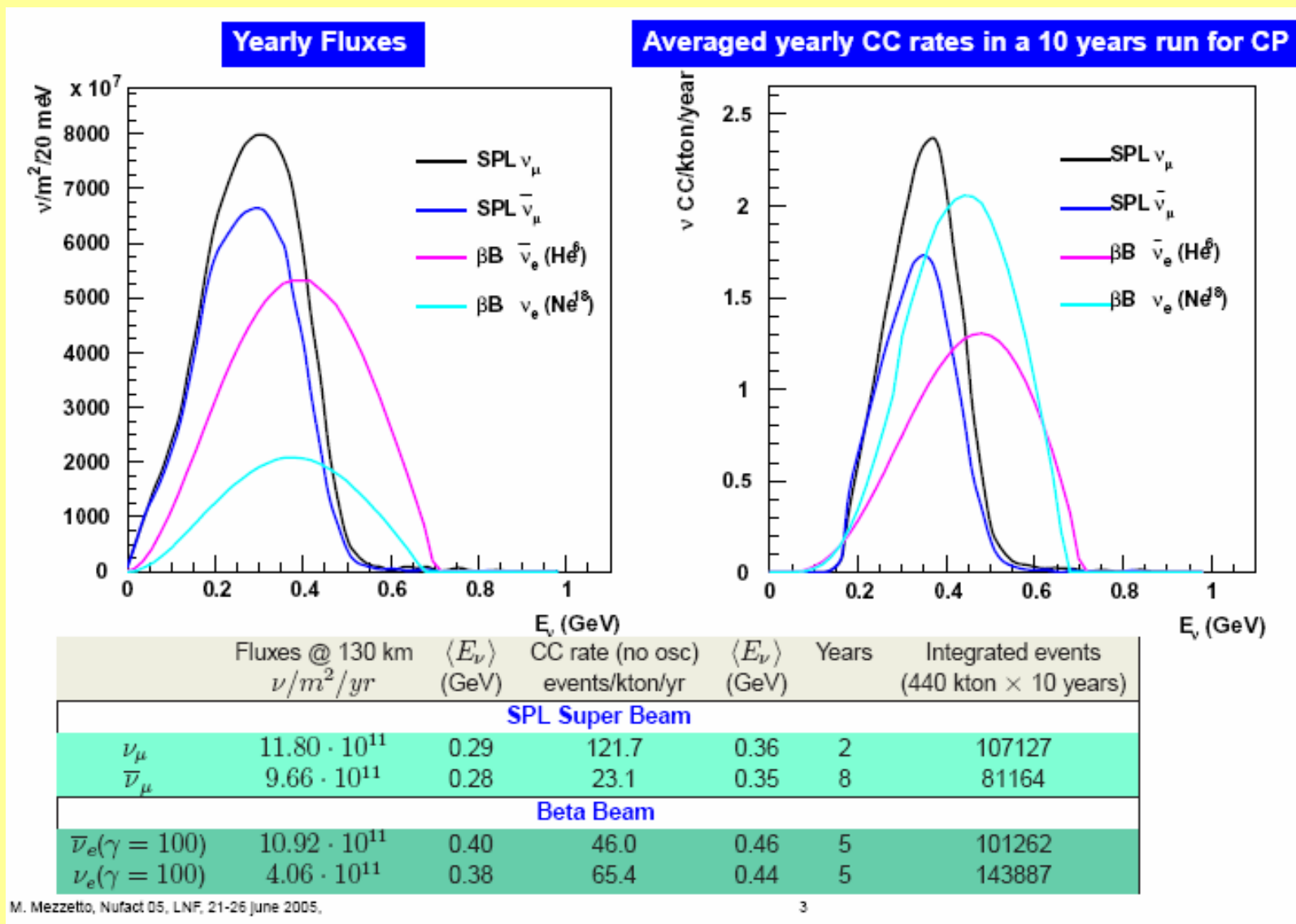
$\langle E_{\nu} \rangle \sim 275 \text{ MeV}, 4.5 \cdot 10^{11}/m^2/yr$

Old ν Fact optimum

Reflector: 50% of the Flux

*: Lipari x-sect. (see later)

β B and SB fluxes



Analysis: GLoBES + M. Mezzetto's parameterization file

440kT x 5yrs: **2,2 Mt.yrs (+)**

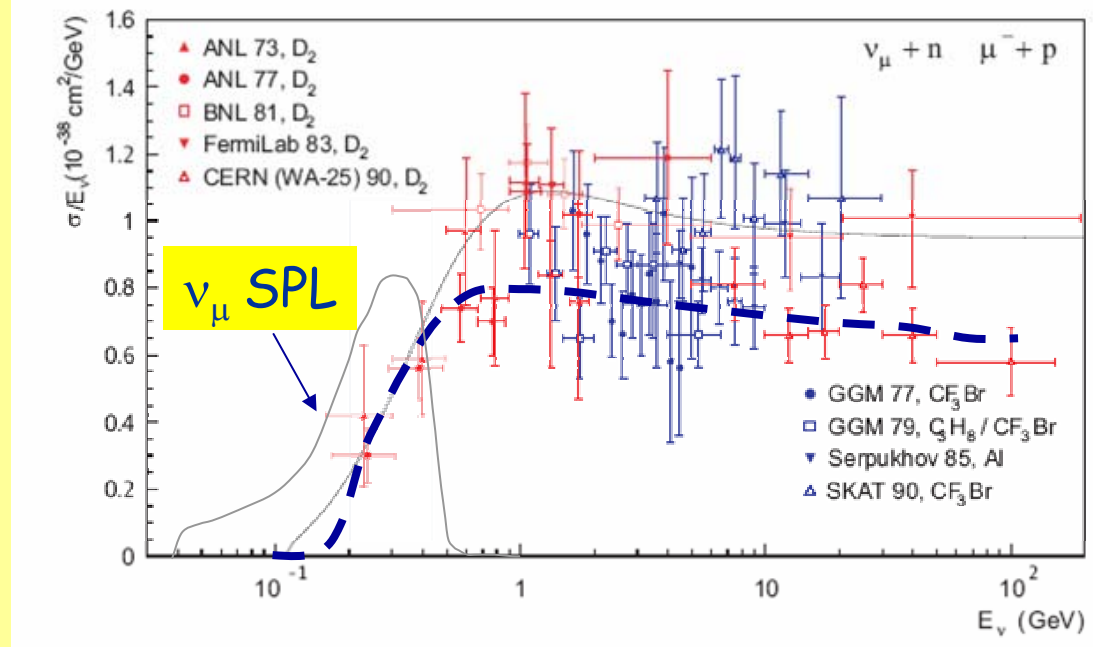
	$\theta_{13} = 1^\circ$	$\theta_{13} = 3^\circ$	$\sin^2 2\theta_{13} = 0.05$	
$\nu_\mu \rightarrow \nu_e$ (Sig)	33 ($\delta = \pi/2$)	330 ($\delta = \pi/2$)	2200 ($\delta = \pi/2$)	3670 ($\delta = 0^\circ$)
$\nu_\mu \rightarrow \nu_e$ (Bkg)	1500			
	$\nu_e \rightarrow \nu_e$ CC	π^0 from NC	$\nu_\mu \rightarrow \nu_\mu$ CC (μ missId)	$\bar{\nu}_e \rightarrow \bar{\nu}_e$ CC
Frac. of Bkg	90%	6%	3%	1%
Reduction Factor	0.707(1060)	$6.5 \cdot 10^{-4}$ (90)	$5.4 \cdot 10^{-4}$ (45)	0.677(15)
$\nu_\mu \rightarrow \nu_\mu$ (Sig)	64950 ($\delta = \pi/2$)		64414 ($\delta = 0^\circ$)	
$\nu_\mu \rightarrow \nu_\mu$ (Bkg)	3 (4.310^{-5} $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$ CC)			

