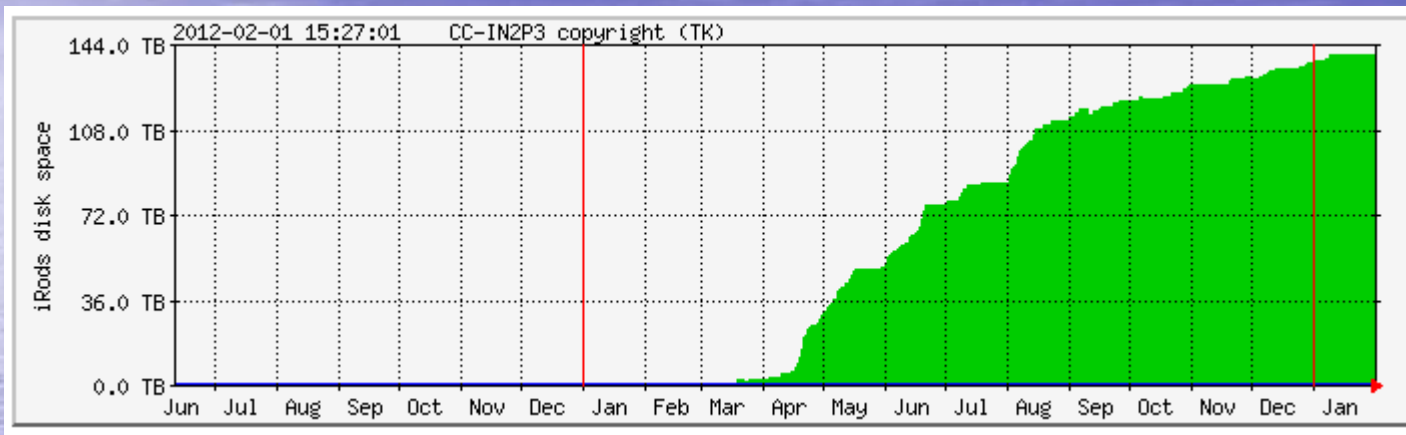


Amas Processing Status

J.E Campagne (LAL) 3/2/12-Meudon

- Data proc. Status
- Calibrator sources analysis
- Abellxyz: zoom freq. band for HI signal
- Noise: introduction to A.S Torrento's talk

Data



- 500 cycles ON-OFF Abell85
- 240 cycles ON-OFF Abell2440
- 900 cycles ON-OFF Abell1205
- Few 3C161, 3C273, NGC4383 runs both ON-OFF & DR

1 cycle ON-OFF 2min total: 30sec x 30% ON + 30sec x 30% OFF

1 cycle DR ~3min total: 170sec x 30% x corr. fact

Data processing on Irods @ CCIN2P3 (short list)

1. Median ov. 5120 BAO paq. ON/Gain or OFF/Gain
2. Mean of 5 medians => med. spectrum (FITS files)
3. Mean on each cycle
4. For each cycle (ON-OFF)/OFF_f
 1. Result of (ON/G-OFF/G)/(OFF/G)_f
 2. G=Gain med. Filtered 3MHz
 3. (OFF/G)_f= (OFF/Gain) med. Filtered 2MHz
5. For each cycle too $< (ON-OFF)/OFF_f >$ ov. $[v_1, v_2]$

Limitation due to batch queue parameters although we are using LONG & HUGE !
In particular: median filtering limited ~ 5120 paq; cycles with 170sec (Drift) are the maximum but for ON-OFF it could be non-particable as then we should split the run into individual cycles and we are limited by < 20 simultaneous jobs accessing Irods.

Calibrator Sources

3C273, 3C261, NGC4383

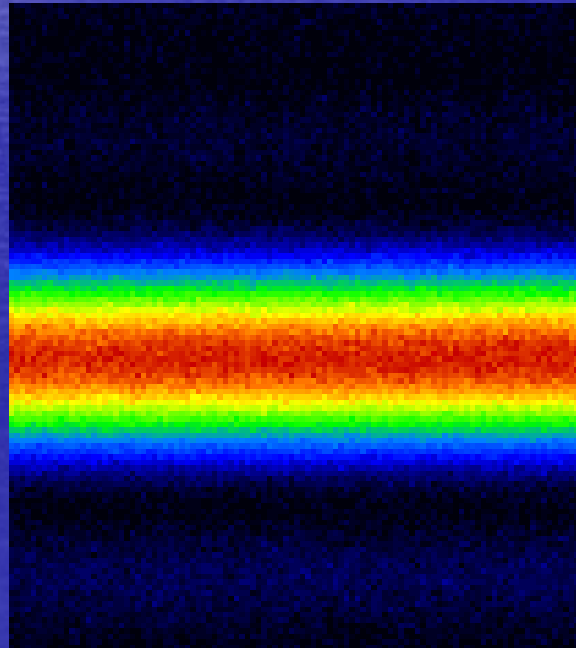
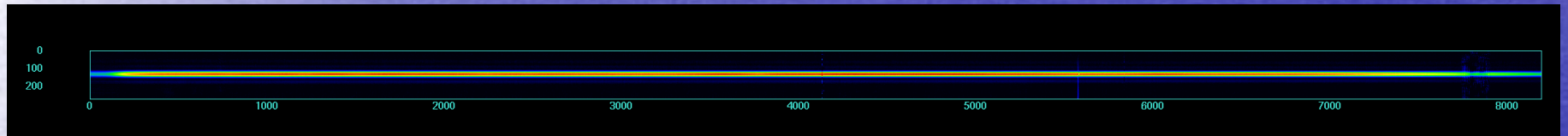
3C161 Drift Scan

158451.171 & 158452.171

Data

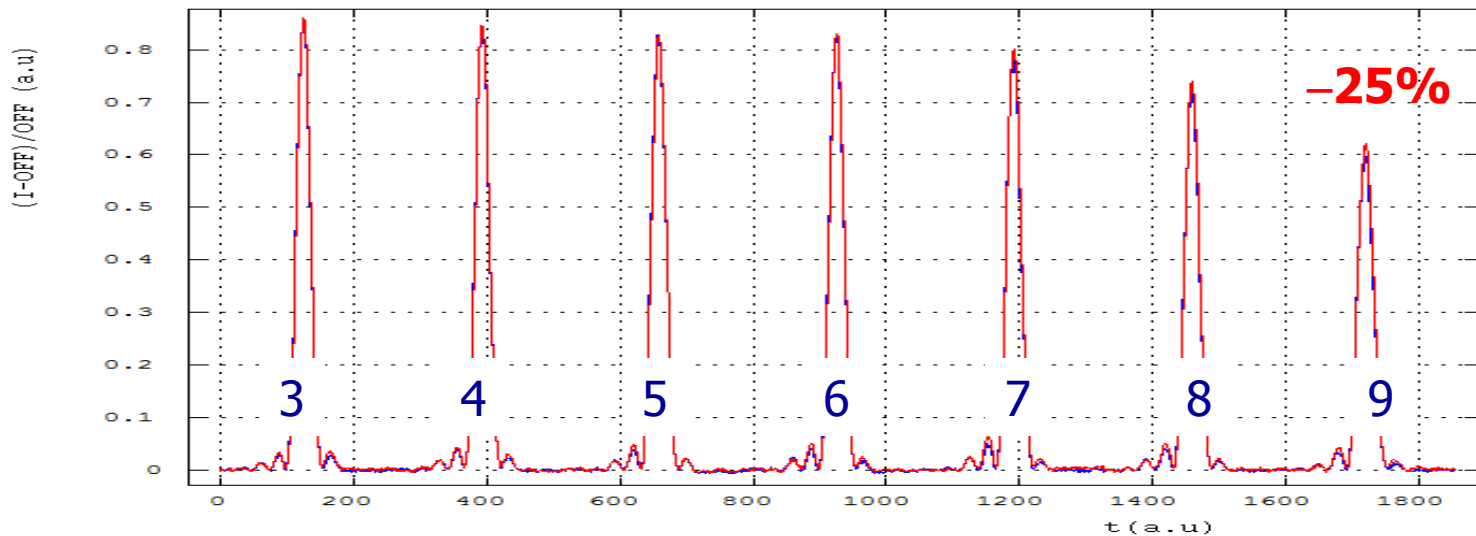
- 2 runs 2011-12-09
 - 3C161 : 0:57:21 start
 - 3 cycles/4 available 2-4
 - 3C161B : 1:11:50
 - 7 cycles/9 available 3-9
- 1 cycle = 170 sec ON-like
- Add the images time-freq. and use the first 30sec as "0"
- No use of DAB

Image time-Freq.

