

# NNN08 - Paris

## International Workshop on Next generation Nucleon decay and Neutrino detectors - 2008

September 11-13, 2008

Laboratoire APC  
AstroParticule et Cosmologie  
Paris, France

<http://nnn08.in2p3.fr>  
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NNN08

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a summary  
(focused on  
Europe)

Alessandra Tonazzo  
APC -  
Université Paris 7

on behalf of the LOC

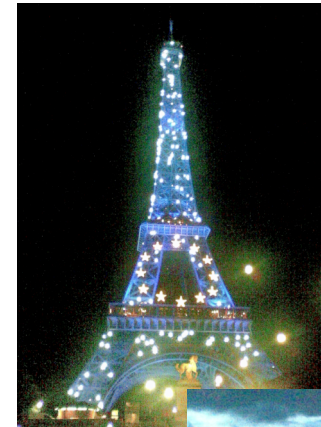
# NNN08 - Paris

9th of a series of International Workshopn on Next generation  
Nucleon decay and Neutrino detectors

**NNN05 (SUNY), NNN00 (UCI/FNAL), NNN01 (LSU), NNN02 (CERN),  
NNN05 (Aussois), NNN06 (Seattle), NNN07 (Hamamatsu)**



~100 participants from 12 countries  
Research institutes + Private firms



Seine river  
cruise -  
Conference  
dinner



# Scientific Program

- Thursday, 11 September 2008

  - 09:10-10:30 **Proton Decay**

  - 11:00-13:10 **Supernova Neutrinos**

  - 14:30-19:10 **Current & Near Future Neutrino Experiments**

- Friday 12 September 2008

  - 09:00-12:45 **Towards next generation projects**

  - 14:00-18:25 **R&D towards Large Scale Detector**

- Saturday 13 September 2008

  - 09:00-10:00 **"Interdisciplinary" topics**

  - 10:00-11:00 **Engineering**

  - Discussion**

    - H.Murayama *"Neutrino and proton decay what can we learn FOR FUNDAMENTAL PHYSICS AND COSMOLOGY"*
    - Totsuka memorial talk - H.Kajita

Apologies to the speakers I will not mention explicitly....

# Recent results and near future

<b>Borexino</b>	>1yr data: $N_{\text{nu}}(^7\text{Be})=49\pm 3_{\text{stat}} \pm 4_{\text{sys}}/100\text{t}$ ; $\mu_{\nu} < 5.4 \times 10^{-11} \mu_{\text{B}}$
<b>MINOS</b>	CC $3.36 \times 10^{20}$ POT: $\Delta m^2 = (2.43 \pm 0.13) \times 10^{-3} \text{eV}^2$ , $\sin^2(2\theta) > 0.9$ ; $\nu_e$ disappearance $\theta_{13}$ soon
<b>SuperK</b>	after 2yrs of SK-III, SK-I+II+III $> 25000 \nu_{\text{atm}}$ . $\theta_{23} = 45 \pm 5^\circ$ , $f_s < 23\%$
<b>KamLAND</b>	solar $\Delta m^2 = 7.59 \pm 0.21 \cdot 10^5 (\text{eV}^2)$ , geo- $\nu$ $73 \pm 27$ . New physics opportunities after purification

F.vonFeilitzsch

A.Sousa

M.Fechner

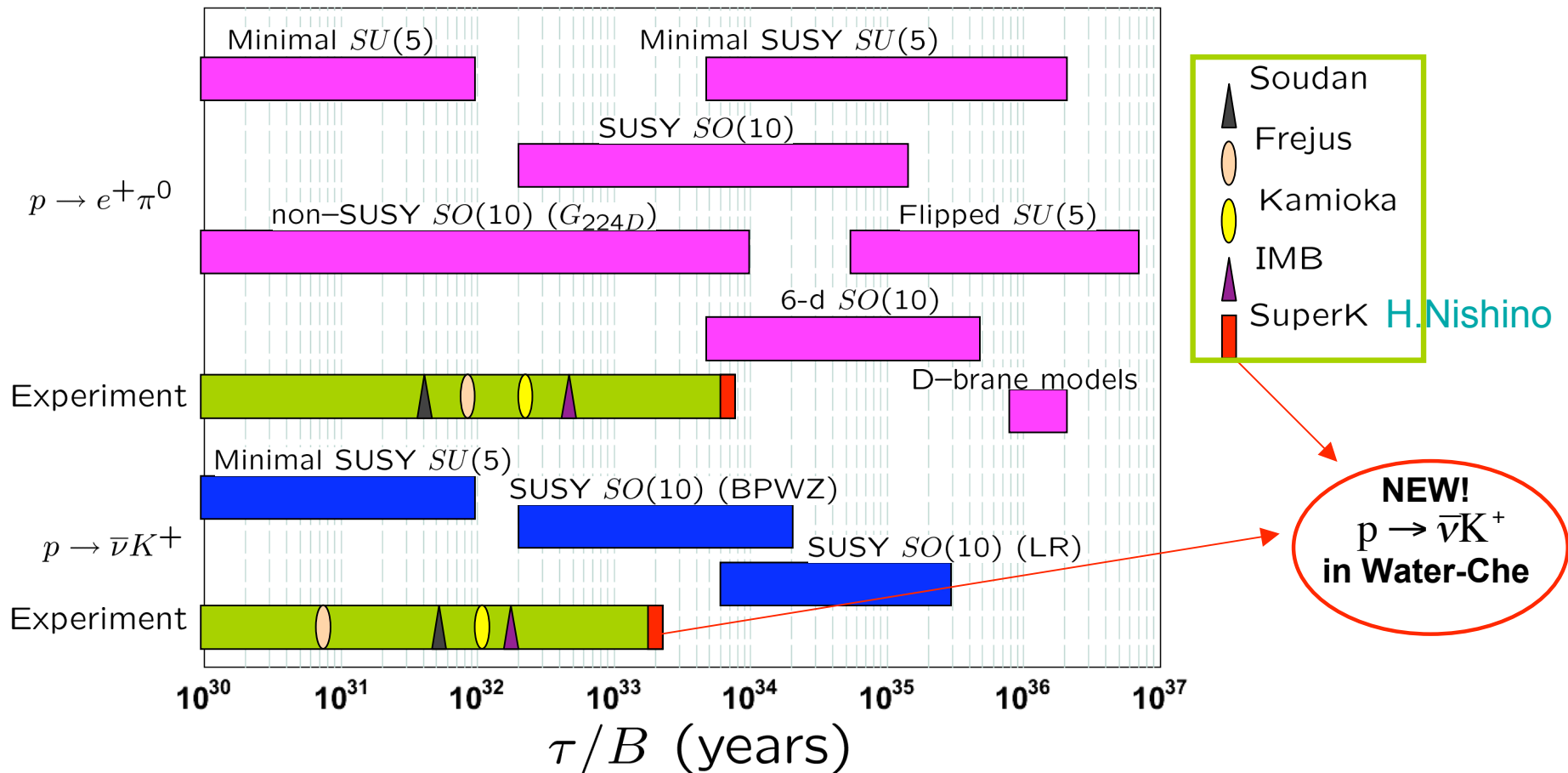
A.Kozlov

K.Morishima	<b>OPERA</b>	Taking data, 123d, $6 \times 10^{18}$ POT
S.Centro	<b>ICARUS T600</b>	Installed underground at LNGS, filling soon
T.Lachenmeier	<b>Double Chooz</b>	Building detector, data in 2009 with Far Det., in 2010 with both detectors, $\sin^2 2\theta_{13} < 0.03$ after 3y
K.Lau	<b>Daya Bay</b>	Civil construction ongoing, detector building start in 2009, data taking with full setup in Dec.2010. Goal $\sin^2 2\theta_{13} < 0.01$
L.Kormos	<b>T2K</b>	Apr.2009 first nu's, then install ND and start full run. Results in 2010. Address $\sin^2 2\theta_{13}$ , $\delta_{\text{CP}}$ , $\theta_{23}$
P.Shanahan	<b>NOVA</b>	Back on track! FD construction 2011-13

Crucial impact on future  
neutrino beam strategy !

# Physics at future underground detectors (1): Nucleon decay

K.Babu



Challenge for next generation detectors:  
 → Sensitivity  $10^{34}$ - $10^{36}$ y in  $p \rightarrow e^+\pi^0$  and  $p \rightarrow \bar{\nu}K^+$

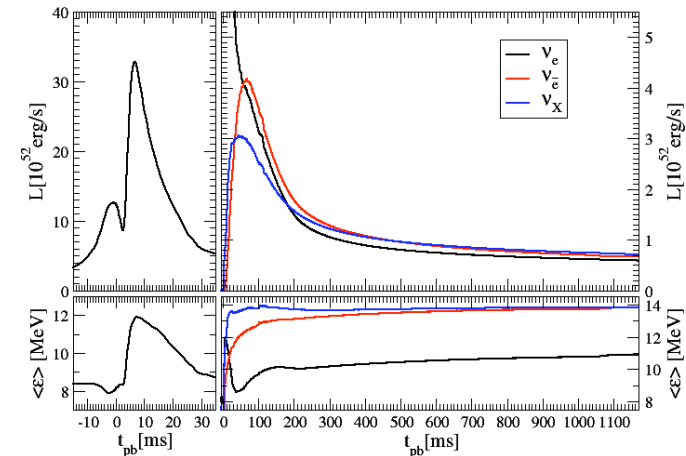
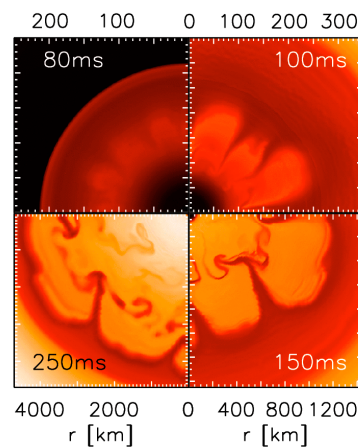
# Physics at future underground detectors (2) : SuperNovae

Th.Janka

## Core Collapse

- Neutrino signals from CC SN are unique probes of the physics inside the SN core and nascent SN
- Measuring SN neutrinos (flux, E and t) could help us understand explosion dynamics/mechanism and properties of SN matter

new 2-D simulations



Challenge for future detectors:  
→ Sensitivity even for distant SN  
→ Spectral measurement in Energy and time

# Physics at future underground detectors (2bis) : SuperNovae

C.Lunardini

## Diffuse SN neutrinos

- The feeble signal of all past SN probes deep in stars' interior
- Alternative to a galactic SN. Continuous, no waiting time
- Might become standard "everyday" physics in the future