$$\sin^2 2\theta_{12} = 0.82, \theta_{23} = \pi/4, \Delta m^2_{21} = 8.1 \cdot 10^{-5} eV^2, \Delta m^2_{31} = 2.2 \cdot 10^{-3} eV^2$$

Reduction factor and efficiencies taken from SK simulation (D. Casper) and a tight cut for e/ μ misId. (cf. hep-ph/0105297)

The X-sections

V.V. Lyubushkin et al., internal NOMAD memo



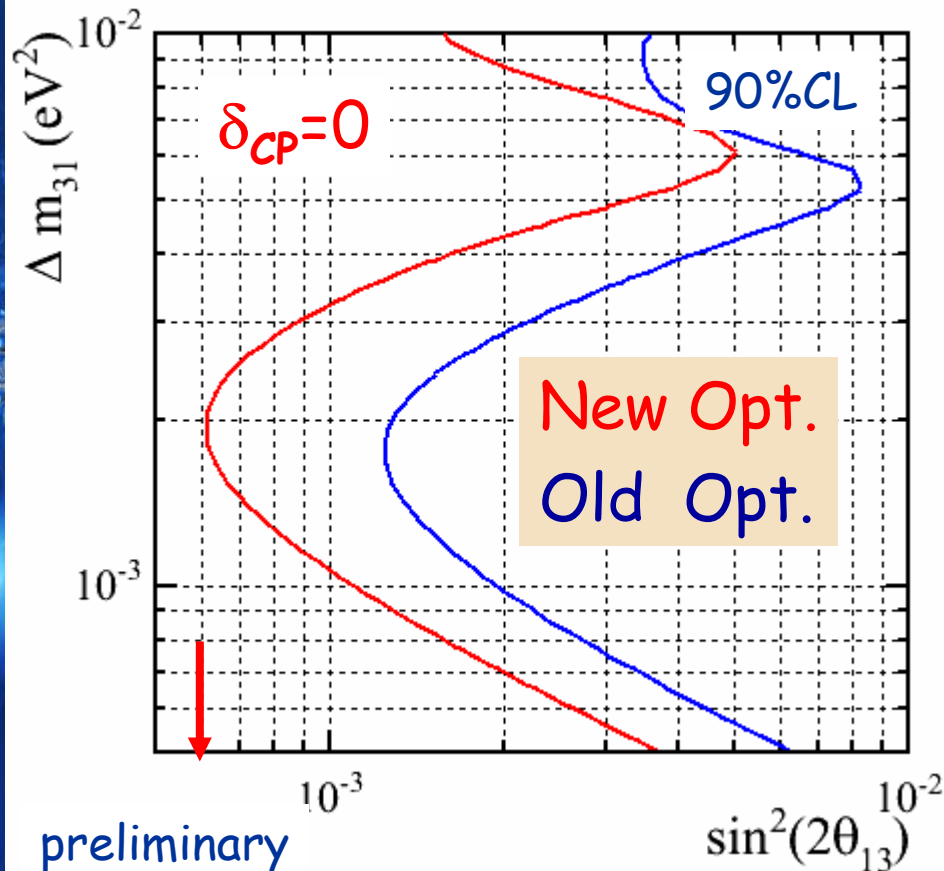
---: Lipari et al.
PRL74(95)4384
on H₂O

βB is an ideal tool to measure these cross-sections and a 2% systematic error on both signal and background are used.

Require close position

Some physics performances

440kT water Č, 4MW SPL, GLoBES



5yrs (+)

True values: $(\Delta m^2_3, \sin^2 2\theta_{13})$
 $\sin^2 2\theta_{12}=0.82, \theta_{23}=\pi/4, \Delta m^2_{21}=8.1 \cdot 10^{-5} eV^2$
5% external precision on θ_{12} and Δm^2_{21} and
use SPL disappearance channel and
spectrum analysis*