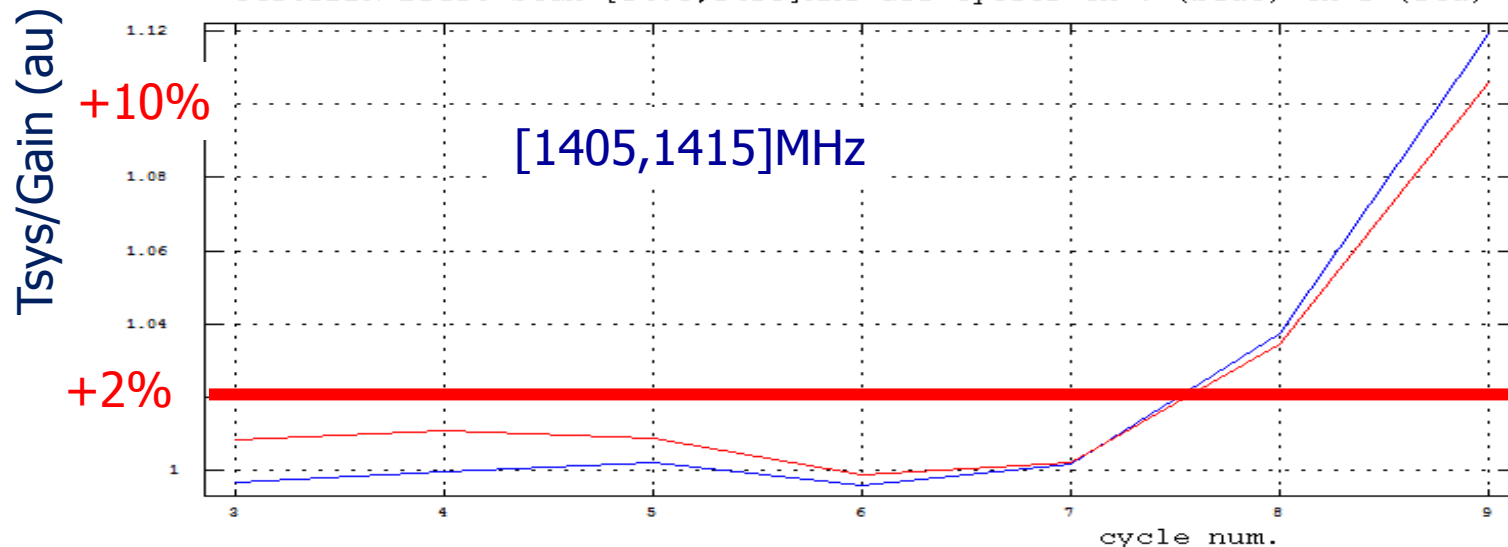


Peak int. & Tsys variations: 3C161B

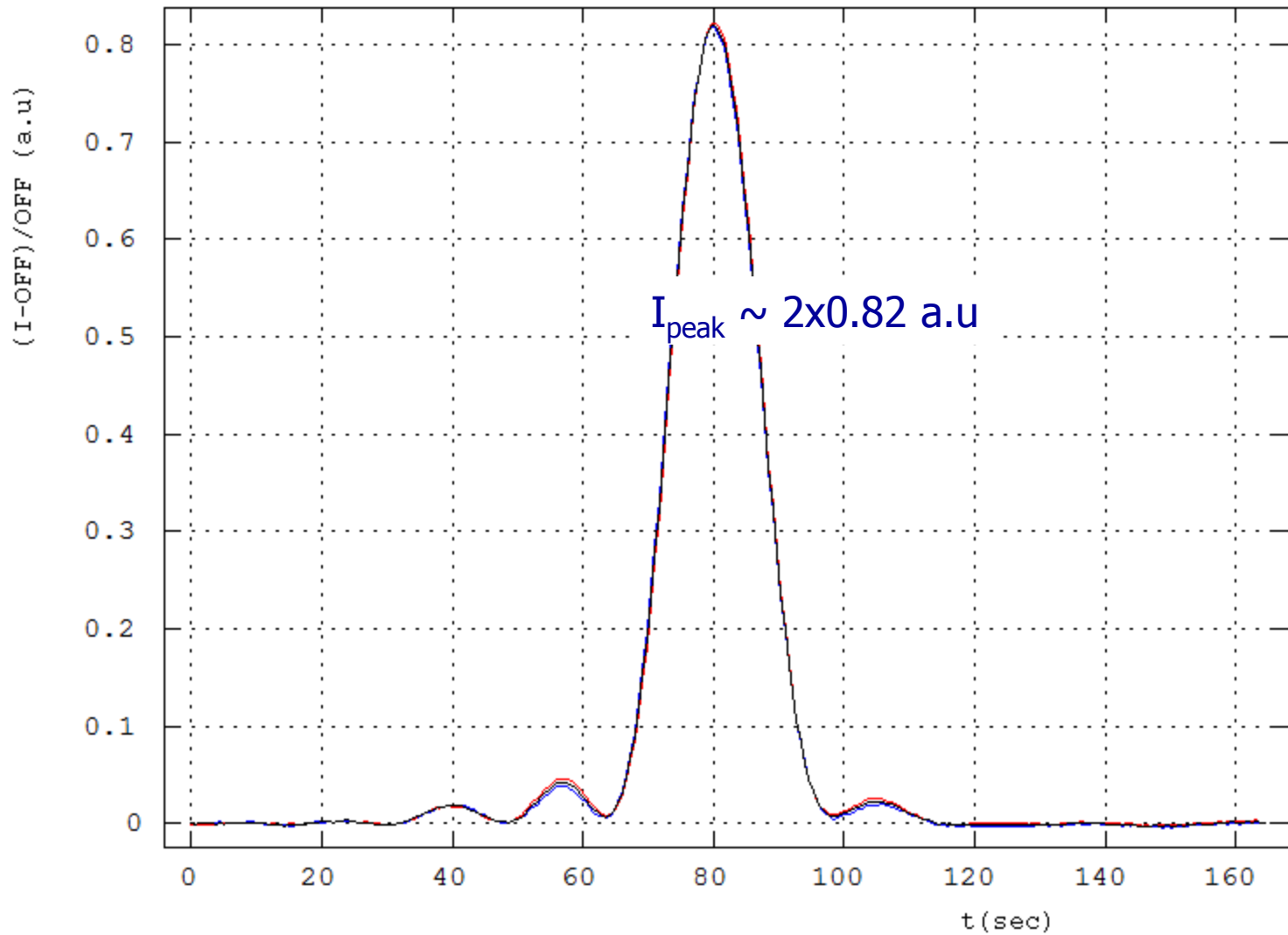
3C161BDR Drift Scan [1405,1415]MHz all cycles Ch 0 (blue) Ch 1 (red)



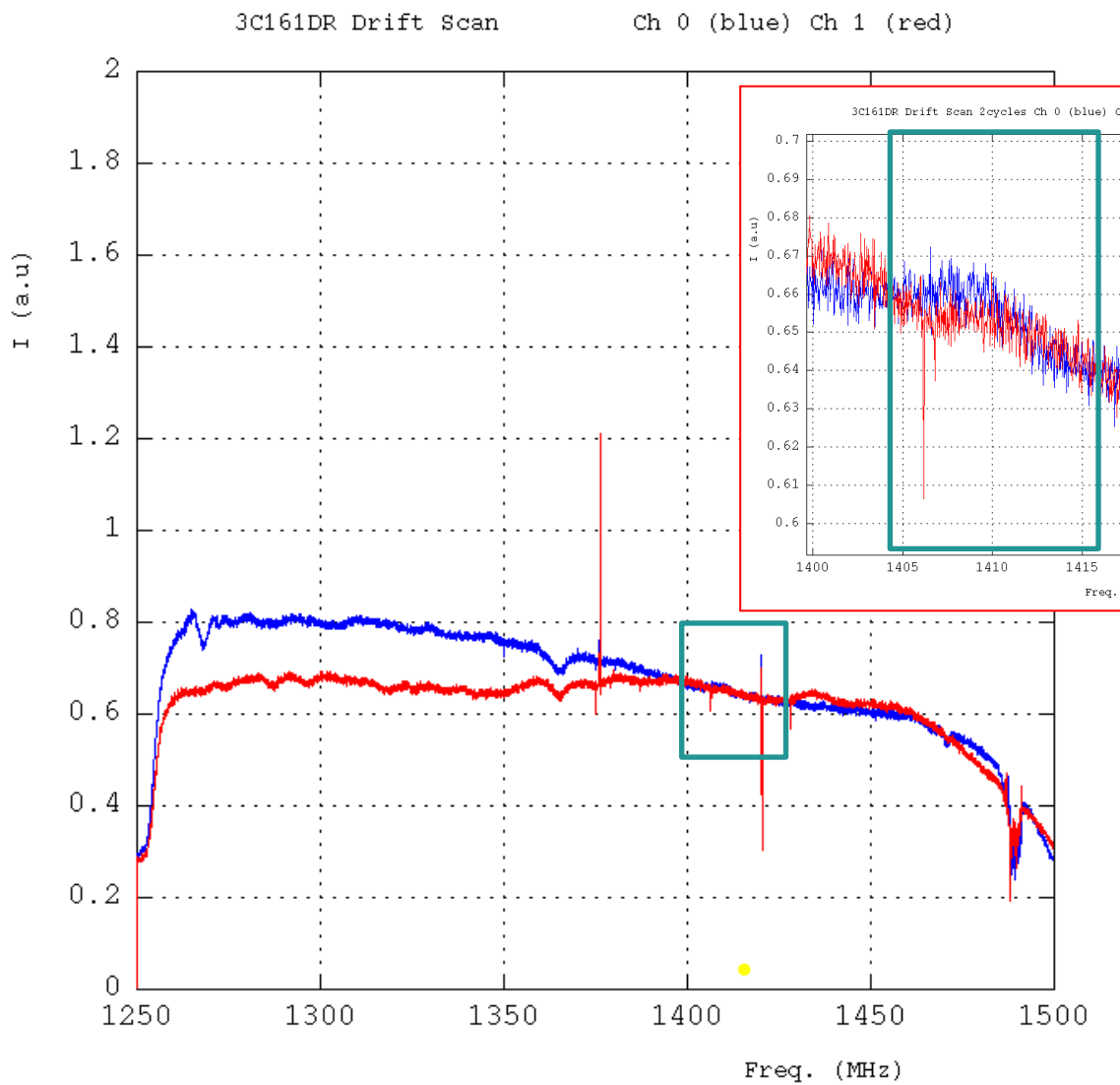
3C161BDR Drift Scan [1405,1415]MHz all cycles Ch 0 (blue) Ch 1 (red)



3C161BDR Drift Scan [1405,1415]MHz cycles [3,7] Ch 0 (blue) Ch 1 (red) Mean (black)



Timing: 8.22kHz



Spectra integrated during the FWHM of the drift.

Differences wrt the 3C273 spectra Polarization?

Calibration

$$(I_{max}^{Pol1} + I_{max}^{Pol2})(a.u) \times C_{mean\ polar} = I_{tot}(Jy)$$

3C161 0624-05 (Baars et al 1977)

Freq MHz	1250	1280	1380	1410
Flux Jy	20,636	20,294	19,233	18,937

Ott et al. , 1994

Freq. MHz	Flux Jy	date Obs.
1408	18.58 (0.09)	Feb. 1990

1410MHz

$$2 \times 0.819 \times C = 18.58$$

$$11.3 \pm 0.3 \text{ Jy/a.u (*)}$$

3% from litt. Meas.

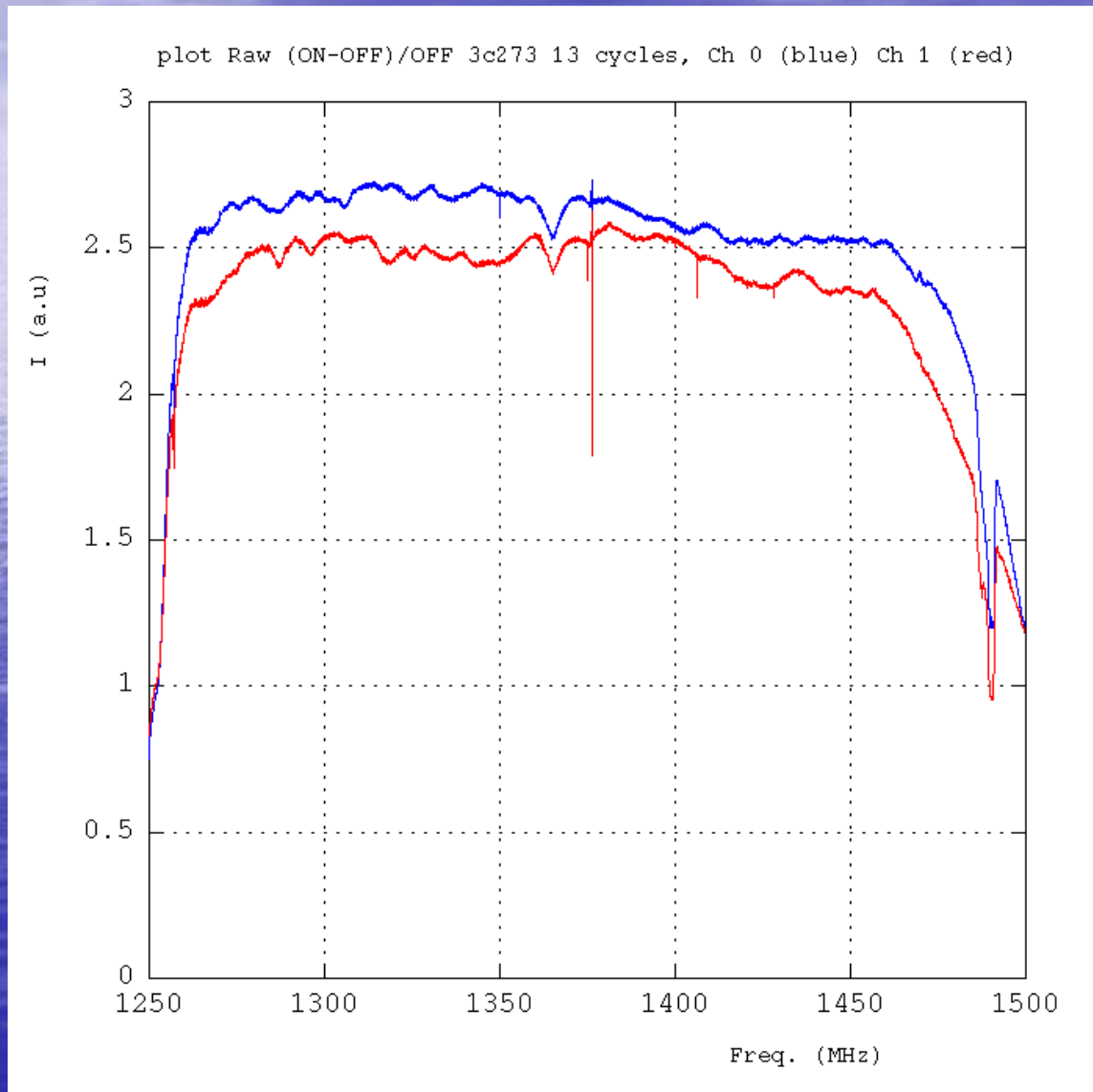
3C161

(*): be careful this is to be applied to the SUM of the 2 polarizations, not the MEAN

