

HICluster/PAON2/Test

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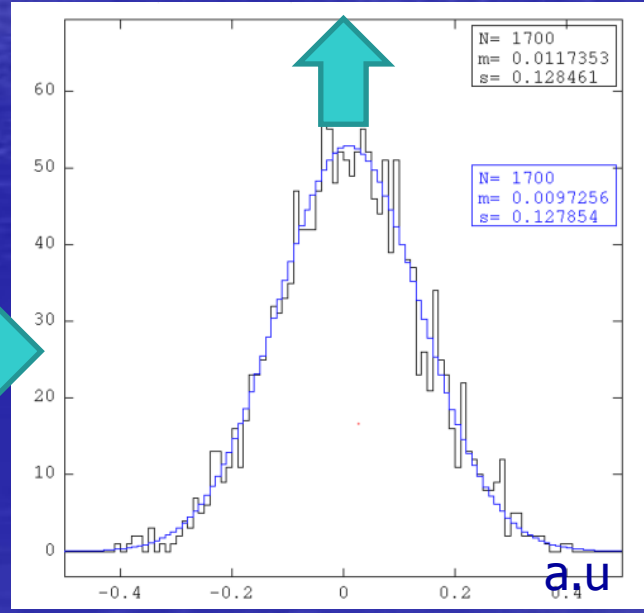