2% syst. on signal & bkg

$$\sin^2 2\theta_{13}(90\%CL) = 6 \cdot 10^{-3} (0.7^\circ)$$

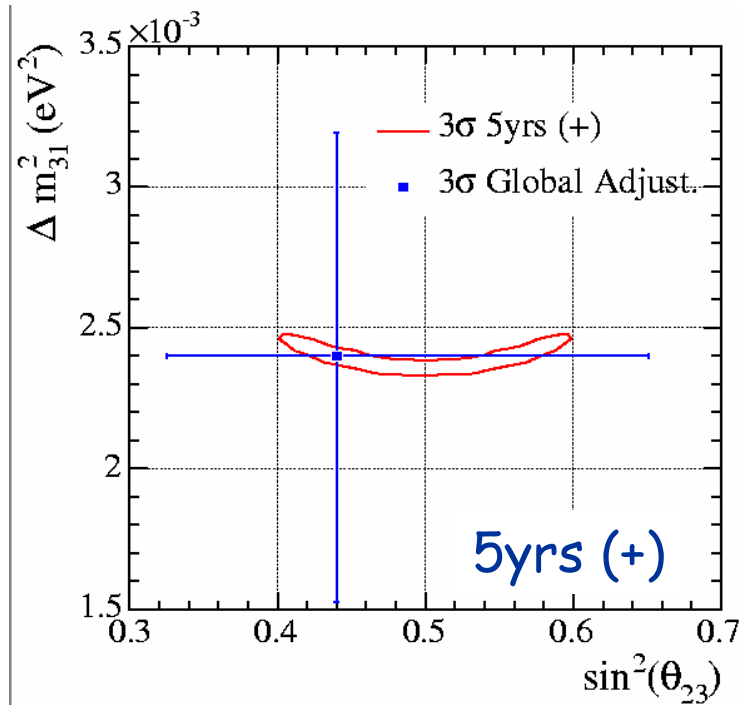
sizeable improvement

*: 5 bins [0.08,1.08] GeV
ISS CERN 05 J.E Campagne LAL

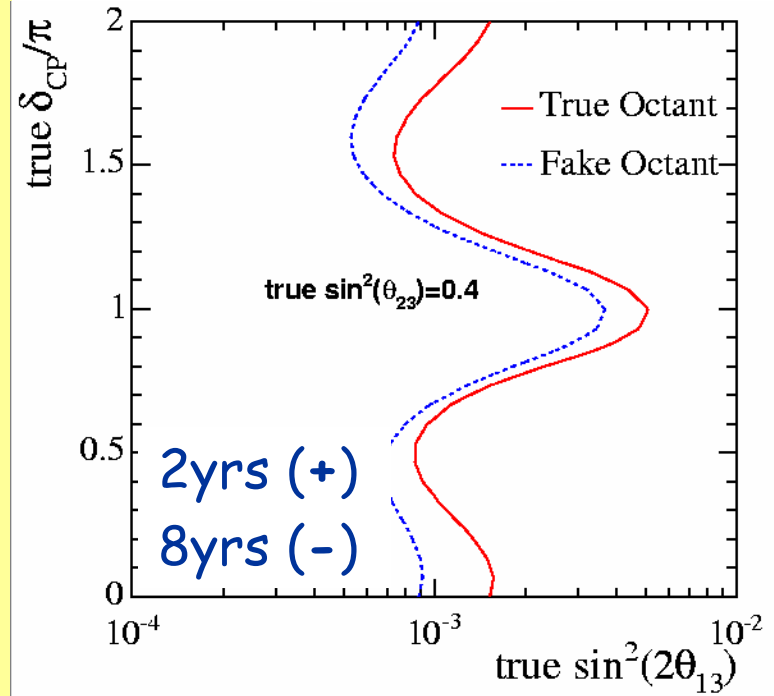
$(\chi^2(2dof)=4.6 \text{ or } 11.83)$

Other expected performances

Δm_{31}^2 & $\sin^2\theta_{23}$ measurements



$3\sigma \theta_{13} \neq 0$



Comparison with other facilities

Everything computed with the identical program.
Thanks to the GLoBES experiment library.

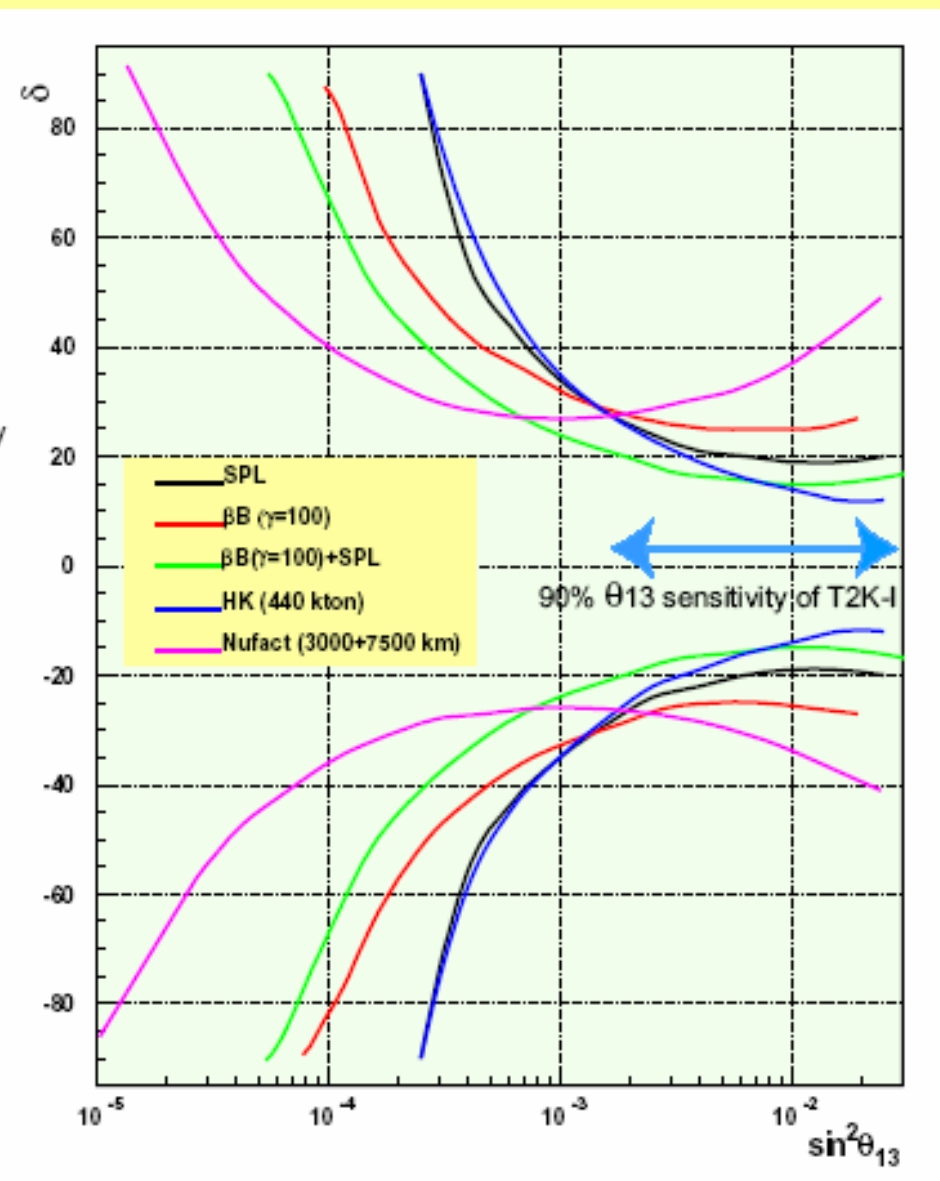
HK taken from Huber, Lindner and Winter, hep-ph/0204352, with a fiducial of 440 kton (it was 1 Mton), 2% systematics on QE signal and backgrounds (it was 5%) and 2+8 years running (it was 2+6).