3C273 – 2011/12/09
ON/OFF & Drift

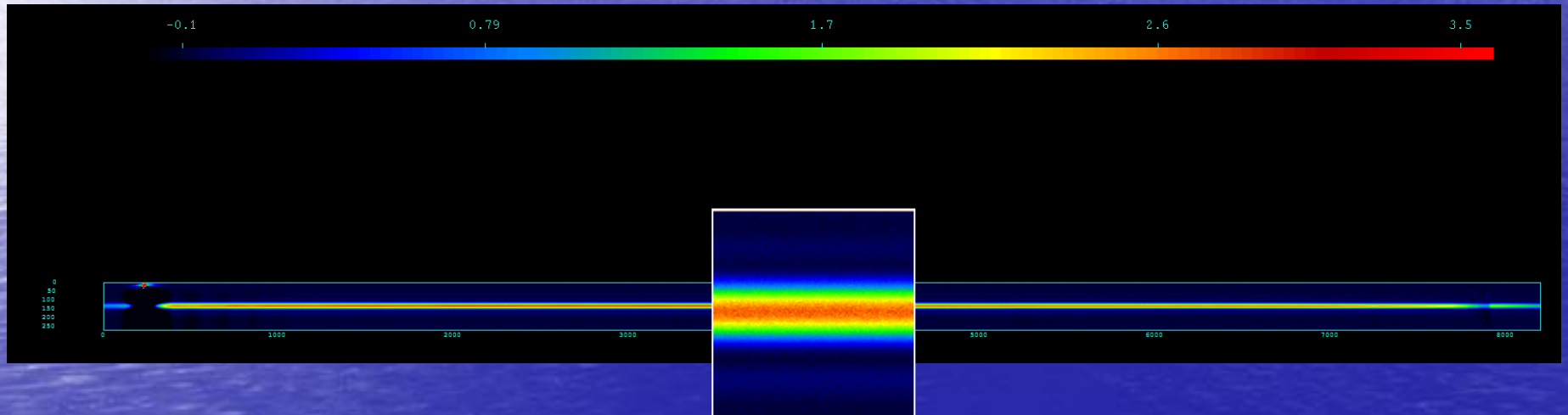
158461.171 & 158462.171

(ON-OFF)/OFF-filtered

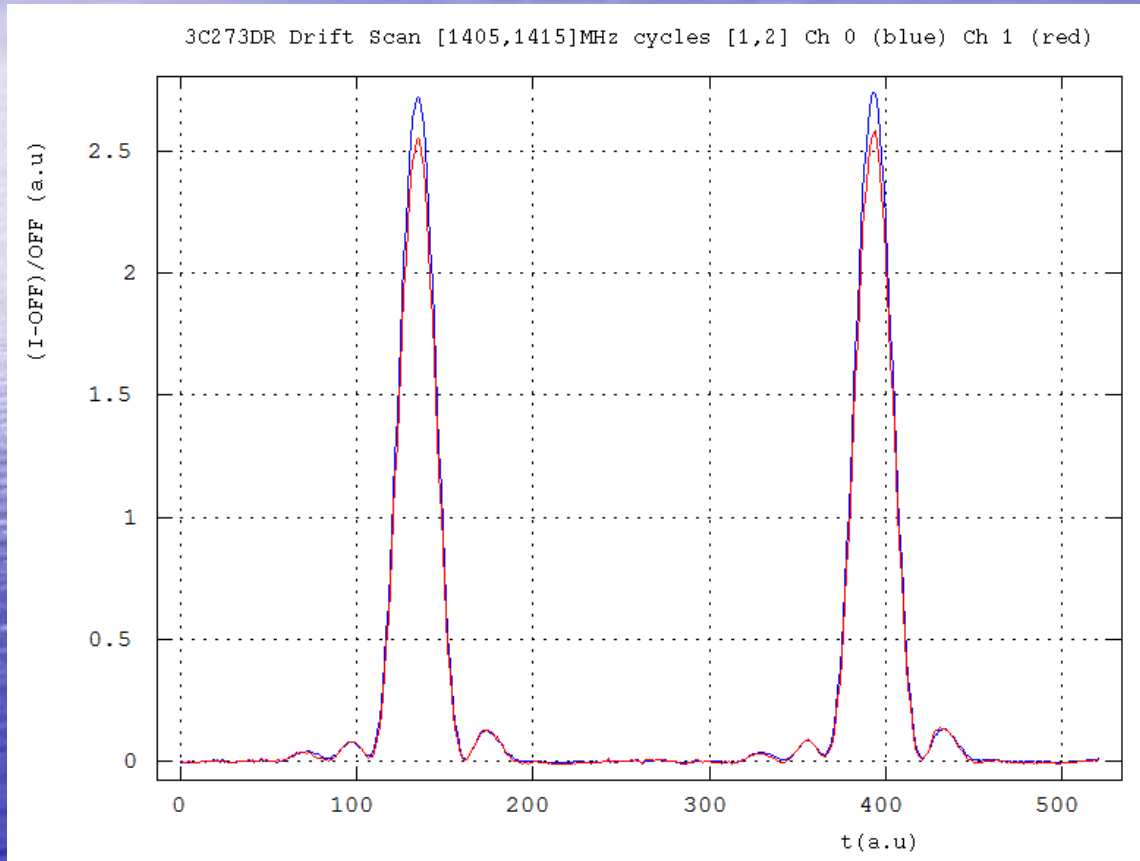


13 cycles 30sec
each (30% eff.)

Drift Scan: $I(\text{time, frequency})$



Drift Scan: time evolution

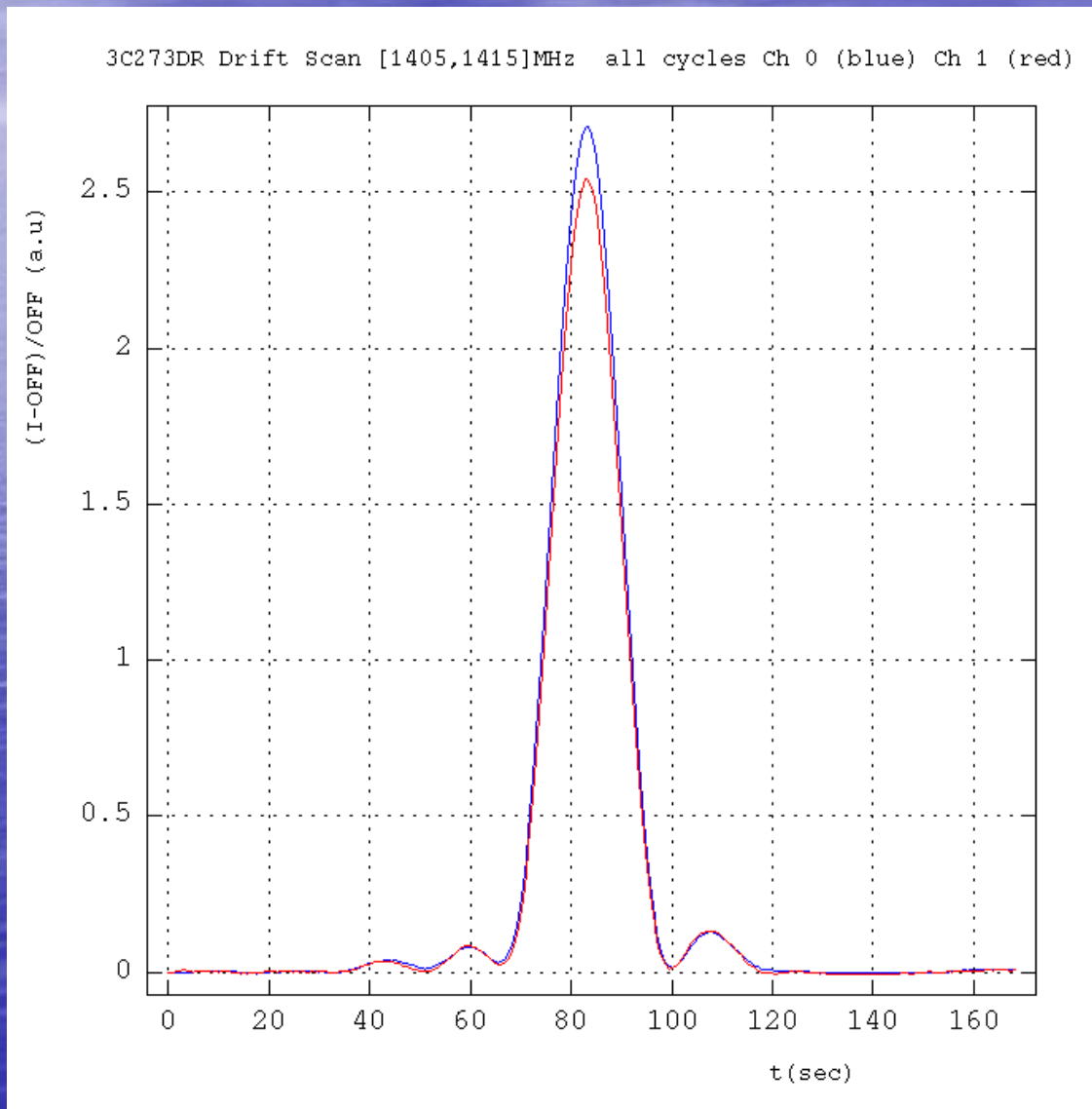


2 cycles 170sec (*)
(30% eff.)

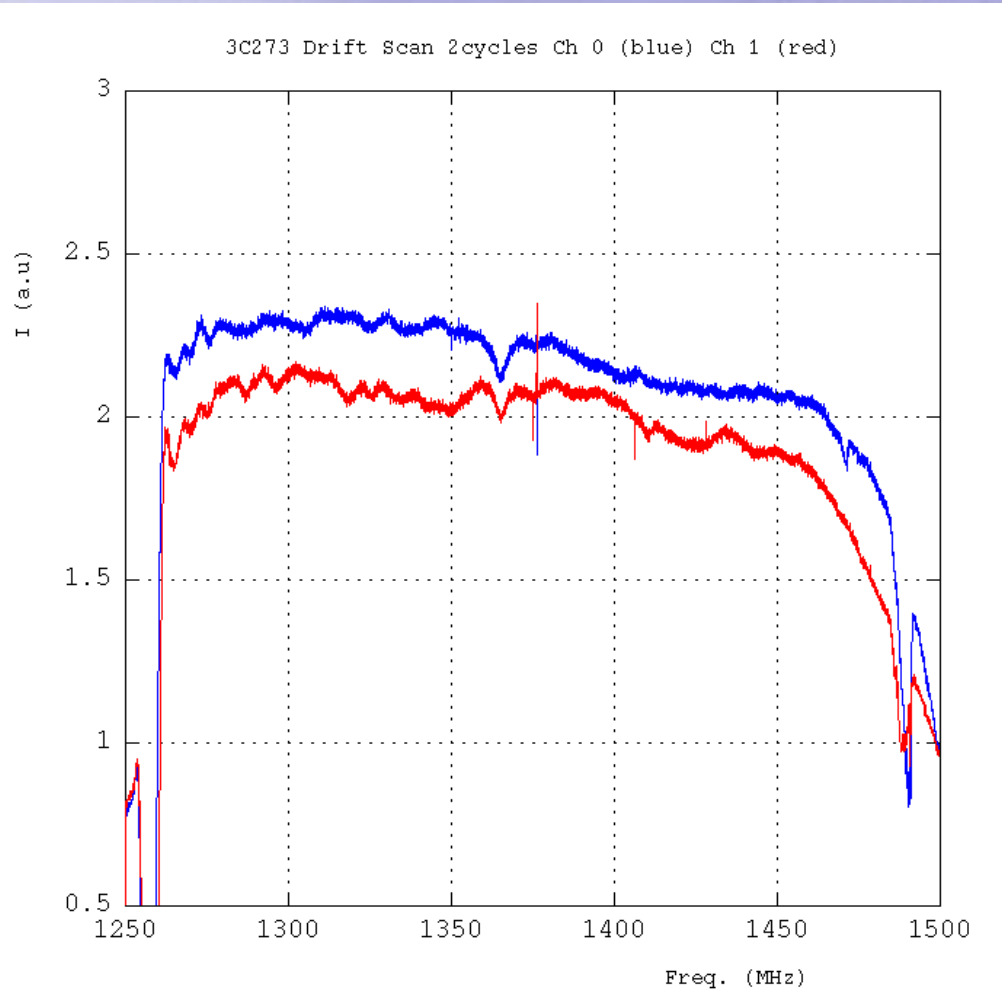
"OFF" done with the
first 30sec

Mean of the band
[1405,1415]MHz

Drift Scan: time evolution



Drift Scan: Frequency



Mean over the time FWHM
(absolute value not to be
compared to ON-OFF)

Same behavior as ON-OFF

3C273

-Bridle et al., 1972: I(1400 MHz) : 38.84 (0.70) Jy

- Kühr et al., 1981 & NED

Freq.(MHz) Flux (Jy)

1400	41.28	1.23	Witzel et al. 1971
1400	46.30	2.30	Kellerman et al. 1969
1410	45.17	1.07	Wills 1975
1410 MHz	42.00	...	Jy -> Wright et al. 1990, Parkes
1400 MHz	45.0	+/- 5 %	Jy -> PAULINY-TOTH et al., 1966
1400 MHz	39.62 ± 0.38		Jy -> idem
1.4 GHz (ATCA)	35.82	...	Jy -> Tingay et al., 2003, PASJ
1.40 GHz	50100	...	milliJy -> White et al., 1992, 300ft, -> Condon et al. 1985,1986
1.4GHz	54992.1 ± 1900.3		milliJy -> Condon et al. 1998, NVSS