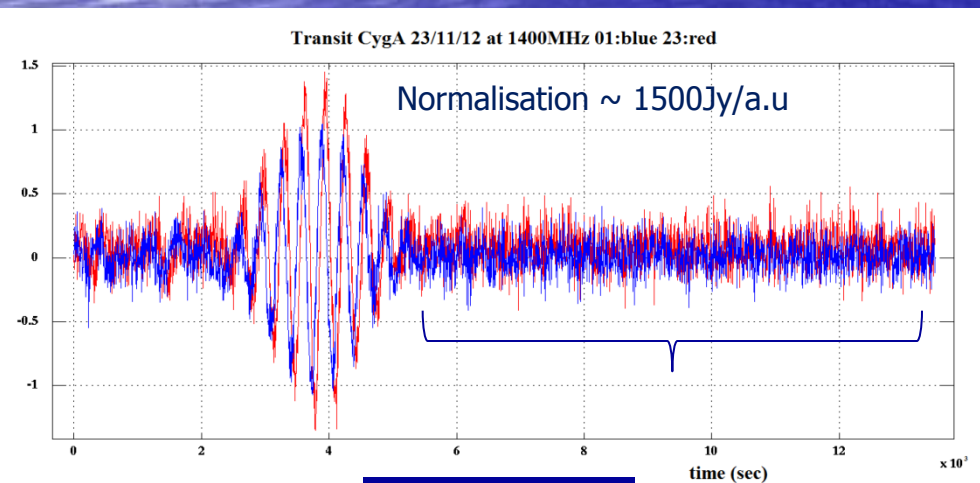
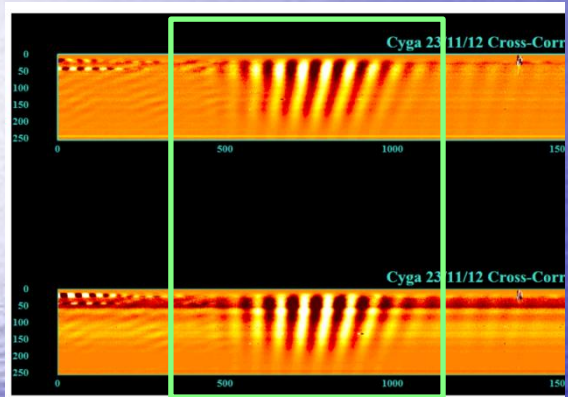
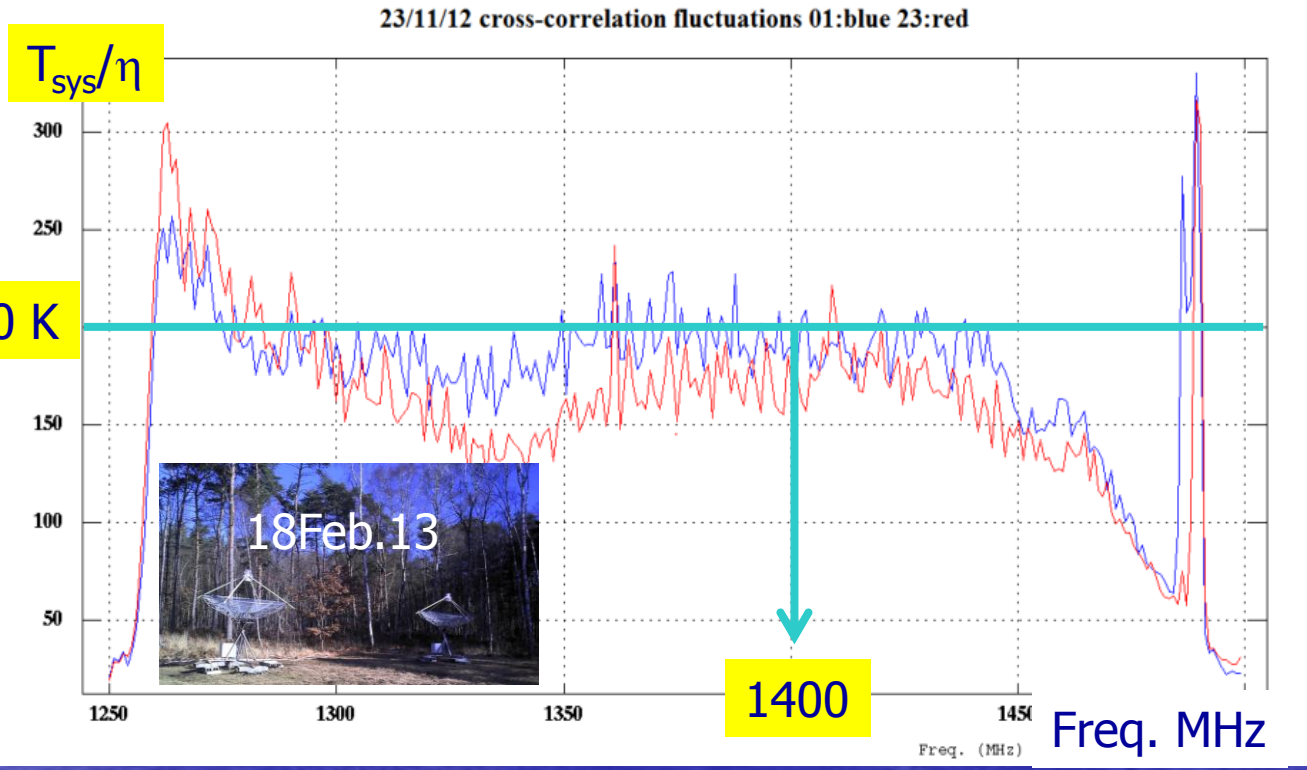
28/02/13