NUFACT taken from Huber, Lindner and Winter, hep-ph/0204352, changing the systematics from 0.1% to 2% and the running time to 5+5 years (it was 4+4).

Other parameters: two iron magnetized detectors, 50 kton, at 3000 and 7500 km, 50 GeV muons, $1E21$ useful decays/year, 5% systematics on matter profile, threshold at 4 GeV, 20 bins from 4 to 50 GeV.

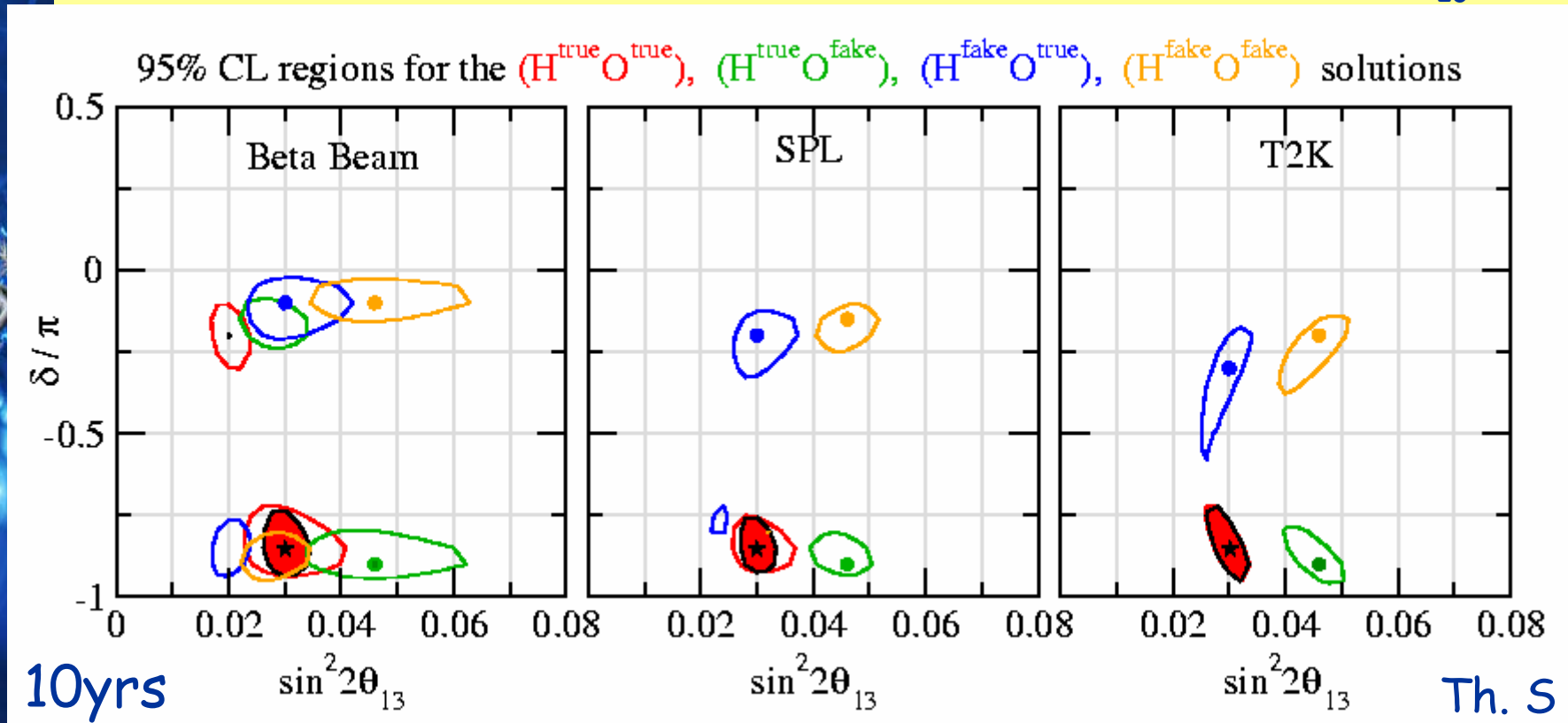
SPL 3.5 GeV (see J.E. Campagne talk) with $2 \nu + 8 \bar{\nu}$ years, 2% systematic error, 200 MeV binning, 440 kton fiducial.