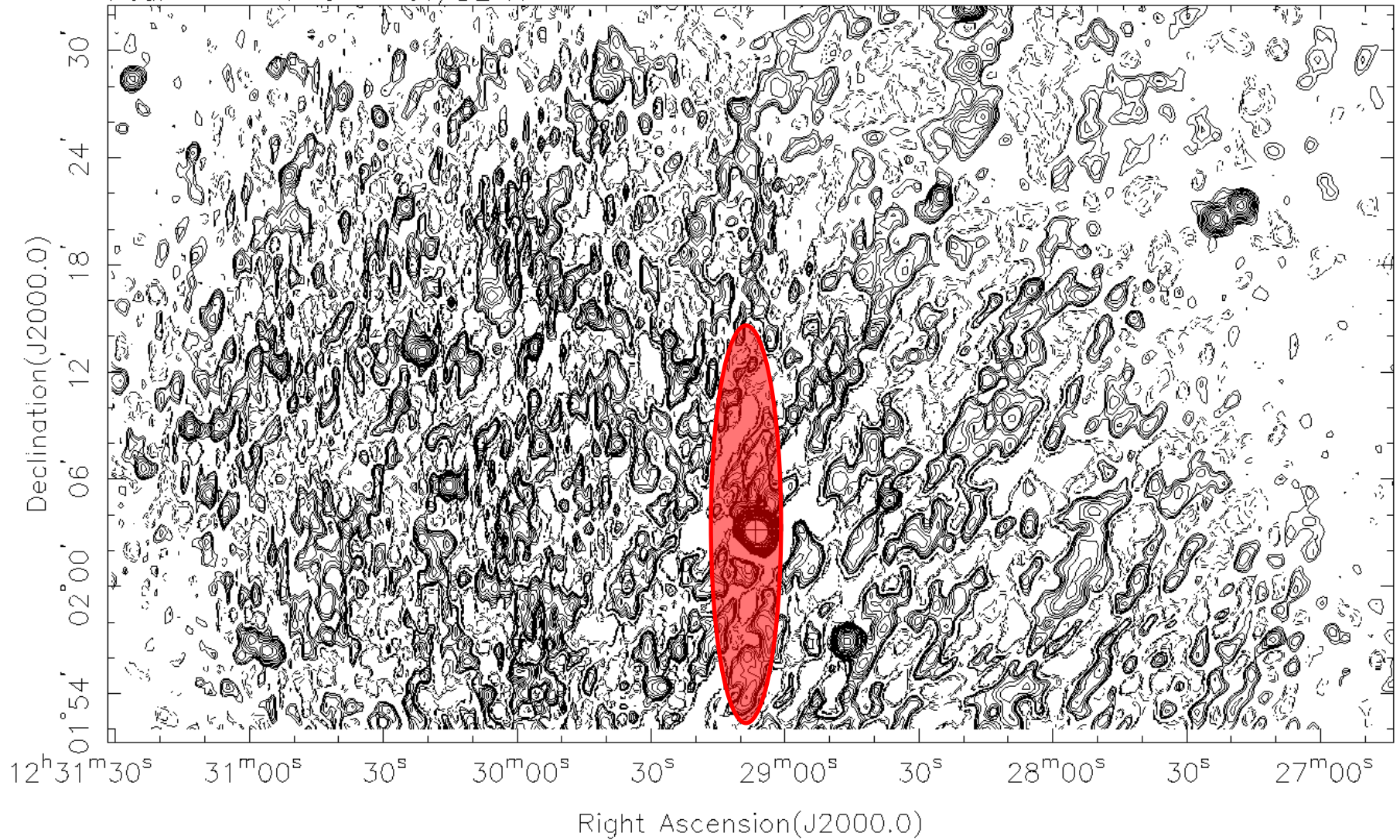
Our measurement : $\sim 2 \times 2.62 \text{ au} \times 11.3 \text{ Jy/au} \sim 59.2 \text{ Jy}$

3C161



NVSS: 3C273 (levs= $\pm 1, 1.4, 2, 2.8, 4 \dots$ mJy/b)

Peak = 47.0 JY/BEAM



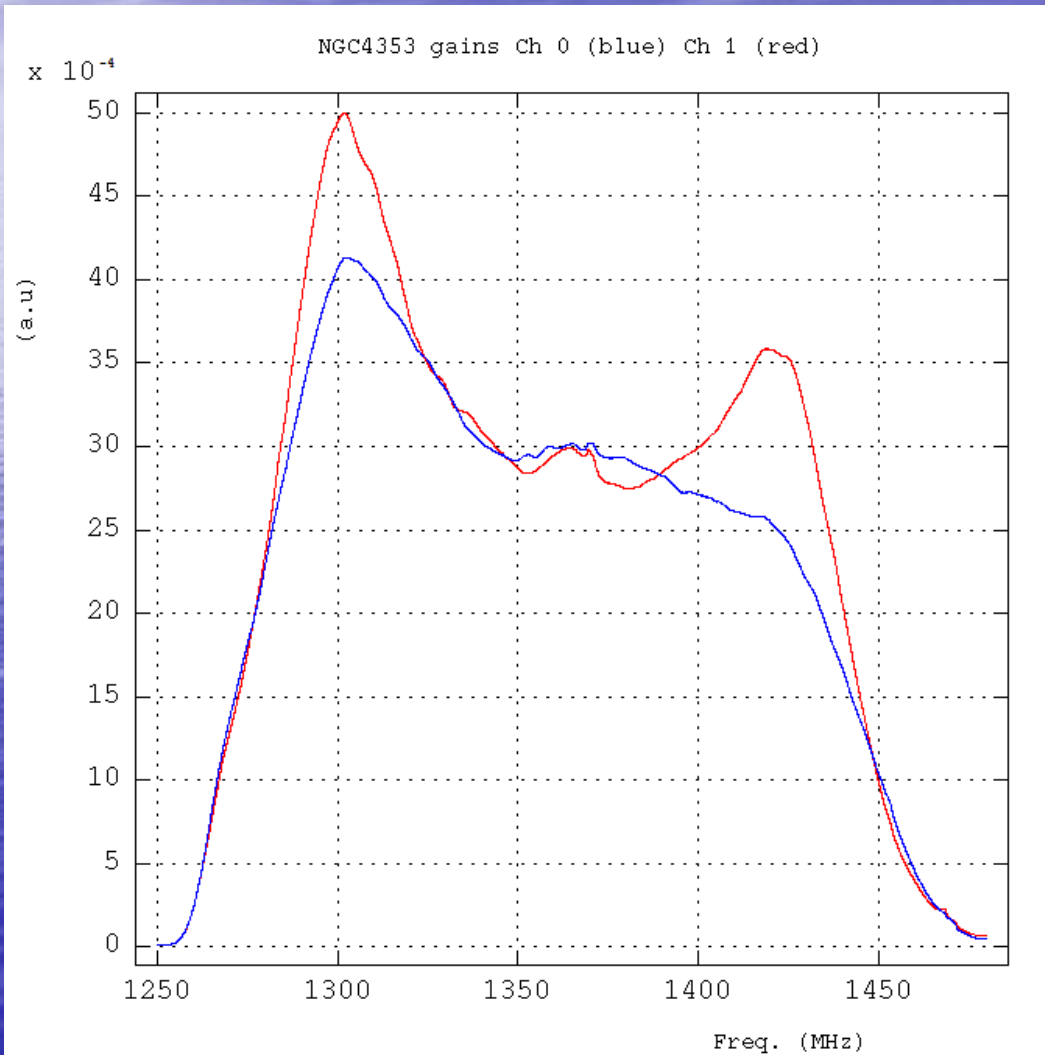
NGC4383 ON-OFF

SCA156855.171

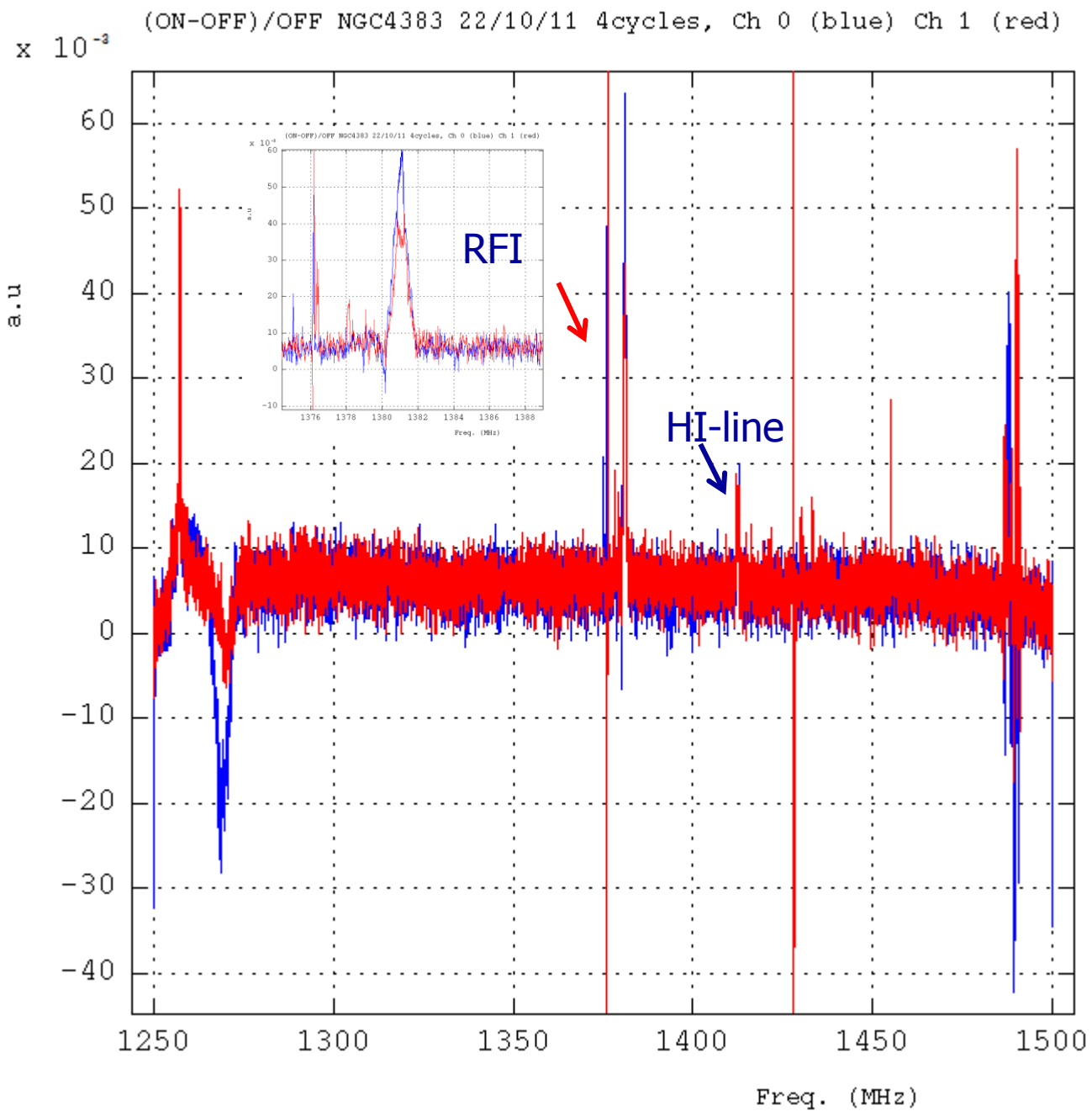
Date 2011-10-22

Cycles 1-15 foreseen but 12-15 only !

Gains cycle #13



Gains different shape as I was expecting. But confirmed by Ana that they have changed since 7th July...