Tianlai video-meeting

PAON2

Cyga 23/11/12

200 K

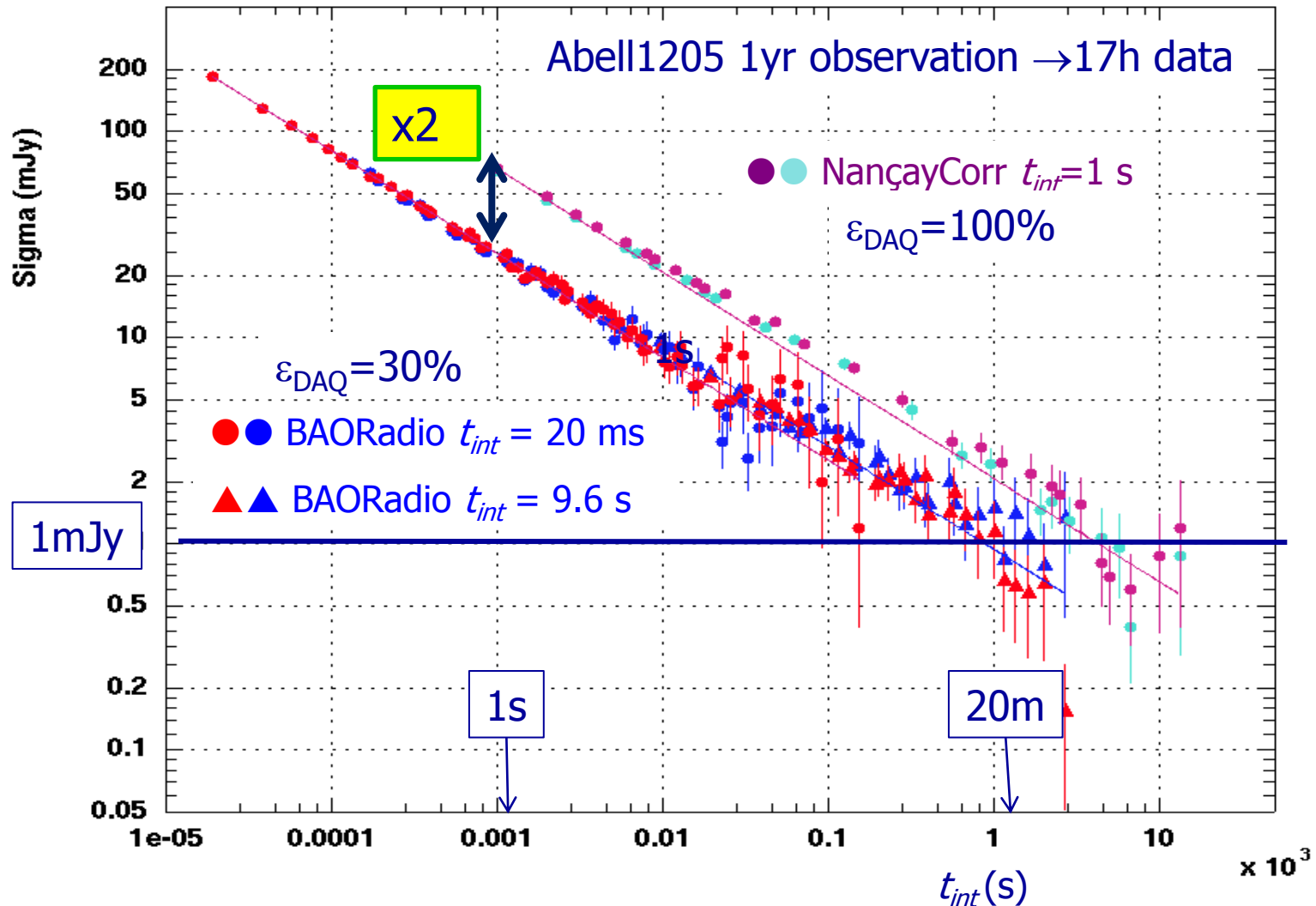


4h @ 1kHz

HI cluster: Sensitivity vs. t_{int}

A.S.T & J.EC Nançais/Amas-Corr/1.03.13

abell1205 (ON-OFF)/OFF [13203,13213]MHz abell1205 bank1,Ch0 (blue) bank2,Ch1 (red)



HI in clusters: sources

- Search for potential HI signals in nearby galaxy clusters
 - Abell 85 (HI?), Abell 1205, Abell 2440

Source	RA	Dec	ν_{HI} (MHz)	z	Observation period	Number of observations	T_{obs}
Abell 85	00 ^h 43 ^m 16.99 ^s	-9°09'46.99"	1353	0.05	April – Dec 2011	35	8.1 ^h
Abell 1205	11 ^h 15 ^m 08.37 ^s	2°33'01.39"	1316	0.08	March 2011 – January 2012	77	16.7 ^h
			1317.7	0.078			
Abell 2440	22 ^h 24 ^m 33.30 ^s	0°53'18.59"	1320	0.076	March – June 2011	15	3.5 ^h
			1292	0.099			
3C161	06 ^h 24 ^m 43.09 ^s	-5°51'14.00"	1301	0.091	9 th Dec 2012	1	6 ^m