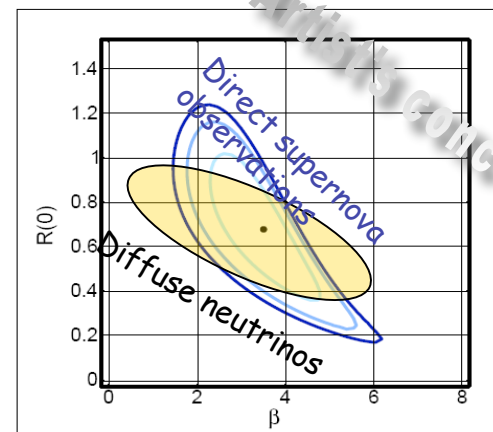
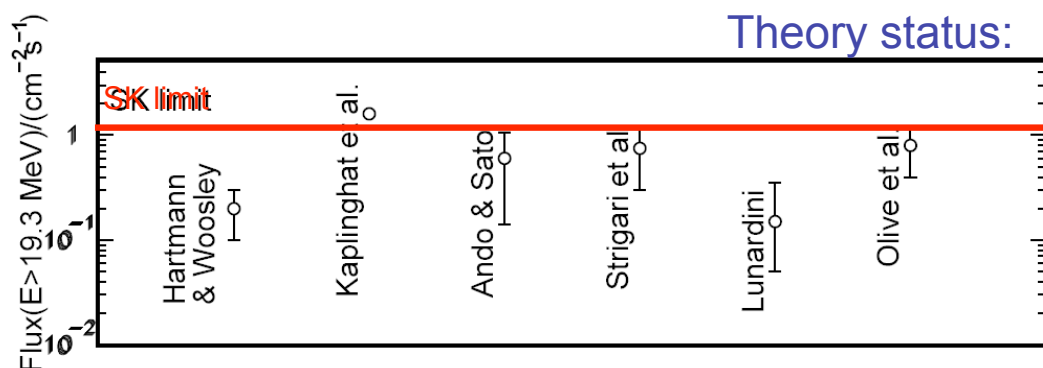


Fig. 5. Best fit point and isocontours of  $\chi^2$  in the space of the parameters describing the SNR function,  $R_{SN}(z)$ . These are the intercept,  $R_{SN}(0)$  (in units of  $10^{-4} \text{ yr}^{-1} \text{ Mpc}^{-3}$ ) and the power,  $\beta$ . The contours refer to 68.3, 90, 95.4% C.L..

Challenge for next generation detectors:

- increase statistics (= size)
- reduce backgrounds, also to gain spectral sensitivity (e.g.: Gd in  $\text{H}_2\text{O}$ )
- improve detection of other than anti- $\nu_e$ : e.g.  $\nu_e + {}^{40}\text{Ar} \rightarrow {}^{40}\text{K} + e^-$

# Physics at future underground detectors (3) : Neutrino Beams

- Goals: "ultimate" measurement of neutrino oscillation parameter, in particular the unknown ones:  $\theta_{13}$ ,  $\text{sign}(\Delta m_{13})$ ,  $\delta_{CP}$

3-family  $\nu$  oscillations in matter => "magic" baseline  $\frac{2\pi}{\sqrt{2}G_F n_e} \sim 7500$  km

- up to 8 possible degeneracies => need different baselines and/or different types of beam to resolve them

➔ SuperBeam: very high intensity "traditional" Beam  $\nu_\mu \rightarrow \nu_\mu$

➔ Beta Beam: pure  $\nu_e$  or anti- $\nu_e$  from accelerated  $\beta$ -decaying ions  $\nu_e \rightarrow \nu_\mu$

➔  $\nu$ -Factory: pure  $\nu_\mu + \bar{\nu}_e$  or  $\bar{\nu}_\mu + \nu_e$  from decay of accelerated  $\mu^\pm$ s

$\nu_e \rightarrow \nu_\mu$  (wrong sign)

Challenge for next generation detectors:

- ➔ large size
- ➔ lepton flavour identification
- ➔ (charge measurement)



# Future detectors

General requirements for next generation nucleon decay, neutrino astrophysics and beam neutrino detectors:

- Large mass (MegaTon...)
- Low Energy threshold
- Low background: shielding from cosmic rays (deep underground) or excellent cosmic bkg rejection capabilities
- Particle identification

Techniques:

- Water Cherenkov (+ Gd)
- Liquid Scintillator + Gd
- Liquid Argon
  - all need large nb of PMTs
  - all need large size underground excavations

# What / Where ? @NNN08

	Beams	Detectors			PMTs	Excavation
		Water Cherenkov	Liquid Scint.	Liquid Argon		
<b>Europe</b>	P.Soler	A.Tonazzo: MEMPHYS G.Martin-Ch.: PARISROC	T.Marrodan LENA C.Buck: R&D	A.Marchionni: GLACIER F.Pietropaolo: MODULAR	B.Combettes Photonis massive production	G.Nuujten: Phyasalmi R.Margineau: Slanic
<b>USA</b>	M.Diwan FNAL to Homestake	K.Lande: R&D M.Vagins: GADZOOKS		K.Soderberg: TPC R&D D.Cline: 100kt LAr R&D		K.Lande: DUSEL
<b>Japan</b>	T.Hasegawa JPARC	M.Shiozawa: R&D			H.Aihara: HPD	

(more projects ongoing but not presented at this conference)

# Europe's plans

Following the request by the UDiG Workshop organizers, I will concentrate on Europe

- Beam plans
- R&D for Liquid Scintillator [LENA]
- R&D for Water Cherenkov [MEMPHYS]
  - PMm<sup>2</sup> program
  - MEMPHYNO prototype
- R&D for Liquid Argon
  - [GLACIER]
  - [MODULAr]
- Underground site excavation studies
  - Phyasalmy (Finland), Unirea (Romania), [Fréjus (France)]
- EU-funded common Projects:
  - LAGUNA, EUROnu

# European Beam Plans

P.Soler

- The **International Scoping Study** looked at physics, accelerator and detector prospects for future  $\nu$ -oscillation facilities  
[ArXiv 0712.4129 0802.4023 0712.4129]
  - for a  $\nu$ -Factory: baseline 2 detectors, 4000 and 7500 km
    - for "golden channel" Magnetised Iron [cfr INO, I.Dharmavaram] or Magnetised Emulsion Cloud chambers
    - for other channels, Magnetised Liquid Argon or Magnetised Totally Active Scint. Detector (TASD)
  - for Super-Beam or Beta-Beam, no magnetisation needed
    - Water Cherenkov (baseline 500kt), Liquid Argon or non-magn.TASD
  - Beams:
    - Super-Beam: CERN-LNGS (730km) or CERN-Fréjus (130km)
    - Beta-Beam: CERN(SPL)-Fréjus (130km)
    - $\nu$ -Factories: several R&D projects on accelerator concept (MERIT, MICE, EMMA, MuCool)

Note: **Th.Schwetz**: A low-E  $\nu$ -Factory with non-magnetic detectors  
[0805.2019] *"Making  $\nu$ -Factory and b-beams talk to each other!"*