M.M@NuFact05



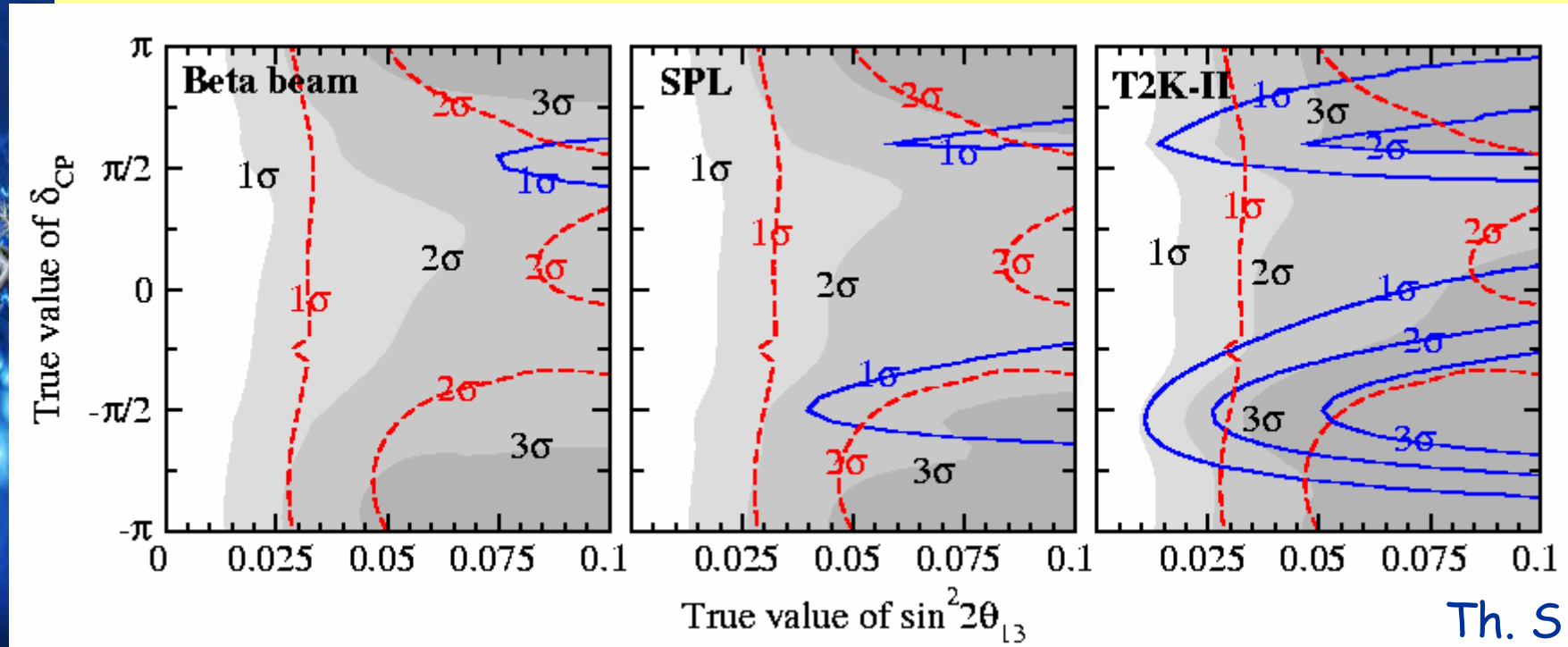
Remove ambiguities with ATM ν

Favorable case $\sin^2\theta_{23}=0.6$



- Contour after ATM combination
- *: true value

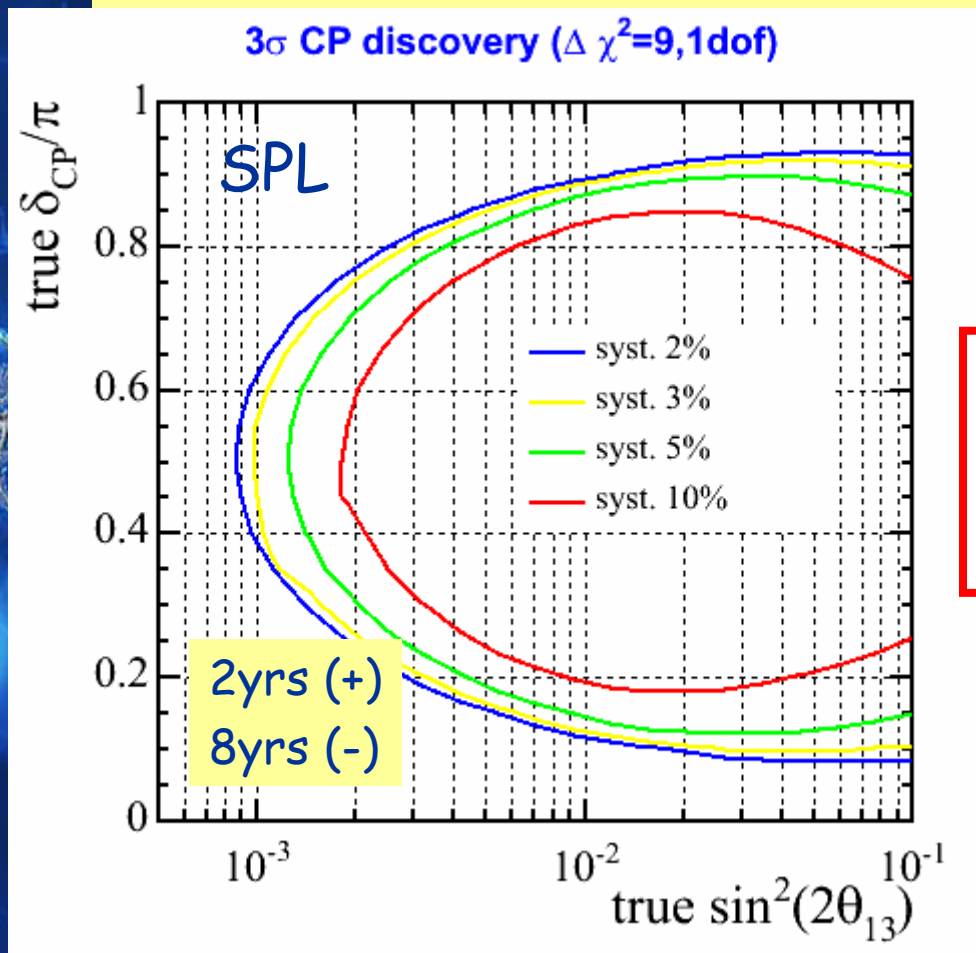
Mass hierarchy sensitivity



- - - Atm.
 ——— LBL
 ■ Atm+LBL

T2K-II takes benefit of more matter effect

Effect of the systematic (sig. & bkg)