HI in clusters: acquisition

- 2 acquisition chains:
 - BAORadio: LAL-IRFU electronics
 - NRT Correlator (NançayCorr): acquisition standard du NRT,
« black box »

	Bandwidth	Freq bin	# Polars	Minimum t_{int}	Efficiency on-sky	Output
BAORadio	[1250,1500] MHz	30 kHz	2	30 μ s	30 %	Signal vs. time (FFT offline)
NançayCorr	25 MHz including HI line	12 kHz	2	1 s	100 %	Spectrum / P_{tot}

HI cluster: HI signal in Abell85

- No HI signal has been detected

After 2h of integration (eff. Included) $\sigma \sim$

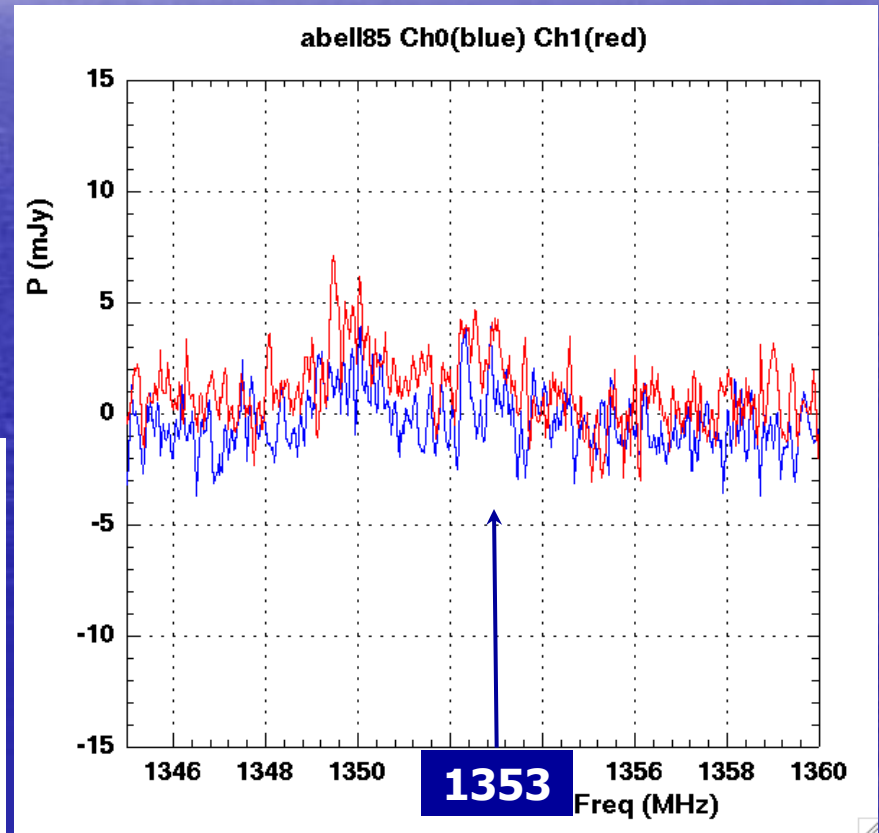
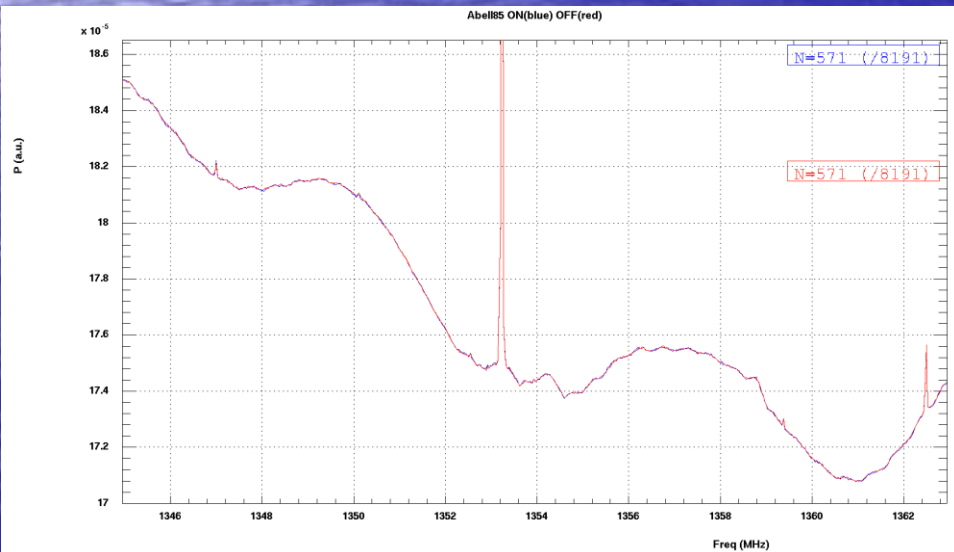
$1 \div 2 \text{ mJy}$