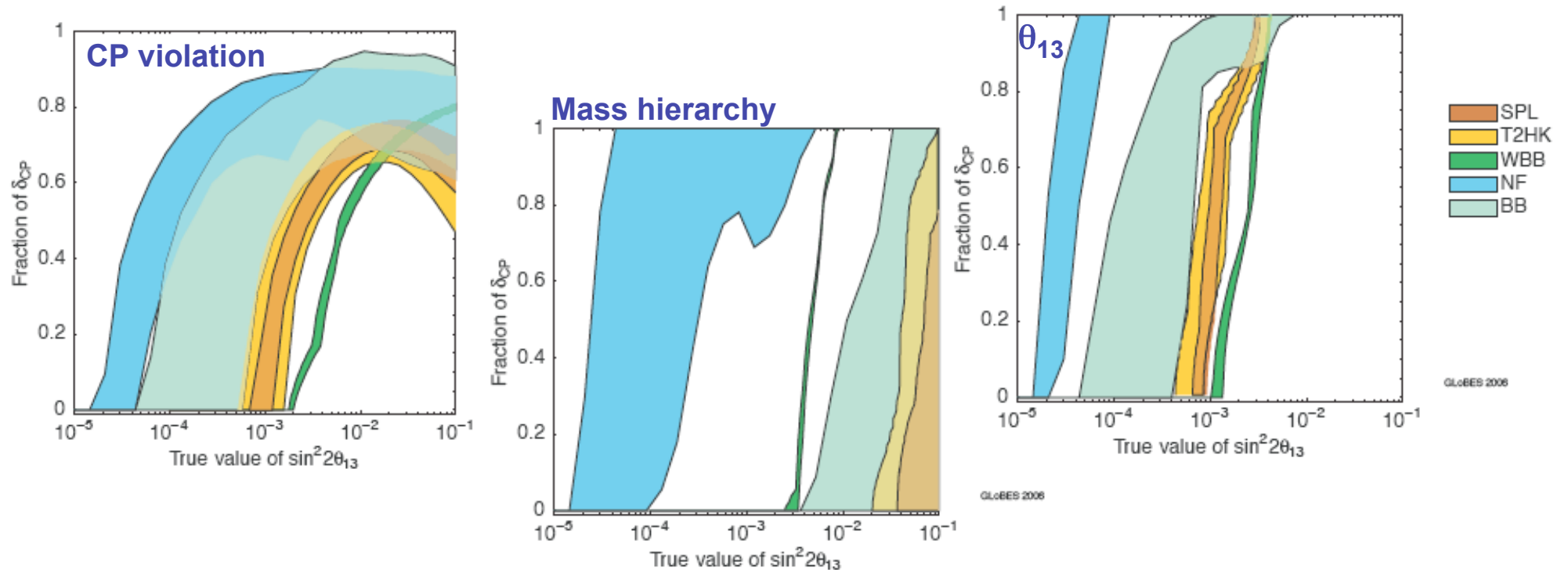
# European Beam Plans

P.Soler

## International Scoping Study conclusions

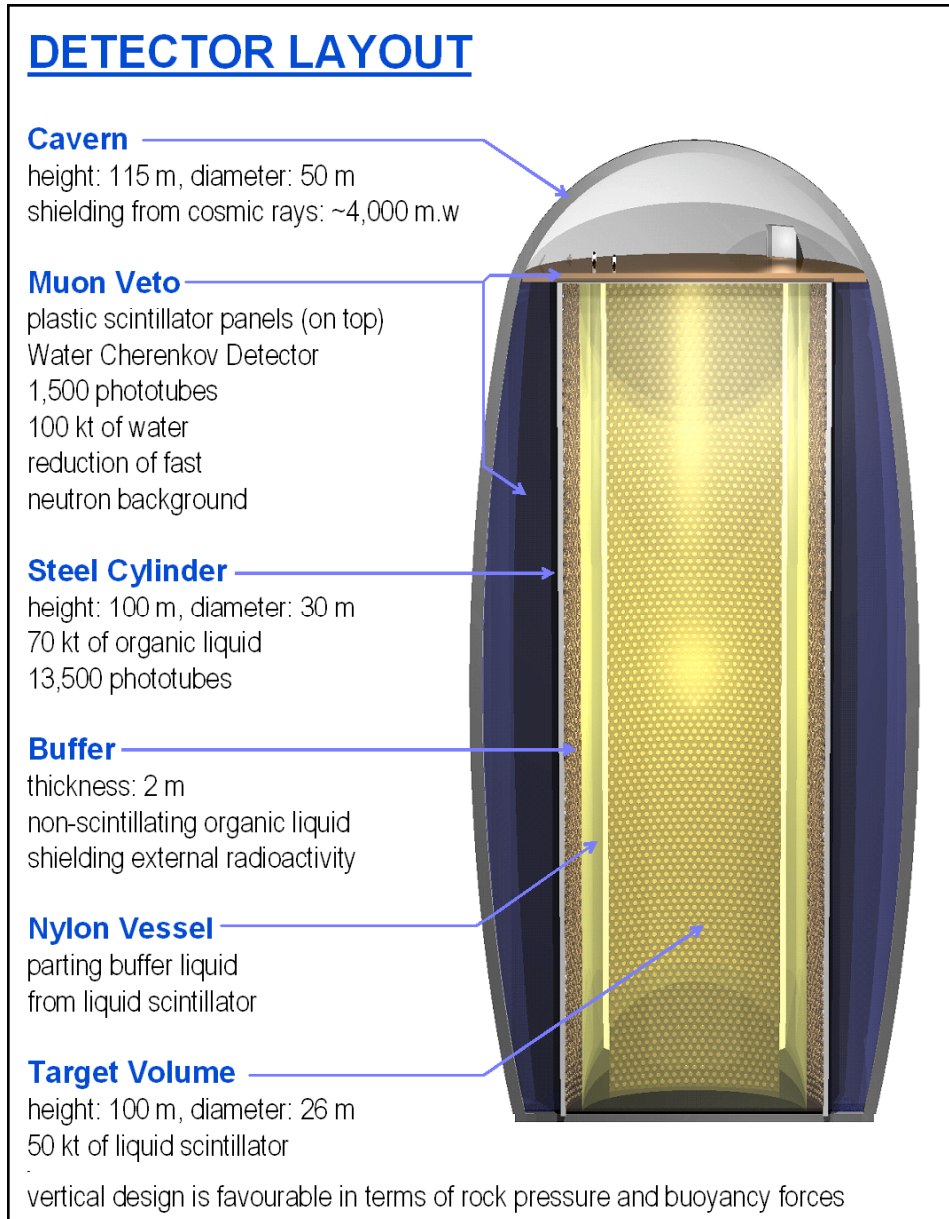
### Comparison of facilities from ISS:

- If  $\sin^2 2\theta_{13} > 10^{-2}$  super-beam and beta-beam facility compatible with neutrino factory to explore CP violation but accuracy might be issue
- If  $\sin^2 2\theta_{13} < 10^{-2}$ , a neutrino factory with two detectors at  $\sim 7500$  km and  $\sim 4000$  km gives optimal CP violation coverage

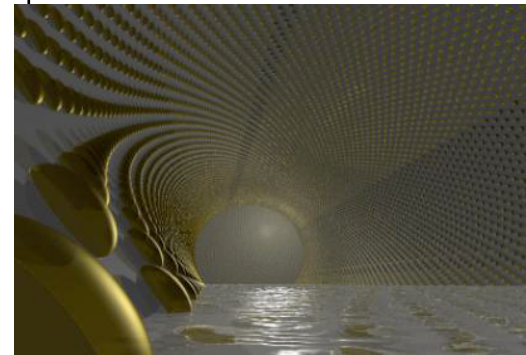


# Liquid scintillator : LENA

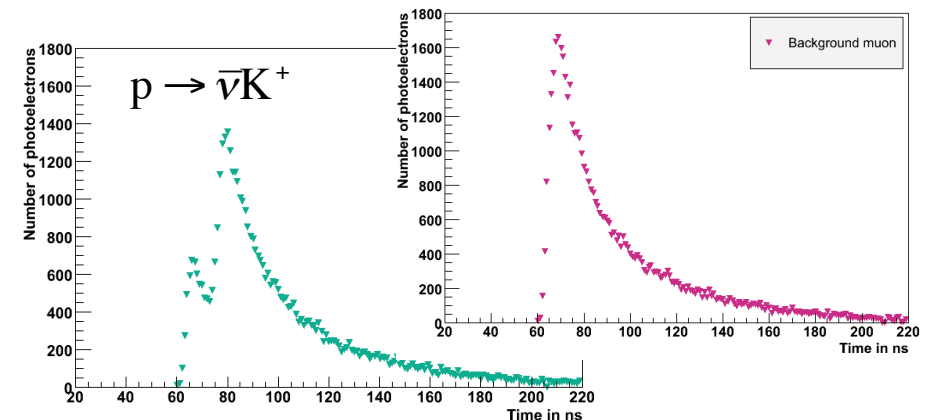
T.Marrodan



Favourite location:  
Phyasalmyi (Finland)  
other locations being studied



## Pulse-shape discrimination:



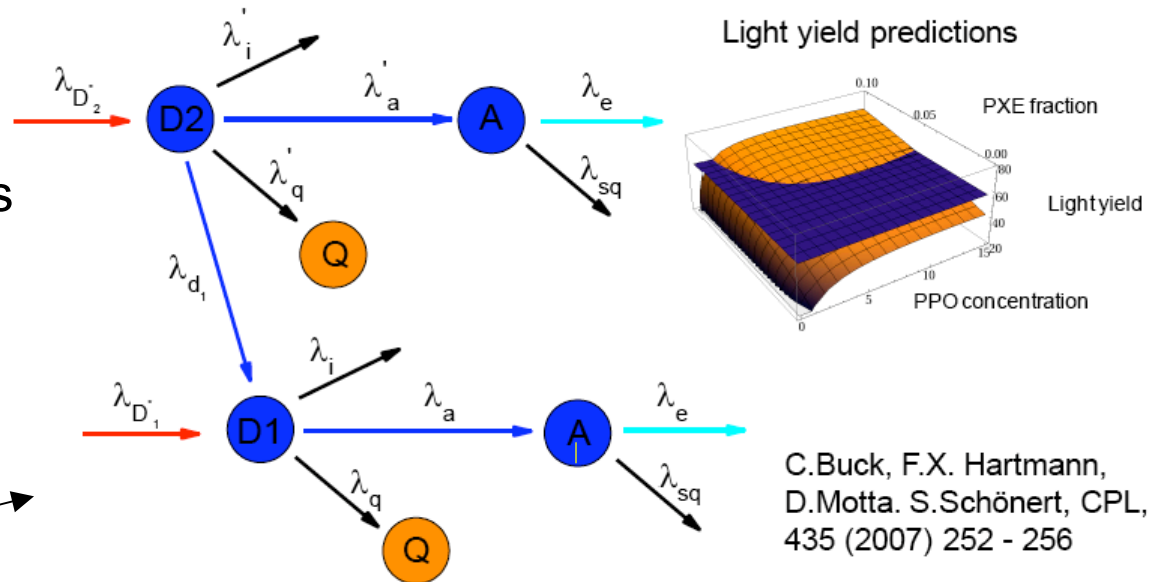
# Liquid scintillator R&D

C.Buck

Requirements for scintillator:

- High energy resolution
- Low energy detection threshold
- high purity (Borexino)
- fast signals (timing)
- moderate cost
- Improved stability of metal loaded scintillators

- ➔ detailed study of many candidates for individual components
- ➔ comparison of purification methods
- ➔ timing measurements
- ➔ detailed light-yield model for simulations



C.Buck, F.X. Hartmann, D.Motta, S.Schönert, CPL, 435 (2007) 252 - 256  
 C.Aberle, diploma thesis, MPIK Heidelberg (2008)

# Liquid scintillator R&D

C.Buck

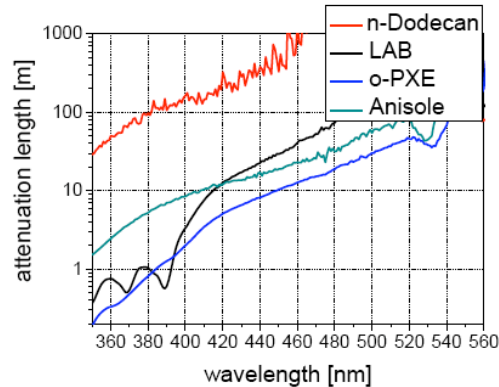
## Comparison solvents

### Scintillation yield

solvent	Yield
PC	1.00
Anisole	0.81
PXE	0.88
LAB	0.74
n-Dodecan	0.40
Oil	0.33

MPIK measurements (6 g/l PPO)

### Attenuation length

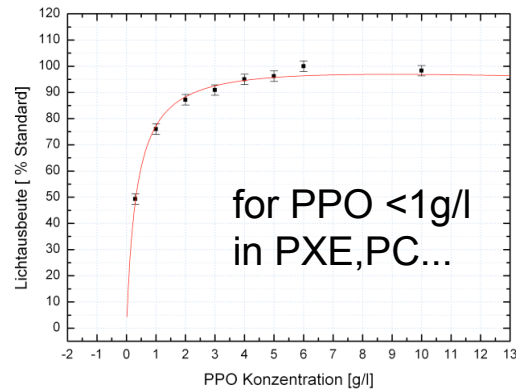


- UV/Vis (MPIK): (absorbance + scattering)
- Single contributions determined at TU Munich

## Comparison fluors

<p><b>PPO</b></p> <p>☺ • transparent</p> <p>☺ • well established</p> <p>☺ • high quantum yield</p> <p>☹</p>		<p><b>BPO</b></p> <p>☺ • high(est) light yield</p> <p>☺ • emission around 400 nm</p> <p>☹ • absorption properties</p> <p>☹ • limited availability</p>	
<p><b>(Butyl-)PBD</b></p> <p>☺ • high light yield</p> <p>☹ • poor quantum yield</p> <p>☹ • costs</p>		<p><b>pTP</b></p> <p>☺ • fast</p> <p>☺ • overlap with bis-MSB</p> <p>☹ • low solubility</p> <p>☹ • poor quantum yield</p>	