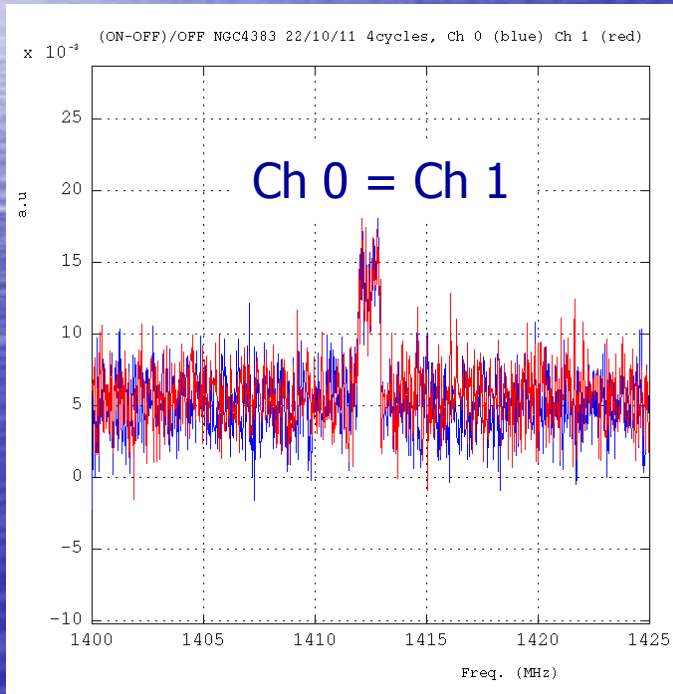
Continuum & line

Continuum: $\sim 44.3 \pm 4.1$ mJy

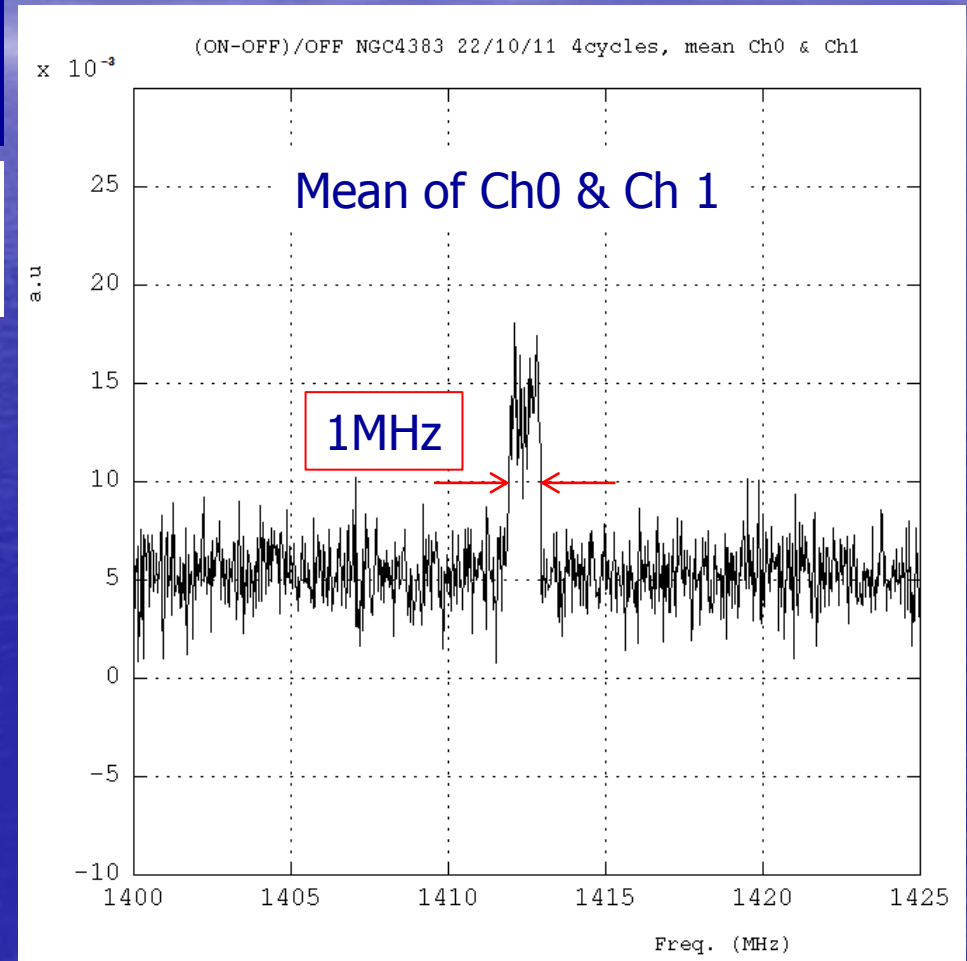
HI Line:

- width ~ 213 km/s ~ 1 MHz
- int. 48.4 ± 5.1 Jy km/s ~ 230 mJy

A. Chung et al, VLA IMAGING OF VIRGO SPIRALS IN ATOMIC GAS (VIVA). I., The Astronomical Journal, 138:1741–1816, 2009 December
THE ATLAS AND THE HI PROPERTIES



Zoom of HI line



$$1420.20/(1+0.0057) = 1412.15 \text{ MHz} = 1700 \text{ km/s} \quad \text{OK}$$

The HI line intensity

Integral in the range [1412-1413] MHz

$$\begin{aligned}\int HI(\nu)d\nu &= 2_{polar} \times \left(\sum_i (S_i - C_{ont}i_i) \right) \Delta\nu \\ &= 2 \times (447 - 175) 10^{-3} [a.u.] \times 30 10^{-3} [MHz] \times 213 [km/s/MHz] \times C_{mean polar} \\ &= 3.48 [a.u.km/s] \times C_{mean polar} \\ &= 48.4 \pm 5.1 [Jy.km/s]\end{aligned}$$

Then

$$C_{mean polar} = 13.9 \pm 1.5 [Jy/a.u.]$$

To be compared to 11.3 Jy/au

Or

$$3.48 \times 11.3 = 39.3 [Jy.km/s]$$

To be compared to 48.4 Jy.km/s

3C161



NGC4383 Drift Scan

SCA157756.171

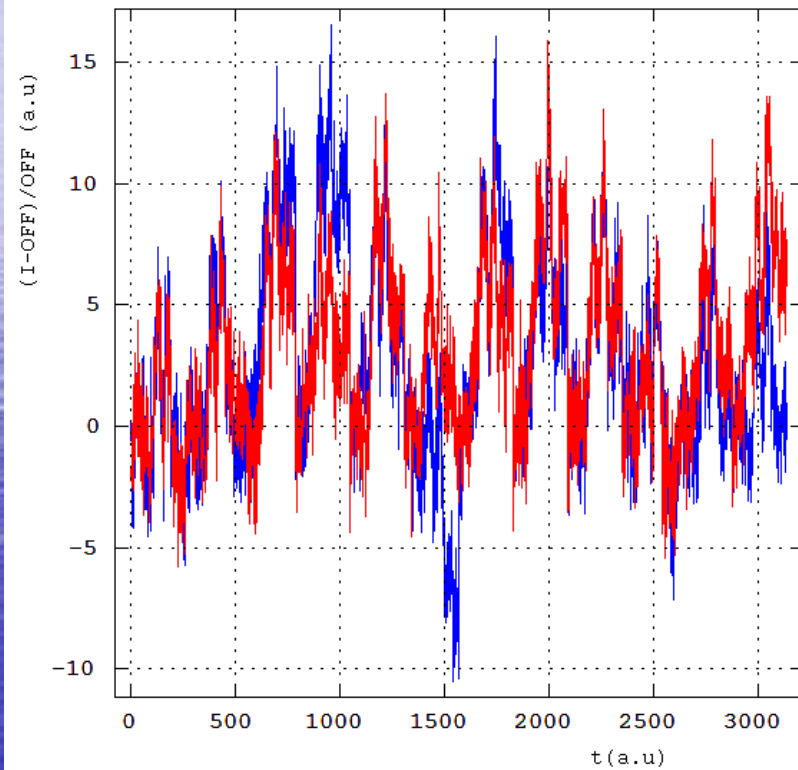
Date 2011-11-19

Cycles 2-15

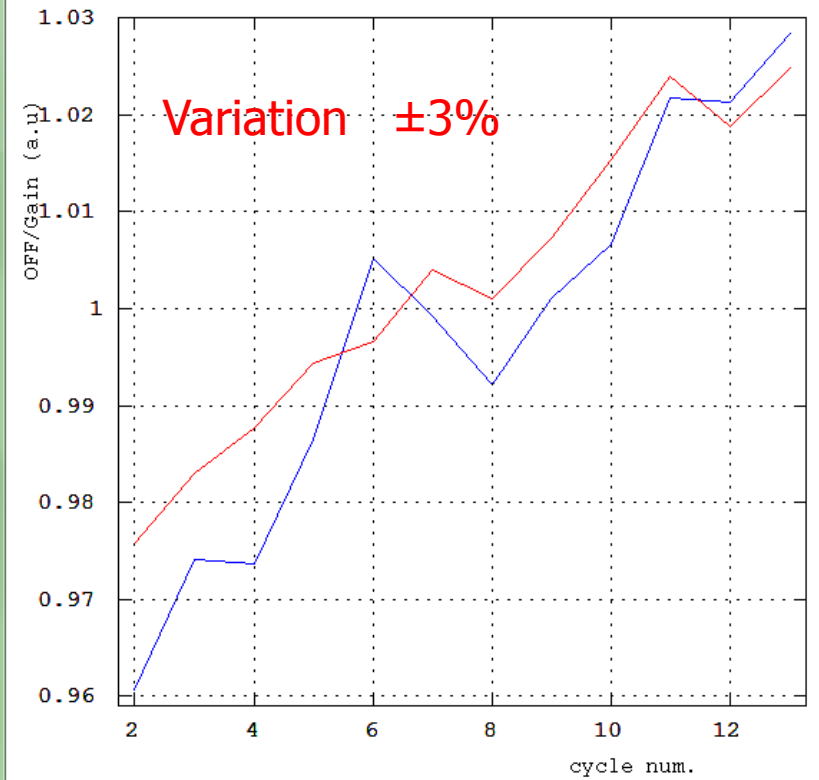
(I-OFF)/OFF

OFF/Gain

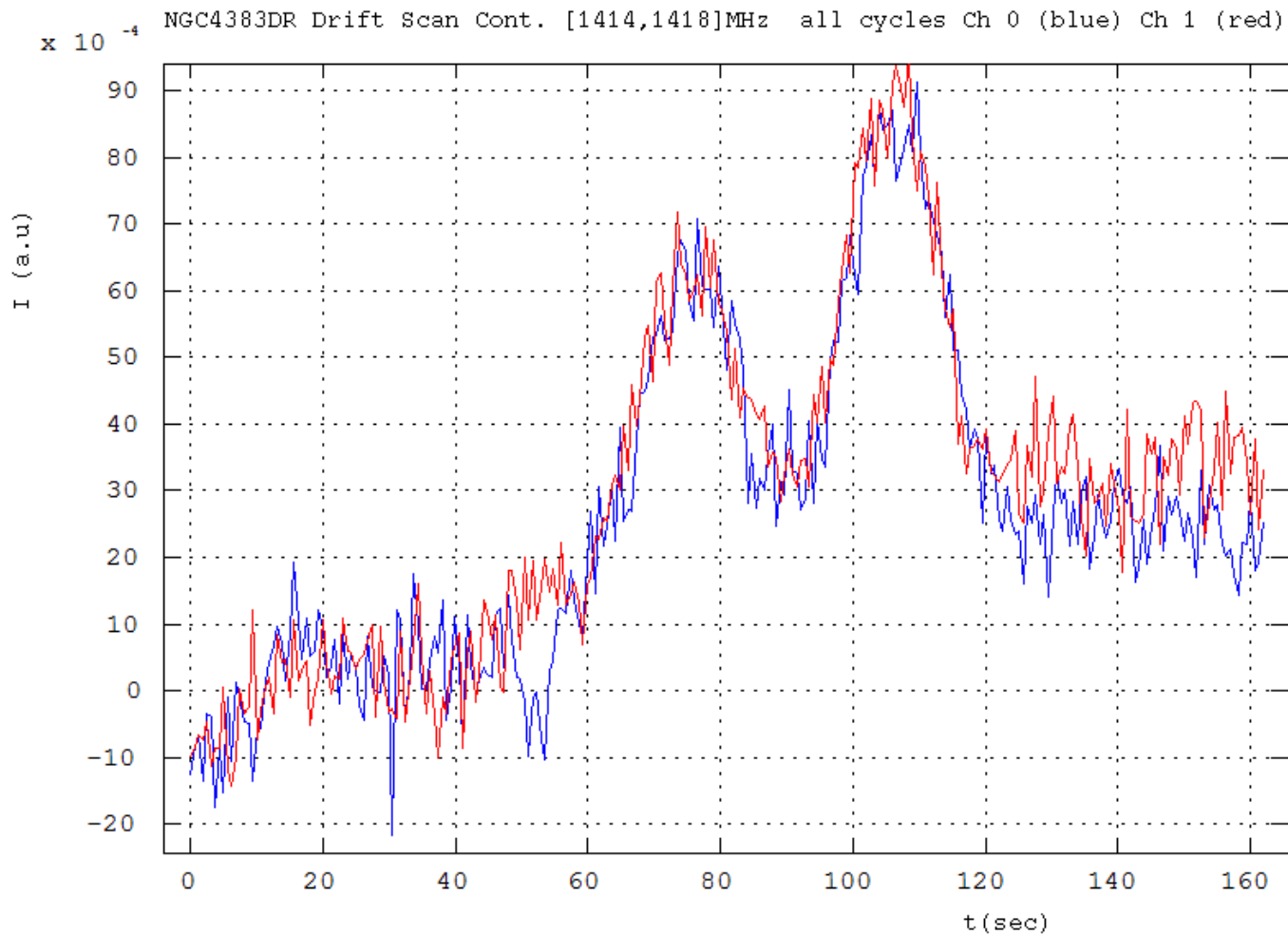
NGC4383DR Drift Scan [1405,1415]MHz all cycles Ch 0 (blue) Ch 1 (red)
 $\times 10^{-3}$