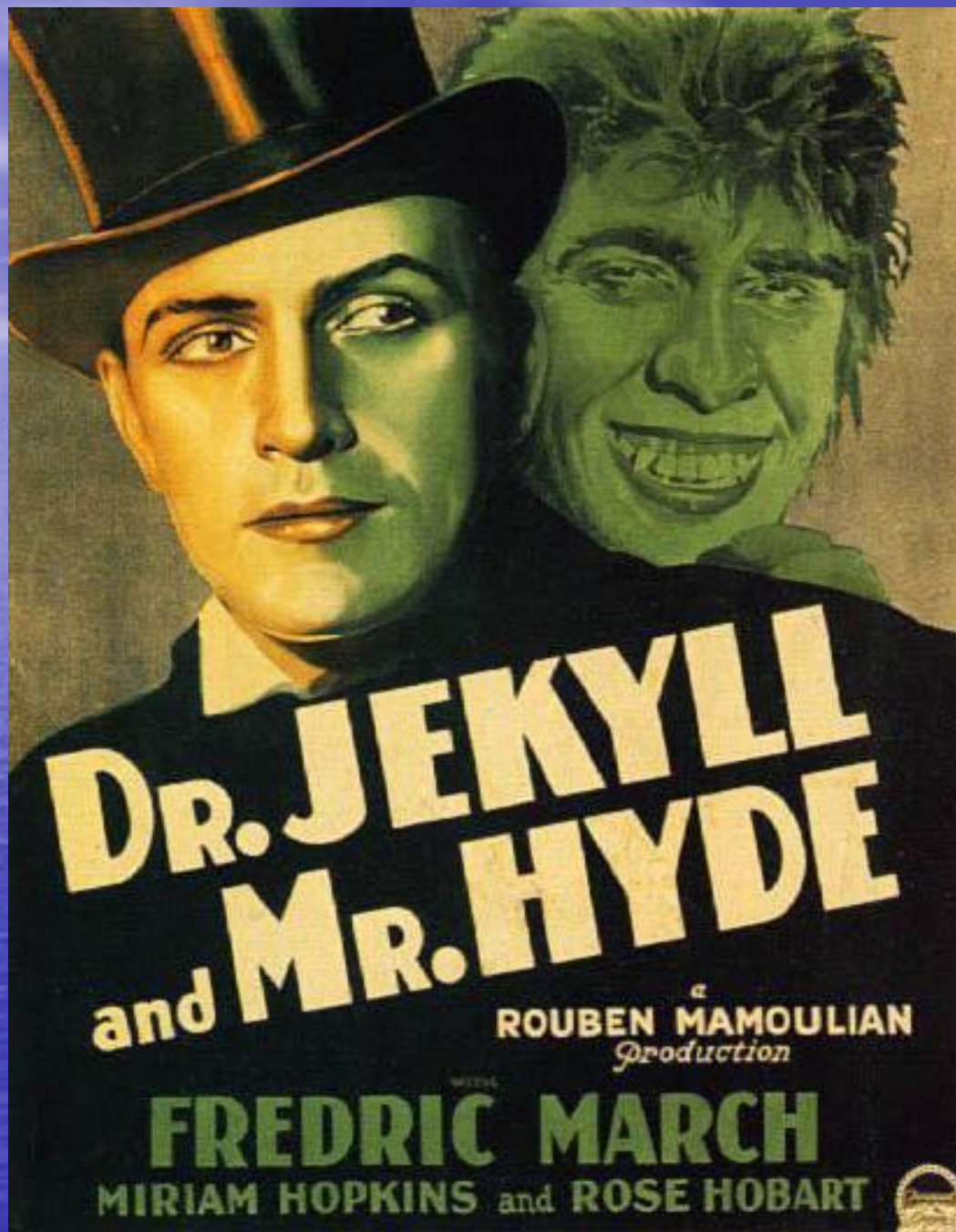
(Abell 85, claim detection by