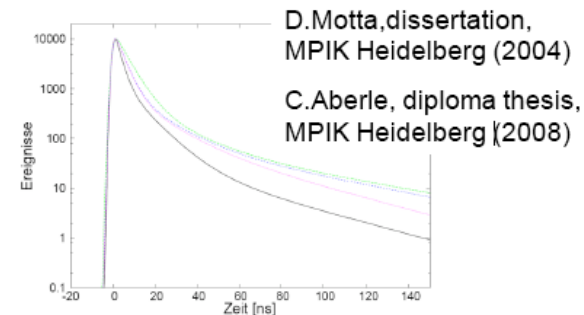
## Critical concentration



for PPO <1g/l  
in PXE,PC...

for PPO ~3g/l in dodecane

## Timing properties



D.Motta, dissertation,  
MPIK Heidelberg (2004)

C.Aberle, diploma thesis,  
MPIK Heidelberg (2008)



# Liquid scintillator R&D

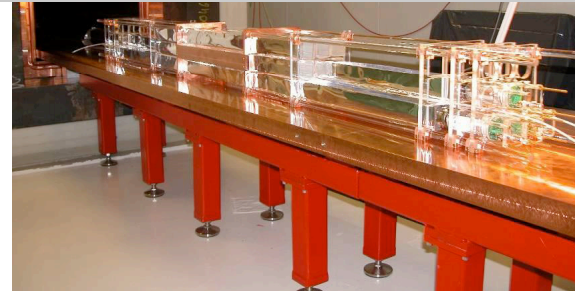
C.Buck

## Metal loaded scintillators

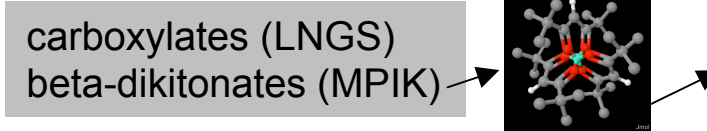
- Solar neutrinos (LENS, SIREN)
  - Yb, In, Gd
  - challenge: high loading

Indium-loaded scintillators at LLBF > 1 year

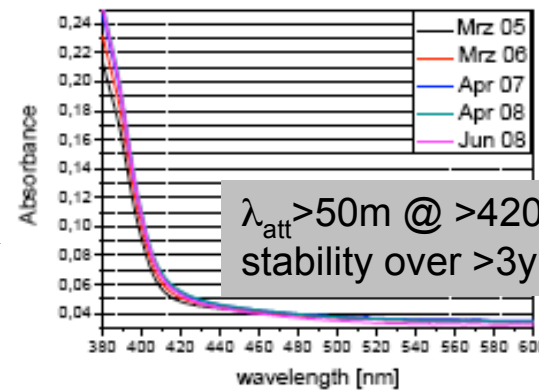
➤ MPIK: $\text{In}(\text{acac})_3$ (F.X.Hartmann et al.)	} > 50 g/l Indium
➤ INR/LNGS: Carboxylic acid version	



- Reactor (Double Chooz (LENA))
  - Gadolinium
  - Challenge: stability



Gd not sublimed, Jan 05, no fluors, PXE isomers mixed



100 kg produced  
(for both DC detectors)



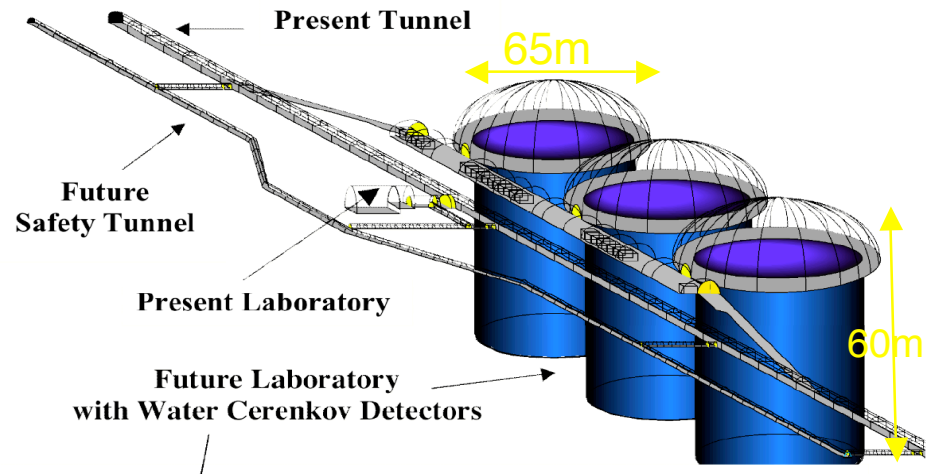
- bb decay (SNO+)
  - Neodymium
  - Challenge: transparency, purity

tests done on BDK and carboxylates  
@ MPIK / LNGS

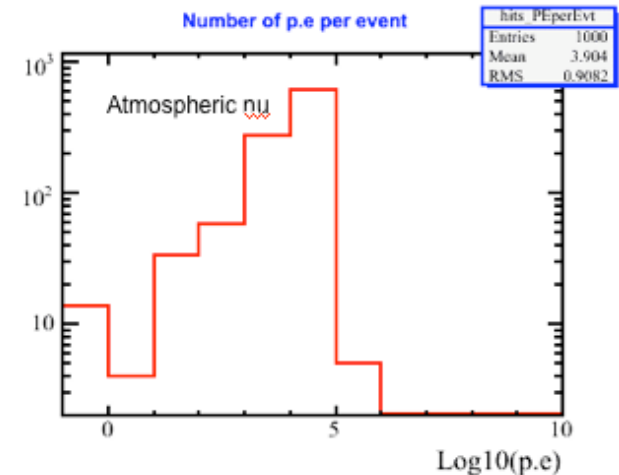
# Water Cherenkov: MEMPHYS

## Megaton Mass PHYSics @Fréjus

- Water Cherenkov (“cheap and stable”)
- Total fiducial mass: 440 kt
- Baseline:
  - 3 Cylindrical modules 65x65 m
    - Size limited by light attenuation length ( $\lambda \sim 80\text{m}$ ) and pressure on PMTs
    - Readout: 12” PMTs, 30% geom. cover (#PEs = 40% cov. with 20” PMTs)
- Installation in extension of LSM in the Fréjus tunnel, b/w France and Italy
  - 4800 m.w.e overburden
  - 130 km from CERN



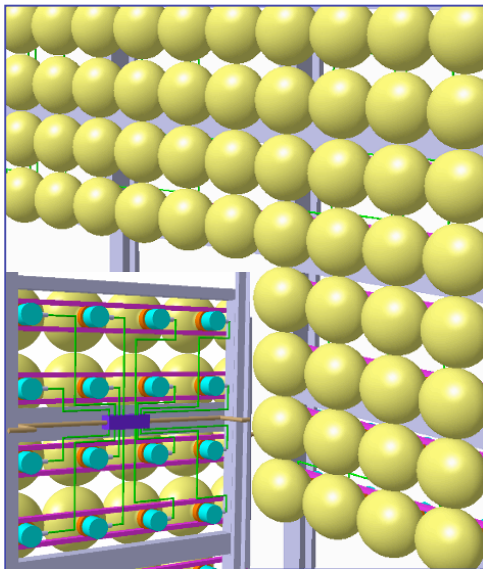
## Simulation: Nuance + Geant4 + OpenScientist + AIDA/root



[http://www.apc.univ-paris7.fr/APC\\_CS/Experiences/MEMPHYS/](http://www.apc.univ-paris7.fr/APC_CS/Experiences/MEMPHYS/)  
**arXiv: hep-ex/0607026**

# (Water Cherenkov) PMm<sup>2</sup> R&D

G.Martin-Chassard



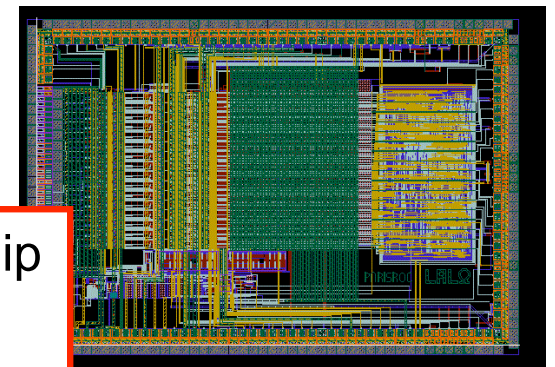
*“Innovative electronics for array of photodetectors used in High Energy Physics and Astroparticles”.*

R&D program funded by French national agency for research (LAL, IPNO, LAPP and Photonis) (2007-2010)

**Basic concept:** very large photodetection surface  
→ macropixels of PMTs connected to an autonomous front-end electronics.

Replace large PMTs (20”) by groups of 16 smaller ones (12”) with central ASIC :

- Independent channels
- charge and time measurement
- water-tight, common High Voltage
- Only one wire out (DATA + VCC)



I. studies on 12" PMTs design

- parameter correlation
- potting
- pressure resistance

(collaboration with BNL since NNN07)

II. **PARISROC** readout chip

- complete front-end chip with 16 channels
- testboard now in layout, soon available

# Water Cherenkov: MEMPHYNO

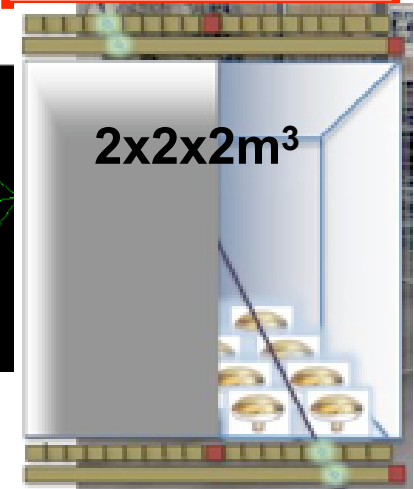
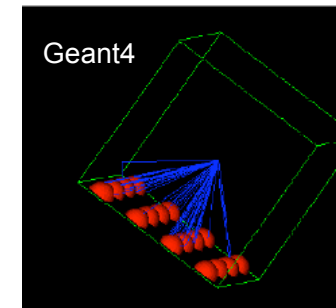
## Goals:

1. full test of electronics and acquisition chain with actual physics events
2. trigger threshold studies
3. self-trigger mode
4. Track reconstruction performances
5. Gd doping: feasibility and performance (if studies still needed...)