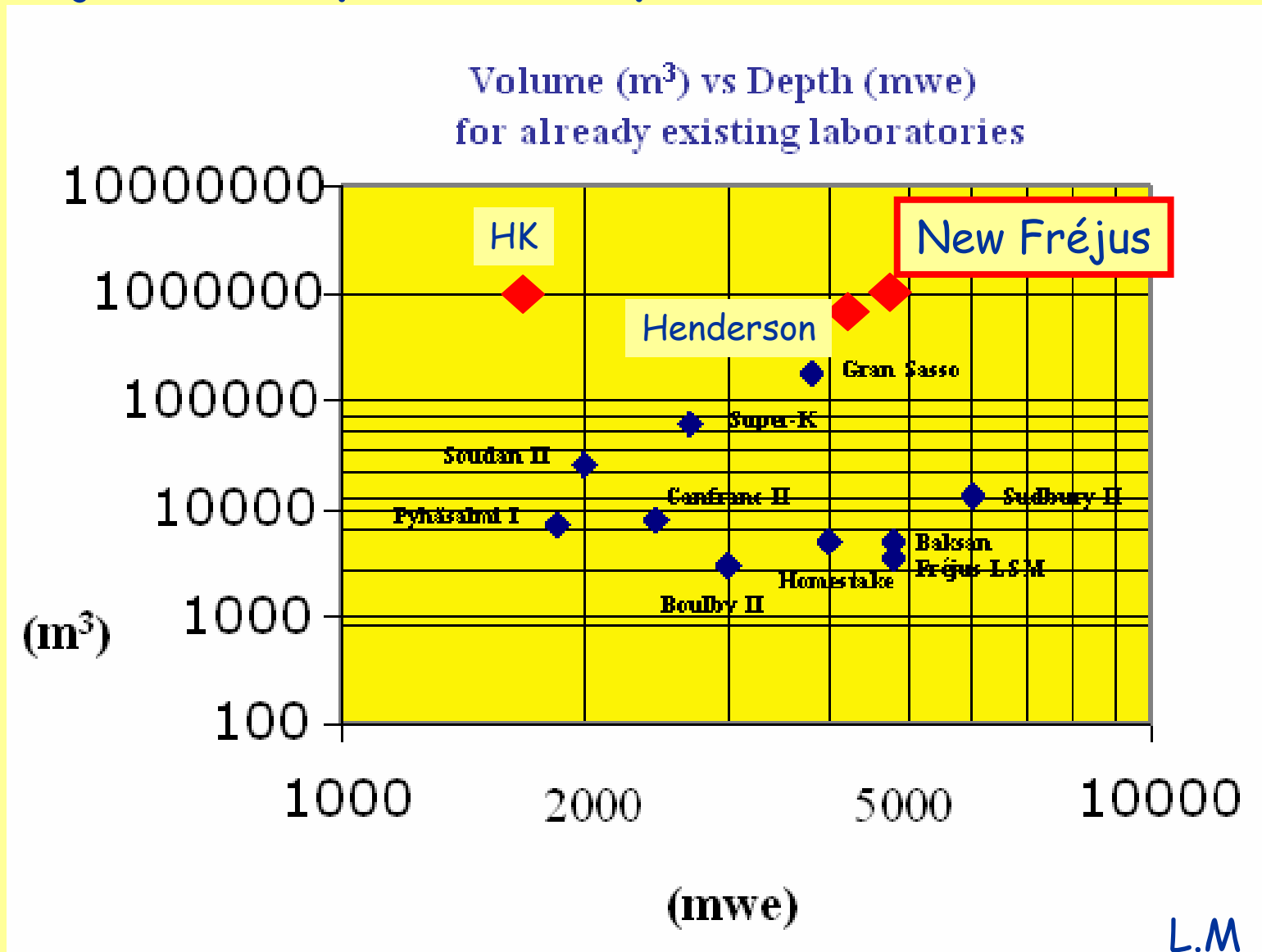


Much more dramatic than ambiguities

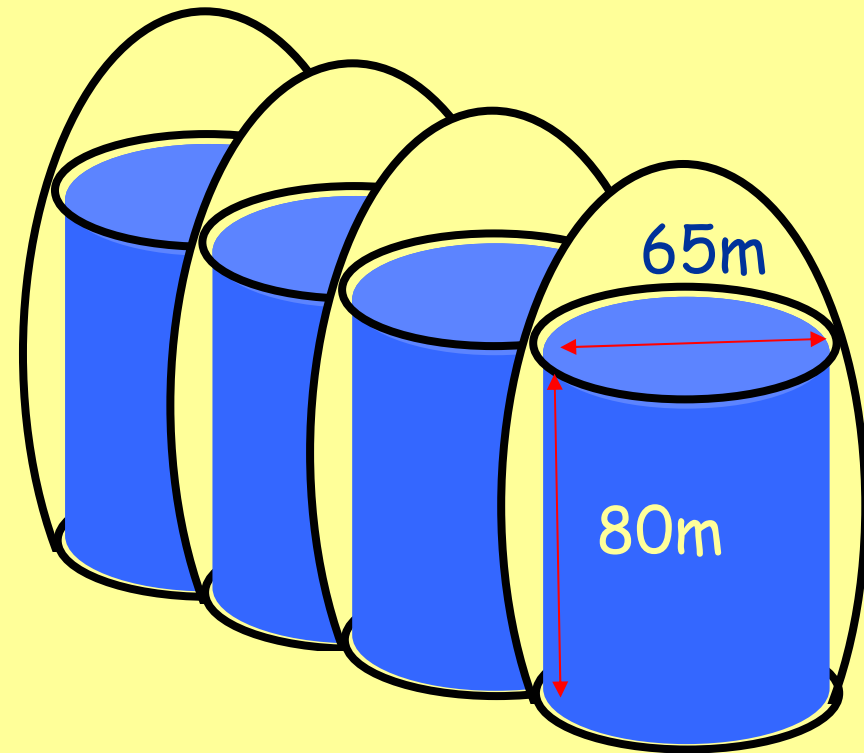
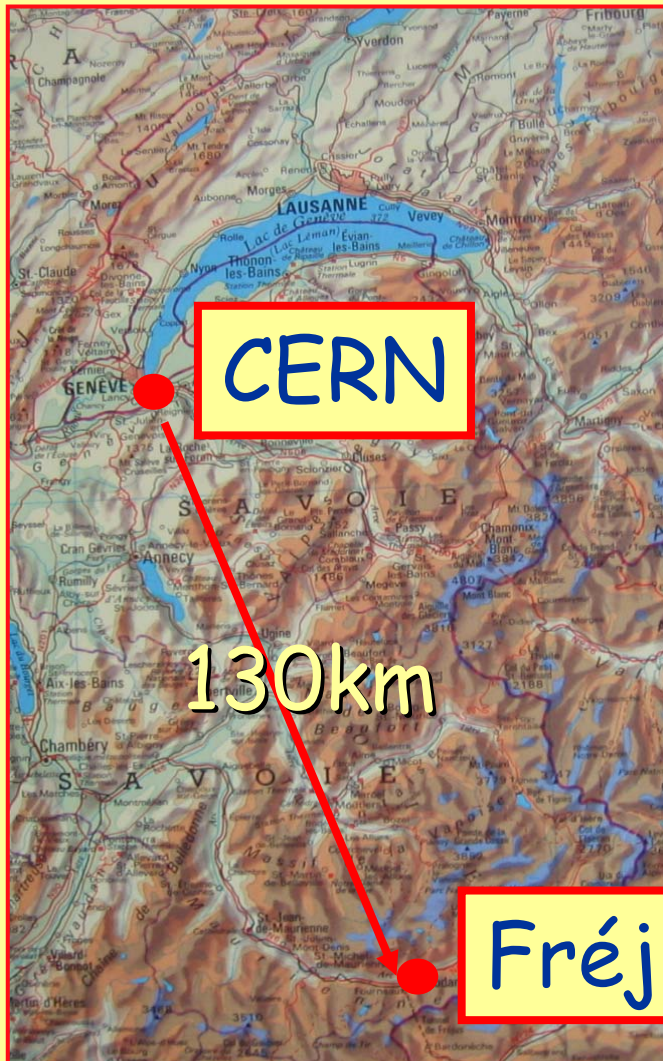
Old opti. is 3x worse