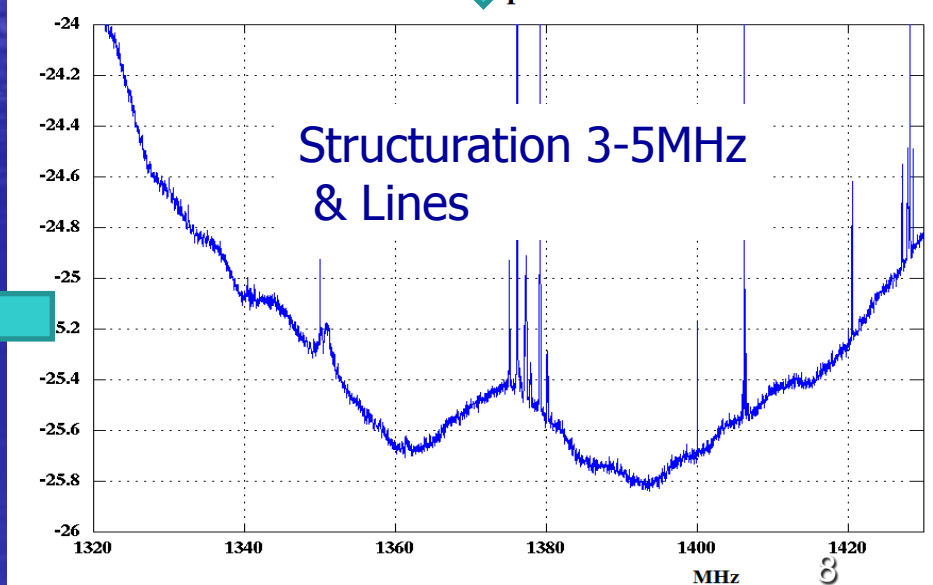
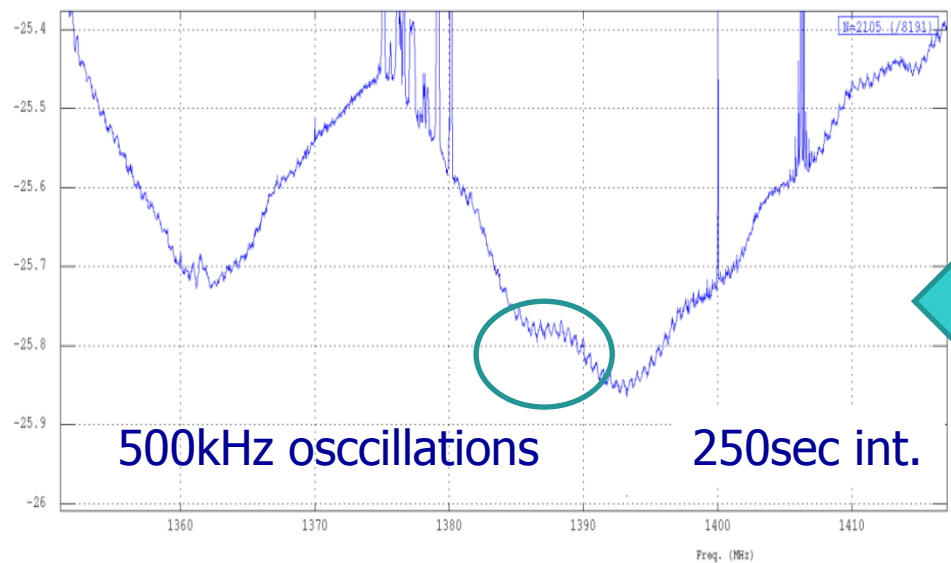