Test bench for photodetection solutions for large detectors

## Plan:

1. Install at APC lab ("hall de montage") => cosmic muon data
  - direction selected with hodoscope
  - test timing, track reconstruction
2. Transport to Fréjus LSM lab
  - Measure backgrounds at underground site
  - Test trigger and thresholds
3. Expose to CERN beam [with DevDet European program]
  - Check electron vertex reconstruction and e/pi separation
4. Back to APC
  - Test Gd-doping



- for DSN detection in MEMPHYS
- for non-proliferation studies :

MEMPHYNO is a "small-scale", "portable" detector, easy to install near reactor power plants

# Liquid Argon: GLACIER

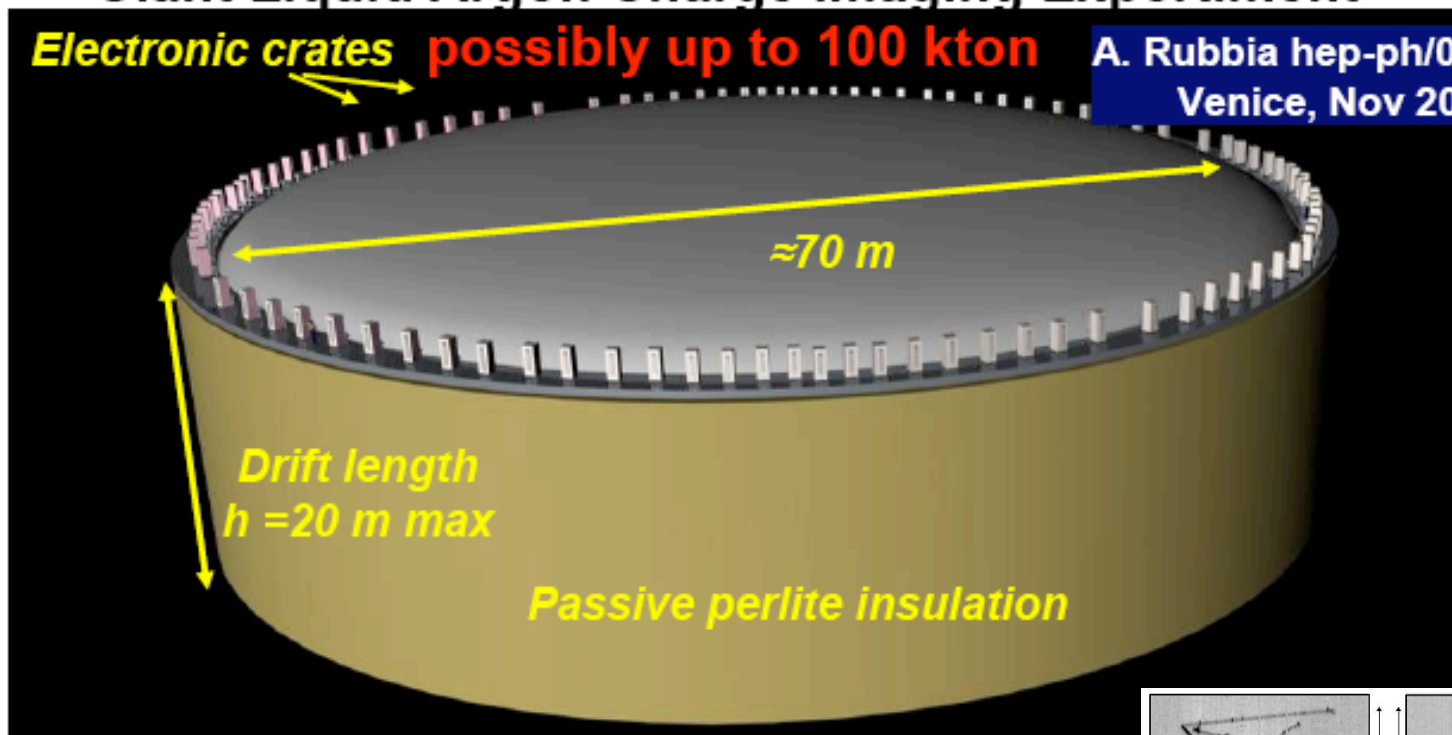
A. Marchionni

A scalable detector with a non-evacuatable dewar and ionization charge detection with amplification

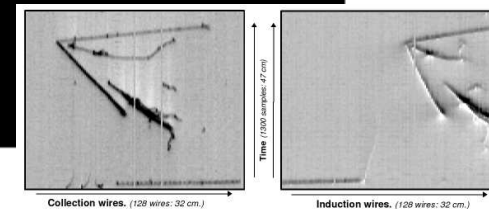
## Giant Liquid Argon Charge Imaging Experiment

**Electronic crates possibly up to 100 kton**

A. Rubbia hep-ph/0402110  
Venice, Nov 2003



Single module cryo-tank based on industrial LNG technology



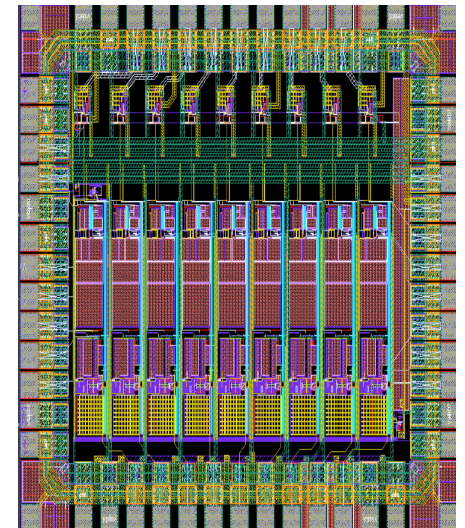
**LAr-TPC @ CERN-WANF**  
Phys. Rev. D 74, 112001 (2006)

# Liquid Argon: GLACIER

A.Marchionni

Critical issues for construction → R&D items:

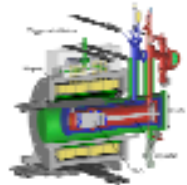
- Drift over long distances in Ar: Ok if high purity → purification system
- Dewar → (non evacuabile?) R&D with Technodyne in LAGUNA
- HV system
- ReadOut system → Novel techniques, other than wires, possibly with charge multiplications (double-phase with Large Electron Multiplier)
- Electronics → Aggressive R&D on warm/cold solutions (IPNL+ETHZ)
  - analog ASIC amplifier working at cryo temperature
  - Gygabit Ethernet readout chain + network time distribution PTP
- Detector engineering



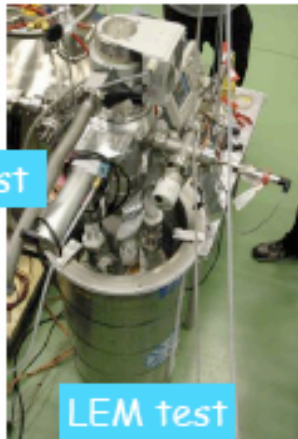
# GLACIER R&D - Prototyping

A. Marchionni

**Small prototypes → ton-scale detectors → 1 kton → ?**



B-field test



LEM test

proof of principle double-phase LAr LEM-TPC on 0.1x0.1 m<sup>2</sup> scale

LEM readout on 1x1 m<sup>2</sup> scale UHV, cryogenic system at ton scale, cryogenic pump for recirculation, PMT operation in cold, light reflector and collection, very high-voltage systems, feed-throughs, industrial readout electronics, safety (in Collab. with CERN)

ArDM ton-scale



direct proof of long drift path up to 5 m

we are here

Argon Tube: long drift, ton-scale



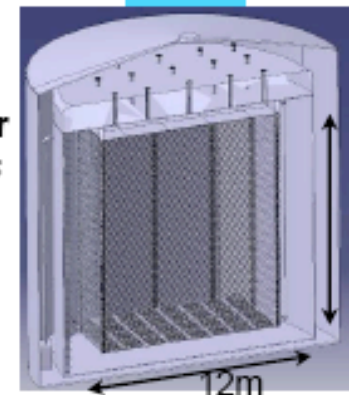
Test beam 1 to 10 ton-scale

Application of LAr LEM TPC to neutrino physics: particle identification (200-1000 MeV electrons), optimization of readout and electronics, cold ASIC electronics, possibility of neutrino beam exposure



full engineering demonstrator for larger detectors, acting as near detector for neutrino fluxes and cross-sections measurements, ...

1 kton

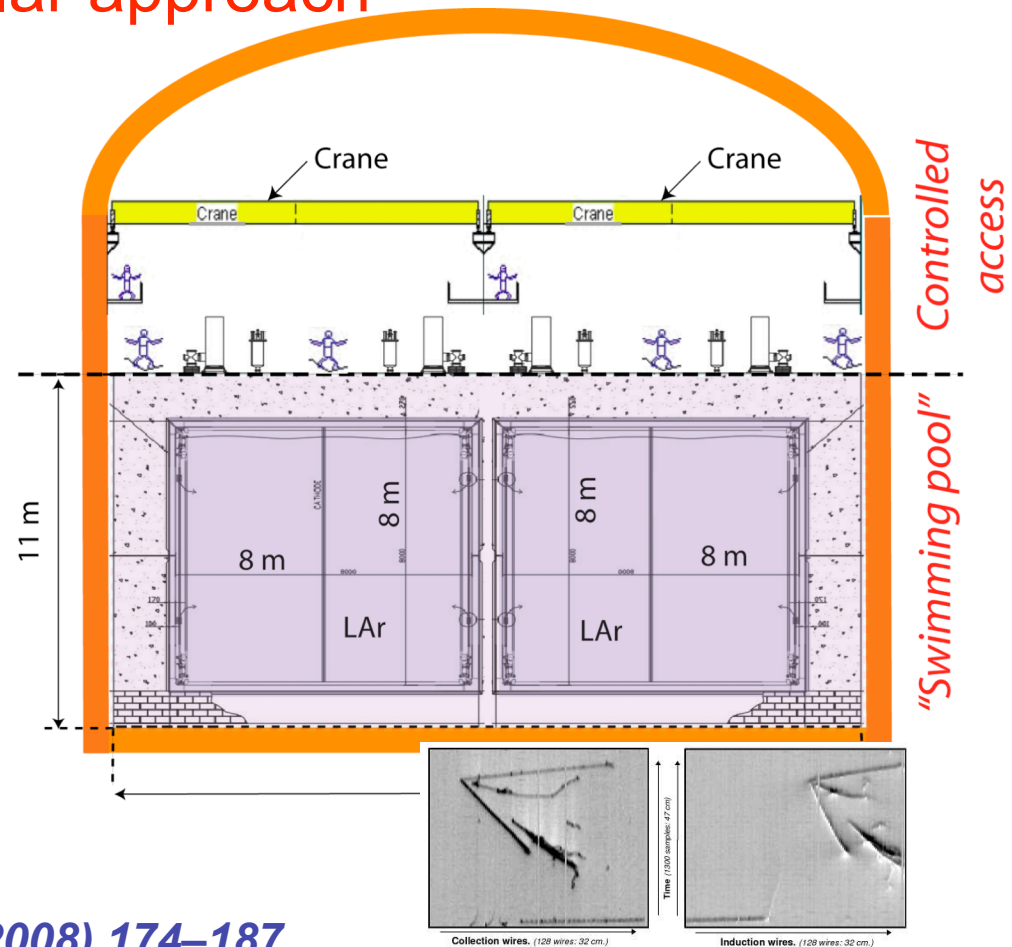


# Liquid Argon: MODULAR

F.Pietropaolo

## From ICARUS-T600 to a multi-kton LAr-TPC with a modular approach

- MODULAR will be initially composed by four identical module located in the new shallow-depth cavern
- Each module is a scaled-up version of the T600 ( $\times 2.66^3$ ):
  - 8 x 8 m<sup>2</sup> cross section and about 60 m length
  - LAr active mass: 5370 ton
  - 4 m electron drift
  - 3-D imaging similar to T600 but 6 mm pitch (three wire planes, ~50000 channels)