True values: ($\delta/\pi, \sin^2 2\theta_{13}$)
 $\sin^2 2\theta_{12}=0.82, \sin^2 \theta_{23}=0,4$
 $\Delta m^2_{21}=7.9 \cdot 10^{-5} \text{eV}^2, \Delta m^2_{31}=2.4 \cdot 10^{-3} \text{eV}^2$
 5% external precision on θ_{12} & Δm^2_{21}
 use SPL disappearance channel and spectrum analysis

Fréjus site possibility



New Fréjus Cavern (MEMPHYS)



$4 \times 250,000 \text{m}^3 \text{H}_2\text{O}$

Based on well experienced engineer studies. First cost estimate soon

*: Modane

Photodetector R&D in France

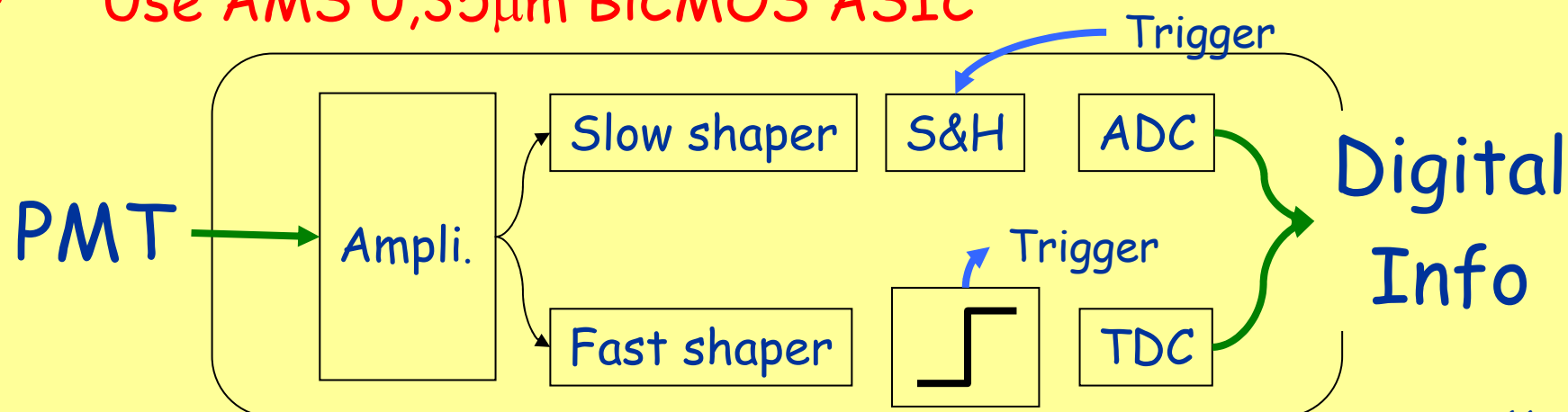
- R&D launched after NNN05 but based on on-going R&D with Photonis
- **IPN-Orsay, LAL & Photonis** together in an official GIS to develop **Smart-Photodetectors** (*ie electronic up to ADC/TDC included*): 6 engineers + 2 post-docs + Photonis engineers
- 200k€/3yrs has been asked at the new National Research Agency (ANR)

Photonis @ NNN05: 500,000 PMT -12"- 800€/u
Target electronics + HV: 200€/channel

Electronics

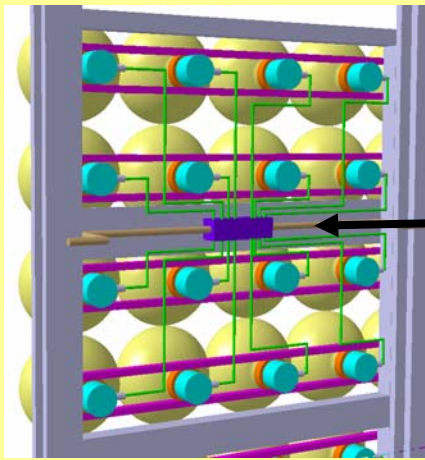
Taken in charge by **LAL**: from amplifier up to ADC/TDC based on past experience with similar state of the art front-end electronics developed for OPERA, W-Si ILC prototype, LHCb...