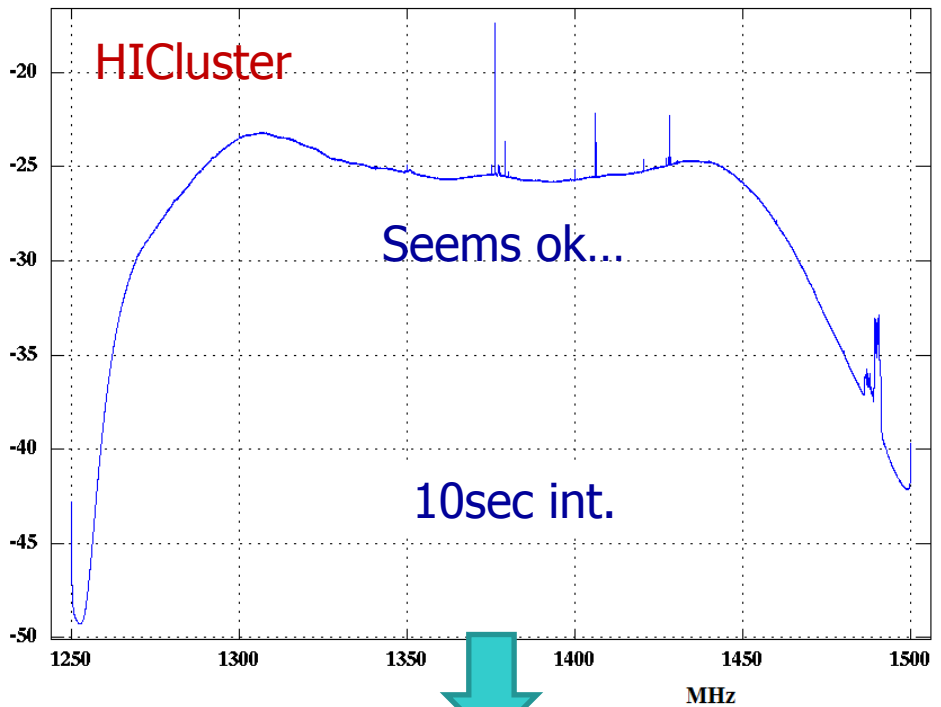
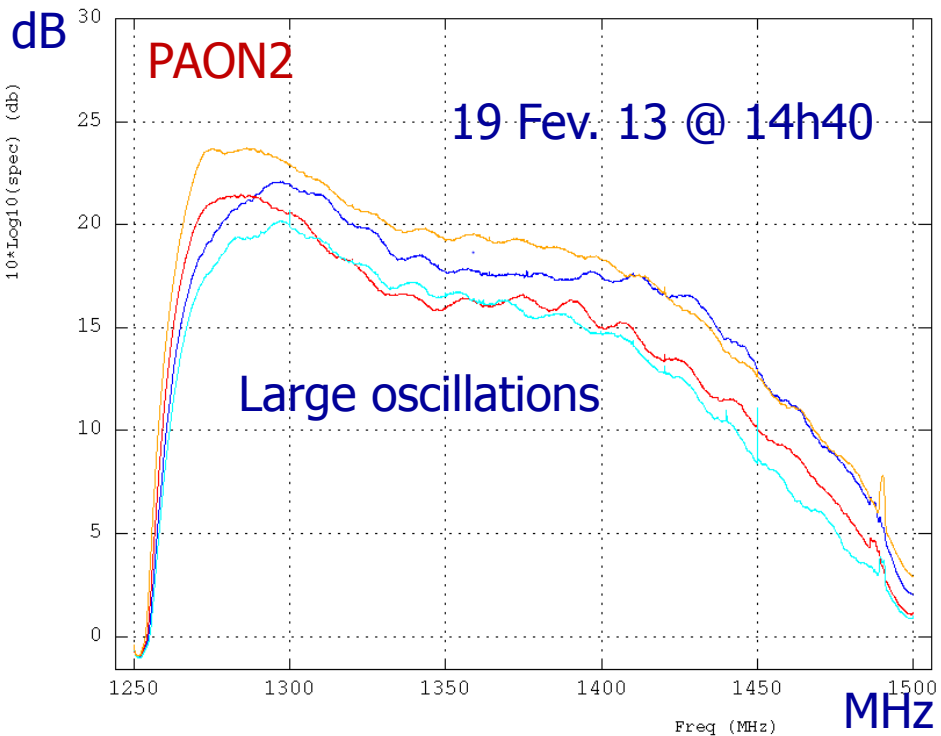
Bravo-Alfaro A&A 495,