**Proposal: Astroparticle Physics 29 (2008) 174–187**

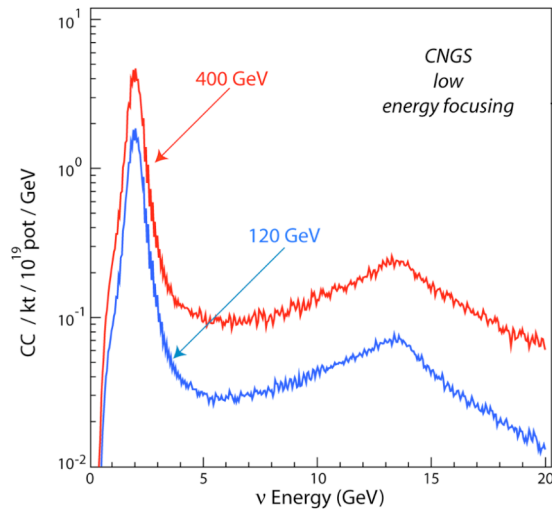
**LAr-TPC @ CERN-WANF**  
*Phys. Rev. D 74, 112001 (2006)*



# MODULAR at LNGS off-axis

F. Pietropaolo

732km CERN-LNGS, 10km off-axis

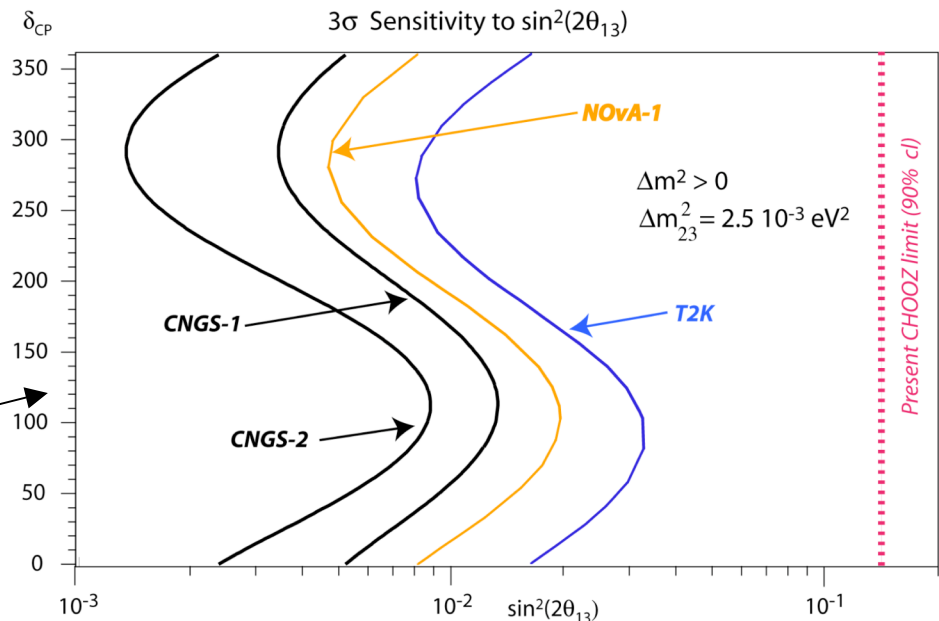


There is no near detector @CERN, but ICARUS-T600 will be operational on-axis (on  $\rightarrow$  off axis normalisation is straightforward)

Thanks to LAr properties:  
 MODULAR (20 kt) + CNGS ( $1.2 \cdot 10^{20}$  pot/y) ~  
 NOvA + NUMI ( $6.5 \cdot 10^{20}$  pot/y)

Sensitivity to  $\theta_{13}$  and  $\delta_{CP}$

CNGS-1(-2):  $1.2$  ( $4.3$ )  $10^{20}$  pot/y  
 Exposure: 20 kt x 5 years  
 GLOBES: 5% beam syst.,  $\Delta E/E = 15\%$



# MODULAR R&D

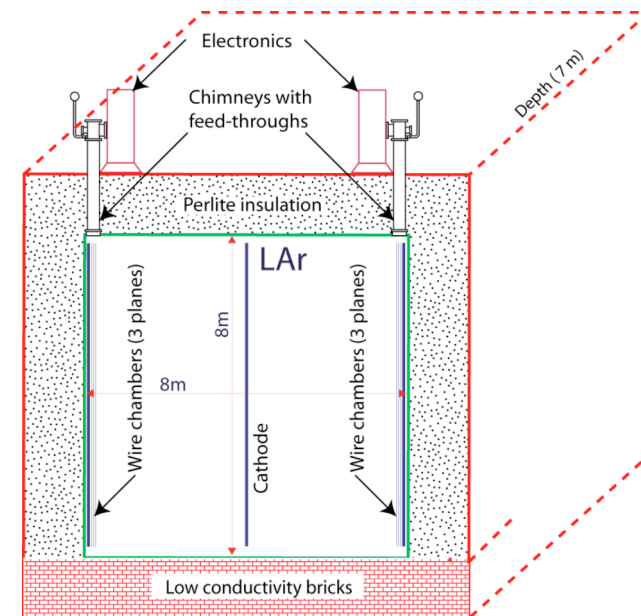
F.Pietropaolo

Scaling by only 1 order of magnitude w.r.t. ICARUS600 requires only some "not very substantial" R&D items:

- the **filling process starting from air to pure LAr**, taking into account the motion of the gas, optimizing the inlet and outlet geometries and minimizing the number of cycles;
- the **thermal convection of the LAr**, in order to optimize the temperature gradients and to ensure circulation in all regions of the dewar, both in the cool down and stationary phases;
- the **out-gassing rate and the re-circulation processes** necessary to achieve the required electron lifetime;
- the **geometry of the compact re-circulators** both in the liquid and in the gaseous phases.
- finally, also the **electronics and DAQ** may require some specific developments to improve the layout of the analogue front-end and the DAQ architecture.

## SLICE prototype

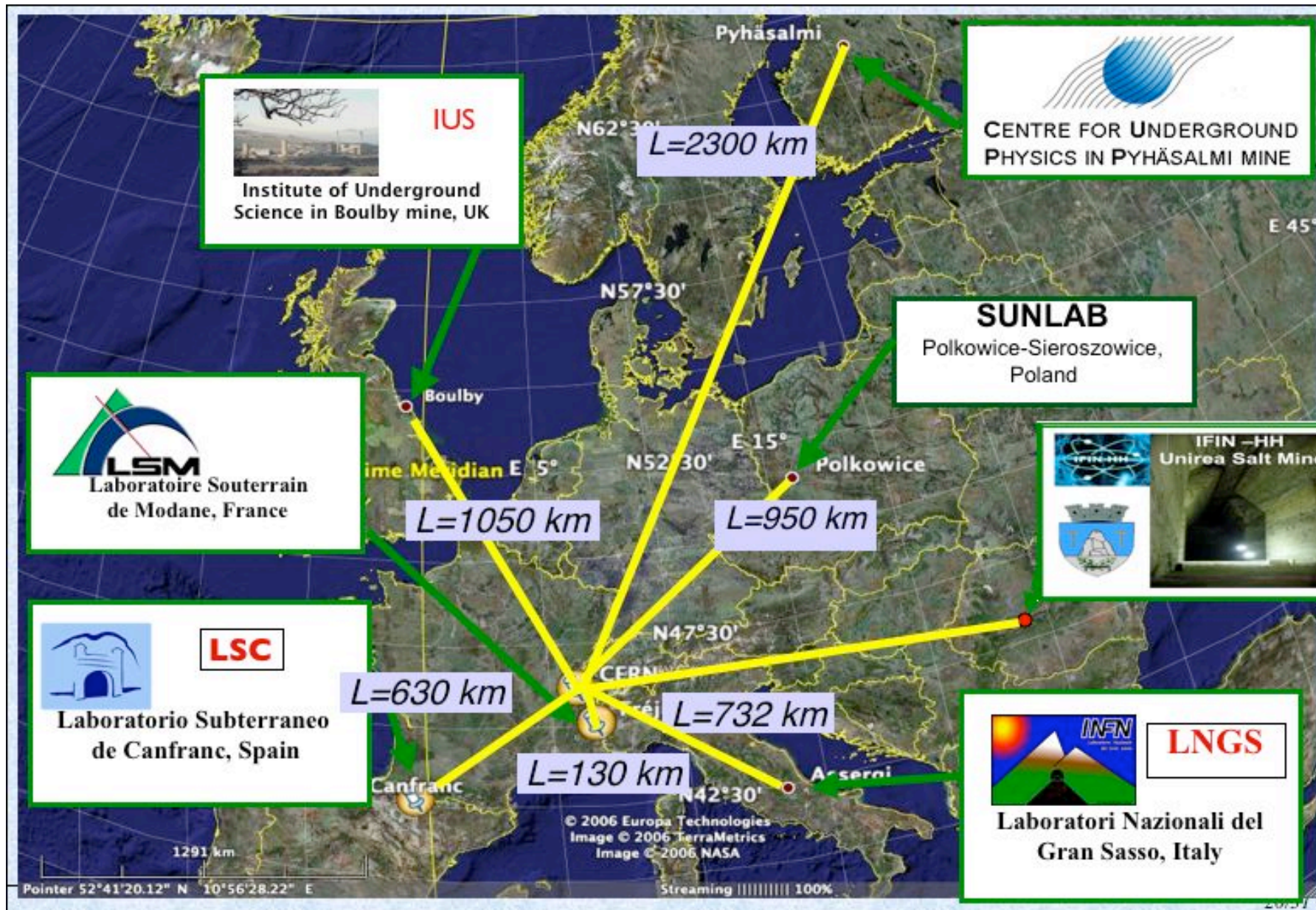
sensitive area 8x8m<sup>2</sup>, depth 4m