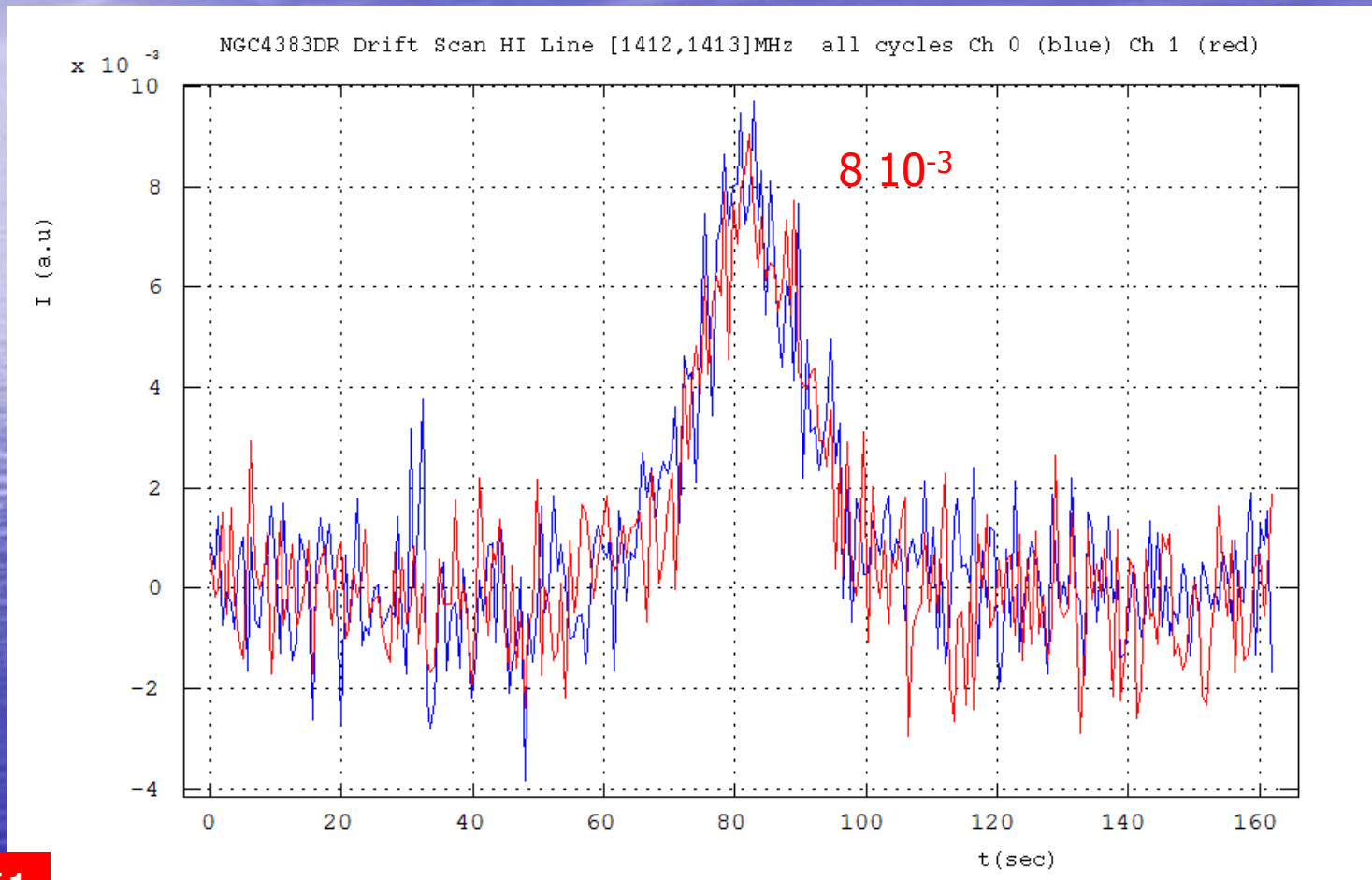
NGC4383DR Drift Scan [1405,1415]MHz all cycles Ch 0 (blue) Ch 1 (red)



Continuum evolution



HI Line evolution



3C161

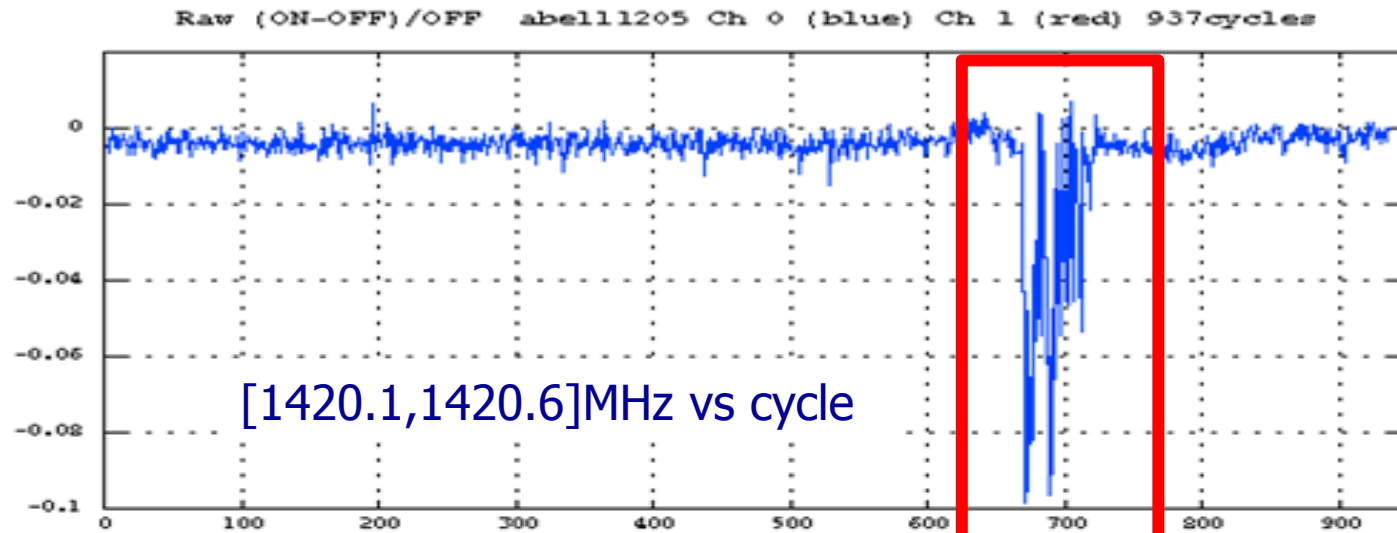
$11.3 \text{ Jy/au} \times 2 \times 8 \times 10^{-3} \text{ (au)} = (181 \pm 5) \text{ mJy}$ to be compared to $(230 \pm 24 \text{ m}) \text{ Jy}$

Summary

- 3C161 @ 1410MHz gives 11.3 ± 0.3 Jy/au
- Overestimation of 3C273 continuum
- Underestimation of HI NGC4383
- Close to Ch 1 DAB (Est) calibrated thanks to PKS1127-14 WIBAR DAQ.
- How to calibrate another freq. although the BAO spectra are quite stable in freq.