- Trigger @ $\frac{1}{4}$ p.e (3kHz from SK)
- TDC: 12bits 0,4ns/c
- ADC: 12bits 0,15pC/c with 1 p.e @ 20-30 adc channels.
- High speed digital readout
- **Cost reduction** thanks to high level of integration
- **Use AMS 0,35 μ m BiCMOS ASIC**



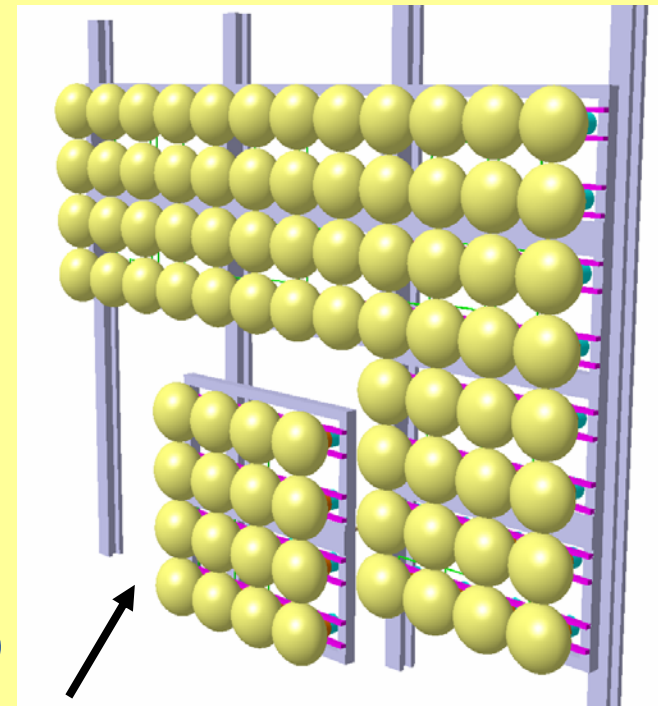
Mechanics

Taken in charge by **IPNO**: well experienced in photodetectors (last operation: Auger). With **PHOTONIS** tests of **PMT** 8", 9" → **12"** and Hybrid-PMT and HPD



Electronic box
water tight

Basic unit that we want to
build and test under water



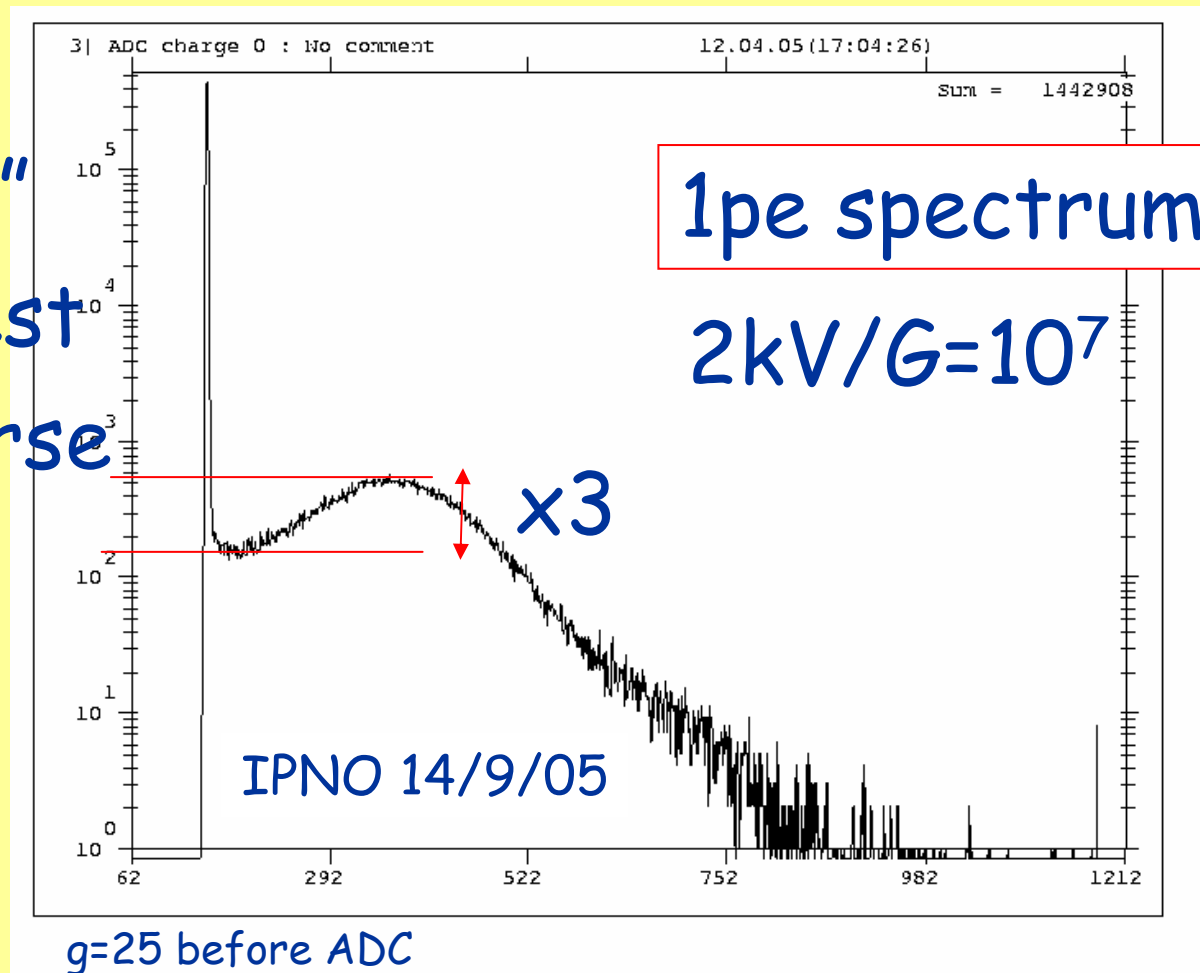
IPNO

Some PMT characteristics measurements

XP1806 8"

Not the best

Not the worse



No diff. 5",8",10" so 12" should be identical