# Physics scope

	Water Cerenkov	Liquid Argon TPC	Liquid Scintillator
Total mass	500 kton	100 kton	50 kton
$p \rightarrow e \pi^0$ in 10 years	$1.2 \times 10^{35}$ years $\epsilon = 17\%$ , $\approx 1$ BG event	$0.5 \times 10^{35}$ years $\epsilon = 45\%$ , $<1$ BG event	$\sim 10^{32}$ years in 1year $\epsilon = 12\%$ , BG under study
$p \rightarrow \nu K$ in 10 years	$0.15 \times 10^{35}$ years $\epsilon = 8.6\%$ , $\approx 30$ BG events	$1.1 \times 10^{35}$ years $\epsilon = 97\%$ , $<1$ BG event	$0.4 \times 10^{35}$ years $\epsilon = 65\%$ , $<1$ BG event
SN cool off @ 10 kpc	194000 (mostly $\nu_e p \rightarrow e^+ n$ )	38500 (all flavors) (64000 if NH-L mixing)	20000 (all flavors)
SN in Andromeda	40 events	7 (12 if NH-L mixing)	4 events
SN burst @ 10 kpc	$\approx 250$ $\nu$ -e elastic scattering	380 $\nu_e$ CC (flavor sensitive)	$\approx 30$ events
SN relic	250(2500 when Gd-loaded)	50	20-40
Atmospheric neutrinos	56000 events/year	$\approx 11000$ events/year	5600/year
Solar neutrinos	91250000/year	324000 events/year	?
Geoneutrinos	0	0	$\approx 3000$ events/year

# Underground sites



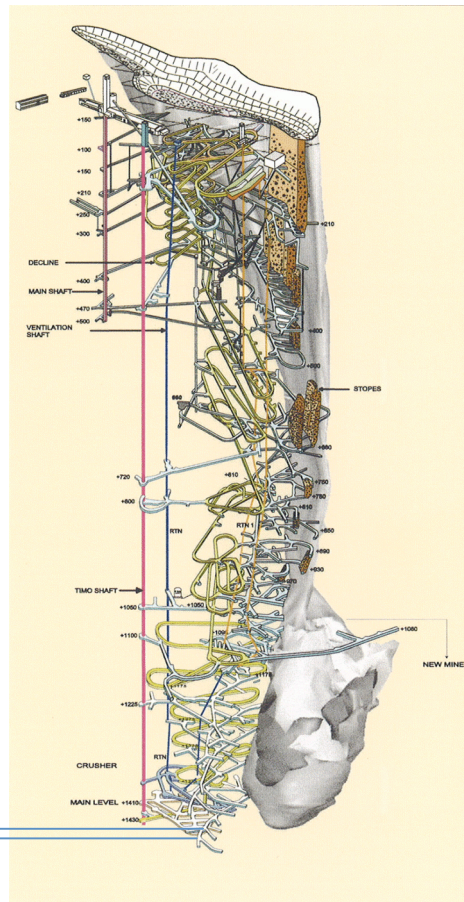
# Pre-feasibility studies

G.Nujiten

## Phyasalmi rock mine (Finland)

very old bedrock  
~1400m depth

vertical  
cylindrical  
cavern (LENA)  
can be  
constructed  
~0,5km from  
present mine

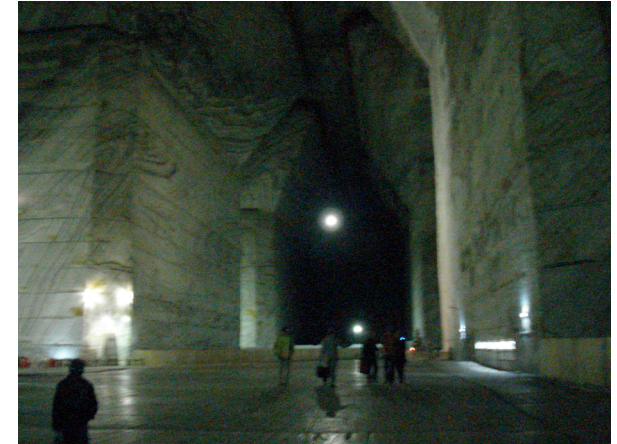


LENA

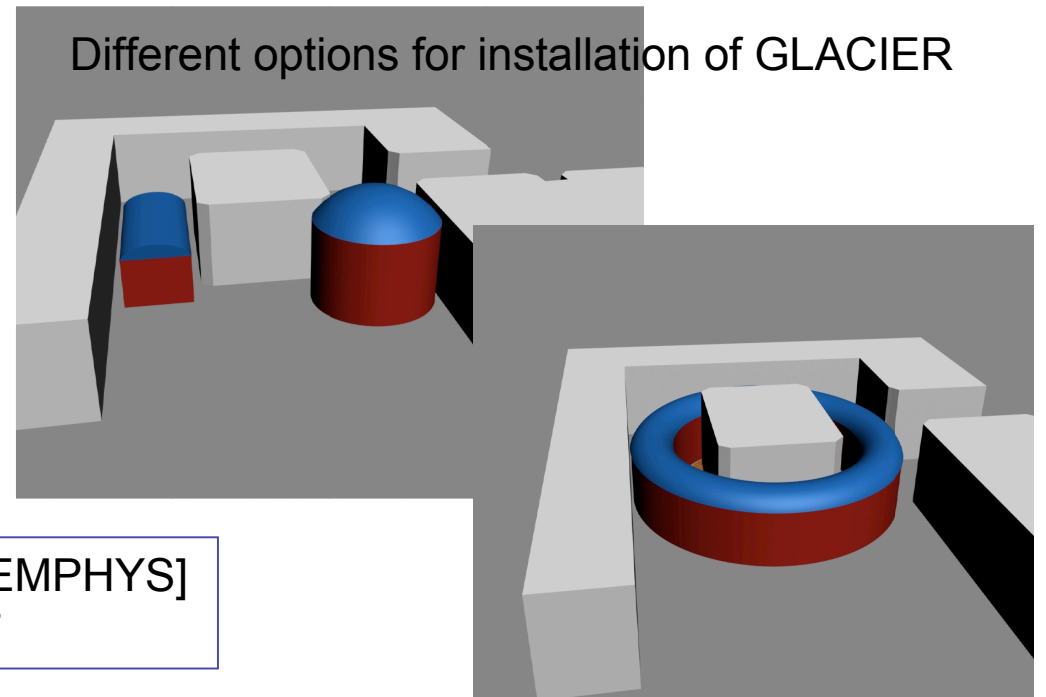
R.Margineau

## Unirea salt mine (Romania)

under  
~150m hill



Different options for installation of GLACIER



+ **Fréjus** (France) [MEMPHYS]  
cfr L.Mosca @NNN07

# Europe FP7 funded projects

**ASPERA** (from ApPEC) is a network of national government agencies responsible for coordinating and funding national research efforts in Astroparticle Physics



recommendation in the roadmap:

*"We recommend that a **new large European infrastructure** is put forward, as a future **international multi-purpose facility on the  $10^5$ - $10^6$  ton scale** for improved studies of **proton decay and of low-energy neutrinos from astrophysical origin.***

*The three detection techniques being studied for such large detectors in Europe, **Water-Cherenkov, Liquid Scintillator and Liquid Argon**, should be evaluated in the context of a common design study which should also address the underground infrastructure and the possibility of an eventual detection of future accelerator neutrino beams.*

*This design study should take into account worldwide efforts and converge, on a time scale of **2010**, to a common proposal.*

The **European Union** gives funding for collaborative infrastructure studies, provided institutions contribute with proper funds (personnel)

Two EU-FP7 funded programs are related to NNN08 subjects:

- LAGUNA and EUROnu

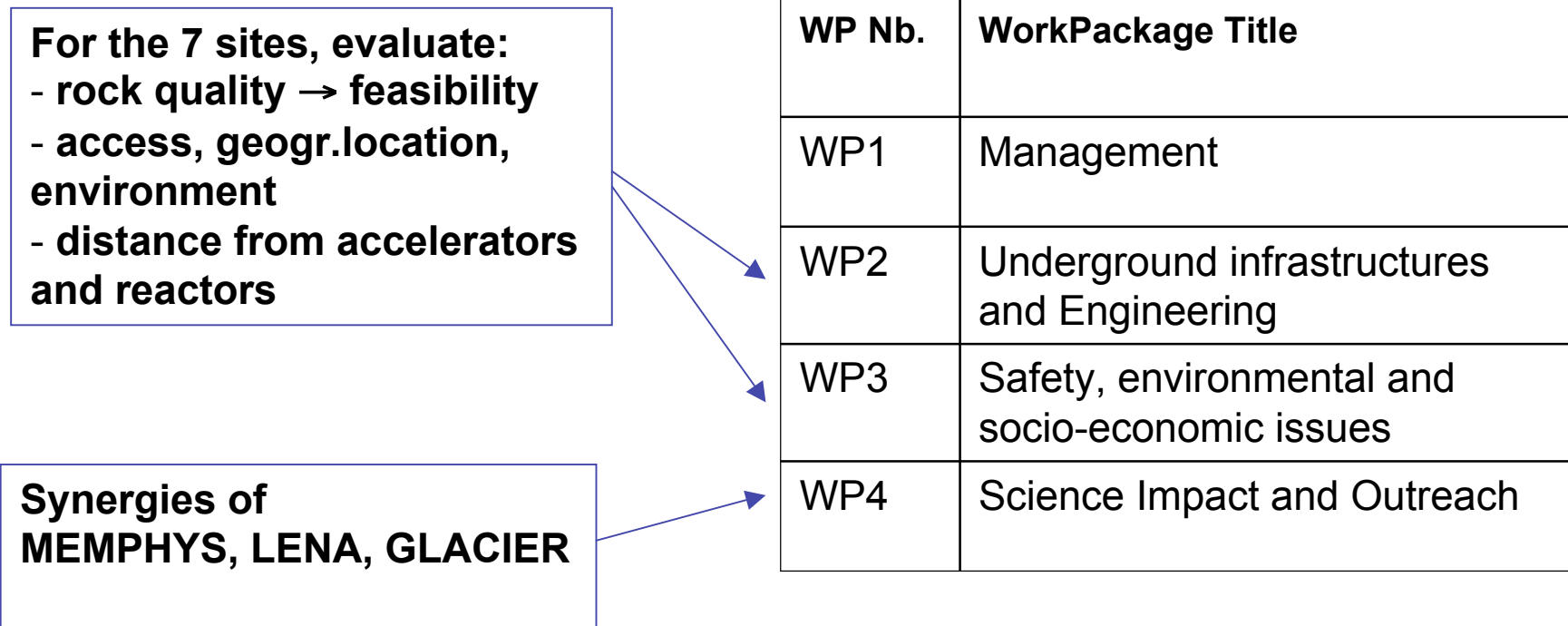
# LAGUNA

Th.Patzak

## Large Apparatus for Grand Unification and Neutrino Astrophysics

**1.7 M€** to be mainly devoted to **the sites infrastructure studies**

21 beneficiaries in 9 countries: 9 higher education entities, 8 research institutes, 4 private companies (+4 additional universities)