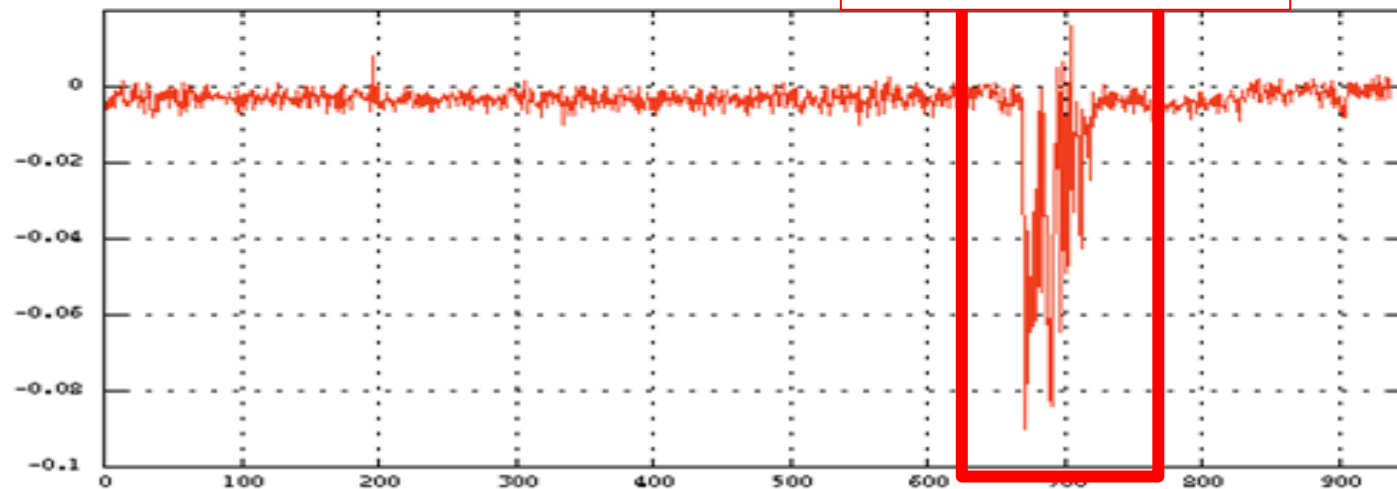
Abell85, 2440, 1205
zoom on HI

Abell 1205: DAQ stability

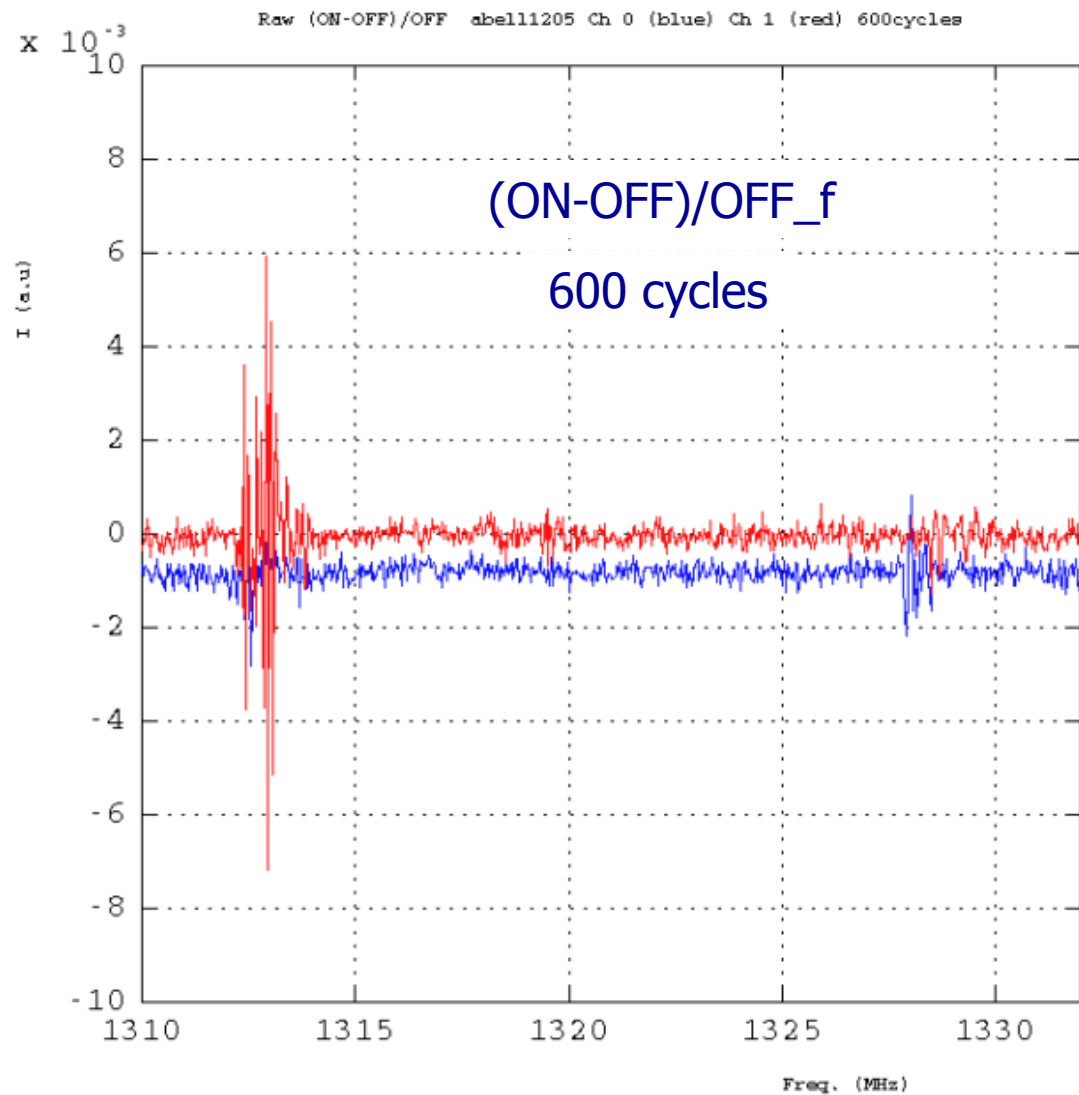


22/08/11-22/09/11

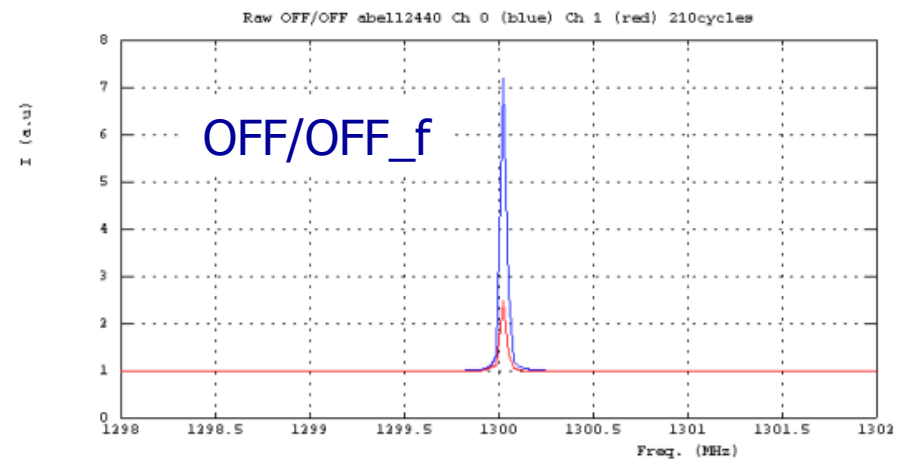
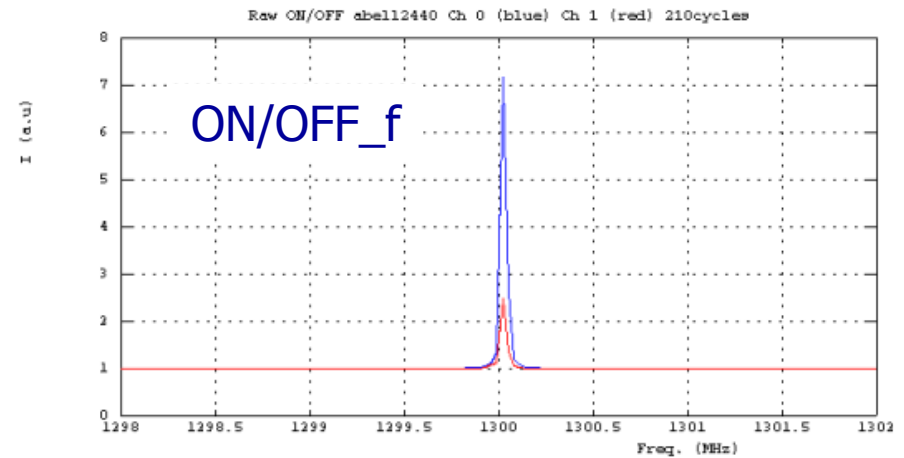
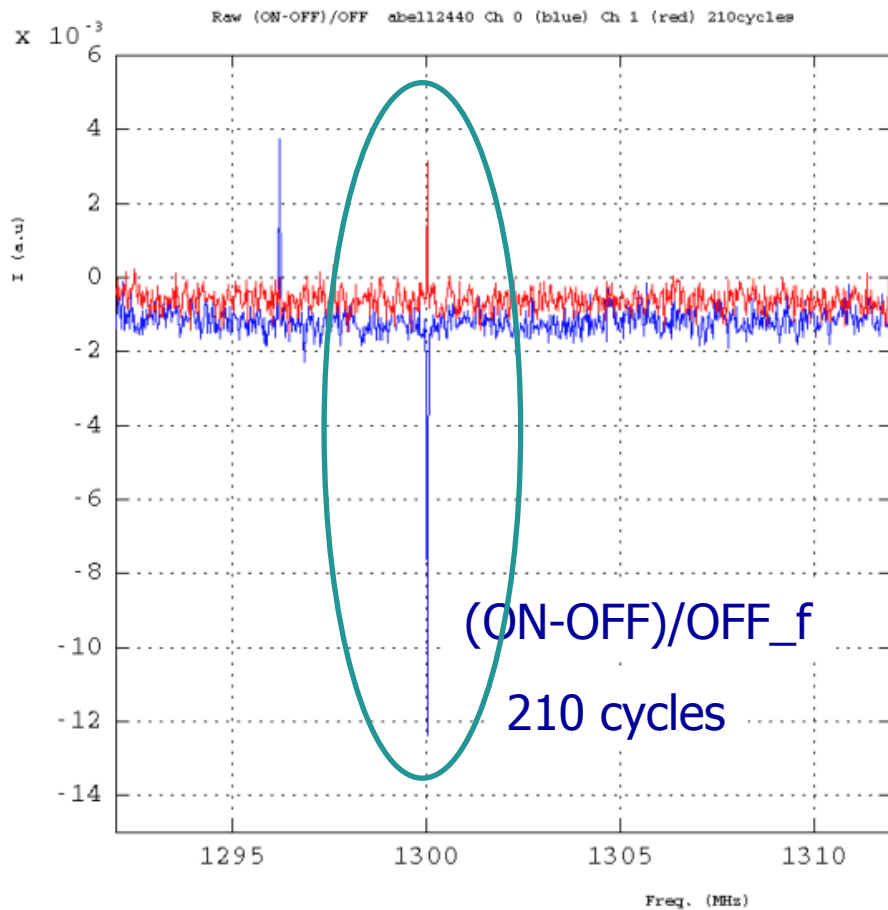
I[1420.1,1420.6]MHz Raw (a.



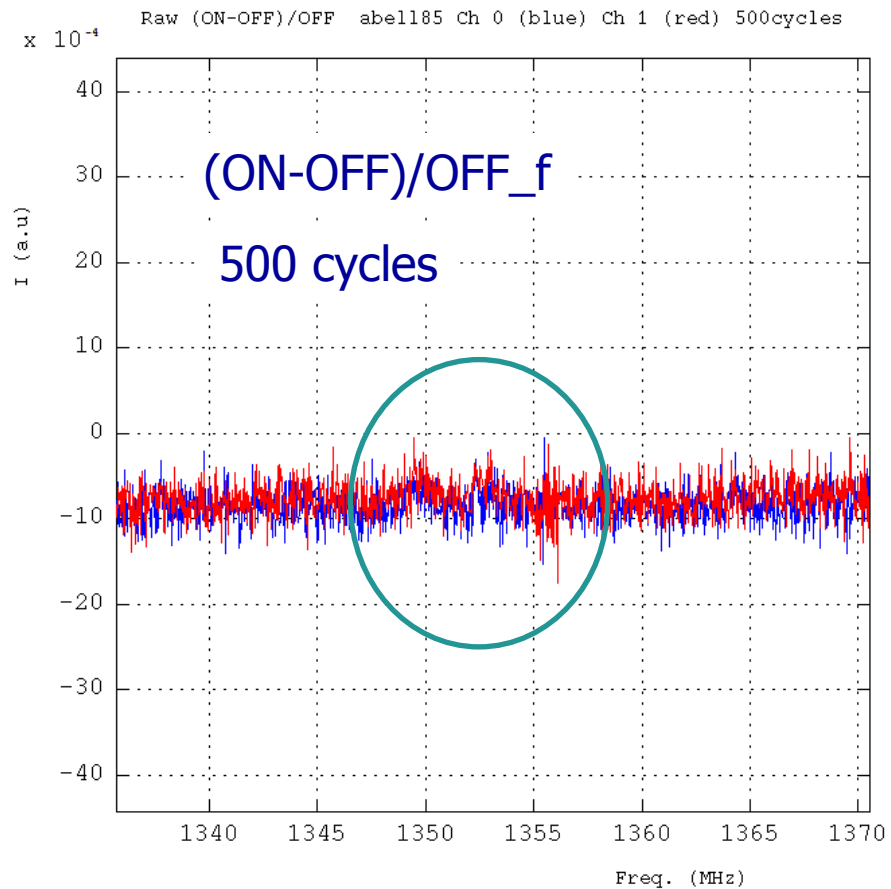
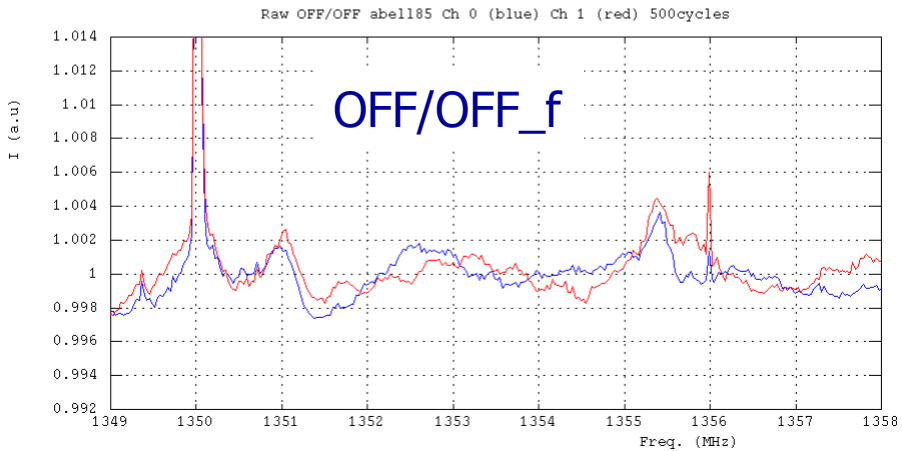
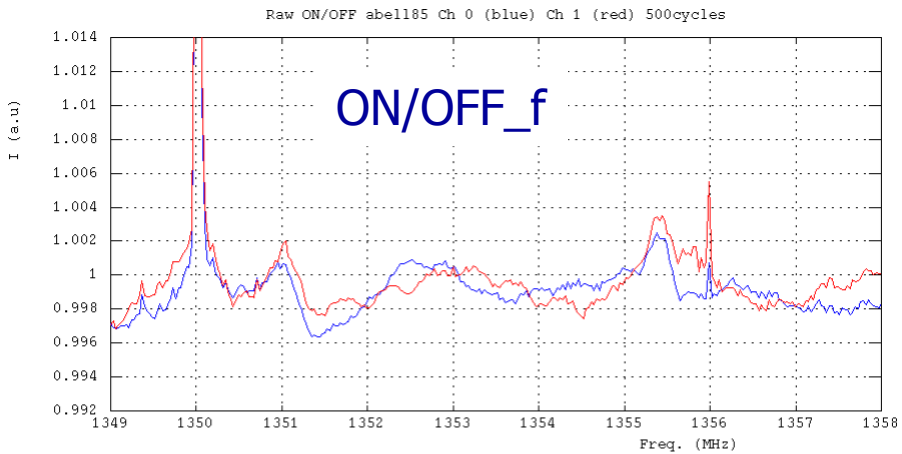
Abell1205