Issue 2, p.379-387,2008

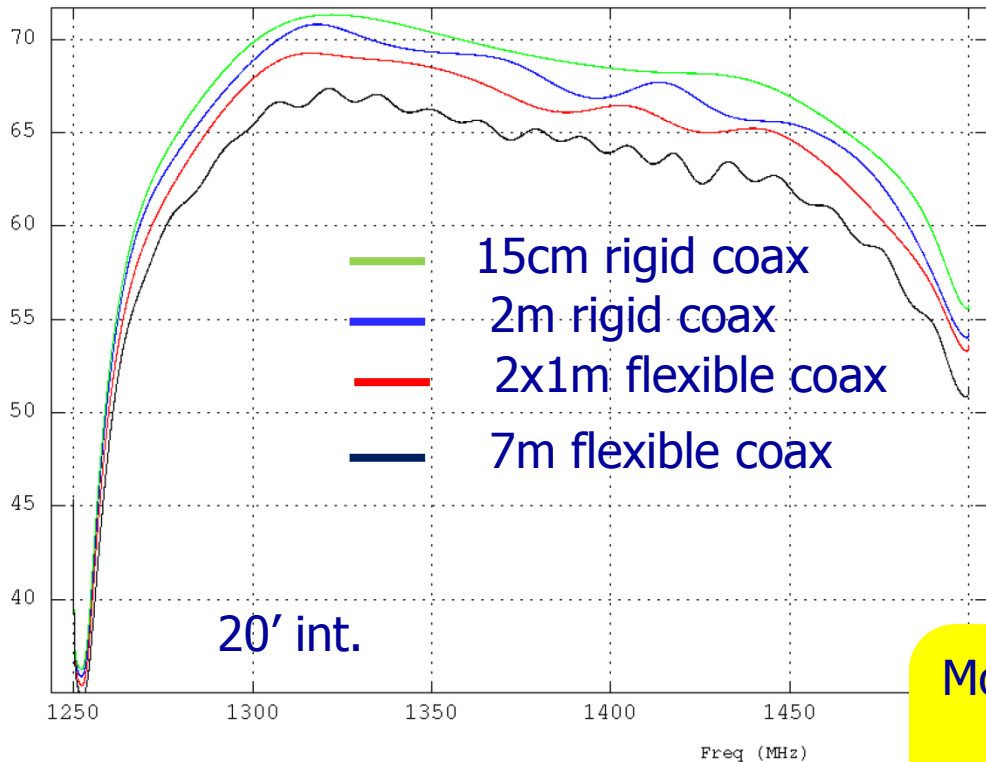
HI flux $\sim 1 \div 3 \text{ mJy}$)







LAL Test bench



Empirical formula

$$\Delta v L_{cable} \sim 96 \text{ MHz.m}$$

Good also to explain the 500kHz with 66% velocity Factor. 300m ~ NRT secondary mirror to chariot mirror.

Motivation for new electronic R&D
AS CLOSE AS POSSIBLE
 to the feed for digitization and optical transfert to visibility calculator.