# EUROnu

P.Soler

“A High Intensity Neutrino Oscillation Facility in Europe”  
4.5MEuros, 15 beneficiaries in 9 countries



- **WP1: Management**
- **WP2: Super-Beam**: design of a 4 MW proton beam (SPL), target and collection system for a conventional neutrino beam
- **WP3: Neutrino factory**: define design for muon front-end, acceleration scheme, spent proton beam handling and component integration in an end-to-end neutrino factory simulation
- **WP4: Beta beam**: following from EURISOL, study production, collection and decay ring of beta beam for high Q isotopes ( $^8\text{Li}$ ,  $^8\text{B}$ )
- **WP5: Neutrino detectors**: study Magnetised Neutrino Iron Detector (MIND) performance for golden measurement at neutrino factory, water Cherenkov detector for beta and super beams and near detectors for all facilities
- **WP6: Physics**: comparison of physics performance, systematic errors and optimisation for all facilities



# Summary and outlook

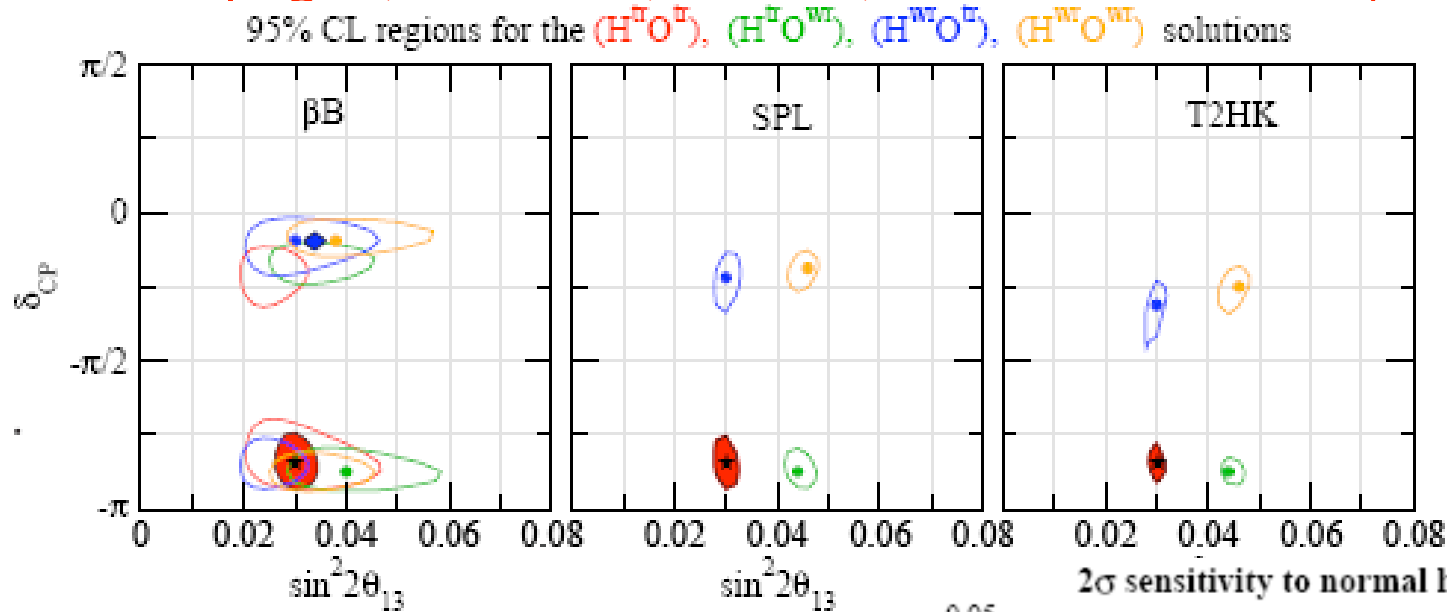
- NNN08 was a very dynamic workshop
- We had an exhaustive review of the physics scope for next generation detectors for proton decay, supernovae neutrinos and neutrino oscillations
- We got a vast overview of the plans of Europe, US and Japan on Beams, Detectors, R&D and Infrastructures
- Reminder on Europe:
  - ISS on future accelerators
  - Detectors: LENA, MEMPHYS, GLACIER and MODULAR
  - 7 possible underground labs
  - ASPERA priority: Large underground detectors for proton decay, Particle Astrophysics and neutrino oscillations
  - 2 EU-funded projects: LAGUNA and EUROnu
- New international collaborations are being started - looking forward to strengthening them at this UDiG workshop

optimal strategy depends on  $\theta_{13}$  :  
current / near-future experiments are crucial

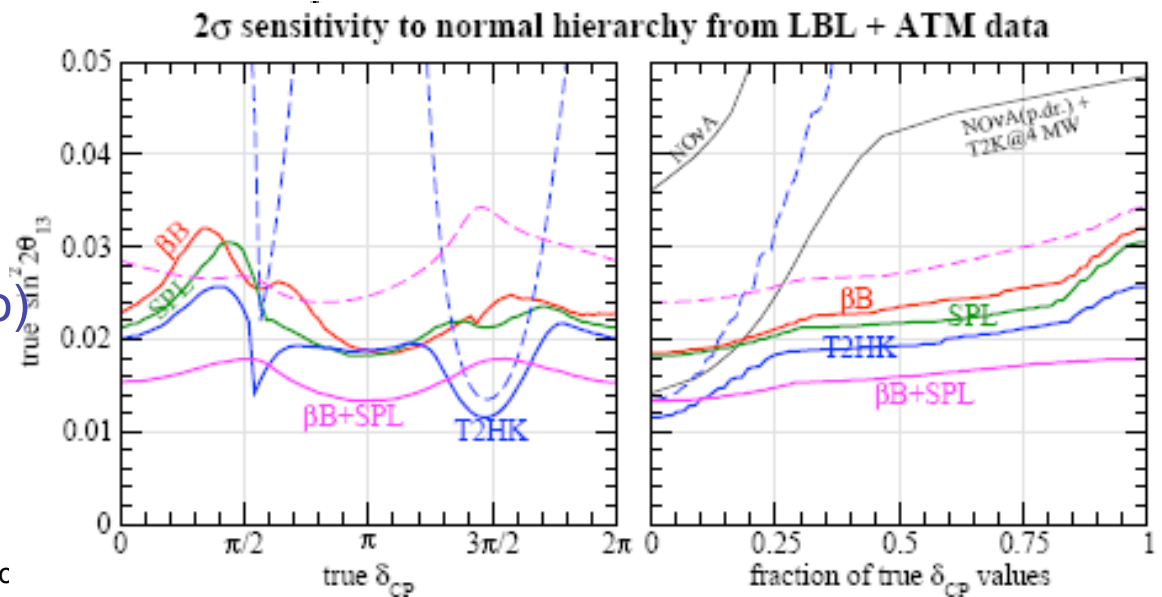
**BACKUP**

# Resolving degeneracies

- Physics reach of Memphys+SPL; Betabeam+SPL and T2HK:  
Campagne, Mezzetto, Maltoni, Schwetz JHEP 04 (2007) 003.



- Resolve degeneracies by combining with atmospheric
- $\beta B$ : cannot solve degeneracies (no  $\nu_{\mu}$  and insufficient spectral info)
- $\beta B+SPL$ ,  $SPL+ATM$  &  $T2HK$  can resolve degeneracies



# PArISROC description (I)

- Complete front-end chip with 16 channels
  - Sent in fabrication in **June 2008**
  - Technology : **AMS SiGe 0.35  $\mu\text{m}$**
- Characteristics :
  - 16 inputs preamplifier
    - Variable gain : 1  $\rightarrow$  8 (4bits) (common on 16 channels)
    - PMTs gain adjustment by a factor 4 (8 bits) (channel by channel)
    - Input dynamic range : 0  $\rightarrow$  300 pe (0  $\rightarrow$  50pC)
    - Good linearity (1%)
  - 16 trigger outputs:
    - Fast shaper ( $\tau=15\text{ns}$ )
    - Low offset discriminator
    - Threshold provided by common 10bit DAC +4bit DAC/ch. (1/3 pe)
    - “OR” of 16 triggers output
  - 1 digitized and multiplexed charge output :
    - Dynamic range : 0  $\rightarrow$  300 pe
    - Slow shaper with variable shaping time ( $\tau=50\text{ns}, 100\text{ns}, 200\text{ns}$ )
    - SCA with depth 2



# PArISROC description (II)

- Coarse time measurement (timestamp) :
  - 24-bit counter @ 10MHz
  - Step : 100ns
  
- 12-bit ADC for charge and fine time measurement :
  - Wilkinson type ADC
  - T&H on slow shaper for charge measurement
  - T&H on TDC ramp (100ns) for fine time measurement
  - 2 discriminators with 12 bit ramp (100 $\mu$ s) as threshold
  
- Serialization of digital output information :  
Channel number - time stamp – charge - fine time  
4bits                                      24bits                      12bits                      12bits



## CONCLUSIONS (1/2)

- The hard and very old bedrock ( $> 2 * 10^9$  yr) of Finland provides by far one of the best locations in Europe to locate the LENA laboratory deep under the ground.
- Within Finland the Pyhäsalmi Mine offers the best location for this purpose, as it is the deepest present location in Finland 1400m below surface.
- The depth of the laboratory is at -1400m (top) to -1500m (bottom level).
- The location of the cavern is about 0,5km west from the present mine.
- Construction of the access tunnels can be started directly at -1400m due to good infrastructure at that level.



## CONCLUSIONS (2/2)

- The shape of the cavern is vertical and ellipse to cope best with in-situ stresses
- The scintillator tank is also constructed vertically
- The infrastructure needed for LENA is already present on the surface as well as at -1400 m.
- The mine has one access tunnel, one main transport and one ventilation shaft. (probably one extra shaft has to be built)
- The mine can take over all excavated rock (~half a million m<sup>3</sup>) until the year 2015.
- The layout of the underground tunnels is based on a double exit strategy that in case of accident there are two directions to flee from any location.
- Otanmäki and Vihanti are other possible locations