Abell2440



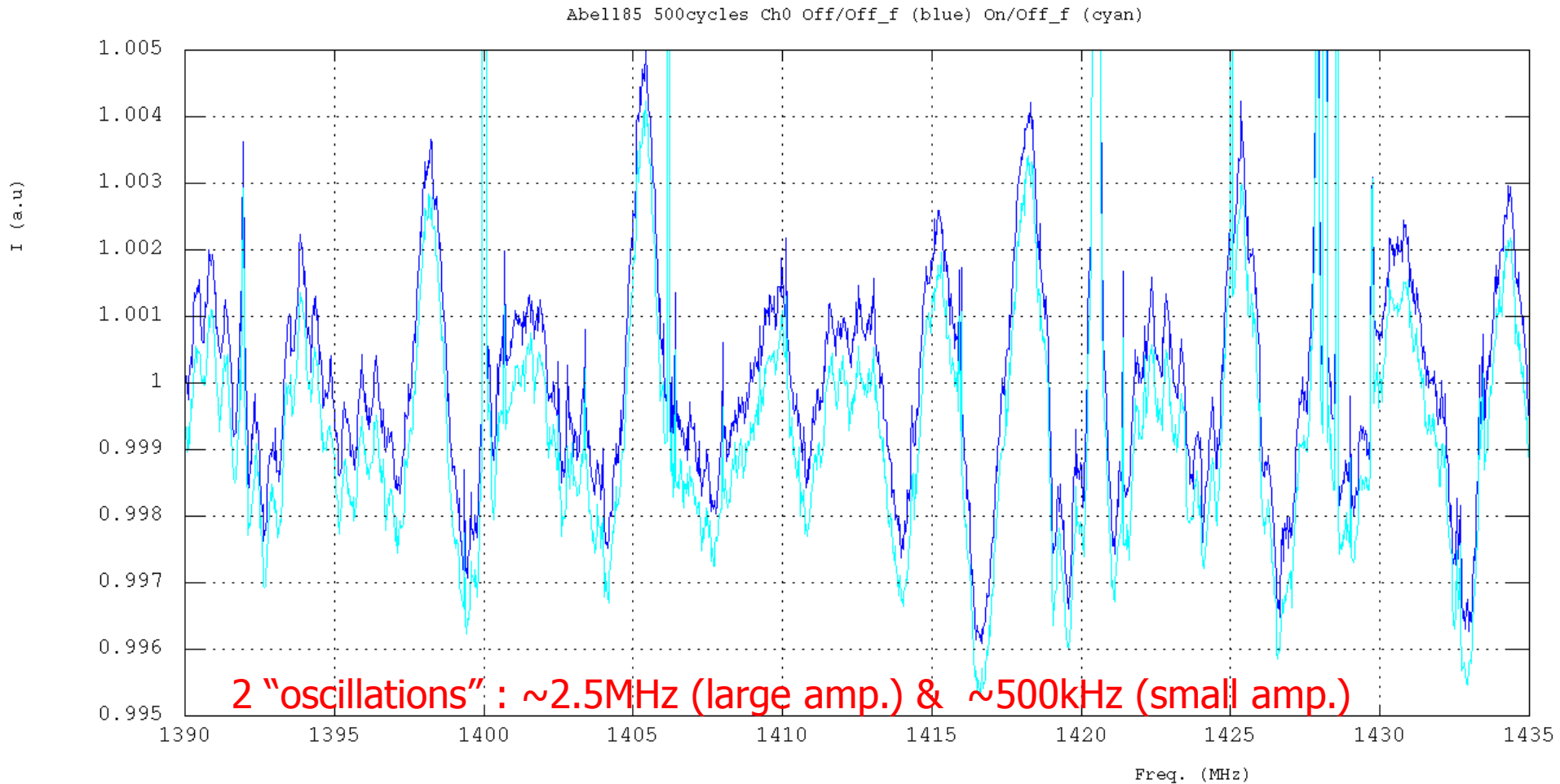
Abell85



Noise: introduction

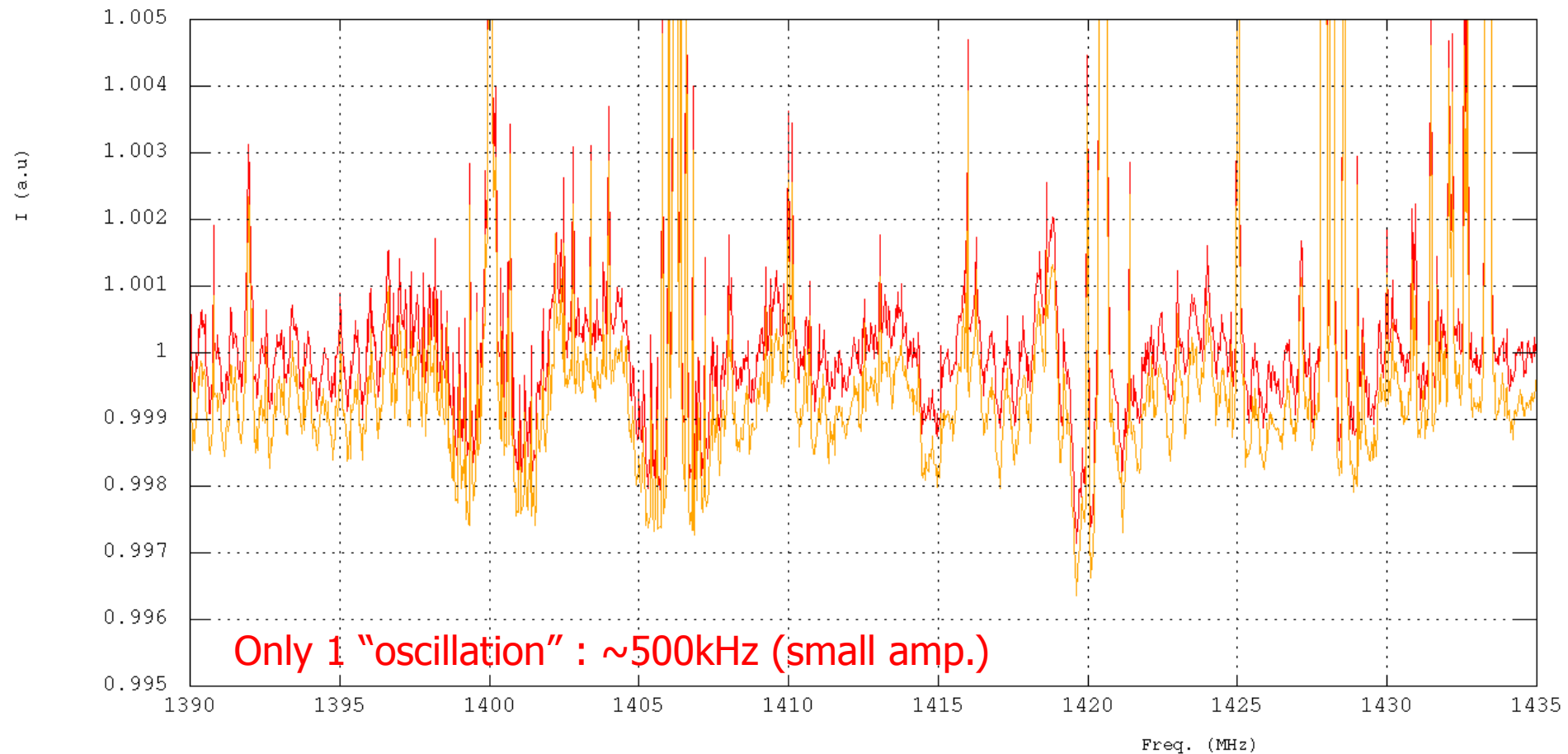
Zoom of the Abell85 normalized spectra obtained with 500cycles

Ch. 0 (On & Off)/Off_filtered



Ch. 1 (On & Off)/Off_filtered

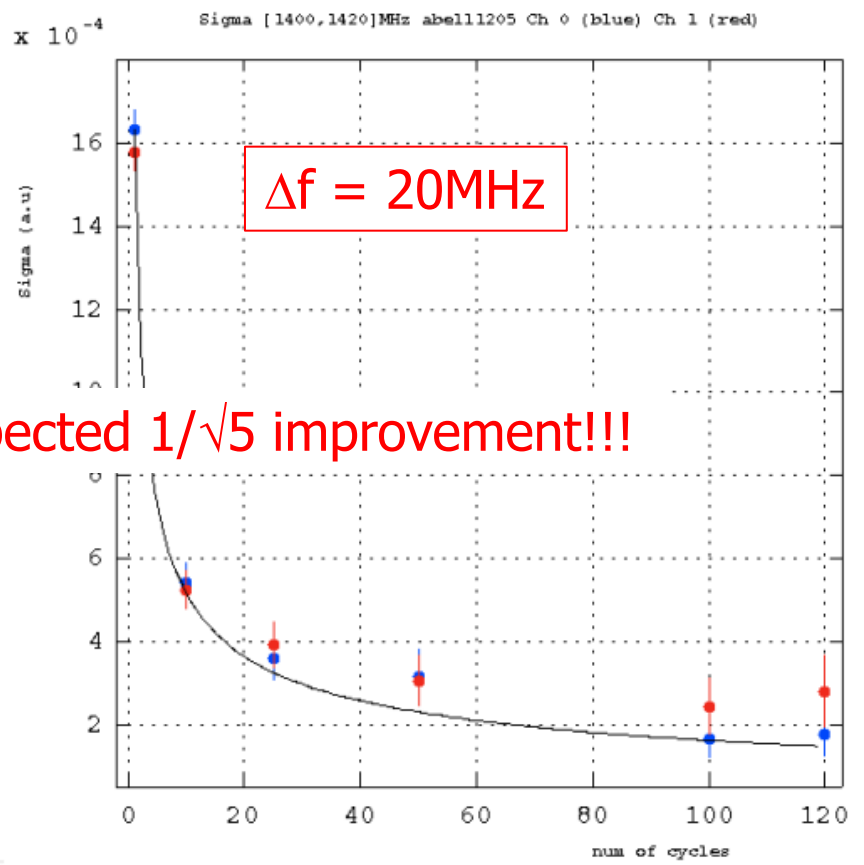
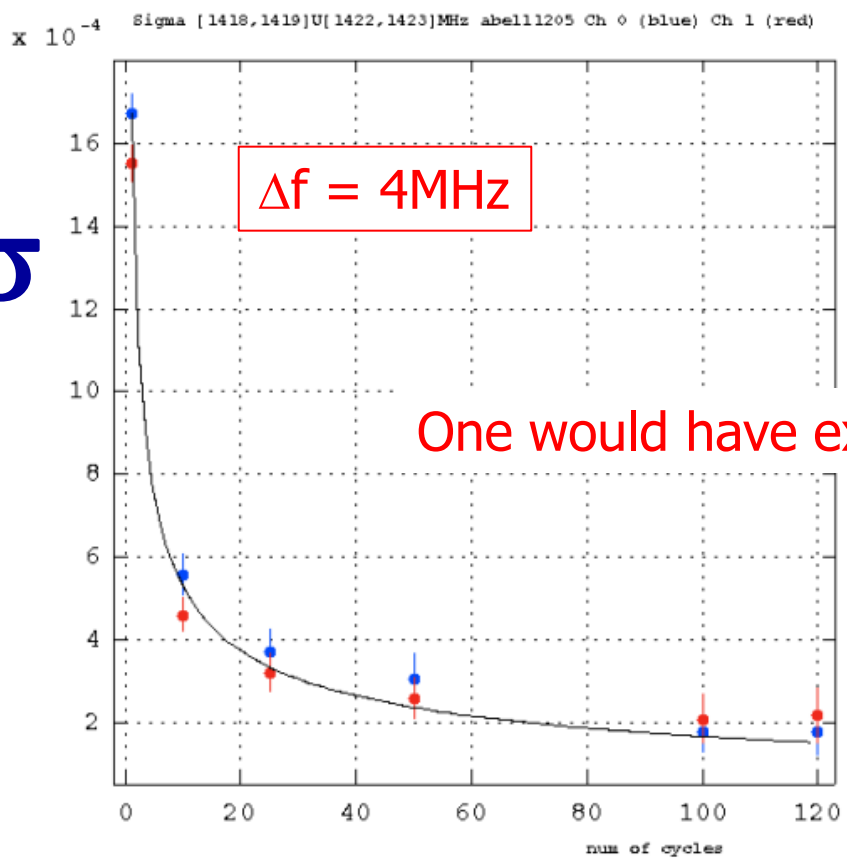
Abell185 500cycles Ch1 Off/Off_f (red) On/Off_f (orange)



Some facts

- The large oscillations are visible with only few data taking $< 1\text{min}$.
- The small ones need summation so normalized spectra are the gain change...
- The noises are present on all data: Abell, 3C, NGC.
- The 500kHz noise was in fact present since Jul. 2010 in the UGC4358 on non normalized spectra for both channels.

9



One would have expected $1/\sqrt{5}$ improvement!!!

Figure 9. Baseline standard deviation as a function of number of cycles per mean value (see text), for the ON-OFF signal integrated in the HI-line side-bands ([1418,1419] \cup [1422,1423] MHz) (left) and the RFI-protected band [1400,1420] MHz.

Data processing on Irods @ CCIN2P3

1. Median ov. 5120 BAO paq. ON/Gain or OFF/Gain
2. Mean of 5 medians => med. spectrum (FITS files)
3. Mean on each cycle
4. For each cycle $(ON-OFF)/OFF_f$
 1. Result of $(ON/G-OFF/G)/(OFF/G)_f$
 2. $G=Gain$ med. Filtered 3MHz
 3. $(OFF/G)_f = (OFF/Gain)$ med. Filtered 2MHz
5. For each cycle too $< (ON-OFF)/OFF_f >$ ov. $[v_1, v_2]$

Sigma

Estimation

1. $\langle \text{ON/Gain} \rangle(v) \sim \langle \text{OFF/Gain} \rangle(v) \sim 1$ (no signal & def.)
2. $\sigma[\text{OFF/OFF}_f] \ll \sigma[\text{ON/OFF}_f]$ (at least it is a fact)
3. $\sigma_{\text{med}} = \text{med} / (\ln 2 \sqrt{N_{\text{paq}}})$ and $\text{med}[\text{ON/Gain}] \sim \text{med}[\text{OFF/Gain}] \sim 1$

$$\sigma \approx \frac{\sqrt{2}}{\ln 2} \times \frac{1}{\sqrt{N_{\text{med/medspec}} \times N_{\text{medspec/cycle}} \times N_{\text{pad/med}} \times \Delta f}} \quad (\text{a.u.})$$

1 cycle

5

~9

5120

$$\sigma(1 \text{ cycle}) \propto 1/\sqrt{\Delta f} \quad \text{and} \quad \sigma(n \text{ cycles}) \propto \sigma(1)/\sqrt{n}$$

Evolution of sigma computation

- The sigma may be affected by the noise (see Ana's talk)
- To avoid artifact we use now the inter quartile range IQR normalized

$$\sigma \equiv \frac{Q_3 - Q_1}{1.34898}$$

IQR for Normal dist. with $\sigma=1$

